# STRUCTURALAND GEOCHRONOLOGICAL STUDY OF THE LOWEST KOHISTAN COMPLEX, INDUS KOHISTAN REGION IN PAKISTAN, NW HIMALAYA 

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# STRUCTURAL AND GEOCHRONOLOGICAL STUDY <br> OF THE LOWEST KOHISTAN COMPLEX, INDUS <br> KOHISTAN REGION IN PAKISTAN, NW HIMALAYA 

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"TMenn ich nun aber eben diese Spalten und Risse als Bucbstaben bebandelte, sie $z u$ entziffern bä̈te, sie zu 7 Morten bildete und sie fertig zu lesen lernte, bättest du etwasdagegen?"

Johann Wolfgang von Goethe (aus Wilhelm Meister)

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#### Abstract

A northward dipping subduction active during Mid Cretaceous times began producing a volcanic arc prior to 100 Ma . This volcanic arc was installed on the oceanic crust somewhere in the Tethys around the equatorial zone: The Kohistan Arc Complex (KAC), now located between the Eurasian and Indian Plates, is bounded by two major suture zones, the Indus Suture Zone (ISZ) at its southern border and the "Northern Suture" at its northern border, and thus part of the NW Himalayan collisional system.

This research work aimed at elucidating two main questions: (1) Is there a fabric in the mid-crustal plutons that is related to the kinematics of the subduction system and (2) which deformation mechanisms were involved during the arc evolution. To constrain the geological setting, detailed mapping and structural traverses were carried out in the Indus region of NW Pakistan. Petrological and geochemical analyses were conducted on key lithologies to identify the magma origin and the metamorphic evolution. Geochronological measurements were carried out to provide time constraints on the magmatic $(\mathrm{U} / \mathrm{Pb}$ on zircon) and exhumation (Fission Track on zircon and apatite) histories. Structural analysis of shear zones and paleostress analysis on faults document the deformation history of the southern (and former lower) part of the KAC.

The base of the arc is characterized by the mantle - crust transition between an ultramafic sequence at its bottom and a granulitic gabbro above. The overlying Metaplutonic Complex comprises sheared gabbros to diorites (e.g. Sarangar Gabbro in the Patan region) and a sequence composed of imbricate (meta-) gabbroic, dioritic, granitic and tonalitic rocks intruded in a crustal sequence. The Granulitic and Sarangar gabbros intruded the arc at its base at pressures of at least ca. 1.0-1.2 GPa under high temperature $>$ $1000{ }^{\circ} \mathrm{C}$. Strain localisation took place continuously from magmatic emplacement to solid state deformation during cooling of the plutons and formed 3 successive sets of shear zones. Set 1 , composed of associated discrete Riedel and shear zones developed above solidus conditions during SW-ward thrusting. Continuous deformation from solidus (granulite facies conditions at ca. $800{ }^{\circ} \mathrm{C}$ and $\sim 1.1 \mathrm{GPa}$ ) to amphibolite facies conditions of $\sim 600^{\circ} \mathrm{C}$ and $\sim 0.8 \mathrm{GPa}$ between 99 and 83 Ma formed the set 2 shear zones. The lower amphibolite facies set 3 shear zones (at ca. $550{ }^{\circ} \mathrm{C}$ and $\sim 0.4 \mathrm{GPa}$ ) are differentiated by strain within thicker and retrograde mylonitic zones. The anastomosing pattern of shear zones described here probably represents arc-related deformation during subduction of the Tethys oceanic lithosphere below the KAC, because it is older than the collision between the arc and the Indian plate and most likely older than the collision of the arc with Eurasia along the "Northern Suture".

The subduction related magmatism produced during 99-92 Ma the series of the above mentioned gabbros, diorites and granites with a calc-alkaline to tholeiitic character. The geochemical signature is similar to that of modern arc systems (e.g. Tonga, New Hebrides and Mariana). They were generated by partial melting of mantle with MORB-type Hf isotopic characteristics $(\varepsilon \mathrm{Hf}=+14)$. The formation of granites could be the result of partial melting at the base of the arc, a process which would efficiently remove and


transport upward granitic components from the root of the arc. Extensional magmatism, 85 Ma ago (emplacement age of the Chilas Complex) tapped a more fertile mantle source $(\varepsilon \mathrm{Hf}=+10)$. Remelting of the deep base of the arc 82 Ma ago resulted in felsic peraluminous dykes, which again show MORB-type Hf isotopic compositions inherited from the initial rocks. A diorite intruded ca. 1865 Ma ago in the Indian Plate (at Duber Bazar) and represents a within plate pluton, unrelated to the arc history. The isotopic and geochemical results of analysed rocks indicate that the arc development is characterized by changes in melt source region and by processes of melt extraction and re-enrichment, leading to a maturation from a basic towards a more felsic composition of the arc crust.

After ca. 83 Ma further cooling and decompression (exhumation) took place. The zircon FT ages at the Indus valley are, in the footwall of today’s ISZ geometry $\sim 23 \mathrm{Ma}$ and in the hanging wall $\sim 42 \mathrm{Ma}$. This indicates, that the change from reverse to normal faulting at the ISZ was not later than 23 Ma . Normal faulting outlasted the termination of extension of ISZ segments elsewhere as is suggested by apatite FT ages $(\sim 3.7 \mathrm{Ma}$ for the footwall and $\sim 10 \mathrm{Ma}$ for the hanging wall). This relatively young differential movement here across the ISZ is a local phenomenon caused by the growth of the "Besham Syntaxis", a possible small-scale equivalent to the Nanga Parbat "Syntaxis". This local phenomenon is most likely "decoupled" from the dynamic evolution during late, brittle deformation within the southern KAC. There, the dynamic evolution is characterized by a general $\mathrm{N}-\mathrm{S}$ compression, disturbed by an $\mathrm{E}-\mathrm{W}$ compression, tentatively related to the formation of the nearby Nanga Parbat "Syntaxis" and subsequent radial extension.

## ZUSAMMENFASSUNG

Der Kohistan-Inselbogen-Komplex (KAC), ein ehemaliger vulkanischer Inselbogen, hat sich vor mehr als 100 Ma über einer in der Mittleren Kreidezeit initiierten Subduktionszone entwickelt. Dieser vulkanische Inselbogen bildete sich auf der ozeanischen Kruste der Tethys in äquatorialen Breiten. Der KAC befindet sich heute zwischen dem Eurasischen Kontinent und der Indischen Platte und ist durch zwei grosse Suturzonen begrenzt. Die Indus-Sutur-Zone (ISZ) bildet die südliche, die sogenannte Nördliche Sutur die nördliche Begrenzung des KAC. Der KAC ist ein in das NW Himalaja Kollisionssystem mit einbezogener ehemaliger vulkanischer Inselbogen.

Diese Studie beabsichtigt zwei wichtigen Fragen zu erörtern: (1) gibt es in den Plutonen des unteren bis mittleren Krustenstockwerks Strukturen, die im Zusammenhang mit der Kinematik einer Subduktion stehen und (2) welche Deformationsmechanismen wirken bei der Inselbogenentwicklung. Neben einer detaillierten geologischen Kartierung des Indus Gebiets im Nordwestens Pakistans wurden auch strukturgeologische Profile angefertigt, um den geologischen Aufbau festzulegen. Anhand petrologischer und geochemischer Analysen wird der Ursprung der Magmen und die metamorphe Geschichte dargestellt. Geochronologische Untersuchungen zeigen die magmatische Geschichte ( $\mathrm{U} / \mathrm{Pb}$ an Zirkon) und Exhumation (Spaltspurendatierung an Zirkon und Apatit) des südlichen KAC. Mit Hilfe von Strukturanalysen der Scherzonen und Paläospannungs-Analysen an Sprödbrüchen wurde die Deformationsgeschichte im südlichen (und ehemals unteren) Teil des KAC rekonstruiert.

Die Basis des Inselbogens ist charakterisiert durch einen Mantel - Kruste Übergang mit einer ultramafischen Sequenz im unteren und einem granulitischen Gabbro im oberen Bereich (Jijal Komplex). Der darüber folgende Metaplutonische Komplex besteht aus duktil verformten Gabbros und Dioriten (z.B. Sarangar Gabbro in dem Gebiet von Patan) und einer Abfolge imbrikierter (Meta-) gabbroischer, dioritischer, granitischer und tonalitischer Gesteine, die in eine Krustenabfolge intrudiert sind. Der granulitische und der Sarangar Gabbro sind unter Druckbedingungen von mindestens $1.0-1.2 \mathrm{GPa}$ und Temperaturen $>1000{ }^{\circ} \mathrm{C}$ an der Basis des Inselbogens intrudiert. Hier erfolgte eine kontinuierliche Lokalisierung der Verformung während der Abkühlung der Plutone. Von der magmatischen Inplatznahme bis hin zu Deformation im festen Zustand bildeten sich drei aufeinanderfolgende Scherzonen-Sets. Set 1, gebildet aus assoziierten diskreten Riedel- und Überschiebungs- Scherzonen, entwickelte sich oberhalb der Solidus-Bedingungen während einer SW gerichteten Überschiebung. Set 2 wurde durch eine kontinuierliche Verformung vom festen Zustand unter granulitfaziellen Bedingungen von ca. $800{ }^{\circ} \mathrm{C}$ und $\sim 1.1 \mathrm{GPa}$ hin zu amphibolitfaziellen Bedingungen von ca. $600{ }^{\circ} \mathrm{C}$ und $\sim 0.8 \mathrm{GPa}$ zwischen 99 und 83 Ma gebildet. Die Scherzonen der unteren Amphibolitfazies (ca. $550{ }^{\circ} \mathrm{C}$ und $\sim 0.4$ GPa), Set 3 sind durch eine Verformung innerhalb von mächtigen und retrograden Mylonitzonen gekennzeichnet. Das hier beschriebene anastomosierende Muster der Scherzonen ist älter als die Kollision zwischen dem KAC und der indischen Platte und höchstwahrscheinlich auch älter als die Kollision
entlang der Nordlichen Sutur und spiegelt folglich die Subduktion ozeanischer Lithosphäre der Tethys unter dem KAC wider.

Der mit der Subduktion zusammenhängende Magmatismus produzierte zwischen 99 und 92 Ma die zuvor bereits genannte Serie von Gabbros, Dioriten und Graniten mit kalk-alkalischen bis tholeiitischen Charakter. Ihre geochemische Signatur ist vergleichbar der moderner Inselbogensystemen (z.B. Tonga, Neue Hebriden und Marianen). Die Gabbros, Diorite und Granite wurden durch teilweise Aufschmelzung des Mantels, der eine MORB typische Hf Isotopensignatur von $\varepsilon \mathrm{Hf}=+14$ anzeigt, generiert. Die Bildung der Granite kann aber auch das Ergebnis einer teilweisen Aufschmelzung der Basis des Inselbogens sein, ein Prozess, der sehr effizient granitische Komponenten von der Basis des Inselbogens entfernt und sie im Inselbogen weiter nach oben transportiert. Extensionaler Magmatismus (Intrusion des Chilas Komplexes) hat eine fertilere Mantelquelle $(\varepsilon \mathrm{Hf}=+10)$ angezapft. Wiederaufschmelzung der Basis des Inselbogens vor ca. 82 Ma resultierte in felsischen, peraluminischen Gängen, die wieder eine vom initialen Magmatismus vererbten MORB typische Hf Isotopensignatur von $\varepsilon H f=+14$ anzeigen. Erste geochemische Ergebnisse deuten darauf hin, dass die Inselbogenentwicklung durch Wechsel in den Ursprungsregionen der Schmelzen und durch Prozesse wie Schmelzextraktion und Wiederanreicherung charakterisiert ist. Die beiden letztgenannten Prozesse führten zu einer Reifung des urpsrünglich von basischen Gesteinen dominierten Inselbogenkruste hin zu einer vermehrten felsischen Zusammensetzung.

Ab ca. 83 Ma erfolgte eine weitere Abkühlung und Dekompression (Exhumation) des Metaplutonischen Komplexes. Die Spaltspurendatierung an Zirkonen ergab im Indus Tal im Liegenden der heutigen Geometrie der ISZ $\sim 23 \mathrm{Ma}$ und im Hangenden $\sim 42 \mathrm{Ma}$. Das deutet darauf hin, dass die abschiebende Bewegung entlang der ISZ spätestens vor 23 Ma einsetzte. Die Abschiebung überdauerte hier in ihrer Aktivität die Extension entlang anderer Teilstücke der ISZ, wie aus der Spaltspurendatierung an Apatit ( $\sim 3.7 \mathrm{Ma}$ im Liegenden, $\sim 10 \mathrm{Ma}$ im Hangenden) hervorgeht. Diese junge differentielle Bewegung an der ISZ ist hier eine lokale Erscheinung, die mit dem Wachstum eines wahrscheinlich kleinräumigen Äquivalents der Nanga Parbat „Syntaxe" einhergeht. Diese lokale Erscheinung ist wahrscheinlich von der dynamische Entwicklung während der späten, spröden Deformation innerhalb des südlichen KAC unabhängig. Die dynamische Entwicklung ist hier durch eine generelle $\mathrm{N}-\mathrm{S}$ Einengung gekennzeichnet, die durch eine, wahrscheinlich auf die Bildung der Nanga Parbat „Syntaxe" zurückzuführende E- W Kompression und darauffolgender radialer Extension gestört wurde.

Die vorliegende Studie zeigt, dass es Deformationsstrukturen an der Basis eines Inselbogens gibt, die im Zusammenhang mit der Kinematik einer Subduktion stehen. Die Lokalisierung der Verformung ist dabei verantwortlich für das anastomosierende Muster der Scherzonen.

## 1. INTRODUCTION

### 1.1 Aims and Project Background

This thesis is part of the research project \# 20-49372.96 supported by the Swiss National Science Foundation. The project was a three-year interdisciplinary investigation to understand the deformation processes in volcanic arc systems during subduction. Two main questions were raised: 1) Is the mid-crustal fabric in deformed plutons related to the kinematics of the subduction system rather than collision and 2) which deformation mechanisms were involved during the arc evolution?

To answer these questions, the Kamila Shear Zone (Treloar et al., 1990) was chosen as a case history. The aim of this thesis is to provide the regional setting and specify the geological history of the southern part of the Kamila shear zone in the Kohistan Arc Complex (KAC). Several methods were used, including detailed mapping, structural traverses, microstructural analysis, petrological analysis and geochronology.

### 1.2 GeOGRaphic outline of the study area

The investigated area is located in NW Pakistan along the Indus, Kandiah and Natai valleys in the Kohistan area, N of Besham (Fig. 1.1). The central part covers a ca. 50 km long NE - SW area stretched from Duber Bazar (ca. 12 km N of Besham) in the SW to Dasu in the NE along the Karakorum Highway, which provides extraordinarily good and unweathered outcrops and a complete section through the crust of a former island arc (Tahirkheli et al., 1979; Bard, 1983; Coward et al., 1986). With extension of mapping into the smaller tributary valleys to the Indus River, the mapped area covers $2000 \mathrm{~km}^{2}$. The morphology is characterised by steep slopes and gorges. Only few small plains, build up of quarternary river alluvium occur. Vegetation is generally sparse and localised on gentler slopes. Dense forests may be found in an altitude range of 2000 to 3500 m . In populated regions terrace cropping, mainly corn, is intensive.


Fig. 1.1: Location of investigated areas (squared) in NW Pakistan. The main part extends along the Indus River north of Besham.

### 1.3 Tectonic Gross structure of the Himalaya and Geological Overview

### 1.3.1 Tectonic Evolution of the Himalaya

The break-up of Gondwana and subsequent opening of the Tethys was followed by subduction and northward drift of the Indian plate. The Indian subcontinent separated ca. 130 Ma ago from Gondwana (Johnson et al. 1976). The northward drift of the Indian plate can be divided into three different phases (Patriat \& Achache, 1984): 1) relative drift velocity between Asia and the Indian plate of $15-20 \mathrm{~cm} / \mathrm{yr}$ between $83-52 \mathrm{Ma}$; 2) relative drift velocity of less than $10 \mathrm{~cm} / \mathrm{yr}$ between $52-36 \mathrm{Ma}$ and 3) relative drift velocity of less than $5 \mathrm{~cm} / \mathrm{yr}$ since 36 Ma . For the time span between 52 and 36 Ma the drift direction changed several times. The reduction of drift velocity around 52 Ma is attributed to the collision between the Indian plate and Asia. Precise timing of the collision is under discussion and ranges from $40-55 \mathrm{Ma}$ based on paleomagnetic data
(Royer et al., 1992; Klootwijk et al., 1985) and goes up to ca. 65 Ma indicated by intercontinental migration of fauna (Jaeger et al., 1989).

The paleogeographic position of the Kohistan Arc (Fig. 1.2), an intraoceanic island arc developed above the subduction zone, is conjectural. Nevertheless, existing data constrain the mid-Cretaceous position of the Kohistan Arc around the equatorial zone and close to the southern marginal terrane (Karakoram block) of the Eurasian continent (Yoshida et al., 1996). At the same time, the Tajikistan and Tarim basins within Eurasia were located more than 2000 km to the north of the Kohistan Arc and Karakoram block (Yoshida et al., 1996). Therefore a large convergence occurred between the terranes forming today the southern Eurasian continent (Molnar \& Chen, 1978; Chen et al., 1993). The Indian plate was, during this time, in the southern hemisphere (Klootwijk et al., 1986a,b). In Middle Eocene times there is no significant difference in paleolatitudes between the Indian plate and the Kohistan Arc (Yoshida et al., 1996) which points to complete suturing along the Indus Suture Zone.


Fig. 1.2: a) Northward drift of India from 70 Ma (after Patriat \& Achache, 1984) and b) paleo-latitudes of the Eurasian continent, Kohistan Arc Complex and India calculated and estimated by paleomagnetism (after Yoshida et al. 1996). The stippled lines represent the positions of (from $S$ to $N$ ) northern Indian Plate, Kohistan Arc Complex, Tarim and Tajikistan (part of the Eurasian continent). The exact paleoposition of the Karakoram Block $N$ of Kohistan is not known and not displayed. The Karakoram Block was separated from Kohistan by an ocean of unknown extent until 102 $\pm 12$ - 75 Ma (Petterson \& Windley, 1985).

The structure of the resulting mountains has been summarised in several publications (e.g. Gansser 1964). From N to S the major tectonic contacts and units in between are:

- Indus-Yarlung Tsangpo Suture Zone (IYTSZ)
- The Tethyan Himalaya
- Southern Tibet detachment system (STDS)
- The High Himalaya
- Main Central Thrust (MCT)
- The Lesser or lower Himalaya
- Main Boundary Thrust (MBT)
- The sub-Himalaya

Since the Paleocene a significant amount of shortening, estimated to a maximum of 2000 km (Molnar \& Tapponier, 1975; Molnar, 1984) took place between the Tethyan Himalaya and the Indian craton. Estimates based on balanced cross sections give significantly lower amounts of $500-700 \mathrm{~km}$ of shortening for northwestern Pakistan (Butler, 1986). Malinconico (1989) based his estimates on the remaining crustal volume in the orogen. His crustal model based on gravity profiles can account for a shortening of 570 to 1140 km . Whether the 2000 km of shortening are an overestimate or whether part of the shortening was accommodated by diffuse deformation and/or erosion is still under discussion. It is generally accepted, that a major part of the shortening was absorbed by large scale thrusts like the MCT, MBT and MFT which are parts of the fold-and-thrust wedge within the Indian Plate. The MBT forms, along with the MFT (Main Frontal Thrust) and the MCT a northward dipping decollement horizon which is traced to at least 225 km north of the Himalayan Mountain front at a depth of ca. 50 km beneath South Tibet (Seeber et al., 1981). In Pakistan the major tectonic units are slightly different and are, from N to S :

- Karakoram
- Northern Suture
- Kohistan Arc Complex (KAC)
- Indus Suture Zone (ISZ)
- Indian Plate.


### 1.3.2 Karakoram

The Karakoram range, of Gondwana affinity (Gaetani, 1997), forms the southern part of Eurasia. It was sutured along the Rushan-Pshart suture zone (Shvolman, 1978; Ruzhentsev \& Shvolman, 1981) to the Pamir range during the Late Triassic. Three major subunits are distinguished: 1) The northern Sedimentary belt, 2) The Karakoram Batholith and 3) The southern metamorphic complex. The northern part is formed by unmetamorphosed Ordovician to Cretaceous sediments (Zanchi, 1993; Gaetani et al., 1995). The central part is comprised of calc-alkaline plutons yielding Jurassic to Miocene ages (Searle, 1991). The mid Cretaceous plutons with ages of $95 \pm 5$ (LeFort et al., 1983) and $105.7 \pm 0.5 \mathrm{Ma}$ (Fraser et al., 1999) are related to northward subduction of the north Tethys lithosphere below the Karakoram (Debon et al., 1987). The metamorphic rocks to the S of the Karakoram batholith record 3 phases of metamorphism: Jurassic (?) to lower Cretaceous (andalusite garnet grade), Barrovian type with a minimum age of 37 Ma (Searle, 1991) and contact metamorphism around the 21 Ma old Baltoro plutonic unit (Searle \& Tirrul, 1991).

The closure of the Northern Suture (Tahirkheli et al., 1979) is constrained to a minimum age of 75 Ma by the age of undeformed basic dikes which crosscut closure-related fabrics along the Northern Suture zone (Petterson \& Windley, 1985).

### 1.3.3 Kohistan Arc Complex (KAC)

The Kohistan Arc Complex (Fig. 1.3) is sandwiched between the Karakoram terrane to the north and the Indian Plate to the south (Tahirkheli et al., 1979; Bard, 1983; Coward et al., 1986). The eastern limit is formed by the N-S trending Nanga Parbat syntaxis, which is a crustal antiform where the Indus Suture Zone (ISZ) is folded around a halfwindow of Indian crust (e.g. Treloar et al., 1991a; Burg \& Podladchikov, 2000). The syntaxis underwent rapid tectonic denudation and growth during the last 5 Ma (Zeitler et al., 1989, 1993). To the east of the syntaxis, the Ladakh Arc is the eastern equivalent to the Kohistan Arc system. Both Arc Complexes may have represented a single island arc which collided with the southern margin of Eurasia before India collided with the system (Petterson \& Windley, 1985; Coward et al., 1986; Pudsey, 1986; Searle, 1991).

The Kohistan Arc Complex comprises Cretaceous and Tertiary igneous and subordinate sedimentary rocks which are divided into several distinct units and groups. They are,
roughly from N to S: Yasin Group, Utor and Chalt Volcanics, Oceanic series, Kohistan Batholith, Chilas Complex, Metaplutonic Complex and Jijal Complex. This succession is interpreted as calc-alkaline plutons intrusive into an oceanic crust and overlain by the calc-alkaline lavas and associated sediments. Accordingly, the interpretation is an intraoceanic arc that developed during the Cretaceous somewhere in the Tethys ocean (Tahirkheli et al., 1979; Bard 1983; Coward et al., 1986).

### 1.3.3.1 Yasin Group Sediments

Just to the south of the Northern Suture, the Yasin group consists of interlayered volcanoclastic sediments, volcanites and rather immature turbidites deposited in a deepwater environment (Tahirkheli, 1979, 1982; Pudsey, 1986). At the type locality, Pudsey et al. (1985) divided the group into an upper 2000 m thick sequence of shales with interbedded greenstones, probably tuffs, a middle 500 m thick unit with grey slates and distal turbidites and a lower part, consisting of volcano-lithic conglomerate and sandstones, tuffs, slates and rudist limestones with an Albian-Aptian fauna. The lithologies vary laterally. Sedimentation occurred most likely in an intra-arc and/or back arc basin (Pudsey, 1986; Khan et al., 1995).

### 1.3.3.2 Chalt Volcanics

The Chalt Volcanic Group comprises calc-alkaline andesites to rhyolites succeeding to andesitic lavas, tuffs and agglomerates of Early Cretaceous age (Petterson et al., 1991a). The Chalt volcanics are locally interbedded with the Yasin group. Low grade metamorphism is characteristic.

### 1.3.3.3 Oceanic series

To the southwest and within the Kohistan Complex, metasedimentary sequences of marine origin (Dir, Utror and Kalam Groups, Tahirkheli, 1979) yield Eocene fossils in upper-level limestones (Kakar et al., 1971; Khan, 1979; Sullivan et al., 1993). The Utror volcanics show low grade metamorphism and a variety of rocks similar to Chalt volcanics. Ar-Ar hornblende age for a basaltic flow within the Utror volcanics is $55 \pm 2$ Ma (Treloar et al., 1989a).

### 1.3.3.4 Kohistan Batholith

The Kohistan Batholith is the principal unit of the Kohistan magmatic arc. It shows at least 3 formative stages (Petterson \& Windley, 1991; Petterson et al., 1991b): an early
suite of gabbro-diorites and trondhjemites formed within an island arc setting between $110-85 \mathrm{Ma}$; the main bulk was formed within an Andean type plate margin and comprises an older gabbro-diorite suite at 85 to 60 Ma and a younger granitic suite, emplaced at $60-40 \mathrm{Ma}$; and a postcollisional magmatism with leucogranitic sills and dikes between 40 and 26 Ma (Petterson \& Windley, 1985). Stage 1 and early stage 2 plutons have isotopic signatures characteristic of a mantle derivation. The isotopic signatures of younger plutons show evidence for an increasing crust to mantle ratio, with the latest magmas being crustal derived. This evolution is interpreted as the result of arc thickening and lower arc melting following suturing to Asia (Petterson \& Windley, 1991; Petterson et al., 1993).

### 1.3.3.5 Chilas Complex

The Chilas Complex consists of a large, 8 km thick and 300 km long body of mafic ultramafic rocks. The bulk rock type is noritic gabbro, locally layered with subordinate diorites. Several lenses of ultra-mafic-anorthosite associations (dunite, pyroxenite and peridotite) are interpreted as tops of mantle diapirs (Burg et al., 1998). Timing for the emplacement of the Chilas Complex gabbronorites relative to the collision of the Kohistan Arc with the Karakoram plate is under discussion. The gabbronorites have intruded the Jaglot schist group to the north and the Metaplutonic Complex to the south. Treloar et al. (1996) suggest an emplacement of the Chilas Complex "after suturing" of the Kohistan Arc to the Karakoram Plate due to the low deformation of the Chilas Complex compared to the intruded older units. Granulite facies re-equilibration occurred short after emplacement of gabbronorites at $600-800{ }^{\circ} \mathrm{C}$ and $0.6-0.8 \mathrm{GPa}$ (Jan \& Howie, 1980; Bard, 1983). Amphibolite facies shears, related to the southwestward thrusting on a pile of imbricated calc-alkaline laccoliths, have taken place shortly after re-equilibration as indicated by cooling below $500^{\circ} \mathrm{C}$ at around 80 Ma (Treloar et al., 1989a) The Chilas complex may have provided the thermal energy for regional metamorphism within the Kohistan arc. Conventional zircon U-Pb age yields an age of ca. 84 Ma reported by Zeitler et al. (1981).


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### 1.3.3.6 Metaplutonic Complex

A thick pile of imbricated calc-alkaline laccoliths, metavolcanics and remnants of the Tethyan oceanic crust, variably sheared in amphibolite facies conditions forms the socalled Kamila amphibolite belt located between the Jijal and Chilas Complex (Tahirkheli et al., 1979; Bard, 1983; Jan, 1988; Treloar et al., 1990). The variety of lithologies summarized under the name Kamila amphibolite belt is not well represented by this term. In addition, the Kamila amphibolites s.s. (Treloar et al., 1990), exposed in the region of Kamila, are only one part of the belt. In the Indus area metaplutonic rocks are dominant. The proposed name Metaplutonic Complex summarizes three subunits (Sheared gabbro-diorites, Kiru Sequence and Kamila Sequence) and emphasizes the essentially plutonic origin. Amphibolite facies conditions ended around $80 \mathrm{Ma}(\mathrm{Ar}-\mathrm{Ar}$ cooling ages on hornblendes cluster, Treloar et al., 1989a, Wartho et al., 1996).

### 1.3.3.7 Jijal Complex

The Jijal Complex, is composed of more than 3 km thick ultramafic rocks overlain by garnet-plagioclase granulites. The sharp contact between the ultramafic rocks and the overlying granulites, with well preserved igneous structures, is the intrusive contact of lower crustal, calc-alkaline garnet-gabbros (the Granulitic Gabbro) within upper mantle rocks. The contact is also the lower boundary of the arc crust, i.e. the arc-Moho (Burg et al., 1998). The Granulitic Gabbro underwent metamorphism at 750 to $1150{ }^{\circ} \mathrm{C}$ and 1.2 to 1.9 GPa (Ringuette et al., 1999) and has later re-equilibrated at $>700^{\circ} \mathrm{C}$ and $1.5 \pm 0.4$ GPa (Jan \& Howie, 1981), which are pressure conditions similar to those calculated from the underlying, unmetamorphosed ultramafic rocks. $\mathrm{Sm}-\mathrm{Nd}$ age of the granulite facies assemblage is around 95 Ma (Anczkiewicz \& Vance, 2000; Yamamoto \& Nakamura, 2000).

### 1.3.4 Indian Plate

The contact between the Kohistan Arc Complex and the Indian Plate can be observed close to Jijal. In many parts the ISZ comprises a "mélange" with tectonic blocks of ophiolite, blueschists, greenschists, metavolcanics and metasediments (Jan, 1980; Kazmi et al., 1984, Anczkiewicz et al., 1998). The Indian Plate rocks comprise gneisses and Mid Proterozoic diorites belonging to the Besham Group (Tahirkheli, 1979). The Besham group forms a N-S trending antiform, a small-scale equivalent to the Nanga

Parbat Syntaxis (Coward, 1986; Treloar et al., 1989c). Radiometric ages were obtained for an amphibolite sheet within a granite near Duber with Ar-Ar age on hornblende at $1920 \pm 24 \mathrm{Ma}$ and for a granite near Duber with K-Ar age on biotite at $550 \pm 20 \mathrm{Ma}$ (Treloar et al., 1989a).

### 1.3.5 Study area: Indus Kohistan Region

The studied area provides a "section" through the lower units of the Kohistan Complex. It also includes part of the northernmost Indian plate. The main focus was given to the geology in the Indus valley and tributary valleys because of the good outcrop situation, accessibility and, the most important fact, the valley cuts the relevant units almost perpendicular to strike (Fig. 1.4).


Fig. 1.4: Morphology of the Indus valley and mapped areas. $3 D$ view of Landsat satellite image (Band 1, 2 and 3) was generated using interpolated Gtopo 30 elevation data (original ca. 1000 meter grid was interpolated to a ca. 333 meter grid with kriging method). Altitude is vertically exaggerated by 1.5 .

Two more areas were investigated and partly mapped to provide additional information on the units bounding the area of interest: 1) to the N, the Kandiah valley, which screens a central part of the Chilas Complex and 2) to the $S$ the Allai and Natai valleys for getting information on a part of the Indus Suture Zone, where a "mélange" is involved, in contrast to the ISZ at Jijal, where the contact is characterized by a narrow fault zone.

### 1.3.5.1 Kandiah Valley

The Kandiah is a ca. 30 km E-W striking valley within the Chilas Complex. At the E end the valley joins the Indus valley through a steep gorge, almost at the location where the Indus river bends from an E-W direction to a N-S direction. To the W the valley is terminated by a N-S trending up to 6257 m high mountain chain that also builds the natural barrier to the Swat valley and Kalam.

The geology of the Kandiah valley is not mapped in detail. Only few publications and maps describe or provide data from this area. Yamamoto (1993) shows a retrograde metamorphic overprint for the Chilas gabbronorites. P-T estimates are $0.61-0.8 \mathrm{GPa}$ and $607-720{ }^{\circ} \mathrm{C}$ (Yamamoto, 1993) for amphibolite facies assemblages which were later altered to greenschist facies assemblages. The metasediments intruded by gabbronorite in the central part of the Kandiah valley are interpreted as equivalent to the Gilgit metasediments (Searle \& Khan, 1993).

### 1.3.5.2 Natai and Allai Valley

The Natai and Allai valleys are located to the ESE of Besham. The Allai valley forms a ca. 22 km long NW-SE valley northeast to the Indus valley with its NW part parallel to it. The confluence between the Indus valley and the Allai valley is ca. 4 km S of Besham. The ca. 3 km long Natai valley branches off perpendicular to the Allai valley (ca. 14 km upstream) towards the NE to the small village of Pastho. This valley provides a section through the eastern part of the Besham syntaxis.

Detailed studies of the mineralogy and geochemistry of the volcanic series involved in the Indus Suture Zone "mélange" in the Natai valley (Majid \& Shah, 1985; Shah \& Majid, 1985) describe a variety of lithologies ranging from ultramafic rocks, pillow lavas with associated meta-chert and limestone, greenschist and blueschist facies graywackes and lavas with a tholeiititc affinity. The ISZ is defined by the fault contact (named Shergarh fault by DiPietro et al., 2000) between the "mélange" units and the

Indian Plate rocks and the fault contact between the "mélange" and the Kohistan Arc Complex, referred as Kohistan Fault. Treloar et al. (1989b) described the inverted metamorphic zoning in the footwall of the ISZ. The Map presented in chapter 2 is based on own observations completing an unpublished map from Lahmeyer International (Allai Khwar Feasability study, Lahmeyer International, 1998).

### 1.3.5.3 Jijal Dasu Region

The Jijal Dasu Region is described by several authors and a reasonable quantity of geochemical (e.g. Jan, 1988; Miller et al., 1991; Jan \& Windley, 1990) petrological (e.g. Yamamoto \& Yoshino, 1998; Yoshino et al., 1998; Ringuette et al., 1999) and geochronological (Treloar et al., 1990; Wartho et al., 1996; Zeitler 1985; Yamamoto et al., 2000; Anczkiewicz \& Vance, 2000) data. Mapping is so far concentrated either on small areas like the Jijal Complex (e.g. Miller et al., 1991) or limited to sections along the Karakoram Highway (e.g. Treloar et al., 1990). The area is characterised by the steep Indus gorge and up to 4000 m high mountains. Several tributaries provide additional N-S sections in the Patan area and dominantly E-W stretching sections in the Dasu area.

The area covers two units of the Kohistan Arc Complex: 1) Jijal Complex and 2) "Kamila Amphibolite Belt" (Jan, 1988) or Metaplutonic Complex (defined and treated later for more details). A central feature, from the structural point of view, is the intense shear deformation in the Granulitic gabbro of the Jijal Complex and within the Metaplutonic Complex. The Kamila Shear Zone identified by Treloar et al. (1990) separates the Chilas Complex from the ultramafic rocks of the Jijal Complex. This up to 38 km wide zone is well exposed along the Karakorum Highway between Patan and Dasu.

### 1.3.6 Tectonometamorphic evolution of the Lower Kohistan Arc Complex

The early arc history is recorded in the Lower Kohistan Arc Complex. The following will give a brief summary about the main tectonometamorphic events.

Remnants of the Tethys crust are most likely preserved as fine grained amphibolites for which geochemical analyses suggest that the high Ti content has similarity to N - and Etype MORB (Kahn et al., 1993; Treloar et al., 1996). The subduction zone formed
during Mid Cretaceous times and produced the volcanic arc. Dating of the granulite facies assemblage provided $\mathrm{Sm}-\mathrm{Nd}$ ages of $95 \pm 2.7$ (Anczkiewicz \& Vance 2000), $118 \pm 12,91.0 \pm 6.3$ and $83 \pm 10$ (Yamamoto \& Nakamura 2000). These ages give the time of cooling below $700-800^{\circ} \mathrm{C}$. Anczkiewicz \& Vance (2000) also show that parts of the Metaplutonic Complex underwent cooling below $700-800{ }^{\circ} \mathrm{C}$ at the same time ( $94.6 \pm 5.3 \mathrm{Ma}$, garnet amphibolite, Swat valley). Based on PT calculations (Yamamoto 1993; Ringuette et al. 1999) a major decompression affected the Granulitic gabbro as well as the Metaplutonic Complex (Anczkiewicz \& Vance 2000). For the petrogenetic evolution of the Granulitic gabbro essentially two models are proposed: Burg et al. (1998) interpreted the Granulitic gabbro as intrusion within mantle rocks, now displayed as the Jijal ultramafics. Ringuette et al. (1999) give a similar interpretation, where cumulates emplaced at the base of the oceanic arc-type crust and underwent subsequently sub-solidus isobaric cooling. Yamamoto (1993) instead suggests heating and burial of two-pyroxene granulites during crustal thickening.

N of the Granulitic gabbro a sequence of metabasic rocks (so-called Kamila amphibolites) is representing the lower to mid crustal level of the arc. The origin of these sheared metabasic rocks is discussed. They represent either the oceanic crust on which the island arc was built (Bard, 1980), or they are a highly deformed, arc-type plutons and volcanite with possible ocean floor relics (Coward et al., 1986). Jan (1988) also suggested that essentially metavolcanic amphibolites include rare sediments. Yamamoto (1993) proposed that the Kamila amphibolites are the hydrated equivalents of the Chilas Complex. Geochemical data indicate, that the Kamila amphibolites is comprised of subduction related intrusions, remnants of the oceanic crust and metasedimentary (-volcanic) portions (Khan et al., 1993). The relative proportion of these different rocks varies strongly along strike. In the Indus area the Kamila amphibolites is built up mainly by metamorphosed plutonic rocks (Kazmi \& Jan, 1997). Therefore the new term Metaplutonic Complex is suggested to emphasize the essentially plutonic origin. The Metaplutonic Complex includes the Patan sheared Gabbros and Diorites, the Kiru sequence and the Kamila sequence. Only the Kamila sequence is equivalent to the Kamila amphibolites s.s. (Treloar et al., 1990). Metamorphic conditions were $550-650{ }^{\circ} \mathrm{C}$ and 0.9 to 1.0 GPa (Treloar et al., 1990).

The Metaplutonic Complex was cooled below $500{ }^{\circ} \mathrm{C}$ at ca. $76 \pm 4 \mathrm{Ma}$ based on $\mathrm{Ar}-\mathrm{Ar}$ dating of hornblende (Wartho et al., 1996).

The metamorphosed plutonic rocks (hereafter called Metaplutonic Complex) and the Granulitic gabbro were intensely sheared, by an impressive anastomosing array of amphibolite facies mylonites within which fabric intensity varies. They were related to the southward propagation of shortening across the Kohistan Complex following suturing with Asia (Treloar et al., 1990). Deformation developed under different temperature conditions. High temperature deformation occurred probably in a range of $500{ }^{\circ} \mathrm{C}$ to $650{ }^{\circ} \mathrm{C}$ (Treloar et al., 1990; Yamamoto, 1993). The anastomosing shear pattern began probably to develop during granulite facies conditions short after the magmatic emplacement. A minimum age for the shearing is given by $83 \pm 1 \mathrm{Ma}$ (Treloar et al., 1990), based on Ar-Ar dating on hornblende grains in a sheared gabbro of the southern Metaplutonic Complex. A K-Ar muscovite age of 66 Ma from a pegmatite that cuts the main shear fabric indicates that post shearing cooling had occurred by 60 Ma (Treloar et al., 1989a).

Later deformation below $450{ }^{\circ} \mathrm{C}$ shows cataclastic behaviour and greenschist facies rock assemblages. After the collision between the Kohistan Complex and subsequent thrusting onto the Indian Plate a major extensional phase took place. It is recognized from folds that indicate NE directed normal movement. These folds and normal faults are interpreted as collapse structures related to the backsliding of the Kohistan Arc (Burg et al., 1996; Vince \& Treloar, 1996). The dynamic evolution during later, brittle deformation is represented by four stages chronologically ordered from field evidence as SSE-NNW compression, E-W compression, radial extension and SSW-NNE compression (see chapter 6 for details). The latter corresponds to the present-day stress field defined from seismic activity. The earlier faulting events were active during the Miocene, when convergence-related stresses were disturbed by the formation of the nearby Nanga Parbat and Besham syntaxes.

## 2. LITHOLOGIES AND STRUCTURES

### 2.1 Techniques

Mapping was carried out during 18 weeks of field work (summer 1997, spring 1998 and summer 1998) on Pakistan topographic sheets at a scale of 1:50.000. Landsat 7 ETM images, areal photographs and digital elevation model Gtopo 30 were additionally used for the recognition of lineaments. Outcrop and sample locations were defined with handheld GPS and additional field/map verification. All these data were combined for processing, analysis and map production in the Geographic Information System ArcInfo8/ArcView3.

### 2.2 Mapped Units

The mapped units are described according to their structural position from the bottom to the top. The units refer to 1) Jijal Dasu region map (Fig. 2.1 and enclosed map), 2) the Allai and Natai valleys (Fig. 2.2), 3) sections for the Jijal Dasu map (Fig. 2.3, Fig. 2.4 and enclosed sections) and the Swat valley (Fig. 2.5). The reconnaissance geological map of the Kandiah valley is provided under subchapter 2.2.3.3 Chilas Complex (Fig. 2.12 and enclosed map).

### 2.2.1 Indian Plate

The Indian Plate lithologies were mapped along the Indus valley south of Jijal (Fig. 2.1) and in the Allai valley (Fig. 2.2). Immediately south of the northward dipping Indus Suture (see section A-A' in Fig. 2.3, section C-C' in Fig. 2.4 and Swat section, Fig. 2.5), the Indian plate is comprised of Proterozoic, strongly foliated and folded gneisses (IPGns). Folds in the mylonitic gneisses (Fig. 2.6a) are NNE trending, recumbent ESE vergent. Their axial plane is sub-parallel to the pervasive foliation and their axes (sub-) parallel to the SW trending stretching lineation. Kinematic determination from rotated feldspar clasts shows a dominant southwestward sense of shear. The gneisses belong to the Besham Group (Fletcher et al., 1986), in a N-S trending open antiform oblique to the SW-ward directed older stretching lineation (Treloar et al., 1989c). In places, the gneisses contain dark xenoliths of amphibolitic composition (Fig. 2.6b). Meta-(grano-) diorites (IP-Dr) are incorporated in these gneisses.


Fig. 2.1: Geological map of the Jijal Dasu region, reduced scale (original scale map is enclosed).


Fig. 2.2: Geological map of the Allai and Natai valleys.

NE



Fig. 2.4: Geological Sections $B-B^{\prime}$ and $C-C^{\prime}$ for the Jijal Dasu region, reduced scale (original scale sections are enclosed).


Fig. 2.5: Geological Section Swat valley (ages from Anczkiewicz, 1998; Anczkiewicz \& Vance, 2000 and this study in bold).

The dioritic body (Fig. 2.6c) at Duber Bazar (so called Duber Diorite) is composed of:
K-feldspar + hornblende + plagioclase + quartz $\pm$ opaque oxides $\pm$ zircons $\pm$ epidotes.
The magmatic fabric is represented by a weak NE-ward dipping foliation defined by aligned hornblende and in part by a minor compositional layering. A stronger metamorphic fabric is concentrated at the borders of the up to 1000 m thick body. Internally the weak magmatic fabric is bend into occasional steep northeastward dipping reverse sense mylonitic shear zones. $\mathrm{Ca} .30-80 \mathrm{~cm}$ thick pegmatites are generally undeformed and cut the magmatic foliation and the shear zones, across which they are occasionally slightly bent. Assuming that the shear zones developed during thrusting of the KAC over the Indian plate, intrusion of the pegmatites is late to post kinematic. The pegmatites were cut in places by steep, N -dipping normal faults, possibly related to the backsliding of the KAC along the ISZ (Anczkiewicz et al., 2001).

In the Ranolia Khwar small scale recumbent folds of 1-2 m wavelength (northward plunging fold axis, eastward dipping axial planes) are frequent. Folding affects the gneisses beneath the Indus Suture.

The Allai and Natai Khwar provide a section from the core of the Besham antiform to its eastern limit. Descriptions for the regional geological units are given by several authors (e.g. Calkins et al., 1975; Tahirkheli, 1979). The regional structure is described by e.g. Treloar et al. (1989c, 1991a) and DiPietro et al. (1999). The following description of units mapped in the Allai and Natai valley is based on own observations and Allai Khwar Feasability study (Lahmeyer International, 1998).

Marble beds are exposed in Natai Khwar directly below the lower ISZ fault contact and mark, together with the inter-layered amphibolites, the northern boundary of the Indian Plate. To the $S$ of the lower ISZ fault contact black phyllites (IPH-Bph) intercalated with Limestone bands (IPH-Lst) are dipping ca. $70^{\circ}$ ESE. They are members of the Banna group (Treloar et al., 1989b). The amphibolites (IPH-Am) are strongly foliated (Fig. 2.6d) and strike NNW-SSE, dipping ca. $60^{\circ}$ towards NE. The contacts between amphibolites and marbles (IPH-Mb) is locally sheared, and contains calcareous schist layers (IPH-Cs), probably retrograde, finely intercalated amphibolites and marbles. To the west, a series of gneisses, undifferentiated as IPH-Gn is underlying the amphibolites and marbles. The gneisses comprise paragneisses (IPH-Gnp) and orthogneisses (IPHGng). The paragneisses are mainly banded with distinct feldspar and biotite rich bands. They are well exposed along the Pashto road, where they dip ca. $65^{\circ}$ E. The orthogneisses are variously strained. The mylonitic gneisses (IPH-Mga, IPH-Mgg and IPH-Mgt) are fine grained and contain amphibole, epidote and garnet. Migmatized gneisses (IPH-Mgm) are usually fine-grained and developed in thin bands at the contacts of porphyritic granites. The porphyritic granites (IPH-Grp) exhibit K-feldspars phenocrysts in medium to coarse grained groundmass of feldspar, quartz and biotite. The tourmaline granite is medium grained, equigranular and contains feldspar, quartz, muscovite, biotite and tourmaline with other accessory minerals. The granites are well foliated in shear zones. N-S trending bands of garnet mica schist (IPH-Gs) with a maximum thickness of about $50-60 \mathrm{~m}$, highly sheared and fractured metavolcanics (IPH-Mvg) and metavolcanics mainly of basaltic composition (IPH-Mv) are interlayered with gneisses in the Allai Khwar. Towards Besham, a N-S trending, 60-70 $m$ thick sequence of garnet bearing quartzites (IPH-Qtz) belongs to the metasedimentary portions of the Precambrian Besham gneisses.

### 2.2.2 Indus Suture Zone (ISZ)

Lithologies of the Indus Suture Zone were mapped along the Natai valley (E of Besham) and Ranolia valley (NE of Duber Bazar).

Along the Natai valley the ISZ is comprised of a ca. 2 km thick series of tectonically imbricated sheets, separated by a ca $40^{\circ} \mathrm{E}$ dipping lower fault contact from the Indian plate and by a ca. $70-80^{\circ}$ NE dipping upper fault from the Kohistan Arc Complex.


Fig. 2.6: a) E-verging folds beneath the Indus Suture, 300 m SW of Jijal at KKH; b) Amphibolitic pod in Indian Plate gneisses, 700 m E of Duber Bazar at KKH; c) Sample of Duber Diorite, a body intrusive in Indian Plate gneisses at Duber Bazar; d) Amphibolite with quartz band beneath lower ISZ-fault, Natai valley, along road, 2 km $W$ of Pastho; e) Pillow lava in Natai valley, along road, 1 km S of Pastho; f) Phyllites of the ISZ "mélange" at Kopu Khwar with N-side down folds, ca. 5,5 km NW of Duber Bazar; g) Metadiorite of the Metaplutonic Complex at the hanging wall of the Kohistan Fault at Pastho Kalai; h) Garnet and hornblende bearing dike in gabbros 500 m S of Pastho Kalai, along road.

Upsection, the series comprises greenstone (ISZ-Gst), metagraywackes (ISZ-Mug), blueschists (ISZ-Bs) and meta-volcanics with pillow lavas (ISZ-Muv) (Fig. 2.6e). The greenstone consists of hornblende, epidote and chlorite rich layers interlayered with thin plagioclase bands. In the field they appear as well foliated, dark green rocks with occasionally stretched plagioclase and minor quartz bands. They are followed by a series of green to brownish metagraywackes. In the upper part the metagraywackes contain glaucophane (Shah \& Majid, 1985). The metavolcanics have a dark-green to yellowish green colour on the fresh surface. They are mostly fine grained and contain plagioclase, hornblende, quartz, epidote and chlorite. Ellipsoidal pillows are a tens of centimetres to more than $1 / 2$ meter in diameter and are enclosed in a yellowish green matrix. Chilling effects at the surface are glassy in appearance. The succession of lava flows and pillows is followed by low grade, fine grained and layered metavolcanics. They were originally tuffs, ashes and volcanic agglomerates. Their colour depends on the relative amounts of quartz, plagioclase, epidote, chlorite and hornblende. Ripple marks and graded bedding in metasediments prove that the sequence is not overturned.

In the Ranolia Khwar and Kopu Khwar (NW of Duber Bazar) the ISZ is comprised of an up to 100 m thick sequence of garnet bearing schists (ISZ-Sch), black at the lower part where they are dominated by micas, whereas in the upper part they are light green, epidote-rich with stretched feldspar clasts. They show a strong schistosity parallel to the northward dipping $\left(40-70^{\circ}\right)$ lower and upper fault contacts.

### 2.2.3 Kohistan Arc Complex

The rocks of the Kohistan sequence exposed close to Pastho Kalai are hornblendites (KHA-Hbt) and garnet-bearing gabbros and metadiorite sheets (KHA-Gb) (Fig. 2.6g), intruded by garnet - hornblende bearing dikes, similar to the rocks of the lower Metaplutonic Complex (Fig. 2.6h). They are variably foliated (ca. $45^{\circ}$ towards E).

Within the lower part of the KAC, several lithologies were observed which are not limited to a specific group. They are: Small lenses of hornblendite ( $\mathrm{KH}-\mathrm{Hbl}$ ). Hornblende pegmatites (KH-HblP) occur sporadically in the area N of Kiru, where they yield various degrees of deformation, close to Kiru, where they are boudinaged along shear zones (Fig. 2.7a) and N of Dasu, where only little to no ductile deformation
affected them (Fig. 2.7b). Pegmatites (KH-Pegm) are frequent throughout the whole area and comprise:

$$
\text { quartz }+ \text { feldspar }+ \text { muscovite }+ \text { tourmaline } \pm \text { garnet } \pm \text { hornblende } \pm \text { biotite } .
$$

A high density of pegmatites is observed close to the granite sheets between Kiru and Dasu. Thickness ranges from centimeter to 20 m . A pegmatite cutting the main shear fabric south of Dasu yielded a K-Ar muscovite age of $66 \pm 3$ Ma (Treloar et al., 1989a). Ductile deformation within the pegmatites varies from nearly undeformed to intensely foliated (Fig. 2.7c,d).

In the SW to NE sinistral zone, bordering the mapped area to the NW, brecciated serpentinite (KH-Serp) occurs (Fig. 2.7e). Within this zone black phyllites and schists (KH-Phy) form "layers" up to 40 m thick, parallel to the deformation zone. They are observed in the northern part of Ranolia Khwar (Fig. 2.7f). The origin is not defined, but possible sources are the metasedimentary sequences of marine origin within the Kohistan Arc Complex, which were incorporated in the deformation zone.

### 2.2.3.1 Jijal Complex

The Jijal Complex is the lowest level of the Kohistan Arc Complex and is limited by the Indus Suture in the S and SW, by the serpentinitic zone in the NW and an intrusive contact to the Metaplutonic Complex in the NE (Fig. 2.1). It comprises essentially two units: an ultramafic dominated sequence and a mafic one. The ca. 3-4 km thick ultramafic sequence (KHJ-Um) is not differentiated on the map. The section has been described in details by Jan \& Howie (1981), Bard (1983), Miller et al. (1991) and Burg et al. (1998).

The lowest part is a peridotite unit composed of dunites and wherlites interlayered with cm - to m-thick layers of fine-grained pyroxenites (Fig. 2.8a,b). Coarse grained websterites successively dominates up-section the peridotites. Hornblende is sparsely intergrown with websterite, but becomes abundant further up-section, where layers of coarse grained hornblendite occur. In the upper part of the ultramafic-mafic section garnet-hornblende pyroxenites are dominant (see section Fig. 2.3).


Fig. 2.7: a) Hornblende Pegmatite stretched along shear zones in the Kiru amphibolites, 200 m S of Kiru; b) Undeformed Hornblende Pegmatites close to Duga, $10 \mathrm{~km} N$ of Dasu; c) Pegmatites in fine banded and strongly folded amphibolites ca. 1,7 km NNW of Mandraza; d) Pegmatites in amphibolite with variable degree of boudinage, $2,5 \mathrm{~km}$ NNE of Dasu; e) Brecciated serpentinite in the sinistral fault zone, 2 km SW of Duber Kale; f) Phyllites within the sinistral fault zone, close to Kichar Banda, 10 km NW of Duber Bazar.


Fig. 2.8: a) Dunite layers intercalated with pyroxenites, ca. 1,5 km E of Jijal along KKH; b) Coarse grained pyroxene layers with dunite and chromite lenses, ca. $6 \mathrm{~km} N$ of Duber Bazar; c) Hornblendite lenses in Granulitic gabbro, view from the southern Indus River bed ca. 2.5 km SW of Patan towards NW; d) Interpretative sketch of c), see also hornblendite lenses in enclosed section $A-A^{\prime} ;$ e) Layering of garnetite/Granulitic gabbro, ca. 1 km SW of Patan along KKH; f) Initial shear zone in Granulitic gabbro, 2 $k m$ SW of Patan; g) Feldspar segregation veins with aligned cm large garnets, ca. 500 $m S$ of Patan along Indus; h) Melt extraction in Garnet-Diorite indicating "normal" layering, ca. $60 \mathrm{~km} N$ of Patan in Chawa Khwar.

The gently northward dipping contact between garnet-hornblende pyroxenites and the overlying Granulitic gabbro (KHJ-GtGrlt) is interpreted as the exhumed petrological "Moho" (Burg et al., 1998). The contact locally displays interdigitations and coincides with last occurrence of dunite flames in websterites and hornblendites. Hornblendite (KHJ-Hbl) and garnetite bodies are found within the Granulitic gabbro. Two intrusion relations are observed in the field: the Granulitic gabbro intruding Hornblendite lenses and Hornblendites intruding the Granulitic gabbro. The Hornblendites, forming large enclaves and bodies with sharp boundaries (Fig. 2.8c,d) are intrusions that may have stemmed from the underlying mantle (Burg et al., 1998). Some hornblendite enclaves were intruded by the Granulitic gabbro and associated pegmatites, indicating that the Granulitic gabbro is intrusive into the hornblende-rich upper mantle.

The Granulitic gabbro, divided by Yamamoto \& Yoshino (1998) into a two-pyroxene granulite and a garnet-clinopyroxene granulite is characterized by:

$$
\text { (orthopyroxene) }+ \text { clinopyroxene }+ \text { plagioclase }+ \text { garnet }+ \text { hornblende } \pm \text { quartz } \pm \text { ilmenite } \pm \text { rutile. }
$$

Magmatic textures are locally preserved. Plagioclase is the dominant mineral in whiter layers (up to several tens of centimetres thick). The darker, mafic layers are pyroxenerich (Fig. 2.8e).

The two-pyroxene granulite is locally replaced by coarse-grained garnet-clinopyroxene granulite (Yamamoto \& Yoshino, 1998). Large garnets (up to a few cm in diameter) are concentrated along feldspar segregation veins (Fig. 2.8g). Foliation in the Granulitic gabbro dips generally towards NE (Fig. 2.8h). Ductile deformation within the Granulitic gabbro is expressed by anastomosing shear zones that show a dominantly top to the SW shear sense (Fig. 2.8f).

### 2.2.3.2 Metaplutonic Complex

The Metaplutonic Complex forms a thick pile of imbricated calc-alkaline laccoliths variably deformed in amphibolite facies conditions (Tahirkheli et al., 1979; Bard, 1983; Treloar et al., 1990). Remnants of the Tethyan oceanic crust are possibly preserved in enclaves. Field observations show that the definition of the Kamila amphibolite belt as a single unit (Jan, 1988) is inadequate. A new separation into 3 different zones is proposed, from S to N , i.e. bottom to top: a) Sheared gabbro-diorites (Patan region) comprising retrograde metagabbros with locally preserved igneous layering and
showing a heterogeneously distributed deformation forming an anastomosing pattern of shear zones; b) Kiru sequence consisting of more than hundred meters thick imbricate intrusions of gabbroic to dioritic composition, intensely deformed along East-West trending shear zones and c) Kamila sequence, an amphibolitised sequence composed of gabbroic, dioritic, granitic and tonalitic rocks intruded in a crustal sequence, whose remnants are represented by fine layered amphibolites with epidote-rich lenses, enclaves of carbonates and quartzite bands.

### 2.2.3.2.1 Sheared gabbro-diorites (Patan region)

The Sheared gabbro-diorites, mapped in the north-eastern Patan area are directly lying on the Granulitic gabbro. The Garnet-Diorite (KHA-GtDr) (Fig. 2.4) is a layered, coarse grained garnet-rich meta-diorite well exposed on the W side of Chawa Khwar. The layering is constituted by garnet-rich, plagioclase-rich and hornblende-rich cm to m thick layers. Quartz and epidote complete the assemblages. The sharp contact to the Granulitic gabbro dips in the S with ca. $10^{\circ}$ towards north and steepens to ca. $25^{\circ}$ northward. The contact plane is irregular, forming bulges in a meter to tens of meter scale in the Granulitic gabbro. Up to a few meter sized, elongated enclaves of Granulitic gabbro occur above the contact indicating that the Garnet Diorite is intrusive into the Granulitic gabbro. Finely layered, tens of meter thick meta-diorite sills (KHA-dDr) are parallel to the north-dipping contact (Fig. 2.9a). The relation of the Garnet Diorite to the Sarangar gabbro (KHA-GbS) and the dioritic/gabbroic units to the N could not be determined, because the Garnet Diorite is cut by faults to the N and the E . The Sarangar gabbro, a coarse grained, locally layered gabbro (Fig. 2.9b,c) make up the outcrops E of Patan Banil Khwar and Gulbagh Khwar. The mineralogical composition within the main metamorphic fabric is dominantly:

$$
\text { hornblende }+ \text { clinopyroxene }+ \text { plagioclase } \pm \text { quartz } \pm \text { rutile } \pm \text { garnet } \pm \text { epidote. }
$$

Compositional layering is characterised by a varying amounts of plagioclase. The gabbro shows spectacular shear zones anastomosing around coarse grained, weakly deformed pods with well preserved plagioclase-clinopyroxene assemblages. The contact to the Granulitic gabbro partly sheared and bended by anastomosing shear zones, is observed at the S side of the Indus close to Patan.


Fig. 2.9: a) Garnet Diorite (bright) with intrusive dioritic sills (dark), ca. $4 \mathrm{~km} N$ of Patan in Chawa Khwar; b) Sarangar gabbro, sheared along feldspar vein, ca. 1.7 km E of Patan; c) Sarangar gabbro with compositional layering, ca. 2 km E of Patan along KKH; d) Granulitic gabbro xenolith enclaved in sheared Sarangar gabbro, ca. 2,5 km $N$ of Patan in Banil Khwar; e) Magmatic breccia with hornblendite (old), gabbro and diorite (young), 5 km ENE of Patan along KKH; f) At lower side sheared hornblende dike cutting diorites/hornblendites, ca. 300 m NE magmatic breccia, along $K K H$; g) Top to the S sheared Kayal Gabbro; h) Layered diorite (KHA-Drb), WSW of Kayal Indus confluence at KKH.

The Sarangar gabbro is impregnating the Granulitic gabbro along a dense, but diffuse vein network, forming a $1-2 \mathrm{~m}$ wide transition zones. Above the contact small hornblende lenses (few m thick) occur. N of Patan a Granulitic gabbro pod is sheared within the Sarangar gabbro (Fig. 2.9d). The gabbro is intruded by hornblendite veins and garnet-bearing feldspar veins which are locally folded. A meta-carbonate enclave was incorporated in the gabbro (at the KKH ca. 500 m S of its northern contact), providing evidence that the gabbro was intrusive into metasediments under mid-crustal conditions of ca. 0.8-1.1 GPa (Yoshino et al., 1998).

In the Chor Nala valley the Sarangar gabbro is underlain by a coarse grained, well foliated Hornblende Gabbro (KHA-cGbHbl) with sporadic cm large hornblende crystals. The assemblage contains also garnet and mica as retrograde metamorphic mineral. Thin, fine grained basic dikes cut the gabbro. The diffuse contact of this gabbro to the Granulitic gabbro dips gently N to NE and is characterised by an up to 3 m wide transition zone, where the gabbro impregnates the Granulitic gabbro along a network of small veins. In the Sarangar gabbro two ca. 50 m thick sill-like bodies of the same rock type are observed in the Chor Nala. The upper contact of the Sarangar gabbro is comprised in the eastern part (Indus valley) of a magmatic breccia (KHA-GbDrHbl) (Fig. 2.9e), where hornblende-rich rocks were intruded and brecciated by gabbros (Sarangar gabbro?) and diorites. N of Patan the Sarangar gabbro is bounded at its upper contact by finely layered and deformed diorites/hornblendites (KHA-DrHbl). Above the sharp contact up to ca. 1 m small enclaves of Sarangar gabbro were incorporated in the strongly SW-ward sheared diorites/hornblendites at its bottom parts. This could represent a sheared equivalent to the above mentioned magmatic breccia. To the top diorite layers are coarser grained, showing a plagioclase, hornblende, quartz, epidote and minor chlorite assemblage. They are folded and sheared along conjugated shear zones.

The diorites/hornblendites were intruded by hornblendite dikes and pegmatites. The hornblendite dikes are locally sheared but cut the foliation in the diorites in other places (Fig. 2.9f), suggesting a syn-shearing intrusion age. At the Indus section the Kayal Gabbro (KHA-GbKa) is intruded by several diorites/hornblendites sills. The Kayal Gabbro is coarse grained and similar to the Sarangar gabbro, but without layering. A
dense pattern of anastomosing shear zones deforms the gabbro (Fig. 2.9 g ) and the contacts. Both sides of the lower Kayal valley along the KKH screen diorite bodies with a hornblende, plagioclase, quartz and accessorily minerals paragenesis. The basal diorites (KHA-Drb) are layered (Fig. 2.9h) and strongly sheared. Mylonitic anastomosing shear zones developed around lense shaped pods. The foliation is bent into the mylonite zones showing SW-ward shear. The bottom part of diorite was intruded by 1) the Kayal gabbro, as indicated by up to 10 m large gabbro enclaves in the diorite, by 2) garnet bearing veins and later 3) by hornblende - plagioclase - garnet pegmatites. They are distinguished from the upper diorite (KHA-Drt) comprised of an interlayered sequence of coarse grained and fine grained diorites intruded by hornblendite dikes. Some steeply northward dipping shear zones have normal sense (top to N ) movement. The ca. $30^{\circ}$ northward dipping upper contact is irregular and in places sheared in normal sense. Where un-sheared, small enclaves (a few tens of cm ) of diorite are within the lowest meter of a coarse grained gabbro which is part of coarse grained interlayered gabbroic to dioritic intrusions (KHA-GbDr). The intrusions are interlayered in various thicknesses, ranging from a few m to hundred m and are distinguished by their colour contrast between dark and bright in the field, where the gabbros often have a higher amount of hornblende and therefore are darker than the diorites. This sequence is sporadically intruded by garnet bearing veins. At the top the gabbros interfinger with the Kiru sequence in a dike like manner forming an irregular contact.

### 2.2.3.2.2 Kiru sequence

Several steeply northward dipping mylonitic shear zones (with top to N movement) are observed at the lower contact to the sheared gabbro-diorites. The Kiru sequence consists of imbricate "amphibolites" intensely deformed along East-West trending shear zones. The lowest part is constituted by the Kiru Amphibolites (KHA-AmKi) comprised of up to hundred meters thick laccolithic bodies of dominantly fine grained rocks. The compositions range from gabbroic with hornblende, plagioclase, garnet, epidote and quartz to dioritic with hornblende, plagioclase, quartz, garnet and epidote. In places black, hornblende dominated layers of metaplutonic origin occur. Locally coarse grained rocks (Fig. 2.10a) are preserved. Lithologic boundaries are parallel to the main foliation and are generally dipping north. Steeply northward dipping to sub-horizontal mylonitic shear zones enclose blocks of variable size ( $<50 \mathrm{~m}$ ).


Fig. 2.10: a) Fine layered Kiru Amphibolite including a small dioritic sill (tip of pen), at Kiru; b) Hornblende Diorite intruding diorites/hornblendites of the Kiru Amphibolites, $2 \mathrm{~km} S$ of Kiru at KKH; c) Steep $N$-side down shear zone in Kiru Amphibolites, $4 \mathrm{~km} S$ of Kiru at KKH; d) Small hornblende lenses and basic dikes cutting a gabbro of the KHA-mDrGb unit, 1,5 km S of Kiru at KKH; e) Carbonate lense in gabbro (boulder) from the area $500 \mathrm{~m} S$ of Kando Banda; f) Interlayered diorites (bright) and gabbros (dark), $4,5 \mathrm{~km} N$ of Kiru.

A coarse to medium grained hornblende-rich diorite (KHA-HblDr) is in sharp contact and contains enclaves of Kiru amphibolites (Fig. 2.10b). It is composed of plagioclase, greenish hornblende, quartz and epidote, chlorite and titanite. The metamorphic fabric has overprinted the magmatic fabric and foliation is defined by the orientation of metamorphic hornblende. The contact at its base dips irregularly $40^{\circ}$ northwards. Locally anastomosed shear zones are spaced by $50-100 \mathrm{~m}$. The ca. $50^{\circ} \mathrm{N}$ dipping upper contact is more regular and parallel to main foliation. Just to the S of Kiru again, subvertical amphibolites with pegmatites ( $\mathrm{qtz}+\mathrm{fsp}+\mathrm{msc}$ ) and hornblende pegmatites appear. Partial remelting sporadically produced thin plagioclase veins within mylonitic zones. In the Kiru amphibolites top to the $S$ shear zones are dominant. However, steep N -dipping, lower amphibolite facies (hornblende + plagioclase + epidote + quartz), shear zones show a top to N relative shear sense (Fig. 2.10c).

A series of (meta-) gabbro and (meta-) diorite bodies N of the Kiru amphibolites is summarized in a mixed unit (KHA-mDrGb). Hornblendite and basic veins intersect them sporadically (Fig. 2.10d). The good outcrop situation along the KKH permitted mapping several of these gabbros and diorites. Within the gabbros, lenses of metacarbonates, which were not observed in the diorites, occur infrequently. The interlayered gabbros and diorites (Fig. 2.10f) were heterogeneously affected by amphibolitisation and deformation, some of them having preserved magmatic textures. Most of these imbricated bodies are hornblende-rich diorites and meta-diorites (KHADrmet). Towards the N the number of pegmatite and small granitic sheets increases.

The contact to the Mandraza amphibolites (KHA-AmMa) is set where dark, fine grained and strongly foliated amphibolites become dominant. They are cut by a swarm of SWdipping, up to a few meters thick pegmatites. The Mandraza amphibolites are comprised of a series of coarse to medium grained meta-gabbros, hornblende-rich meta-diorites and granites of variable thickness not exceeding 100 m . They are interlayered with fine grained, hornblende dominated, strongly foliated dark hornblende-plagioclase-quartzepidote rocks. Layering is due to alternating changes in hornblende and plagioclase contents. Occasionally, elongated ( $<20 \mathrm{~cm}$ ) green to yellowish epidote lenses occur within the dark layers. Top to the SW shearing is expressed by flat to steep N dipping shear zones around $50-100 \mathrm{~m}$ wide lenses with asymmetrical shape (Fig. 2.11a,b). In the

Sangri Khwar, N of Mandraza, a unit of more massive, coarse grained gabbros and diorites appears. It probably is cut in its eastern part by the big granite body (the largest one on the map), as indicated by the bending of the ca. $60^{\circ} \mathrm{N}$ dipping foliation towards S. The uppermost metadiorite contains schlieren of fine grained, well layered, epidoterich Dasu Amphibolites (Fig. 2.11c) within a meter thick zone below the contact.

### 2.2.3.2.3 Kamila sequence

The Kamila sequence is composed of amphibolitised gabbros, diorites, granodiorites granites and tonalites intruded in fine layered amphibolites with epidote-rich lenses, enclaves of carbonates and quartzite bands, interpreted to be remnants of a upper crustal sequence. The Kamila sequence is, compared to the Kiru sequence, characterized by the increased occurrence of granitic rocks. A nearly 1 km thick granite body (KHA-Gr) overlies the Mandraza amphibolites and bends at its base the steep N -dipping foliation of the amphibolites. These coarse grained granite is composed of:

$$
\text { quartz }+ \text { plagioclase }+ \text { muscovite }+ \text { garnet } \pm \text { epidote } \pm \text { pyrite } \pm \text { chlorite } \pm \text { apatite } \pm \text { zircon. }
$$

Below these mentioned granite body only metagranite sheets, up to 1 m thick, are intercalated with the Mandraza amphibolites. There the contact between amphibolite and granite is sharp and foliation/layering parallel. Within the bottom part of the big granite body amphibolite pods (fine grained dark meta-diorite) were stretched. Towards the upper part of the big granite body xenoliths of greenish, fine grained, epidote rich amphibolites of possibly metasedimentary origin are more frequent. Coarse grained, muscovite-rich pegmatites cut the generally ca. $40^{\circ} \mathrm{N}$-dipping foliation. In places they are elongated and foliated parallel to the main foliation. The granite becomes strongly foliated near the contacts. N-dipping, less than 50 cm wide shear zones with a top to SW sense are spaced by ca. 20 to 60 m . A second ca. 500 m thick, E-W trending massive granite sheet occurs ca. 500 m more to the north. Between both sheets, interlayering of fine grained amphibolites (meta-gabbros and a greenish, fine layered variety of possible metavolcanic origin), granites and coarse grained pegmatites sheets make up to $50 \%$ of the outcrop (Fig. 2.11d). Bigger ( $>1 \mathrm{~m}$ in thickness) pegmatites cut the foliation of the amphibolites, but others are stretched, boudinaged and sheared, indicating syn- to postdeformation emplacement. Towards Kamila village the foliation steepens northward and the relative amount of granites and pegmatites decreases.


Fig. 2.11: a, b) Lense shaped block of gabbro enclosed in mylonitic shear zones, view to E side of Indus valley ca. $6 \mathrm{~km} N$ of Dasu; c) Contact between Dasu Amphibolites(dark) and metagabbro, cut by pegmatite, $2 \mathrm{~km} N$ of Mandraza; d) Granitic sheets in Dasu Amphibolites, location between the two granite bodies; e) Granite sills in a series of diorite and gabbro sheets, $N$ of Kamila; f) Basic dikes cutting a diorite of the Dasu Amphibolites, 800 m N of Dasu.

Several thin (up to a few m) gabbro sheets getting more frequent, build up a nearly 1 km wide, E-W extended zone of gabbroic intrusions (KHA-GbmetDa) at Kamila.

The Dasu area is dominated by a strongly layered amphibolitic sequence (KHA-AmDa) composed of rocks with varying hornblende, plagioclase, epidote and quartz contents. Sheet-like gabbros and granites (Fig. 2.11e) contain smaller enclaves of dark, fine
grained hornblende- and epidote-rich amphibolites. Basic and acidic dikes crosscut the amphibolitic sequence N of Dasu (Fig. 2.11f). The layer parallel foliation dips ca. $70^{\circ}$ toward NNE in the vicinity of the northern metagabbro, which is cut by a more gently N -dipping granite on the northern flank of Dongai Gah. The metagabbro and the granite are intrusive into a series of dominantly fine grained, well layered to laminated amphibolites (KHA-AmO). They show green epidote-rich lenses and layers (Fig. 2.12a), most likely produced by a calcite rich (sedimentary?) source, sliced in dioritic magmas.


Fig. 2.12: a) Boudinaged epidote rich lenses of sedimentary origin in amphibolites (light grey left of hammer), $3 \mathrm{~km} N$ of Dasu; b) Laminated epidote- and hornblende-rich layers with weak graded bedding (younging towards top), ca. 3, $1 \mathrm{~km} N$ of Dasu; c) Tonalite intrusion into amphibolites, $4 \mathrm{~km} N$ of Dasu, view towards NE; d) Carbonate lense in diorites along the KKH, $6 \mathrm{~km} N$ of Dasu.

Occasionally the few cm thick metasedimentary layers show graded bedding in originally volcano-detritic sediments (Fig. 2.12b). Several tonalite bodies (KHA-Tn) occur in the center parts of the sequence along the Indus valley (Fig. 2.12c). The
contacts are generally sharp and sheared in places. The amphibolites are overlain by a massive, coarse grained diorite (KHA-DrmetN). Metamorphic foliation is heterogeneously distributed, steep northward dipping and sporadically bent by N dipping, less than 1 m thick shear zones with a top to S sense. Meta-carbonates (KHAmC ) occur as m -sized enclaves in the diorite along the Indus valley (Fig. 2.12d). Vertical to steeply N-dipping sheet-like Quartzites bands (KHA-Qtz) of 10 to 20 m thickness occur ca. 6 km N of Dasu at the Indus valley, indicating a involvement of metasediments in the dominantly plutonic sequence of the Metaplutonic Complex.

### 2.2.3.3 Chilas Complex

The bulk rock type is noritic gabbro, locally layered with subordinate diorites. Gabbronorites and diorites ( $\mathrm{CH}-\mathrm{Dr} / \mathrm{Nr}$ ) and volcanodetritic metasediments ( $\mathrm{CH}-\mathrm{VcS}$ ) were mapped along the Kandiah River (Fig. 2.13). The coarse to medium grained norites show a weak foliation. The gabbronorite is composed of:

$$
\text { clinopyroxene }+ \text { orthopyroxene }+ \text { amphibole }+ \text { plagioclase } \pm \text { quartz } \pm \text { biotite } \pm \text { oxides. }
$$

The magmatic fabric is preserved in parts. In places the gabbronorites are cut by undeformed feldspar dikes. In the western part of the Kandiah valley the gabbronorites unit contain also banded epidote-amphibolites. The size of the epidote-amphibolites bodies could not be determined, but most likely they are xenoliths of variable extend. A dioritic intrusions formed a magmatic breccia close to Zambil. Close to the eastern limit of the metasediments, peridotites outcrop along the road. Most likely they come from nearby lenses of ultra-mafic-anorthosite associations. The metasediments in the central part of the Kandiah valley have a northward dipping bedding. They are mm to cm thin muscovite rich layers (Fig. 2.14a) with quartz and calcite lenses folded in small scale around WNW-ESE trending fold axis.


Fig. 2.13: Geological map of the Kandiah valley, reduced scale (original scale map is enclosed).

### 2.3 Mapped faults

Several distinct fault zones are recognized with different tectonic importance. The Indus Suture is the most important. It divides the crystalline shield rocks of the Indian Plate from the mafic and ultramafic rocks of the KAC. The Indus Suture close to Jijal is marked by a ca $35^{\circ}$ northeastward dipping fault contact. Brittle deformation produced a serpentinite zone.


Fig. 2.14: a) Metasediments, mm to cm thin layered muscovite and quartz rich bands; b) Gentle slopes $N$ of Patan made of fractured Granulitic gabbro and Sarangar gabbro indicating brittle activity; c) Steep dipping gouge zone, 3 km S of Kiru; d) Small scale, SW verging folding in gabbro; e) Folded granite sheet in Mandraza Amphibolites, ca. 1 km S of Mandraza.

Towards the W of Duber Khwar a ca. 100 m thick sequence of schists constitutes the ISZ with mesoscopic folds and kinkbands indicating a N -side down movement during ductile-brittle transition (Fig. 2.6f). The fault contacts between the Kohistan ultramafic rocks, the schists and the Indian Plate gneisses developed gouge zones with a N -side down movement indicated by striations. Immediately below the lower fault contact the Indian gneisses yield a northward plunging stretching lineation and a top to the S shear sense. Within the ISZ at Natai Khwar occasionally interlayered marble bands are sheared towards the SW. However, extensional structures like NE down facing folds and shear bands in schists in the metavolcanics and normal faults are dominant. The upper fault contact is dominantly brittle over a maximum 30 m wide zone. The associated small fault strands cut the metavolcanics and incorporate blocks of KAC rocks (DiPietro et al., 2000).

A sinistral NE-SW trending fault also developed a serpentinite zone. It limits the mapped area to the NW and is traced on satellite images from the Indus River N of Dasu to N of Alpurai in the SW (Fig. 2.15). The relation to the Indus Suture remains open, but most likely the fault joins the Indus Suture N of Alpurai. The SW-ward continuation is a similar trending fault described by Anczkiewicz et al. (1998) in the Alpurai area, which acted first as a thrust and later as a sinistral strike slip fault. If this is the case, the NE-SW trending fault would cut into the ISZ and offset it.

Faulting in the Patan area is expressed by steep N-S and E-W trending faults. In the Patan Banil Khwar the contact between the Granulitic gabbro and the Sarangar gabbro is such a fault. In Patan and in the Gulbagh Khwar the contact between the Granulitic gabbro and the Sarangar gabbro is an intrusive one, excluding that the above mentioned fault is major tectonic contact between them. The widened Indus valley at Patan (Fig. 2.14b) shows fractured rocks and gouges indicative for active faulting.

Two major E-W striking faults are present between Patan and Dasu. The one ca. 3 km S of Kiru is a sinistral fault with a thrust component and developed a 2 m wide gouge zone (Fig. 2.14c). It joins the NE-SW trending fault zone at Duber, where a thin serpentinized zone was developed. The fault is steeply dipping north and offsets in its eastern part hornblende diorite against Kiru Amphibolites. In the western part it cuts into gabbros and diorites and is observed as a narrow brecciated zone in Chawa Khwar.

The fault 5 km S of Dasu is inferred from satellite images but could not be followed in outcrops (Fig. 2.15). Following mainly river beds, the actual fault contact is covered by quaternary sediments. Only at the SE corner, where the Indus turns towards W, a 1 to 2 $m$ wide zone, of brecciated amphibolites is observed. The displacement and the sense of movement could not be determined. The fault limits to a certain extent the granite bodies to the N (southern granite) and to the S (northern granite).


Large scale folding affected the Northern Indian Plate and produced the N-S trending Besham antiform. Folds observed within the KAC are generally small to mesoscale tight to isoclinal S to SW verging folds (Fig. 2.14d). Fold axis are dominantly E-W trending. The Kiru amphibolites show in places mesoscale S vergent, downward facing folds, indicating N -side down extensional movement. In strongly layered and foliated units, intrafolial folds have dominant SW- vergence. The granite and amphibolite sheets in the northern Kiru sequence are occasionally folded with a SW- vergence, like 3 km N of Kiru (Fig. 2.14e, Fig. 2.3).

## 3. PETROLOGICAL AND GEOCHEMICAL ASPECTS OF THE LOWER KOHISTAN COMPLEX

This chapter describes petrological and geochemical aspects and is intended to provide data and interpretations of selected lithologies. A more detailed investigation will require, despite a dense sampling grid covering the area, petrological/geochemical analyses of all various metaplutonics and metavolcanics. The following information is additional to findings previously published by several authors.

### 3.1 Petrological characteristic of sheared and undeformed gabbros, a COMPARISON

Investigations are concentrated on the petrological evolution in the anastomosing shear zones of the gabbros of the Patan region as well as on the pressure and temperature conditions of shearing. The area of interest is comprised of two different gabbros: 1) the Granulitic gabbro of the upper Jijal complex and 2) the Sarangar gabbro (NE of Patan). These samples are described in their successive assemblages and mineral compositions of each undeformed and mylonitic zones. One metadiorite of the Kiru amphibolites is then compared to the gabbros.

### 3.1.1 Methods

Mineral compositions were analysed using a Cameca SX-50 electron microprobe (University of Bern and ETH Zurich), equipped with five crystal spectrometers. Samples were coated with carbon. Operating parameters include an acceleration potential of 15 kV , a beam current of 20 nA and a beam size of $\sim 1 \mu \mathrm{~m}$. Probe analyses were recalculated using the program "Norm" by P. Ulmer (ETH-IMP). PT conditions were calculated using the program "Thermobarometry" (Spear \& Kohn 1995).

### 3.1.2 Samples

Five samples were selected in order to represent a comparison between undeformed and mylonitic gabbros of the lower arc, which were affected by anastomosing shear zones and a metadiorite of mid crustal level: Two samples (13/01/01 and 94/04/01) from the Granulitic gabbro, two samples (16/01/01 and 90/02/01) from the Sarangar gabbro and one metadiorite sample (46/02/01) of the Kiru amphibolites (Fig. 3.1). The Granulitic
and Sarangar gabbros show strong strain gradients from mylonitic to undeformed to weakly deformed rocks. Analyses were done for both deformation facies.


Fig. 3.1: Location map of samples used for electron microprobe analysis. Sample 13/01/01 and 94/04/01 are from the Granulitic gabbro collected along the KKH ca. 500 $m$ SW of Patan. Sample 16/01/01 and 90/02/01 were collected in the Sarangar gabbro ca. 2 km E of Patan along the KKH and at the southern riverside. Sample 46/02/01 was collected ca. 300 m W of the metal bridge (KKH) at Kiru. GPS coordinates are given in appendix 3.1.

### 3.1.2.1 Granulitic Gabbro

The Granulitic gabbro is also labelled garnet granulite (e.g. Jan \& Howie, 1981). The detailed petrography of the Granulitic gabbro has been presented by Jan \& Howie (1981), Bard (1983), Miller et al. (1991), Yamamoto (1993), Yamamoto \& Yoshino (1998) and Ringuette et al. (1999). It is divided into two rock types: a two-pyroxene granulite and a garnet-clinopyroxene granulite (Yamamoto \& Yoshino, 1998). The analysed samples represent the garnet-clinopyroxene granulite.

The mineral content in the undeformed Granulitic gabbro is plagioclase (affected by saussuritization), clinopyroxene, garnet, amphibole, minor or accessory quartz, rutile, epidote, chlorite, opaque oxides and biotite. The mid to fine grained heteroblastic rock shows a granoblastic texture (Fig. 3.2). The original magmatic assemblage is partially preserved and was overprinted successively by granulite, amphibolite, epidoteamphibolite assemblages (Ringuette et al., 1999).

The magmatic assemblage is represented by:

$$
\text { garnet }+ \text { clinopyroxene }+ \text { plagioclase }+ \text { quartz } \pm \text { ilmenite } .
$$

The existence of magmatic garnet is supported by the fact that inclusion poor garnet occurs in textural equilibrium with other cumulate phases such as clinopyroxene and plagioclase (Ringuette et al., 1999). Most garnets and quartz grains in the analysed samples are possibly metamorphic phases. This is supported by the interstitial growth of garnet and quartz (Fig. 3.3a). The garnets have no pronounced zoning indicating a relatively fast growth under isochemical conditions. Crystallisation took place under granulite facies conditions. It is likely that, subsequently to crystallisation the enstatite component in clinopyroxene reacted with plagioclase to garnet and quartz. In this case the plagioclase composition would change to an albite richer composition, which is observed in the undeformed Granulitic gabbro with albite rich plagioclase $\left(\mathrm{An}_{25}\right)$.

In the mylonitic part of the Granulitic gabbro the magmatic assemblage is obscured.
The granulite facies assemblage is characterised by:

$$
\text { clinopyroxene }+ \text { garnet }+ \text { plagioclase }+ \text { quartz } \pm \text { rutile } .
$$

Clinopyroxene is hypidioblastic with lobate to symplectitic contacts to plagioclase. The sub- and anhedral garnets are poikiloblastic intergrown with quartz and are in equilibrium with plagioclase (Fig. 3.3a). Quartz is partly dynamically recrystallised (Fig. 3.3b). During amphibolite facies conditions clinopyroxene was overgrown by brown tschermakitic hornblende (Fig. 3.3c). The replacement of clinopyroxene by amphibole is domain-like in the undeformed parts.


Fig. 3.2: Granulitic gabbro samples 13/01/01 and 94/04/01. The positions of analysed sections are for each sample in the mylonitic part and in the less deformed part.

The latest paragenesis, indicative for epidote-amphibolite facies conditions is:

$$
\text { greenish hornblende }+ \text { epidote } \pm \text { clinozoisite } \pm \text { chlorite. }
$$

Beside a strong decrease in grain size, the decrease in amount of clinopyroxene and the increase of amphibole and quartz are the most striking phenomena when focusing on the mylonites. The garnet is porphyroclastic in the very fine grained matrix $(\sim 0.3 \mathrm{~mm}$, mainly tschermakitic amphibole), partly rotated and fractured with new grown amphibole in fractures. Clinopyroxene fragments are incorporated in the matrix. Recrystallised quartz in the pressure shadows of garnets as well as the amphibole crystals oriented parallel to the foliation point to a high shear strain (Fig. 3.3d). Within the mylonites an amphibolite facies mineral assemblage overprints the granulite facies mineral assemblage. The foliation defined by hornblende crystals and few poikiloblastic garnets implies, that mylonites formed under granulite to amphibolite facies conditions.


Fig. 3.3: Granulitic gabbro microphotographs: a) 94/04/01 P1, idiomorph garnet in contact to plagioclase and quartz; b) 94/04/01 P3, dynamically recrystallised quartz in less deformed part; c) 13/01/01 P2; overgrowth of pyroxene by hornblende; d) quartz in pressure shadow of garnet clast floating in a matrix of hornblende, pyroxene fragments and plagioclase. + crossed nicols.

### 3.1.2.2 Sarangar Gabbro

Previous petrological descriptions of the Sarangar gabbro agree that pods surrounded by anastomosing shear zones preserved pyroxene-bearing metagabbro (Treloar et al., 1990) called pyroxene granulite by Yoshino et al. (1998). Both samples represent the transition of the relatively undeformed pods to the surrounding mylonites (Fig. 3.4). The dominant composition in the Sarangar gabbro is amphibole (mainly pargasite and tschermakite), pyroxene, plagioclase, rutile, ilmenite, epidote and quartz. In few cm long sections from undeformed spots to the centre of the shear zones the changes are significant in term of mineral assemblages, grain size and orientation of minerals.

The magmatic assemblage, crystallised under granulite facies conditions and only preserved in domains of the undeformed parts (Yamamoto, 1993), is:

$$
\text { pyroxene }(o p x+c p x)+\text { plagioclase }+ \text { epidote } \pm \text { ilmenite } \pm \text { quartz } .
$$

The hypidioblastic orthopyroxene is progressively replaced by clinopyroxene. The sporadic occurrence of poikiloblastic garnet may be due to the reaction mentioned by Yoshino et al. (1998):
orthopyroxene + plagioclase $=$ garnet + quartz + clinopyroxene.
The reported granulite facies assemblage (Yoshino et al., 1998) including garnet and rutile is overprinted by the amphibolite facies assemblage:
hornblende + plagioclase + quartz.

Interstitial brown to green hornblende, associated with quartz, occurs as rims between pyroxene and plagioclase (Fig. 3.5a). Yoshino et al. (1998) concluded that hornblende and quartz is probably produced by the reaction:

$$
\text { pyroxene }+ \text { plagioclase }+\mathrm{H}_{2} \mathrm{O}=\text { hornblende }+ \text { quartz. }
$$

Plagioclase and pyroxene are successively replaced by quartz and hornblende (Fig. 3.5b). Plagioclase shows an increase of symplectitic structures with pyroxene closer to the mylonite.


Fig. 3.4: Sarangar gabbro samples 16/01/01 and 90/02/01. The positions of analysed sections are for each sample in the mylonitic part and in the less deformed part.

Within the shear zone, no pyroxene was found. Garnet is abundant showing poikiloblastic growth (Fig. 3.5c) and is generally associated with rutile. Some garnets in the mylonites have helicitic rutile inclusions, suggesting a synkinematic growth of garnet (Fig. 3.5d). The amphiboles underline the fabric and are cut by small quartz veins. Amphibole composition (tschermakitic) is homogenous.

Ilmenite is sparsely preserved as inclusions in brownish hornblende. This points to the retrograde reaction (Bohlen \& Liota, 1986):

$$
\text { ilmenite }+ \text { plagioclase }+ \text { quartz }=\text { rutile }+ \text { garnet } .
$$

The macroscopic porphyroclasts ( $\sim 2 \mathrm{~mm}$ ) are plagioclase and quartz aggregates. The dominant, amphibolite facies assemblage is:

$$
\text { hornblende }+ \text { plagioclase }+ \text { garnet } \pm \text { rutile } \pm \text { quartz } .
$$

Chlorite and muscovite overgrow hornblende as late phases under greenschist facies conditions.


Fig. 3.5: Sarangar gabbro microphotographs: a) 90/02/01 P4, fractured orthopyroxene with hornblende rim indicates retrogression; b) 16/01/01 P3, orthopyroxene replaced successively by hornblende in the undeformed zone; c) 16/01/01 P1, typically poikiloblastic garnet in mylonitic zone with plagioclase, minor quartz and hornblende; d) 90/02/01 P1, new grown garnet in equilibrium to hornblende, plagioclase and quartz, with rutile inclusions in the garnet. + crossed nicols; $\|$ parallel nicols.

### 3.1.2.3 Kiru Amphibolite (Metadiorite)

The Kiru amphibolite sample is a fine grained metadiorite characterised by a well developed foliation defined by hornblende. The alternating plagioclase and hornblende rich layers on the millimetre scale form to the macroscopic foliation (Fig. 3.6a). The garnet free assemblage equilibrated under amphibolite facies conditions is:

$$
\text { hornblende }+ \text { plagioclase }+ \text { quartz } \pm \text { biotite } \pm \text { epidote } \pm \text { titanite } .
$$

Titanite is rare (Fig. 3.6b). Biotite has corroded boundaries and zircons in its vicinity (Fig. 3.6c). The amphibolite facies assemblage was only little affected by retrogression as it is indicated by chlorite replacing sporadically hornblende (Fig. 3.6d).


Fig. 3.6: Kiru amphibolite: a) alternating plagioclase (black when affected by saussuritization) and hornblende rich layers parallel to foliation which is defined by a preferred orientation of hornblende crystals; b) titanite as accessory phase; c) biotite affected by retrogression showing corroded boundaries; d) hornblende retrograde overgrown by chlorite. + crossed nicols; || parallel nicols.

### 3.1.3 Mineral composition

Mineral compositions show differences between the two gabbros and between the mylonitic and less deformed parts.

### 3.1.3.1 Plagioclase

The plagioclase composition in the undeformed Granulitic gabbro is $\mathrm{An}_{25-47}$, clustered around a maximum of $\mathrm{An}_{40}$ for sample 13/01/01, and $\mathrm{An}_{43-55}$ for sample 94/04/01 (Fig. 3.7). Plagioclase grains are homogeneous. However, saussuritization may have changed the original composition. Plagioclase compositions in the mylonites are similar to those in the less deformed rock (Table 3.1).

Plagioclase composition in the less deformed Sarangar gabbro is $\mathrm{An}_{55-70}$. Only few grains show a lower anorthite component, around $\mathrm{An}_{25}$. Within the mylonite plagioclase of two different compositions is observed, one with $\mathrm{An}_{25-45}$ and one with $\mathrm{An}_{70-80}$ respectively. Sample 16/01/01 misses the latter, anorthite rich composition. The composition range between the less deformed parts and the mylonitic ones are not overlapping, pointing to a change of composition related to shearing and/or later alteration along the shear zones.

In the Kiru amphibolites plagioclase composition is in the range of $\mathrm{An}_{17-46}$, similar to the plagioclase compositions measured in the Sarangar gabbro mylonite.


Fig. 3.7: Plagioclase compositions of granulitic gabbro, Sarangar gabbro and Kiru amphibolite.


Table 3.1: Representative (averaged, $n=n r$. of analyses) chemical composition of plagioclase. Single analyses are given in appendix 3.1.

Normalization on the basis of 5 CATIONS and 16 CHARGES

### 3.1.3.2 Garnet

Garnet compositions are in the less deformed Granulitic gabbro $\mathrm{Alm}_{40-53} \mathrm{Prp}_{23-42} \mathrm{Grs}_{15}$ ${ }_{29} \mathrm{Adr}_{01-03}$. In the mylonitic part the composition range is $\mathrm{Alm}_{42-53} \operatorname{Prp}_{28-42} \mathrm{Grs}_{12-28} \mathrm{Adr}_{02-04}$ (Fig. 3.8). Garnet composition does not change between the shear zones and the less deformed rock (Table 3.2).

In the Sarangar gabbro garnet occurs only sporadically in the less deformed parts. An analysis with $\mathrm{Alm}_{48-51} \operatorname{Prp}_{30} \mathrm{Grs}_{18-21}$ (Yoshino et al., 1998; Sample P5) is similar to the garnets in the Granulitic gabbro. Within the mylonitic zone garnet occurs as poikilolitic and euhedral neoblasts associated with rutile. Garnet is syn- to postkinematic as indicated by helicitic rutile inclusions and poikilolitic growth. The garnet composition range is $\mathrm{Alm}_{55-69} \mathrm{Prp}_{15-23} \mathrm{Grs}_{11-25} \mathrm{Adr}_{01-04}$ (Fig. 3.8) and therefore less pyrope pronounced as the analysis by Yoshino et al. (1998). Profiles across garnets are commonly flat. One garnet from the mylonite (13/01/01) has a grossularite richer rim, but weaker than in the profiles given from garnets in the Granulitic gabbro by Ringuette et al. (1999).


Fig. 3.8: Garnet compositions of Granulitic gabbro and Sarangar gabbro.

Table 3.2: Representative (averaged, $n=n r$. of analyses) chemical composition of garnet. Single analyses are given in appendix 3.2.

| Electron microprobe data |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Mineral: garnet |  |  |  |  |  |  |
|  | Granulitic Gabbro |  |  |  | Sarangar Gabbro mylonitic |  |
|  | undeformed |  | mylonitic |  |  |  |
| Sample | 13/01/01 | 94/04/01 | 13/01/01 | 94/04/01 | 16/01/01 | 90/02/01 |
| Spot | P2 | P3 | P1 | P1 | P1 | P1 |
| EMP-Nr. | MP2 | MP6 | MP1 | MP4 | MP11 | MP7 |
| $\mathrm{SiO}_{2}$ | 40.07 | 39.26 | 39.75 | 39.16 | 38.68 | 37.32 |
| $\mathrm{TiO}_{2}$ | 0.06 | 0.06 | 0.11 | 0.07 | 0.06 | 0.07 |
| $\mathrm{Cr}_{2} \mathrm{O}_{3}$ | 0.03 | 0.01 | 0.02 | 0.00 | 0.01 | 0.02 |
| $\mathrm{Al}_{2} \mathrm{O}_{3}$ | 22.45 | 21.63 | 22.19 | 21.65 | 21.77 | 21.30 |
| $\mathrm{Fe}_{2} \mathrm{O}_{3}$ | 0.48 | 0.90 | 0.75 | 1.50 | 0.22 | 1.43 |
| FeO | 20.39 | 20.56 | 21.17 | 20.43 | 25.58 | 25.13 |
| MnO | 0.58 | 0.44 | 0.58 | 0.49 | 3.47 | 1.91 |
| MgO | 9.66 | 7.75 | 9.41 | 8.35 | 4.62 | 4.17 |
| NiO |  | 0.01 |  | 0.01 |  | 0.01 |
| CaO | 7.42 | 9.02 | 6.94 | 8.29 | 6.78 | 7.73 |
| $\mathrm{Na}_{2} \mathrm{O}$ | 0.03 | 0.07 | 0.03 | 0.08 | 0.02 | 0.04 |
| $\mathrm{K}_{2} \mathrm{O}$ | 0.02 | 0.01 | 0.01 | 0.02 | 0.01 | 0.01 |
| Total | 101.20 | 99.73 | 100.98 | 100.07 | 101.22 | 99.14 |
| Si | 2.997 | 3.007 | 2.991 | 2.987 | 3.002 | 2.962 |
| Ti | 0.004 | 0.004 | 0.006 | 0.004 | 0.004 | 0.004 |
| Cr | 0.002 | 0.000 | 0.001 | 0.000 | 0.001 | 0.001 |
| AI | 1.979 | 1.953 | 1.968 | 1.946 | 1.991 | 1.993 |
| $\mathrm{Fe}^{3}$ | 0.027 | 0.052 | 0.043 | 0.086 | 0.013 | 0.085 |
| $\mathrm{Fe}^{2}$ | 1.275 | 1.318 | 1.332 | 1.304 | 1.660 | 1.668 |
| Mn | 0.037 | 0.029 | 0.037 | 0.032 | 0.228 | 0.128 |
| Mg | 1.077 | 0.885 | 1.055 | 0.949 | 0.534 | 0.493 |
| Ni |  | 0.001 |  | 0.001 |  | 0.001 |
| Ca | 0.595 | 0.740 | 0.560 | 0.677 | 0.564 | 0.657 |
| Na | 0.005 | 0.011 | 0.005 | 0.011 | 0.004 | 0.006 |
| K | 0.002 | 0.002 | 0.001 | 0.003 | 0.001 | 0.001 |
| $\Sigma$ Cations | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 |
| Grossular | 0.182 | 0.219 | 0.162 | 0.182 | 0.179 | 0.176 |
| Almandine | 0.427 | 0.441 | 0.446 | 0.437 | 0.555 | 0.565 |
| Pyrope | 0.361 | 0.299 | 0.353 | 0.322 | 0.179 | 0.167 |
| Spessartite | 0.012 | 0.010 | 0.012 | 0.011 | 0.076 | 0.044 |
| Andradite | 0.014 | 0.028 | 0.021 | 0.045 | 0.007 | 0.044 |
| average of ncalculated with "Norm" program by P. Ulmer using:a |  |  |  | 39 | 33 | 94 |
|  |  |  |  | GARNET Norm |  |  |
| Normalization on the basis of 7 CATIONS and 24 CHARGES |  |  |  |  |  |  |

### 3.1.3.3 Pyroxene

Pyroxenes in the less deformed and mylonitic Granulitic gabbro have similar composition. They range from $\mathrm{En}_{35-49} \mathrm{Fs}_{10-21} \mathrm{Wo}_{33-48}$ for the less deformed to $\mathrm{En}_{35-45} \mathrm{Fs}_{11-}$ ${ }_{18} \mathrm{Wo}_{42-48}$ for the mylonitic gabbro (Fig. 3.9 and Table 3.3). The composition is homogeneous at the grain scale. Analyses plot mainly in the diopside and augite fields.

The Sarangar gabbro has pyroxenes only in the less deformed parts. Clinopyroxenes $\left(\mathrm{En}_{32-38} \mathrm{Fs}_{19-26} \mathrm{Wo}_{38-45}\right)$ and orthopyroxenes $\left(\mathrm{En}_{48-60} \mathrm{Fs}_{35-52} \mathrm{Wo}_{01-15}\right)$ are coexisting (Fig. 3.9).

## Clinopyroxene



## Orthopyroxene



Fig. 3.9: Pyroxene compositions of Granulitic gabbro and Sarangar gabbro. Nomenclature after Morimoto (1988).

| Electron microprobe data |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Mineral: clinopyroxene |  |  |  |  |  | orthopyroxene Sarangar Gabbro undeformed |  |
|  | Granulitic Gabbro |  |  |  | $\begin{aligned} & \text { Sarangar } \mathrm{Gb} \text {. } \\ & \text { undeformed } \end{aligned}$ |  |  |
|  | undeformed |  | mylonitic |  |  |  |  |
| Sample | 13/01/01 | 94/04/01 | 13/01/01 | 94/04/01 | 90/02/01 | 16/01/01 | 90/02/01 |
| Spot | P2 | P3 | P1 | P1 | P4 | P3 | P4 |
| EMP-Nr. | MP2 | MP6 | MP1 | MP4 | MP10 | MP13 | MP10 |
| $\mathrm{SiO}_{2}$ | 51.48 | 50.50 | 50.86 | 50.84 | 49.79 | 53.60 | 48.87 |
| $\mathrm{TiO}_{2}$ | 0.59 | 0.55 | 0.55 | 0.54 | 0.35 | 0.07 | 0.08 |
| $\mathrm{Cr}_{2} \mathrm{O}_{3}$ | 0.09 | 0.02 | 0.08 | 0.01 | 0.09 | 0.01 | 0.02 |
| $\mathrm{Al}_{2} \mathrm{O}_{3}$ | 6.00 | 5.90 | 6.06 | 4.68 | 3.79 | 2.96 | 2.61 |
| $\mathrm{Fe}_{2} \mathrm{O}_{3}$ | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.79 |
| FeO | 7.57 | 7.66 | 6.91 | 8.52 | 11.97 | 20.71 | 27.61 |
| MnO | 0.07 | 0.04 | 0.06 | 0.07 | 0.33 | 0.80 | 0.74 |
| MgO | 12.33 | 11.77 | 12.06 | 12.05 | 10.81 | 17.14 | 16.07 |
| NiO |  | 0.02 |  | 0.01 | 0.02 | 0.00 | 0.02 |
| CaO | 20.65 | 21.90 | 21.13 | 21.80 | 20.85 | 2.73 | 0.80 |
| $\mathrm{Na}_{2} \mathrm{O}$ | 1.28 | 1.18 | 1.28 | 1.19 | 0.66 | 0.34 | 0.06 |
| $\mathrm{K}_{2} \mathrm{O}$ | 0.05 | 0.01 | 0.05 | 0.03 | 0.01 | 0.11 | 0.02 |
| Total | 100.10 | 99.55 | 99.03 | 99.75 | 98.67 | 98.47 | $\underline{97.69}$ |
| Si | 1.898 | 1.883 | 1.895 | 1.901 | 1.911 | 2.018 | 1.930 |
| Ti | 0.016 | 0.016 | 0.015 | 0.015 | 0.010 | 0.002 | 0.002 |
| Cr | 0.002 | 0.001 | 0.002 | 0.000 | 0.003 | 0.000 | 0.001 |
| Al | 0.261 | 0.259 | 0.266 | 0.207 | 0.171 | 0.131 | 0.121 |
| $\mathrm{Fe}^{3}$ | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.023 |
| $\mathrm{Fe}^{2}$ | 0.233 | 0.239 | 0.215 | 0.267 | 0.384 | 0.652 | 0.912 |
| Mn | 0.002 | 0.001 | 0.002 | 0.002 | 0.011 | 0.026 | 0.025 |
| Mg | 0.678 | 0.654 | 0.669 | 0.671 | 0.619 | 0.962 | 0.946 |
| Ni |  | 0.001 |  | 0.000 | 0.001 | 0.000 | 0.001 |
| Ca | 0.816 | 0.875 | 0.844 | 0.873 | 0.857 | 0.110 | 0.034 |
| Na | 0.091 | 0.086 | 0.092 | 0.086 | 0.049 | 0.025 | 0.004 |
| K | 0.002 | 0.001 | 0.002 | 0.001 | 0.001 | 0.005 | 0.001 |
| $\Sigma$ Cations | 4.001 | 4.014 | 4.003 | 4.024 | 4.017 | 3.930 | 4.000 |
| xMg Fe(tot) | 0.744 | 0.732 | 0.757 | 0.716 | 0.617 | 0.596 | 0.503 |
| Wollastonite | 0.365 | 0.388 | 0.377 | 0.397 | 0.391 | 0.055 | 0.017 |
| Enstatite | 0.339 | 0.323 | 0.334 | 0.328 | 0.304 | 0.478 | 0.456 |
| Ferrosilite | 0.117 | 0.118 | 0.107 | 0.130 | 0.189 | 0.324 | 0.439 |
| average of $n$ | 17 | 33 | 8 | 9 | 18 | 20 | 23 |
| cpx calculated with "Norm" program by P. Ulmer using: CPX Norm (Cawthorn \& Collerson) |  |  |  |  |  |  |  |
| opx calculated with "Norm" program by P. Ulmer using: OPX Norm (Wood \& Banno) cpx normalization on the basis of 12 CHARGES and $\mathrm{Fe}^{3+}=$ Acmite |  |  |  |  |  | *OPX Norm (Lindsley) |  |

Table 3.3: Representative (averaged, $n=n r$. of analyses) chemical composition of pyroxene. Single analyses are given in appendix 3.3.

### 3.1.3.4 Amphiboles

In the less deformed parts of the Granulitic gabbro the amphiboles are ranging from magnesio-hornblende, actinolitic hornblende to actinolite in the $\mathrm{Si}-\mathrm{Mg} / \mathrm{Mg}+\mathrm{Fe}^{2+}$ diagram (Fig. 3.10). In the mylonitic parts most of the analysed amphiboles plot in the tschermakitic hornblende and pargasite field, but also in the tschermakite field. Magnesio-hornblende is also observed in the matrix.

Amphiboles in the Sarangar gabbro are dominantly tschermakite and pargasite (Table 3.4). In the less deformed parts two distinct groups of compositions are present, one plotting in the fields of tschermakitic hornblende, tschermakite, ferro-tschermakite and ferroan pargasite and a less common one plotting in the field of actinolitic hornblende (Fig. 3.10). The actinolitic hornblende composition is typical for epidote-amphibole and greenschist facies conditions.


Fig. 3.10: Amphibole variation plots. The uppermost diagram shows the ranges of the measured amphiboles expressed as numbers of $\mathrm{Na}+K_{A}$ and Si atoms per formula unit, whereas to two lower diagrams use the Si versus $M g / M g+F^{2+}$ plot. Nomenclature after Leake (1978).

Amphibole composition in the mylonitic part shows less variations and plot into small areas in the tschermakite field and ferroan pargasite field (Fig. 3.10). Few analyses are within the magnesio-hornblende and actinolitic hornblende fields (Fig. 3.10). Pargasite and tschermakite are the main phases of the mylonitic shear zones matrix. There is an overlap of composition for amphiboles from the less deformed and mylonitic part in the hornblende field.

Amphiboles in the Kiru amphibolites are homogenous tschermakitic hornblende and ferro-tschermakitic hornblendes (Fig. 3.10 and Table 3.4).


Table 3.4: Representative (averaged, $n=$ $n r$. of analyses) chemical composition of amphiboles. Single analyses are given in appendix 3.4.

### 3.1.3.5 Biotite, Muscovite and Chlorite

Biotite and muscovite occur in the mylonites and sparsely in their vicinity. Biotite occurs frequently in contact to garnet in the Granulitic gabbro (Fig. 3.11a), whereas in the Sarangar gabbro biotite is a minor accessory mineral. In the Kiru amphibolite biotite is corroded at its boundaries.

White mica is concentrated to the plagioclase and quartz rich parts of the Sarangar gabbro mylonites, overgrowing amphibolite facies assemblages (Fig. 3.11b), thus indicating a greenschist facies overprint.

Chlorite, often filling the fractures in pyroxenes or replacing amphiboles, was formed during retrogression in the Granulitic and Sarangar gabbros and the Kiru amphibolite (Fig. 3.11c, analyses in Table 3.5). Chlorite is generally present in small amounts in the bulk rock, but in certain domains the concentration is noticeably higher. This indicates a localised greenschist facies overprint.


Figure 3.11: a) biotite as secondary phase grown in fractured garnet in a mylonitic granulitic gabbro (13/01/01); b) muscovite overgrows hornblende and plagioclase in a mylonitic Sarangar gabbro (16/01/01); c) chlorite as fracture fill of an orthopyroxene surrounded by clinopyroxene in an undeformed Sarangar gabbro (90/02/01); d) epidote with rutile inclusion in contact to hornblende and plagioclase in a mylonitic Sarangar gabbro (90/02/01). + crossed nicols.

| Electron microprobe data |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| biotite |  |  | muscovite | chlorite |  |  |
|  | Gr. Gabbro | Sarangar Gb. | Sarangar Gb. |  |  |  |
|  | mylonitic | undeformed | mylonitic | undeformed | mylonitic | Amph. |
| Sample | 13/01/01 | 16/01/01 | 16/01/01 | 90/02/01 | 90/02/01 | 46/02/01 |
| Spot | P1 | P3 | P1 | P4 | P1 | P1 |
| EMP-Nr. | MP1 | MP13 | MP11 | MP10 | MP7 | MP19 |
| $\mathrm{SiO}_{2}$ | 33.71 | 31.77 | 45.97 | 29.81 | 24.28 | 33.55 |
| $\mathrm{TiO}_{2}$ | 1.72 | 2.39 | 0.06 | 0.10 | 0.06 | 2.02 |
| $\mathrm{Cr}_{2} \mathrm{O}_{3}$ | 0.63 | 0.17 | 0.01 | 0.03 | 0.02 | 0.02 |
| $\mathrm{Al}_{2} \mathrm{O}_{3}$ | 17.01 | 18.06 | 36.71 | 15.12 | 20.84 | 17.63 |
| $\mathrm{Fe}_{2} \mathrm{O}_{3}$ | 1.03 | 15.44 | 2.02 | 0.00 | 0.00 | 0.00 |
| FeO | 15.61 | 6.89 | 0.00 | 26.20 | 25.84 | 14.85 |
| MnO | 0.04 | 0.16 | 0.02 | 0.19 | 0.25 | 0.18 |
| MgO | 12.46 | 13.56 | 0.81 | 14.31 | 14.66 | 14.81 |
| NiO |  |  |  | 0.02 | 0.01 | 0.03 |
| CaO | 0.16 | 0.89 | 0.02 | 0.31 | 0.10 | 0.78 |
| $\mathrm{Na}_{2} \mathrm{O}$ | 0.21 | 0.12 | 1.18 | 0.03 | 0.05 | 0.12 |
| $\mathrm{K}_{2} \mathrm{O}$ | 7.06 | 2.57 | 8.66 | 0.20 | 0.03 | 4.79 |
| $\mathrm{H}_{2} \mathrm{O}$ | 3.77 | 3.97 | 4.60 | 2.30 | 2.28 | 2.47 |
| Total | 93.40 | 95.97 | 100.06 | 88.59 | 88.40 | 91.24 |
| Si | 2.683 | 2.400 | 2.997 | 3.889 | 3.191 | 4.066 |
| Ti | 0.103 | 0.136 | 0.003 | 0.009 | 0.006 | 0.186 |
| Cr | 0.039 | 0.010 | 0.000 | 0.003 | 0.002 | 0.002 |
| Al | 1.595 | 1.607 | 2.821 | 2.315 | 3.227 | 2.525 |
| $\mathrm{Fe}^{3}$ | 0.061 | 0.873 | 0.099 | 0.000 | 0.000 | 0.000 |
| $\mathrm{Fe}^{2}$ | 1.039 | 0.438 | 0.000 | 2.867 | 2.841 | 1.513 |
| Mn | 0.003 | 0.010 | 0.001 | 0.020 | 0.028 | 0.019 |
| Mg | 1.477 | 1.525 | 0.079 | 2.776 | 2.871 | 2.689 |
| Ni |  |  |  | 0.002 | 0.001 | 0.002 |
| Ca | 0.014 | 0.072 | 0.002 | 0.043 | 0.014 | 0.106 |
| Na | 0.033 | 0.018 | 0.150 | 0.007 | 0.013 | 0.027 |
| K | 0.717 | 0.247 | 0.720 | 0.032 | 0.004 | 0.725 |
| OH |  |  |  | 2.000 | 2.000 | 2.000 |
| H | 2.000 | 2.000 | 2.000 |  |  |  |
| $\Sigma$ Cations | 7.763 | 7.337 | 6.871 | 13.963 | 14.197 | 13.860 |
| xMg Fe(tot) | 0.573 | 0.538 | 0.432 | 0.494 | 0.503 | 0.640 |
| average of $\mathbf{n}$ $\mathbf{3}$ $\mathbf{6}$ <br> bitite calculated with "Norm" program by P. Ulmer using:   CATIONS calculated on the bases of 11 oxygens 7 cations $+\mathrm{Na}+\mathrm{K}+\mathrm{Ca}$ muscovite calculated with "Norm" program by P. Ulmer using: Mica Norm MUS (Standard) CATIONS calculated on the bases of 11 oxygens and 6 cations $+\mathrm{Na}+\mathrm{K}+$ chlorite calculated with "Norm" program by P. Ulmer using: General OXIDE Norm all $\mathrm{Fe}(I I+)$ CATIONS calculated on the basis of 18. Oxygens |  |  |  |  |  |  |
|  |  |  |  |  |  |  |
|  |  |  |  |  |  |  |
|  |  |  |  |  |  |  |

Table 3.5: Representative (averaged, $n=n r$. of analyses) chemical composition of biotite, muscovite and chlorite. Single analyses are given in appendix 3.5 .

### 3.1.3.6 Epidote

Epidote occurs in all analysed samples (see Table 3.6). In the Sarangar gabbro epidote shows vermicular growth and rutile inclusions, indicative for fast crystallization at solidus condition (Fig. 3.11d). In the mylonites epidote concentration is higher in the garnet free domains and vice versa, suggesting that the Fe availability was decisive for the growth of epidote or garnet. In the other samples epidote is poikiloblastic or nematoblastic, pseudomorph after plagioclase and infrequently associated with muscovite.


Table 3.6: Representative (averaged, $n=n r$. of analyses) chemical composition of epidote. Single analyses are given in appendix 3.6.

### 3.1.4 PT calculations

Pressure and temperature conditions were calculated using the program "Thermobarometry" by F. Spear \& M.J. Kohn (1995). This relatively easy to use program provides calculations of many published geothermometers and geobarometers and plots the results as lines of constant equilibrium ( $\mathrm{K}_{\mathrm{eq}}$ ) on a PT diagram. $\mathrm{K}_{\mathrm{eq}}$ lines were obtained for different reactions and calibrations. The intersections of $\mathrm{K}_{\text {eq }}$ lines of geobarometer- and geothermometer-reactions are marked with white circles, where compositions from the same minerals were used for calculations (Fig. 3.12).

Results for the undeformed Granulitic gabbro range between $625-905^{\circ} \mathrm{C}$ and $0.65-$ 1.29 GPa (Fig. 3.12, 13/01/01 P2 and 94/04/01 P3). The upper value was obtained with a garnet-plagioclase-clinopyroxene-quartz assemblage (measured in grain centre) and is in partial agreement with published conditions of $813-949{ }^{\circ} \mathrm{C}$ and $1.15-1.7 \mathrm{GPa}$ (Yamamoto, 1993), $713-1134{ }^{\circ} \mathrm{C}$ and $1.53-2.19 \mathrm{GPa}$ (Ringuette et al., 1999) and 860 $-970{ }^{\circ} \mathrm{C}$ and $1.5-3.0 \mathrm{GPa}$ (Anczkiewicz \& Vance, 2000). The calculated pressure of 1.29 GPa is below the pressures presented by Anczkiewicz \& Vance (2000) and Ringuette et al. (1999), possibly because our samples are close to a mylonitic zone. As described before the amphibolite facies assemblage is primarily developed in the mylonite. Therefore amphibolitisation is strong in the mylonite, declining towards the core of the undeformed pods. The sample, just a few centimetres away from the mylonitic zone, is probably more affected by some retrogression than those from the core of undeformed pods. The lower value used a garnet-plagioclase-hornblende-quartz assemblage (measured at grain rims). One garnet-plagioclase-clinopyroxene-quartz assemblage (measured at grain rims) yielded a pressure of 0.89 GPa at a temperature of $625{ }^{\circ} \mathrm{C}$, showing a difference in pressure of 0.24 GPa to the one obtained with the garnet-plagioclase-hornblende-quartz assemblage at the same temperature. This discrepancy is due to applied reaction/calibration and gives an estimate of error. The results show clearly the retrograde character of the successive equilibration. The PT conditions for the mylonitic part (Fig. 3.12, 13/01/01 P1 and 94/04/01 P1) range from $600-900^{\circ} \mathrm{C}$ and $0.54-1.19 \mathrm{GPa}$, with a pronounced pressure range of $0.75-1.0 \mathrm{GPa}$. Temperatures higher than $900{ }^{\circ} \mathrm{C}$ for shear zones are reported by Ringuette et al. (1999). The higher pressure and temperature values are obtained from porphyroclasts and their inclusions. A new grown garnet-hornblende assemblage yields a temperature of ca. $740^{\circ} \mathrm{C}$ and pressure of $0.8-0.9 \mathrm{GPa}$. Small amphiboles in the matrix point to an equilibration temperature of $650-700{ }^{\circ} \mathrm{C}$ and amphiboles newly grown in fractures of garnet clasts yield temperatures of $680-770{ }^{\circ} \mathrm{C}$.

PT conditions for the less deformed parts of the Sarangar gabbro range from 575-800 ${ }^{\circ} \mathrm{C}$ and $0.58-0.81 \mathrm{GPa}$ (Fig. 3.12, 16/01/01 P3 and 90/02/01 P4). Garnet compositions were taken from the centre of garnet porphyroclasts in the mylonitic parts. Therefore the calculated pressures are most likely too low and should not be interpreted without confirmation by garnet independent geobarometers.


Fig. 3.12: PT conditions for Granulitic gabbro and Sarangar gabbro samples of the mylonitic zone and less deformed portions obtained by $K_{\text {eq }}$ lines for specific reactions and calibrations. White circles mark crossings of $K_{\text {eq }}$ lines calculated for same mineral compositions.

Peak conditions obtained with clinopyroxene-plagioclase-quartz assemblages are 760 $870^{\circ} \mathrm{C}$ and $0.82-1.1 \mathrm{GPa}$ (Yoshino et al., 1998). PT estimates from an amphibolite in the Swat valley is $690-860{ }^{\circ} \mathrm{C}$ and $0.7-0.9 \mathrm{GPa}$ (Anczkiewicz \& Vance, 2000), which is similar to the conditions of the Sarangar gabbro. In the mylonitic portion temperatures and pressure range from $525-710{ }^{\circ} \mathrm{C}$ and $0.58-1.0 \mathrm{GPa}$ (Fig. 3.12, 16/01/01 P1 and 90/02/01 P1). New grown garnet-hornblende-plagioclase-quartz assemblage yielded conditions of $640-710^{\circ} \mathrm{C}$ and $0.9-1.0 \mathrm{GPa}$. The lowest obtained temperature ( $525^{\circ} \mathrm{C}$ ) was calculated using the composition at the rims of the minerals (garnet-hornblende-plagioclase).

Metamorphic conditions for the "Kamila Amphibolite Belt" are estimated around 550 $650{ }^{\circ} \mathrm{C}$ and $0.9-1.0 \mathrm{GPa}$ (Treloar et al., 1990) using a garnet-amphibole thermometry. Pressure temperature conditions were estimated for the Kiru amphibolite here with 0.7 1.0 GPa and $640-715{ }^{\circ} \mathrm{C}$ (Fig. 3.13). Temperatures were calculated using the hornblende-plagioclase geothermometer (Holland \& Blundy, 1994). The temperature difference between the above mentioned estimates might be due to the different applied thermometers or differently equilibrated samples. Due to the lack of garnet in the analysed sample the pressure was estimated using the Al content in hornblende. This method is valid for a igneous assemblage (Schmidt, 1992) which is close (but not complete) to the assemblage in the Kiru amphibolites. Pressures are consequently only rough estimates.


Fig. 3.13: PT condition for Kiru amphibolite. Temperature calculated after the hornblende-plagioclase geothermometer (Holland and Blundy, 1994). Pressure estimated by the total Aluminium content in hornblende.

Comparing the PT conditions inside and outside the mylonite of the Granulitic and Sarangar gabbros a pressure difference of ca. 0.2 GPa at the same temperature is noticed. The mylonitic zones of the Granulitic and Sarangar gabbros have similar pressure conditions, whereas the undeformed portions plot in higher pressure (Granulitic gabbro) and lower pressure (Sarangar gabbro) (Fig. 3.14). This effect is most likely the result of an underestimate of pressures in the undeformed Sarangar gabbro due to the applied geobarometers, which require garnet. The lack of applicable geobarometers, that fit the measured assemblage led to the use of an averaged garnet composition from the mylonite zone to complete the assemblages for the applied geobarometers. This does not represent a strict equilibrated (at least for garnet) assemblage. The calculated pressures for the undeformed Sarangar gabbro should therefore be considered only as imprecise minimum pressures.


Fig. 3.14: Comparison between PT conditions of mylonitic and undeformed sample portions. The Granulitic gabbro yields ca. 0.2 GPa higher pressures in the undeformed parts than in the mylonitic zone, whereas the Sarangar gabbro yields higher pressures in the mylonitic zone. Black square in the lower diagram shows the obtained PT conditions of Kiru amphibolites.

### 3.1.5 Interpretation of petrological evolution and PT conditions

The PT conditions obtained in the mylonitic parts suggest that the development of the anastomosing shear zones started under granulite facies conditions at ca. $800{ }^{\circ} \mathrm{C}$ and $\sim 1.0 \mathrm{GPa}$ and continued under amphibolite facies to minimum ca. $550^{\circ} \mathrm{C}$ and $\sim 0.4 \mathrm{GPa}$. The integrated PT path (Fig. 3.15) can be split in three parts: 1) cooling to ca. $750{ }^{\circ} \mathrm{C}$ at ca. $1.0 \mathrm{GPa}, 2$ ) decompression between ca. 1.0 GPa to ca. 0.35 GPa at $550{ }^{\circ} \mathrm{C}$ and 3) cooling. The Kiru amphibolite equilibrated under amphibolite facies conditions at ca.0.7
-1.0 GPa and $640-715^{\circ} \mathrm{C}$ at the beginning of the decompression.


Fig. 3.15: Integrated PT path for the lowest Kohistan Arc Complex (Jijal and Metaplutonic Complex). Black square shows the obtained PT conditions of Kiru amphibolites. Stippled paths are representative for the Granulitic gabbro of the Jijal Complex (after Ringuette et al., 1999; Yamamoto, 1993). Metamorphic facies: GS = Greenschist, $B S=$ Blueschist, EA $=$ Epidote-Amphibolite, $A=$ Amphibolite.

### 3.1.6 Discussion

Early magmatic crystallization of gabbroic magma in both the Granulitic gabbro and the Sarangar gabbro involved temperatures $>1000{ }^{\circ} \mathrm{C}$ (Yamamoto 1993; Yoshino et al., 1998; Ringuette et al., 1999). They agree about the amphibolite and greenschist facies retrograde path that followed this high-temperature stage. Some uncertainty exists over the high-temperature high-pressure evolution of the Granulitic gabbro, which resulted in
different petrogenetic interpretations. Based on geothermobarometry of symplectites and inclusion-host assemblages of garnet, clinopyroxene and plagioclase, Yamamoto (1993) suggested that garnet growth occurred after magmatic crystallization during heating to temperatures of ca. $900{ }^{\circ} \mathrm{C}$ and burial to pressures $<1.5 \mathrm{GPa}$ and before decompression began to amphibolite facies conditions at $700{ }^{\circ} \mathrm{C}$ and 1.0 GPa . He also argued that heating and burial are due to crustal thickening based on $\mathrm{Al}_{2} \mathrm{O}_{3}$ increase from core to rim in orthopyroxene combined with an increase in the grossular content of garnet. Crustal thickening would be caused by accretion of basaltic magma at mid- and upper crustal depths. Yoshino et al. (1998) suggest a similar process where heating and burial formed granulites in the Metaplutonic Complex, reaching conditions of ca. 800 ${ }^{\circ} \mathrm{C}$ and of $0.8-1.1 \mathrm{GPa}$ for the maximum pressure. They estimated, based on the Al zoning of clinopyroxene and plagioclase a prograde path from $\sim 0.8 \mathrm{GPa}$ to $\sim 1.1 \mathrm{GPa}$ along with a slight temperature increase. Ringuette et al. (1999) favour a magmatic crystallization followed by a period of isobaric cooling, after which ca. 1.0 GPa of decompression occurred whilst the rocks remained above ca. $700^{\circ} \mathrm{C}$. They show that some garnet is in textural equilibrium with other cumulate phases such as clinopyroxene, plagioclase and amphibole, which attest for its magmatic origin. Corerim pairs and compositional profiles across garnets and pyroxene indicate a quasiisobaric cooling trend corresponding to the transition from the magmatic to the granulitic stage (Ringuette et al., 1999). Garnet in the Granulitic gabbros then results from magmatic processes and sub-solidus isobaric cooling of cumulates emplaced at the base of the oceanic arc-type crust (Ringuette et al., 1999). Burg et al. (1998) suggested a similar scenario with the Granulitic gabbros interpreted as gabbroic intrusions within mantle rocks. Anczkiewicz \& Vance (2000) conclude that for garnet growth all above mentioned models require temperatures and pressures of over $850^{\circ} \mathrm{C}$ at $1.5-2.0 \mathrm{GPa}$, followed by decompression of $0.5-1.0 \mathrm{GPa}$ while the rocks were still above about 700 ${ }^{\circ}$ C. PT data obtained by Anczkiewicz \& Vance (2000) from the Granulitic gabbro and an amphibolite (Swat valley) of $>1.5 \mathrm{GPa}$ and $900^{\circ} \mathrm{C}$ and of $0.8-1.0 \mathrm{GPa}$ and $800^{\circ} \mathrm{C}$ respectively are consistent with both scenarios. PT conditions presented here most likely missed assemblages indicative for pressures $>1.5 \mathrm{GPa}$ because the analysed samples are close to a mylonitic zone and are probably affected by retrogression. Pressures obtained for the Granulitic gabbro between temperatures of $800-900^{\circ} \mathrm{C}$ plot in general
ca. $0.3-0.4 \mathrm{GPa}$ below the PT path proposed by Ringuette et al. (1999) and scatter between the prograde and retrograde path of Yamamoto (1993). For the Sarangar gabbro, the recorded maximum PT of $0.8-1.0 \mathrm{GPa}$ and $800{ }^{\circ} \mathrm{C}$ is consistent with the maximum pressure conditions in the PT path of Yoshino et al. (1998). The data obtained here cannot exclude nor support a prograde phase in the PT history. The estimated PT conditions for the metadiorite from the Kiru amphibolites ( $0.7-1.0 \mathrm{GPa}$ and $640-715{ }^{\circ} \mathrm{C}$ ) are in partial agreement with the estimates for the "Kamila Amphibolite Belt" (550-680 ${ }^{\circ} \mathrm{C}$ and $\left.0.45-0.65 \mathrm{GPa}\right)$ by Jan (1988) and (550-650 ${ }^{\circ} \mathrm{C}$ and $0.9-1.0 \mathrm{GPa})$ by Treloar et al. (1990). The temperatures estimated for the metadiorite are ca. $100^{\circ} \mathrm{C}$ higher than the previously published estimates. The relatively low pressure estimated by Jan (1988) could not be confirmed.

### 3.1.7 Conclusions

The PT evolution of the Granulitic and Sarangar gabbros is divided into 3 stages: 1) emplacement of gabbroic magmas under high temperature $>1000{ }^{\circ} \mathrm{C}$ (Yoshino et al., 1998) and pressures above or around 1.5 GPa at ca. 100 Ma ; 2) subsequent granulite facies metamorphism under intrusion conditions or with a prograde path of $\Delta \mathrm{P} 0.3 \mathrm{GPa}$ and $\Delta \mathrm{T} 80^{\circ} \mathrm{C}$, as proposed by Yoshino et al. (1998) and 3) amphibolite facies overprint, that ended ca. 83 Ma ago. PT conditions suggested here for the development of anastomosing shear zones started in granulite facies conditions at ca. $800^{\circ} \mathrm{C}$ and $\sim 1.0$ GPa and continued under amphibolite facies conditions to ca. $650^{\circ} \mathrm{C}$ and $\sim 0.7 \mathrm{GPa}$. Cooling of the Granulitic and Sarangar gabbros followed a similar path. Magmatic intrusion happened most likely at the base of an already thickened crust at depths of $>$ 50 km (Ringuette et al., 1999). The metadiorite from the Kiru amphibolite was equilibrated under amphibolite facies conditions at depths of approximately 30 km . This suggests, that the metadiorite was intrusive earlier or at the beginning of regional decompression in the crust of the arc at a level above the gabbros.

### 3.2 ARC MATURATION, INSIGHT FROM GEOCHEMICAL DATA

### 3.2.1 Sampling

Samples are representative of plutonic rocks in the Metaplutonic Complex. They were collected along the KKH, between Patan and Dasu (Fig. 3.16). Sample GPS coordinates are provided in appendix 3. Along a SW - NE profile the samples were collected in the following lithologies (Fig. 3.17): 1) metadiorite from the Indian Plate; 2) Sarangar Gabbro, sampling less deformed and mylonitic portions; 3) magmatic breccia with one darker and one brighter gabbro; 4) metadiorite with sheared and unsheared parts; 5) hornblende diorite, middle and upper part; 6) tonalite dike intrusive into Kiru amphibolites; 7) granites forming veins and massive laccoliths and 8) metagabbro intercalated with fine grained, greenish (metavolcanic?) amphibolites.


Fig. 3.16: Location map of samples used for whole rock and ICP-MS analysis. Sample 00/30/14 is from a metadiorite at Duber Bazar and is not in the arc related magmatism. All other samples were taken across the Metaplutonic Complex along the Indus River valley and represent major lithologies (gabbros, diorites, tonalites and granites).

Fig. 3.17: Samples projected at the $S W$ - NE profile along the Indus river. Photographs for details.

### 3.2.2 Analytical methods

Samples were ground in an agate mill. X-ray fluorescence (XRF) analyses were performed at ETH-IMP. Inductively coupled plasma mass spectrometry (ICP-MS) analyses were performed at ISTEEM, on a VG PQ2 quadrupole ICP-MS following the analytical procedure published by Ionov et al. (1992) for mantle rocks, with the refinements described by Godard et al. (2000 - EPSL Online Background Dataset).

### 3.2.3 Results

The results of XRF and ICP-Ms analysis are given in Table 3.7. The analysed samples range from 45.80 to $74.05 \mathrm{wt} \% \mathrm{SiO}_{2}$ and thus cover a wide range from basic to acid compositions. The major element variations show negative correlations for $\mathrm{TiO}_{2}, \mathrm{MgO}$, CaO and $\mathrm{Fe}_{2} \mathrm{O}_{3}$, positive correlation for $\mathrm{Na}_{2} \mathrm{O}+\mathrm{K}_{2} \mathrm{O}$ and no correlation for $\mathrm{Al}_{2} \mathrm{O}_{3}, \mathrm{MnO}$ and $\mathrm{P}_{2} \mathrm{O}_{5}$ when plotted against $\mathrm{SiO}_{2}$ (Fig. 3.18). The decrease of $\mathrm{TiO}_{2}, \mathrm{CaO}$ and $\mathrm{Fe}_{2} \mathrm{O}_{3}$ with differentiation is characteristic of calc-alkaline rocks (Miyashiro, 1975), confirming the role of plagioclase and magnetite as major fractionating phases in the evolution of the magmas (Wilson, 1989). The positive correlation of $\mathrm{Na}_{2} \mathrm{O}+\mathrm{K}_{2} \mathrm{O}$ is typical for calc-alkaline rocks. The magmas are aligned along a calc-alkaline trend close to the tholeiitic trend in an AFM diagram (Fig. 3.19). The calc-alkaline differentiation trend results in juvenile granites, crystallizing k-feldspar, plagioclase and quartz. The Chilas gabbronorite (data from Jan, 1988; Khan et al., 1993 and Mikoshiba et al., 1999) plots at lower Fe contents on a different calc-alkaline trend. The analysed rocks can be separated essentially in three groups: a) gabbros, samples $16 / 01 / 01,90 / 02 / 01,71 / 03 / 02$, 91/03/02 and 91/03/03, b) diorites, samples RK2, RK1, 35/02/01, 91/07/01 and $00 / 30 / 14$ and c) granites/tonalites: samples $36 / 01 / 01,91 / 08 / 01$ and $00 / 10 / 01$. The calculated CIPW norm nomenclature of the rocks are qtz-gabbro for samples 16/01/01, 90/02/01 and RK2, gabbro for samples 71/03/02, 91/03/02 and 91/03/03, gabbronorite for sample RK1, qtz-diorite for samples 35/02/01 91/07/01 and 00/10/01, tonalite for sample $36 / 01 / 01$ and granite for samples $91 / 08 / 01$ and $00 / 30 / 14$. This nomenclature is substituted in the following in the light of field and petrological descriptions. Samples RK2 and RK1 display geochemical features intermediate between the gabbros and diorites. They are more similar to diorites with regard to trace elements. Plotted in a TAS diagram (total alkalis versus silica) the Sarangar Gabbro is of more "evolved" character than the structurally higher gabbros (91/03/02 and 91/03/03) (Fig. 3.20). The
tonalite (36/01/01) is clearly less "evolved" in term of major elements, than the granites. The tectonic setting under which the magmas were emplaced is an island arc or active continental margin setting as suggested by the comparison of modern volcanic rocks $\mathrm{MgO}-\mathrm{FeO}-\mathrm{Al}_{2} \mathrm{O}_{3}$ composition range with the analysed ones $\left(\mathrm{SiO}_{2}\right.$ range of $51-56$ $\mathrm{wt} \%)$ (Fig. 3.21).

| XRF Analysis |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| facies | gabbu |  |  |  |  | RK2 RK1 |  | $\frac{\text { dioctite }}{35 / 02 / 01}$ | 91/07/01 | 00/30/14 | tonalilite/granite |  |  | $\frac{\text { volcanocl. }}{92 / 07 / 01}$ |
|  | 16/01/01 | 90/02/01 | 91/03/02 | 91/03/03 | 71/03/02 |  |  | 36/01/01 |  |  | 91/08/01 | 00/10/01 |  |
| SiO2 | 52.22 | 53.45 | 45.80 | 48.11 | 49.18 | 53.20 | 51.56 |  | 55.06 | 56.17 | 67.11 | 62.35 | 74.05 | 69.27 | 68.18 |
| $\mathrm{TiO}_{2}$ | 0.72 | 0.78 | 1.27 | 1.06 | 1.08 | 0.90 | 1.00 | 0.73 | 0.55 | 0.89 | 0.45 | 0.05 | 0.21 | 0.51 |
| $\mathrm{Al}_{2} \mathrm{O}_{3}$ | 17.63 | 17.44 | 19.11 | 22.07 | 15.63 | 17.29 | 18.24 | 18.84 | 18.52 | 13.54 | 20.23 | 14.74 | 15.91 | 10.85 |
| $\mathrm{Fe}_{2} \mathrm{O}_{3}$ | 9.62 | 9.28 | 11.84 | 8.17 | 9.24 | 10.18 | 10.86 | 6.21 | 6.31 | 6.57 | 2.67 | 0.85 | 3.10 | 3.56 |
| FeO | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 |
| MnO | 0.17 | 0.15 | 0.14 | 0.08 | 0.16 | 0.18 | 0.18 | 0.11 | 0.14 | 0.06 | 0.02 | 0.15 | 0.05 | 0.05 |
| MgO | 5.28 | 4.62 | 5.13 | 3.14 | 7.82 | 4.48 | 4.70 | 2.89 | 2.66 | 0.62 | 1.07 | 0.17 | 0.91 | 1.58 |
| CaO | 9.69 | 9.45 | 10.49 | 10.45 | 11.21 | 9.35 | 9.74 | 8.17 | 7.93 | 2.91 | 7.16 | 2.01 | 4.38 | 5.06 |
| Na 2 O | 2.52 | 2.87 | 3.32 | 4.32 | 2.63 | 2.79 | 2.90 | 4.58 | 4.27 | 2.98 | 4.52 | 3.64 | 3.73 | 2.36 |
| K2O | 0.40 | 0.30 | 0.19 | 0.18 | 0.37 | 0.30 | 0.28 | 0.36 | 0.36 | 4.35 | 0.17 | 1.77 | 1.14 | 2.11 |
| $\mathrm{P}_{2} \mathrm{O}_{5}$ | 0.12 | 0.11 | 0.34 | 0.26 | 0.13 | 0.16 | 0.19 | 0.30 | 0.22 | 0.26 | 0.14 | 0.06 | 0.10 | 0.09 |
| $\mathrm{LOI}^{1}$ | 0.90 | 0.92 | 1.53 | 1.65 | 2.04 | 0.55 | 0.21 | 1.87 | 1.78 | 0.71 | 0.43 | 1.19 | 1.03 | 4.93 |
| $\mathrm{Cr2O}_{3}$ | 0.02 | 0.02 | 0.01 | 0.01 | 0.04 | 0.01 | 0.01 | 0.01 | 0.01 |  | 0.01 | 0.01 |  | 0.03 |
| NiO | 0.00 | 0.00 | 0.00 | 0.00 | 0.01 | 0.00 | 0.00 | 0.00 | 0.00 |  | 0.00 | 0.00 |  | 0.00 |
|  |  |  |  |  |  |  |  |  |  | (2) |  |  | (2 |  |
| Total | 99.29 | 99.39 | 99.17 | 99.50 | 99.54 | 99.39 | 99.87 | 99.13 | 98.92 | 100.00 | 99.22 | 98.69 | 99.83 | 99.31 |


| ppm |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Zr | 54 | 62 | 41 | 44 | 87 | 88 | 81 | 77 | 104 | 478 | 141 | 48 | 62.4 | 190 |
| V | 172 | 220 | 194 | 110 | 186 | 186 | 170 | 63 | 75 | 15.7 | < 10 | < 10 | 25 | 63 |
| Cr | 40 | 45 | < 6 | < 6 | 136 | < 6 | < 6 | $<6$ | < 6 | 8.4 | 10 | < 6 | 6.7 | 130 |
| Ni | 16 | 14 | <3 | <3 | 85 | < 3 | < 3 | <3 | < 3 | < 5 | < 3 | < 3 | < 5 | 17 |
| Cu | 38 | 55 | 73 | 51 | 8 | 64 | 39 | 32 | 7 | 9.1 | 51 | < 3 | 5.2 | 11 |
| Zn | 74 | 73 | 80 | 53 | 78 | 72 | 74 | 73 | 69 | 108 | 18 | 35 | 35.1 | 44 |
| Ga | 12 | 12 | 15 | 15 | 11 | 14 | 15 | 17 | 14 | 22.5 | 13 | 10 | 14.9 | 7 |
| Sc | 32 | 25 | 35 | 13 | 31 | 27 | 30 | <2 | <2 |  | <2 | <2 |  | <2 |
| S | $<50$ | < 50 | $<50$ | $<50$ | 116 | 189 | 184 | $<50$ | < 50 |  | 703 | $<50$ |  | < 50 |


| ppm | 16/01/01 | na | 91/03/02 | 91/03/03 | 71/03/02 | RK2 | RK1 | 35/02/01 | 91/07/01 | 00/30/14 | 36/01/01 | 91/08/01 | 00/10/01 | na |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Cs | 0.063 |  | 0.080 | 0.088 | 0.126 | 0.007 | 0.011 | 0.086 | 0.104 | 3.480 | 0.275 | 1.080 | 0.550 |  |
| Rb | 2.951 |  | 0.852 | 0.948 | 4.726 | 0.688 | 0.593 | 1.873 | 2.690 | 154.000 | 7.420 | 50.777 | 41.600 |  |
| Ba | 141.970 |  | 66.547 | 88.582 | 52.623 | 86.489 | 83.110 | 111.430 | 108.420 | 1508.000 | 129.040 | 1122.700 | 502.000 |  |
| Th | 0.365 |  | 0.182 | 0.136 | 0.429 | 0.084 | 0.055 | 0.185 | 0.147 | 16.300 | 0.568 | 1.486 | 1.560 |  |
| U | 0.118 |  | 0.067 | 0.046 | 0.128 | 0.028 | 0.023 | 0.055 | 0.039 | 3.380 | 0.119 | 0.418 | 0.380 |  |
| Nb | 1.164 |  | 1.767 | 1.435 | 3.593 | 2.676 | 2.742 | 3.560 | 2.959 | 22.000 | 2.590 | 6.309 | 1.750 |  |
| Ta | 0.066 |  | 0.073 | 0.058 | 0.220 | 0.114 | 0.114 | 0.142 | 0.123 | 1.830 | 0.183 | 0.520 | 0.110 |  |
| La | 5.191 |  | 5.200 | 3.861 | 4.050 | 8.903 | 8.841 | 7.565 | 8.622 | 78.900 | 5.282 | 6.435 | 8.660 |  |
| Ce | 11.951 |  | 14.808 | 10.247 | 11.324 | 22.041 | 22.478 | 19.543 | 21.519 | 144.000 | 11.313 | 13.360 | 15.300 |  |
| Pr | 1.624 |  | 2.313 | 1.544 | 1.682 | 2.999 | 3.092 | 2.815 | 2.885 | 16.100 | 1.469 | 1.512 | 1.700 |  |
| $s$ | 249.800 |  | 352.590 | 480.700 | 203.990 | 274.440 | 287.830 | 582.420 | 427.920 | 212.000 | 518.190 | 194.930 | 327.000 |  |
| Nd | 8.031 |  | 12.900 | 8.299 | 8.665 | 14.028 | 14.571 | 13.687 | 13.305 | 58.400 | 6.531 | 5.826 | 5.550 |  |
| sm | 2.281 |  | 3.717 | 2.405 | 2.516 | 3.556 | 3.614 | 3.275 | 2.950 | 10.400 | 1.341 | 1.228 | 0.804 |  |
| Eu | 0.978 |  | 1.229 | 0.957 | 0.948 | 1.116 | 1.198 | 1.069 | 0.955 | 1.960 | 0.528 | 0.274 | 0.551 |  |
| Gd | 2.737 |  | 4.322 | 2.861 | 3.106 | 3.762 | 3.894 | 3.190 | 2.866 | 8.250 | 1.106 | 1.288 | 0.665 |  |
| Tb | 0.489 |  | 0.747 | 0.481 | 0.545 | 0.662 | 0.678 | 0.509 | 0.474 | 1.210 | 0.146 | 0.216 | 0.084 |  |
| Dy | 3.139 |  | 4.586 | 2.974 | 3.529 | 4.113 | 4.231 | 3.000 | 2.890 | 7.450 | 0.756 | 1.386 | 0.479 |  |
| Y | 21.516 |  | 27.998 | 18.279 | 24.275 | 29.156 | 29.980 | 20.625 | 19.826 | 38.800 | 4.933 | 11.366 | 2.490 |  |
| Ho | 0.724 |  | 1.008 | 0.652 | 0.804 | 0.925 | 0.959 | 0.637 | 0.645 | 1.460 | 0.144 | 0.296 | 0.084 |  |
| Er | 2.115 |  | 2.884 | 1.779 | 2.268 | 2.706 | 2.827 | 1.817 | 1.901 | 3.950 | 0.384 | 0.810 | 0.226 |  |
| Tm | 0.316 |  | 0.406 | 0.248 | 0.333 | 0.412 | 0.426 | 0.265 | 0.291 | 0.599 | 0.054 | 0.125 | 0.031 |  |
| Yb | 2.041 |  | 2.388 | 1.452 | 2.083 | 2.606 | 2.692 | 1.640 | 1.885 | 3.900 | 0.324 | 0.913 | 0.281 |  |
| Lu | 0.336 |  | 0.380 | 0.220 | 0.335 | 0.418 | 0.433 | 0.246 | 0.304 | 0.558 | 0.050 | 0.148 | 0.044 |  |
| Pb | 36.196 |  | 29.710 | 32.194 | 3.904 | 15.478 | 22.145 | 70.814 | 4.391 | 47.300 | 21.976 | 73.337 | 2.620 |  |

${ }^{(1}$ Loss of ingnition
${ }^{(2}$ Emission ICP

Table 3.7: Chemical composition for samples from the Metaplutonic Complex (gabbro: 16/01/01, 90/02/01, 91/03/02, 91/03/03, 71/03/02; diorite: RK2, RK1; 35/02/01, 91/07/01; tonalite/granite: 36/01/01, 91/08/01, 00/10/01). Additional samples of volcanoclastic (92/07/01) and metadiorite from the Indian plate (00/30/14). na: not analysed.


Fig. 3.18: Major element variation diagrams for the whole rock samples from the Metaplutonic Complex. The trends are typical for calc-alkaline magmas.

Samples 35/02/01 and 91/07/01 plot in the spreading centre island field. It must be noticed that the validity of the diagram is limited by the relative mobility of the major elements in basalts and that MgO and $\mathrm{Al}_{2} \mathrm{O}_{3}$ are mobile during greenschist facies metamorphism (Pearce, 1976). However, the gabbros of the Metaplutonic Complex plot around the same position. It is also noticed, that the analyses of Chilas gabbronorites (triangles in Fig. 3.21) cluster around a different position.


Fig. 3.19: AFM plot showing the calc alkaline/tholeiitic differentiation trend of the intrusives. The "Amphibolites" field includes amphibolites of plutonic and sedimentary parentage. The Chilas gabbronorites plot below the Metaplutonic Complex trend. Sample 00/30/14 is the Indian plate diorite and does not belong to the arc plutonics.


Fig. 3.20: Discrimination plot after Cox et al. (1979). The observed trend from gabbros to granites follows a relatively constant $\mathrm{Na}_{2} \mathrm{O}+\mathrm{K}_{2} \mathrm{O}$ Wt. \% value. TAS diagram (total alkalis vs. silica) adapted by Wilson (1989) for plutonic rocks. The Indian plate diorite (00/30/14) does not belong to the arc plutonics.


Fig. 3.21: Tectonic setting based upon the compositional range of recent volcanic rocks (after Pearce et al., 1977). Data points plot in the island-arc and active continental margin field. Labelled data: this study. ( $\mathrm{SiO}_{2}$ range: $51-56$ wt\%).

Insight to the genesis of the magmas represented by the analysed samples is provided by the trace element characteristics. The Zr and Hf values were taken from XRF analyses. Zr and Hf are strongly underestimated in the ICP-MS values because zircon is not dissolved during acid digestion procedure (which is especially adapted to mafic/ultramafic rocks). All ICP-MS analyses display anomalously low Zr values on chondrite- or primitive mantle-normalized diagram, compared with XRF analyses. This is evidence that Zr (and Hf ) hosted by zircon was missed by ICP-MS. On the other hand, the XRF Zr values are about at the same level as MREE on normalized diagrams and display only subtle, positive or negative, anomalies. Therefore, ICP-MS Hf values were deleted from the data set and the Zr values were replaced by the XRF ones.

The trace element patterns of the analysed samples, normalized to primordial mantle abundance after Sun (1980) and McDonough et al. (1992) show similar patterns for the gabbros and diorites, whereas the tonalite/granites are depleted in $\mathrm{Sm}, \mathrm{Ti}, \mathrm{Tb}$ and Y and enriched in Rb, Ba, Th, U and K (Fig. 3.22a) The strong negative spike of Ti in sample $91 / 08 / 01$ can be also caused by an effect of limit of detection ( 0.05 weight $\% \mathrm{TiO}_{2}$ ). The gabbros have some similarity with average MORB, except the strong $\mathrm{Ba}, \mathrm{Sr}$ and P enrichment. The Chilas complex is more enriched in $\mathrm{Rb}, \mathrm{Ba}, \mathrm{Th}, \mathrm{U}$ and K than the average Metaplutonic Complex plutonic (average of gabbros and diorites). For elements right of Nb the pattern is very similar for Metaplutonic Complex and Chilas plutonics with values of 5 to 10 times higher than primordial mantle values (Fig. 3.22b). The Kamila high- Ti and low-Ti amphibolites (Khan et al., 1993) have higher Rb and Ba values than the average Metaplutonic Complex plutonics. The Kamila high-Ti amphibolites are interpreted as remnants of a pre-arc oceanic crust that formed within an intraoceanic setting (Treloar et al., 1996). The Kamila low-Ti amphibolites have geochemical characteristics similar to subduction related origin and are interpreted as most primitive arc-related magmatic rocks within the KAC (Treloar et al., 1996). They show a similar pattern as the average Metaplutonic Complex plutonics, except significant higher Th and K contents (Fig. 3.22b).

Trace element patterns occurring in the KAC are variable for volcanics and plutonics of different areas (Fig. 3.22c). The Jaglot Group volcanics demonstrate geographical and compositional provincialism within calc-alkaline to mildly alkaline compositions (Treloar et al., 1996). The Hunza volcanics have komatiite-high-Mg basalt to high-Mg andesite compositions with typical subduction related trace element patterns (Petterson \& Windley, 1991; Treloar et al., 1996). The Utror and Mankial volcanics yield highly evolved calc-alkaline trace element compositions and show with andesites to rhyolites a wide compositional range typical for mature Andean type margins (Treloar et al., 1996). The western arc volcanics (Petterson et al., 1990) are less evolved than the Utror volcanics and may reflect lateral compositional variations (Treloar et al., 1996). A direct genetic link, meaning that one the above mentioned volcanics reflect the upper crustal volcanic equivalent to the lower crustal plutonics of the Metaplutonic Complex is not evident.

## Trace element plot normalized to primordial mantle





Fig. 3.22: Trace element plots normalized to primordial mantle after McDonough et al. (1992). a) Upper diagram: this study; b) centre diagram: average Metaplutonic Complex, Zeilinger, this work, $n=8$; average Chilas 1, Mikoshiba et al., 1999; Jan \& Howie, 1981; Khan et al., 1993, $n=21$; average Chilas 2, Khan et al., 1989,1993, $n=7$; Kamila HiTi, Khan et al., 1993, n=7; Kamila LoTi, Khan et al., 1993, n=8; shaded field: gabbros and diorites, this work; c) bottom diagram: mean values of: Hu_hMg: Hunza valley high Mg (Petterson et al., 1991, n=6); Hu_hSi: Hunza valley high Si (Petterson et al., 1991, $n=6$ ); Ja_Ma: Jaglot Majne (Khan, 1994, n=5); Ja_Pe: Jaglot Peshmal (Sullivan et al., 1993, n=5); Ja_Th: Jaglot Thelichi (Khan, 1994, n=8); Ma: Mankial (Sullivan et al., 1993, $n=5$ ); Ut int: Utror intermediate (Sullivan et al., 1993, $n=5$ ); We_hSi: western arc high Si (Petterson et al., 1991, $n=13$ ); We_lSi: western arc low Si (Petterson et al., 1991, $n=5$ ); shaded field: gabbros and diorites, this work.

The trace element and REE pattern of the gabbro-diorite sequence (Fig. 3.23a) can be subdivided in two segments: 1) to the right of La: REE, $\mathrm{Sr}, \mathrm{Zr}, \mathrm{Ti}$ and Y show relatively steady evolutions, with only subtle negative anomalies of the HFSE (high field strength elements) Zr and Ti , and positive anomalies of Sr (these anomalies may reflect plagioclase accumulation in some samples or/and they may represent the arc-type signature of parental melts); 2) to the left of La: most of the highly incompatible elements (HIE, notably $\mathrm{Rb}, \mathrm{Th}, \mathrm{U}, \mathrm{Nb}$ and Ta ) are strongly depleted relative to LREE. Exceptions are Ba (enriched) and Cs (enriched, but not for RK1 and RK2). The more LREE-enriched samples (RK2 and RK1) are the most depleted in Rb, Th and U. These elements are probably not direct correlated with LREE. The granites and tonalite are clearly distinguished from the gabbro-diorite sequence by gradual enrichment from moderately incompatible elements (HREE and Y) to the most incompatible elements (HIE, including Cs and Ba ). Positive anomalies of $\mathrm{Cs}, \mathrm{Ba}$ and Sr are yet distinguishable on this overall pattern, reminiscent of those observed in the gabbro-diorite sequence.

The REE patterns of the analysed samples, normalized to continental crust (Fig. 3.23b) and chondrite abundances (Fig. 3.23c) show similar patterns for the gabbros and diorites. The granite and tonalite patterns are relatively flat (except the Eu anomaly) when normalized to continental crust (Fig. 3.23b).

The gabbros and diorites yield a relatively high level of REE abundance ( $>10 \mathrm{x}$ chondrites). The lower gabbros are indistinguishable from the upper diorites with respect to HREE. However, they are less enriched in LREE, with more variable LREE segment. Sample 91/03/03 shows rather flat LREE segment (La-Sm), sample 91/03/02 is slightly depleted in LREE relative to MREE (La-Pr compared with Nd-Sm) and sample 16/01/01 is slightly enriched in La-Ce relative to Pr-Sm.

b) NORMALIZED TO CONTINENTAL CRUST, REE ONLY

c) NORMALIZED TO CHONDRITES - REE ONLY


|  | 16/01/01 |
| :---: | :---: |
|  | 91/03/02 |
|  | 91/03/03 |
|  | 71/03/02 |
| $\bigcirc$ | RK2 |
| $\square$ | RK1 |
|  | 35/02/01 |
|  | 91/07/01 |
|  | 36/01/01 |
|  | 91/08/01 |
|  | 00/10/01 |

Fig. 3.23: Gabbros (filled symbols), diorites (open symbols) and tonalite/granite (cross symbols) presented in normalisation diagrams. a) upper diagram: trace and $R E$ elements normalized to primitive mantle values; b) centre diagram: REE normalized to continental crust; c) bottom diagram: REE normalized to chondrites.

### 3.2.4 Interpretations

The lesser LREE enrichment in the lower gabbros, compared with the upper diorites, may partly result from a more "cumulitic" character. In particular, the flat or slightly depleted LREE segments of samples $91 / 03 / 02$ and $91 / 03 / 03$ may be explained by cumulitic amphibole. In addition, samples 16/01/01 and 91/03/03 show a small but significant positive Eu anomaly indicative of plagioclase accumulation (in sample 91/03/03, plagioclase accumulation is also attested by the anomalously high alumina content $-22.1 \%$, not consistent with a "melt" composition). However, these cumulitic characters are relatively subtle and these gabbros cannot be considered as a proper "cumulate" sequence. The slight enrichment of La-Ce relative to $\mathrm{Pr}-\mathrm{Sm}$ in sample $16 / 01 / 01$ is significant, because the sample is the lowest in the sequence and it is clearly enriched in highly incompatible elements such as $\mathrm{Rb}, \mathrm{Ba}, \mathrm{Th}$ and U and more depleted in Nb and Ta compared with other samples (Fig. 3.23a). This indicates that this gabbro is less depleted than the other gabbros. The selective LREE enrichment in this sample probably results from pervasive infiltration of incipient melt fractions.

The diorites have relatively flat HREE segment (Dy-Lu) and are selectively enriched in LREE relative to HREE, with a narrow range of LREE contents (about 50 x chondrites for La) (Fig. 3.23c). They are devoid of positive Eu anomaly in spite of their relatively high plagioclase content. This indicates that the diorites are not cumulates. They have probably crystallized rapidly in nearly closed system and preserved a "parental" melt composition. The "parental" melt composition was LREE-enriched with a major element composition ranging from quartz-tholeiitic for samples RK2 and RK1 to andesitic for samples 35/02/01 and 91/07/01.

There is yet a clear differentiation trend from the lower gabbros to RK2-RK1 to the upper diorites, especially for major elements. If the sequence derives from a single parental melt, this would imply significant fractional crystallisation, which - in turn would require that at least the less differentiated facies have a cumulitic character. Part of the differentiation by fractional crystallization probably occurred at deeper level in the Granulitic gabbros (J-L.Bodinier, pers. com.). Therefore the studied sequence has probably crystallized from heterogeneous, already differentiated "parental" melts.

The granites are characterized by fractionated REE patterns, with a marked enrichment of LREE relative to HREE. Compared with the gabbro-diorite sequence, the granites have comparable LREE content and mainly differ by a strong HREE depletion. In sample $36 / 01 / 01$, which is rich in alumina ( $20.2 \%$ ) and shows a positive Eu anomaly and a strong Sr spike (Fig. 3.23a), the HREE depletion might partly result from plagioclase accumulation. However, this hypothesis cannot be applied to sample $91 / 08 / 01$ and $00 / 10 / 01$. Therefore HREE depletion in granites is more likely the reflection of their equilibration with HREE enriched residual minerals such as garnet and/or amphibole. Together with field evidence, this suggests that the granites were generated by (hydrous) partial melting of rocks similar to the Granulitic gabbros or/and the hornblendite enclaves.

The trace element composition of the gabbro-diorite sequence normalized to the average oceanic arc volcanics composition (based on New Hebrides and Mariana data) displays an increasing depletion of the Kohistan gabbro-diorites of incompatible elements (Fig. 3.24a). The MREE and HREE ( Sm to Lu ), Nb and Ta are undepleted (Kohistan gabbrodiorites $(\mathrm{KGD}) /$ oceanic arc volcanics $(\mathrm{OAV})=1.06 \pm 0.08)$. The LREE (La to Nd), Zr and Sr are slightly depleted, with increasing depletion degree from Nd to La (KGD / $\mathrm{OAV}=0.9$ to 0.6 ). Ba and Cs are a little depleted (KGD / OAV $=0.4$ to 0.2 ) and Rb , Th and U are strongly depleted ( $\mathrm{KGD} / \mathrm{OAV}=0.13$ to 0.10 ). In contrast with lithophile trace elements, Pb is enriched in the Kohistan gabbro-diorites relative to the oceanic arc volcanics by a factor 5-6 (see Pb spike in figure Fig. 3.24b). The difference is probably the reflection of the siderophile character of this element. Pb is probably mainly hosted by a sulphide component. This can be explained by sulphide precipitation during crystallisation, due to melt saturation, or selective mobilization during a late partial melting event. Sulphide precipitation is unlikely as the lithophile elements indicate that there was very little amount of residual liquid segregated from the gabbro-diorite sequence. The hypothesis that the Kohistan parental melt was substantially different from New Hebrides and Mariana arc volcanics in term of Pb cannot be disregarded. The Tonga arc andesites and basalts are similar to the Kohistan gabbro-diorites (Fig. 3.26b) with patterns overlapping for $\mathrm{Cs}-\mathrm{Th}$ and HREE segments ( Gd to Lu ). The elements from La to Ti are more enriched in the Kohistan gabbro-diorites than in the Tonga arc
andesites and basalts, but follow a similar pattern. U and Nb values are different in the Tonga and Kohistan arc systems.

A comparison between estimated average continental crust composition (Rudnick \& Fountain, 1995) and the Kohistan gabbro-diorite normalized to average island arc volcanics shows that for several elements the continental crust and the Kohistan gabbrodiorite sequence have almost symmetrical, opposite arc-normalized patterns (Fig. 3.24 c ). The element depletion in the gabbro-diorite sequence is mirrored by enrichment in continental crust (e.g., $\mathrm{Rb}, \mathrm{Ba}, \mathrm{Th}, \mathrm{U}, \mathrm{La}$ ). HREE segments ( Gd to Lu ) are remarkably similar in arc volcanics, Kohistan and continental crust. However, some elements show a different behaviour: 1) the strong enrichment of Nb and Ta in arcnormalized continental crust is not coupled with a depletion of these elements in the Kohistan gabbro-diorites; 2) the strong enrichment of Pb in Kohistan (Fig. 3.23a) does not coincide with a negative anomaly in the continental crust and 3) the negative anomaly of Sr in the continental crust is not associated with a positive Sr anomaly in Kohistan. The particular behaviour of these four elements probably reflects the preferential partitioning in sulphides $(\mathrm{Pb})$, Ti-oxides $(\mathrm{Nb}$ and Ta$)$ and plagioclase $(\mathrm{Sr})$. When these elements are disregarded, the symmetrical patterns suggest that the processes involved in continental crust maturation are somehow analogous to the differentiation mechanism responsible for incompatible element depletion in the Kohistan lower crust. However, the degree of crustal maturation which was achieved in the Kohistan island arc prior to accretion with India, is not known.




Fig. 3.24: Comparison between the lower KAC plutonics and recent island arc volcanics: a) normalized to average oceanic arc (New Hebrides and Mariana); b) Tonga arc and KAC gabbros and diorites and c) the lower KAC gabbros and diorites compared to average continental crust (Rudnick \& Fountain, 1995) normalized to average oceanic arc (New Hebrides and Mariana).

Differences of the Kohistan gabbro-diorites to ophiolitic sequences can be demonstrated by comparing them to gabbros from Oman (values provided by Jean-Louis Bodinier, personal communication). When normalized to an olivine gabbro the strong relative enrichment for Cs to Zr , except for Sr , is notable (Fig. 3.25a).


Fig. 3.25: Comparison between the lower KAC plutonics and Oman gabbros. a) upper diagram: normalized to olivine Gabbros; b) centre diagram: REE normalized to upper gabbros; c) bottom diagram: trace and RE elements normalized to upper gabbros. Normalization values provided by Jean-Louis Bodinier, personal communication.

The MREE and HREE are flat with values 3 to 8 times higher for the gabbro-diorites and around 1 ( 0.5 to 2 ) for the tonalite/granites. Comparing the Kohistan gabbro-diorites with upper gabbros from Oman the MREE and HREE are only slightly relative enriched by a factor of 2 . The LREE instead show an increase with decreasing atomic number (Fig. 3.25b). The tonalite/granites are strongly enriched in incompatible elements and depleted in MREE and HREE. The gabbros have a strong Pb spike, the granites a strong Ti negative anomaly (Fig. 3.25c). The diorites RK1 and RK2 have Cs, Rb, Th, U values similar to the upper Oman gabbros. Despite of potential different "parental" melt sources the differences suggest also different processes during "maturation".

One additional sample ( $00 / 30 / 01$ ) representing a diorite from the Indian plate was analysed. In contrast to the diorites of the Metaplutonic Complex the trace element pattern is similar to average upper crust (Weaver \& Tarney, 1984) (Fig. 3.26a).


Fig. 3.26: a) trace element plot of the Indian plate metadiorite normalized to primordial mantle compared with MORB (Saunders \& Tarney, 1984), OIB (Sun, 1980), upper crust (Taylor \& McLennan, 1981), lower crust and average crust (Weaver \& Tarney, 1984); b) REE values of the metadiorite normalized to continental crust.

The REE pattern normalized to continental crust (Fig. 3.26b) is relatively flat and enriched by a factor of 2 to 5 , supporting the similarity to continental crust values.

Using the elements $\mathrm{Nb}, \mathrm{Y}, \mathrm{Ta}$ and Rb for tectonic discrimination after Pearce et al. (1984) the difference between the arc tonalite/granites, plotting in the volcanic arc granite field, and the Indian plate diorite, plotting in the within plate granite field, is emphasized (Fig. 3.27).


Fig. 3.27: Discrimination plots after Pearce et al. (1984) using a) Nb versus Y, b) Ta versus $Y b, c$ ) $R b$ versus $Y b+N b$ and $d$ ) $R b$ versus $Y b+T a$. The granites/tonalites from the Metaplutonic Complex fall in the VAG (volcanic arc granite) fields. The metadiorite emplaced into the Indian plate instead belongs to the WPG (within plate granite) tectonic environment. (syn-COLG: syn-collisional granite; ORG: ocean ridge granite).

### 3.2.5 Discussion

The large ion lithophile elements (LILEs) $\mathrm{Cs}, \mathrm{Ba}$ and Sr display nearly the same variations as the other highly incompatible elements (HIEs). The lowest Cs values and the smallest Sr anomalies are observed in samples RK2 and RK1 which are the most depleted in Rb , Th and U (Fig. 3.23a). The highest Ba concentration is observed in sample 16/01/01 (base of the Sarangar gabbro) which is the most enriched in Rb, Th and U (Fig. 3.23a). This is indicating that $\mathrm{Cs}, \mathrm{Ba}$ and Sr were mobilized together with the other HIEs during incipient melting. Alternatively, similar to Pb , these elements systematically display positive anomalies relative to other HIEs and LREE on primitive mantle-normalized diagrams (with the only exception of Cs in samples RK2 and RK1, Fig. 3.23a). These anomalies are observed in the whole sequence from Jijal to Dasu, including the Granulitic gabbro and the granites (Fig. 3.29a,b). These anomalies represent most likely the primary, arc-type, geochemical signature of the mantle melt parental to the whole sequence.

The strong HIE depletion observed in most of the gabbro-diorite sequence may be considered as a cumulitic character, resulting from the segregation of residual liquid during fractional crystallization. However, this process hardly accounts for the strong depletion of $\mathrm{Th}, \mathrm{U}, \mathrm{Nb}$ and Ta relative to LREE in samples with typical "melt-type" REE patterns (e.g., RK2 and RK1). Conversely, the less differentiated samples (the lower gabbros, which have the lowest LREE contents) are less depleted in HIE than the other samples. Mineral accumulation during fractional crystallisation cannot explain the negative correlation between HIE and LREE. Based on the negative correlation (to samples RK1 and RK2) between HIE enrichment and distance from Patan (Fig. 3.28), a process involving incipient partial melting of rocks similar to the Granulitic gabbro and upward pervasive migration of HIE-rich small melt migrations through the Sarangar gabbro is likely.

Element ratios vs. distance from Patan (in km) (Rb/Ba, Th/Nb, U/La)




| Legend 16/01/01 |
| :---: |
| - 91/03/02 |
| 4 91/03/03 |
| * 71/03/02 |
| O RK2 |
| - RK1 |
| $\triangle$ 35/02/01 |
| $\diamond$ 91/07/01 |
| * 36/01/01 |
| + 91/08/01 |
| $\times 00 / 10 / 01$ |

Fig. 3.28: $U / L a, R b / B a$ and $T h / N b$ ratios versus sample distance to Patan of the gabbros, diorites and granites in the Metaplutonic Complex (see Fig. 3.16 and 3.17 for locations.

The Granulitic gabbro (and the enclosed hornblendites) are strongly depleted in HIE, particularly Th (Fig. 3.29a). Strong depletion of Th and $U$ (and other HIE) is a classic feature in granulitic terranes (Burton \& O'Nions, 1988). HIE depletion in granulitic terranes has been ascribed to prograde granulitic metamorphism (Burton \& O'Nions, 1988), due to dehydration associated with mica breakdown into garnet, as the latter cannot host much Th, U and "large ion lithophile elements" (LILEs). However, this process is unlikely, as there is no mica reported in the protolith of the Granulitic gabbro. Therefore, an explanation involving 1) incipient melting of Granulitic gabbros and hornblendites from the Jijal complex triggered/enhanced by dehydration of the hornblendites, and 2) the formation of silica- and water-rich small melt fractions migrating upwards is favoured. If the "lesser" HIE depletion recorded at the base of the

Sarangar gabbro is attributed to a re-enrichment process, it is then implied that the "most depleted" samples from the top of the Sarangar gabbro are more representative of the "protolith" (i.e., the original gabbro-diorite composition, before re-enrichment). This has to be proven by more analyses from the top of the Sarangar gabbro.

Both the Granulitic gabbro and the studied gabbro-diorite section have been strongly depleted in HIE (Fig. 3.29a) due to (hydrous) incipient melting, upward pervasive migration of small volume melts and local extraction of granites. In this scheme, the HIE/LREE ratio of a given sample will depend on a balance between HIE depletion by incipient melt extraction (earlier) and HIE re-enrichment by migrating melts (later). The Granulitic gabbro was more depleted than the gabbro-diorite sequence (Fig. 3.29a) and notably more than the granites/tonalites (Fig. 3.29b), and not significantly re-enriched. Conversely, the upper diorites were less depleted. The base of the lower gabbro sequence (Sarangar) was probably re-enriched by melts from Jijal complex, as attested by the selective enrichment of $\mathrm{La}-\mathrm{Ce}$ relative to $\mathrm{Pr}-\mathrm{Sm}$ in sample 16/01/01, which cannot be accounted for by a "lesser depletion degree". A direct comparison between the granites and the Granulitic gabbro, the hornblendites and the garnet pyroxenites was plotted (Fig. 3.29b,c) in order to see if the formation of granites could have taken place by partial melting of the Jijal complex. Partial melting would leave a HIE-depleted residue, which can be observed in the Jijal complex. The trace element signature of the granites is consistent with the partial melting hypothesis suggested for their HREE depletion.


Fig. 3.29: Comparison between a) Sarangar gabbro, garnet pyroxenites, garnet hornblendites, and upper granulitic gabbro; b) granite/tonalite, garnet pyroxenites, upper granulitic gabbro; and hornblendites; c) granite/tonalite, garnet pyroxenites, upper granulitic gabbro and hornblendites; REE only. Data of garnet pyroxenites, upper granulitic gabbro and hornblendites provided by Jean-Louis Bodinier and C. Garrido, personal communication.

### 3.2.6 Conclusions

Samples from the Metaplutonic Complex cover calc-alkaline to tholeiitic differentiation trend gabbros, diorites, tonalites to granites, emplaced in a island (volcanic-) arc setting (Fig. 3.19, Fig. 3.21). Certain major element variations are aligned on correlations (when plotted against $\mathrm{SiO}_{2}$ ) underlining the dominant calc-alkaline character of the magmas (Fig. 3.18). The Chilas complex is noticeably calc-alkaline (Khan et al., 1993), indicating a less FeO rich calc-alkaline composition (Fig. 3.19). Trace element patterns of the Chilas complex and other volcanic and plutonic rocks from the Kohistan Arc complex (e.g. Hunza, Jaglot) are incoherent, reflecting the long lasting evolution of magmas, alteration processes and changes in tectonic settings (Fig. 3.22b,c).

The gabbros and diorites from the Metaplutonic Complex yield relatively similar trace element and REE patterns. Assuming they preserved a geochemical signature from the melts and/or melt-source, this indicates a common arc-type melt source. The relatively to the gabbros and diorites HIE enriched and HREE depleted granites and tonalite preserved positive anomalies in $\mathrm{Cs}, \mathrm{Ba}, \mathrm{Sr}$ and Pb , similar to the positive anomalies in the gabbros and diorites. Therefore, they are considered to be genetically related to the gabbros and diorites. The arc-type character of the analysed samples (only the Metaplutonic Complex ones) is evident by comparing them to recent oceanic arcs (Fig. 3.24). Despite small differences (especially for the HIE), the overall geochemical signature is similar to volcanics of the Tonga, New Hebrides and Mariana arc systems. The observed geochemical signature is different from 1) gabbros from the Oman ophiolite complex, excluding that the analysed rocks represent an ocean floor, and 2) from diorites from the Indian Plate (at Duber Bazar), excluding that the analysed rocks have been intruding a continental terrane during a rifting phase.

The Metaplutonic Complex gabbros and diorites, the island arc volcanics and the continental crust have almost identical concentrations of moderately incompatible elements, whereas the HIE depletion of the Metaplutonic Complex gabbros and diorites tends to "balance" the enrichment of the continental crust. The processes whereby a dominantly basaltic/transitional, juvenile arc crust evolves with time towards more acidic and HIE-enriched compositions akin to continental crust can be studied in the KAC. A superposition of primary magmatic, melt perculation and metasomatic
processes are involved in the arc evolution. Based on the preliminary dataset, fractional crystallisation of the plutons/laccoliths seems to be a minor process in the Metaplutonic Complex, because none of the plutonic suites are cumulates with a primary compositional layering. This indicates that there was no common large magma chamber involved in the formation of the gabbros and diorites. Instead, a succession of melt batches with various differentiation degrees, preserving a common "source" signature, is more likely. A process that can partly explain the observed geochemical signatures is the incipient partial melting of lower arc section. Possibly triggered by dehydration of hornblendite facies a pervasive upward migration of small melt fraction would be the consequence. The effect on major elements would be very little and would show no effect on compatible trace elements at all. But a "draining" effect would transport Th, U and LILEs from the lower parts of the arc to shallower levels leading to their selective enrichment relative to slightly less incompatible elements such as LREE. Partial melting at the base of the arc is most likely the process responsible for the formation of granites. If partial melting was a process active continuously at the base of the arc crust or happened as distinct events has to be proofed by dating different granites. Partial melting, probably partly connected with the previous process, will efficiently remove granitic components from the base of the arc and transport them up to shallower levels.

The analysed rocks, even representing only a small part of the KAC, demonstrate a part of a set of spatial (lower section: gabbroic rocks $\rightarrow$ upper section: granitic rocks) and chronological trends (gabbros older than granites; see ages in chapter 4) in the overall geochemical evolution of the arc.

## 4. NEW GEOCHRONOLOGICAL CONSTRAINTS ON THE KOHISTAN ARC COMPLEX

Chapter 4.1 is "Changing mantle sources during island arc magmatism: $\mathrm{U}-\mathrm{Pb}$ and Hf isotopic evidence from the Kohistan arc complex, Pakistan" by U. Schaltegger, G. Zeilinger, M. Frank \& J.-P. Burg (submitted to Geology, 2001).

### 4.1 MULTIPLE MANTLE SOURCES DURING ISLAND ARC MAGMATISM: U-Pb and HF

 isotopic evidence from the Kohistan arc complex, Pakistan
### 4.1.1 Abstract

Supra-subduction volcanism in oceanic island arcs stems from the partial melting of the underlying mantle wedge and possibly from the subducting slab. The parental magmas undergo differentiation and eventually build up the juvenile arc crust. The fossil island arc sequence of the Kohistan Arc Complex (KAC) preserves a record of arc magmatism with a duration of over 20 million years. Thus, the KAC provides a unique opportunity to investigate potential changes in magma sources during arc evolution. Here we present high-precision $\mathrm{U}-\mathrm{Pb}$ ages and initial Hf isotopic compositions of zircons from mafic to felsic rocks of the KAC. Hafnium isotopes in zircon have been shown to be a sensitive isotopic tracer of subtle changes in the source regions of magmatic rocks. Three magmatic pulses tapping geochemically different reservoirs are distinguished from initial to evolved stages of the KAC. Magmatism during initiation of the arc at 99-92 Ma was generated by partial melting of mantle with MORB-type isotopic characteristics. Extensional magmatism 85 Ma ago tapped a more fertile mantle source, most likely consisting of a $>600 \mathrm{Ma}$-old metasomatically enriched mantle, or of mantle contaminated by an old sedimentary component. Remelting of the deep base of the arc 82 Ma ago resulted in felsic peraluminous dykes, which again show MORB-type isotopic compositions inherited from the initial stage. The isotopic results demonstrate that arc development, from initial to evolved stages, is characterized by several changes in melt source region. They show that there was subordinate continental influence and negligible importance of slab components for the Hf budget during the generation of the Kohistan Arc Complex.

### 4.1.2 Introduction

Island arc magmatism, as generated above a subducting oceanic plate, may either be derived from the slab or from the overlying mantle wedge. Evidence from active subduction systems suggests that slab melting is a rare process restricted to very hot and young slabs (Kay, 1978; Sigmarsson, 1998). Melts are commonly formed in the overlying mantle wedge, triggered by fluid release from the subducted slab. The mantle source of primary arc melts is considerably more depleted than typical mid-ocean ridge basalt (MORB)-type reservoirs (e.g. Davidson, 1996) and, therefore, requires unrealistically high temperatures to melt further. Continuous arc magmatism over millions of years implies either (re)fertilization of previously depleted mantle, or the tapping of new fertile reservoirs in the development of the arc. On the other hand we know of arcs with long lifetimes, e.g. 45 million years in the case of the Bonin arc, that apparently show a stable pattern of magma generation throughout their history (Arculus, 1994).


Fig. 4.1: Schematic geological map of the southern part of the Kohistan Arc Complex, after Burg et al. (1998). Analyzed material was sampled along the two cross sections of Indus and Swat valleys, respectively. Rocks of the central and northern Kohistan Complex - metasediments and associated volcanic sequences intruded by plutons of the Kohistan Batholith (Jan \& Howie, 1981; Petterson et al., 1991b) were not investigated.

We investigated this paradox by studying time-resolved melt formation in a long-lived island arc system, using high-precision $\mathrm{U}-\mathrm{Pb}$ dating and determination of initial Hf
isotopes in zircon. The Kohistan Arc Complex (KAC) in northern Pakistan (Fig. 4.1) shows an undisturbed geological record from initial to more mature stages of arc formation; it was formed during Cretaceous north-dipping subduction in the Tethys Ocean between the Indian and Eurasian continental plates (Tahirkheli et al., 1979; Coward et al., 1986). The KAC formed during at least two major tectonic stages (Burg et al., 1998): an initial arc-generating stage starting before 100 Ma and lasting to at least 90 Ma ; and an extensional intra-arc rifting stage, during which huge volumes of noritic gabbros were emplaced between 85 and 80 Ma . Profiles in the Indus and Swat valleys provide a rare opportunity to investigate the source regions of melts emplaced during the first 20 million years lifetime of this island arc (Fig. 4.1).

### 4.1.3 Lithological units of the Kohistan Complex

Rocks of the Jijal Complex represent the sub-arc mantle-crust transition (Burg et al., 1998); peridotites, pyroxenites and hornblendites are overlain by garnet- and 2 pyroxene metagabbroic granulites, representing the deepest exposed crustal rocks (Bard et al., 1980).

The Metadioritic Sequence formed during early arc buildup and consists of a midcrustal association of imbricate and variously metamorphosed gabbros, diorites and granodiorites with a geochemical trend between calc-alkaline and tholeiitic compositions (Treloar et al., 1990; Khan et al., 1993). Diorites and granites, the most evolved lithologies, form intrusive sheets within the metagabbros of the Metadioritic Sequence.

A large body of calc-alkaline noritic gabbros, the Chilas complex, intruded into volcanic and sedimentary rocks in the central part of the arc (Treloar et al., 1996). Structural relationships suggest emplacement during the younger intra-arc rifting stage.

Subsequently, a number of kyanite-bearing pegmatoid dykes of peraluminous geochemical character intruded into the Metadioritic Sequence of the Swat valley (Jan \& Karim, 1995).

### 4.1.4 U-Pb and Hf isotopic results

Sample locations and petrographic descriptions of studied samples are given in Appendix 4. The oldest rock dated by the U-Pb zircon technique is the Sarangar gabbro
with an age of $98.9 \pm 0.4 \mathrm{Ma}\left(\text { all }{ }^{206} \mathrm{~Pb}\right)^{238} \mathrm{U}$ ages; Tab. 4.1; Fig. 4.2a). It intruded into granulite-facies garnet-pyroxene gabbros of the Jijal complex, previously dated by the Sm-Nd pyroxene-garnet technique at $91.0 \pm 6.3 \mathrm{Ma}, 94.0 \pm 4.7 \mathrm{Ma}, 95.7 \pm 2.7 \mathrm{Ma}$, and $94.6 \pm$ 5.3 Ma (Anczkiewicz \& Vance, 2000; Yamamoto \& Nakamura, 1996, 2000). Dated intrusions within the Metadioritc Sequence include a granite sheet (97.1 $\pm 0.2$ Ma; Fig. 4.2a) and a diorite stock ( $91.8 \pm 1.4 \mathrm{Ma}$; Fig. 4.2b), both from the Indus valley. The Metadioritic Sequence is intruded by two different types of lithologies; a kyanitebearing peraluminous dyke with very low-U zircons was dated at $82.8 \pm 1.1 \mathrm{Ma}$ (Fig. 4.2 b ) and the Chilas gabbronorite yielded a precise age of $85.73 \pm 0.15 \mathrm{Ma}$ (Fig. 4.2c).




Fig. 4.2: U-Pb concordia diagrams with age determinations for the Sarangar gabbro and a granite (a) and for a diorite (b) from the Indus Valley; and a kyanite-bearing dyke (b) and the Chilas gabbronorite (c) from the Swat valley. Ellipses denote analytical uncertainty at the 2 sigma level of individual analyses.

The new age for the Chias gabbronorite lends support to the 84 Ma U-Pb zircon age obtained by Zeitler et al. (1980). The new ages of both felsic dykes and Chilas Complex coincide with Ar-Ar hornblende ages for the Metadioritic Sequence ( 85 to 80 Ma ), which were inferred to date cooling of these rocks below temperatures of $550{ }^{\circ} \mathrm{C}$ (Treloar et al., 1989a; Wartho et al., 1996). These ages are older than obduction of the Kohistan Complex onto the Indian plate, which started about 60 Ma ago (Coward et al., 1986).

The Hf isotopic compositions were determined from the same zircon microfractions used for $\mathrm{U}-\mathrm{Pb}$ dating (see Table 4.1). Due to typically very low ${ }^{176} \mathrm{Lu} /{ }^{177} \mathrm{Hf}$ isotopic ratios of 0.005 in zircon, they represent initial Hf isotopic compositions at the time of melt crystallization. The $\varepsilon H f$ values of Sarangar gabbro $(+14.1 \pm 0.1)$, granite $(+14.7 \pm$ 1.6) and diorite $(+14.0 \pm 2.2)$ are analytically indistinguishable and are representative of a MORB-type mantle reservoir (Patchett \& Tatsumoto, 1980; Nowell et al., 1998). The Chilas gabbronorite yielded a significantly lower $\varepsilon H f$ value of $+10.4 \pm 0.05$, pointing to a more fertile mantle source with time-integrated lower $\mathrm{Lu} / \mathrm{Hf}$, whereas the kyanitebearing dyke shows a value at the level of MORB-type mantle $(+13.9 \pm 0.3)$.

| Nr. Concentrations |  |  | Pb <br> non <br> rad. <br> [pg] | $\begin{gathered} \mathrm{Th} / \mathrm{U} \\ \S \end{gathered}$ | Atomic ratios |  |  |  |  |  | $\begin{gathered} \text { Age } \\ { }^{206} \mathrm{~Pb} / \\ { }^{238} \mathrm{U} \\ * \dagger \\ \hline \end{gathered}$ | ${ }^{176} \mathrm{Hf} /{ }^{177} \mathrm{Hf}$ | $\pm 2 \sigma$ | $\varepsilon \mathrm{EHf}$ <br> (T) |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | U | $\begin{gathered} \mathrm{Pb} \\ \text { rad. } \end{gathered}$ |  |  | $\begin{gathered} { }^{206} \mathrm{~Pb} / \\ { }^{238} \mathrm{U} \end{gathered}$ | $\begin{aligned} & \text { Error } \\ & \pm 2 \sigma \end{aligned}$ | $\begin{gathered} { }^{207} \mathrm{~Pb} / \\ { }^{235} \mathrm{U} \end{gathered}$ | $\begin{gathered} \text { Error } \\ \pm 2 \sigma \end{gathered}$ | $\begin{gathered} { }^{207} \mathrm{~Pb} / \\ { }^{206} \mathrm{~Pb} \end{gathered}$ | $\begin{aligned} & \text { Error } \\ & \pm 2 \sigma \end{aligned}$ |  |  |  |  |
|  | [ppm] | [ppm] |  |  | * $\dagger$ | [\%] | * | [\%] | * $\dagger$ | [\%] |  |  |  | ** |
| Sarangar Gabbro, Indus valley |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| 1 | 149 | 2.44 | 4.1 | 0.53 | 0.01553 | 0.37 | 0.1030 | 0.48 | 0.04808 | 0.31 | 99.3 | 0.283120 | 22 | 14.5 |
| 2 | 182 | 2.99 | 2.0 | 0.58 | 0.01544 | 0.33 | 0.1022 | 0.41 | 0.04803 | 0.19 | 98.8 | 0.283106 | 2 | 14.0 |
| 3 | 129 | 2.10 | 3.1 | 0.55 | 0.01545 | 0.33 | 0.1024 | 0.40 | 0.04806 | 0.17 | 98.9 | 0.283110 | 7 | 14.2 |
| 4 | 189 | 3.08 | 1.9 | 0.53 | 0.01549 | 0.39 | 0.1026 | 0.65 | 0.04805 | 0.52 | 99.1 | 0.283311 | 2 | 14.1 |
|  |  |  |  |  |  |  |  |  |  |  |  | Mean: |  | $14.1 \pm 0.1$ |
| Granite, Indus valley |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| 5 | 358 | 5.82 | 1.7 | 0.60 | 0.01517 | 0.41 | 0.1006 | 0.81 | 0.04807 | 0.68 | 97.1 | 0.283096 | 28 | 13.7 |
| 6 | 507 | 8.98 | 8.5 | 0.55 | 0.01517 | 0.50 | 0.1005 | 1.68 | 0.04807 | 1.55 | 97.0 | 0.283114 | 16 | 14.3 |
| 7 | 1377 | 22.25 | 2.2 | 0.57 | 0.01512 | 0.37 | 0.1006 | 0.90 | 0.04797 | 0.78 | 97.3 | 0.283138 | 13 | 15.2 |
|  |  |  |  |  |  |  |  |  |  |  |  | Mean: |  | $14.7 \pm 1.6$ |
| Diorite, Indus valley |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| 8 | 16.26 | 0.25 | 2.8 | 0.68 | 0.01435 | 2.12 | 0.0869 | 31.00 | 0.04392 | 29.50 | 91.8 | 0.283141 | 15 | 15.2 |
| 9 | 38.20 | 0.59 | 5.8 | 0.68 | 0.01426 | 0.86 | 0.0895 | 7.94 | 0.04550 | 7.54 | 91.3 | 0.283112 | 7 | 14.1 |
| 10 | 33.23 | 0.53 | 1.0 | 0.67 | 0.01442 | 0.71 | 0.0964 | 4.19 | 0.04848 | 3.97 | 92.3 | 0.283071 | 13 | 12.7 |
|  |  |  |  |  |  |  |  |  |  |  |  | Mean: |  | $14.0 \pm 2.2$ |
| Kyanite-bearing dyke, Swat valley |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| 11 | 6.74 | 0.08 | 1.1 | 0.05 | 0.01304 | 2.40 | 0.0802 | 16.30 | 0.04459 | 15.70 | 83.5 | 0.283108 | 8 | 13.8 |
| 12 | 2.75 | 0.03 | 1.5 | 0.01 | 0.01299 | 2.16 | 0.0730 | 36.50 | 0.04077 | 35.20 | 83.2 | 0.283121 | 12 | 14.2 |
| 13 | 5.29 | 0.06 | 4.3 | 0.01 | 0.01272 | 2.44 | 0.0739 | 35.44 | 0.04214 | 33.80 | 81.5 | 0.283110 | 3 | 13.8 |
|  |  |  |  |  |  |  |  |  |  |  |  | Mean: |  | $13.9 \pm 0.3$ |
| Chilas gabbronorite, Swat valley |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| 14 | 113 | 1.75 | 5.0 | 0.91 | 0.01337 | 0.36 | 0.0882 | 0.45 | 0.04783 | 0.24 | 85.6 | 0.283012 | 3 | 10.5 |
| 15 | 107 | 1.74 | 4.5 | 1.09 | 0.01337 | 0.35 | 0.0882 | 0.52 | 0.04778 | 0.36 | 85.7 | 0.283010 | 2 | 10.4 |
| 16 | 99 | 1.62 | 8.8 | 1.12 | 0.01337 | 0.37 | 0.0884 | 0.61 | 0.04792 | 0.45 | 85.7 | 0.283011 | 3 | 10.4 |
| 17 | 103 | 1.67 | 1.7 | 1.08 | 0.01340 | 0.36 | 0.0883 | 0.64 | 0.04776 | 0.51 | 85.8 | N.D. |  | N.D. |
|  |  |  |  |  |  |  |  |  |  |  |  | Mean: |  | $10.4 \pm 0.05$ |

$\dagger$ Corrected for initial Th disequilibrium, using an estimated $\mathrm{Th} / \mathrm{U}$ ratio of 4 for the melt
§Calculated on the basis of radiogenic ${ }^{208} \mathrm{~Pb}{ }^{206} \mathrm{~Pb}$ ratios, assuming concordance
*Numbers refer to the last digits
${ }^{* *}$ Epsilon Hf values were calculated with $\left({ }^{176} \mathrm{Hf}{ }^{177} \mathrm{Hf}\right)_{\mathrm{CHUR}(0)}=0.282772$ (Blichert-Toft and Albarède, 1997)
Table 4.1: U-Pb and Hf isotopic data of zircons from the Kohistan Arc Complex.

### 4.1.5 Sources of arc magmatism

From previous age determinations it was assumed that the Kohistan Arc Complex was formed during a prolonged period starting before 100 Ma ago and ending with obduction onto the Indian plate 60 Ma ago or after. Samples of plutonic rocks, however, only span 20 million years from initial to evolved stages. The limited number of samples does not allow a final statement whether this duration is real or a bias of sampling statistics. In any case, the new time-resolved Hf isotopic values provide insights into the changing dynamics of melt sources and different mantle reservoirs within the 20 Myr life-time of this island arc.

The initial stage is characterized by the intrusion of gabbroic, granodioritic and dioritic lithologies. The oldest dated rock, the Sarangar gabbro, intruded into the garnet-pyroxene-granulites at temperatures and pressures in excess of $800^{\circ} \mathrm{C}$ and 1.1 GPa (Yamamoto, 1993). The 96 to $91 \mathrm{Ma} \mathrm{Sm-Nd}$ mineral isochron ages of the granulites indicate system closure during cooling at around 650 to $750{ }^{\circ} \mathrm{C}$ (Mezger et al., 1992; Anczkiewicz \& Vance, 2000). This implies that the high-grade peak conditions recorded by these granulites (Ringuette et al., 1999) were attained prior to 99 Ma .

The granite at 97 Ma is the first evolved melt that differentiated from mafic (ultramafic?) parental magmas. It is followed by the more mafic diorite intrusion 5 million years later. Thus, we speculate that the magmatic rocks were emplaced in several short pulses, each one possibly representing an individual magmatic differentiation cycle, but having the same MORB-type source as indicated by similar $\varepsilon H f$ values. Fluid metasomatism from the subducting oceanic crust leaves the Hf isotopic composition of the mantle wedge virtually unaffected, because slab fluids commonly have negative Hf anomalies (e.g. Pearce et al., 1999). Hf isotopes are therefore believed to reflect the geochemical characteristics of the mantle wedge. We suggest that the mantle wedge is the principal source of the initial pulse of arc melts consisted of normal MORB-type mantle, in agreement with findings from modern arcs (White \& Patchett, 1984).

The intra-arc rifting stage is dated by the 85.7 Ma age of the Chilas gabbronorite intrusion. The gabbronorites originated from the partial melting of a relatively shallow mantle source (Khan et al., 1993) and emplaced at temperatures in excess of $800^{\circ} \mathrm{C}$ and
at pressures of approximately $0.8-0.9 \mathrm{GPa}$ (Yoshino et al., 1998) in an extensional tectonic setting. Renewed melting of previously depleted harzburgitic mantle would, however, require unrealistically high melting temperatures. Therefore, we suggest that either the original magma source was (re-)fertilized, or that the Chilas parental melts tapped a completely different source.

The Hf isotopic composition of the Chilas melts had an $\varepsilon \mathrm{Hf}$ value of +10.4 and is significantly lower than the older magmatic rocks of the Metaplutonic Complex, but well within the range of oceanic basalts (e.g. Nowell et al., 1998). This suggests that an old mantle component with a lower time-integrated $\mathrm{Lu} / \mathrm{Hf}$ ratio than MORB, or even sedimentary or continental crustal contaminants, were involved in the generation of the Chilas parental melts. Assuming melting of a metasomatised mantle with a ${ }^{176} \mathrm{Lu} /{ }^{177} \mathrm{Hf}$ ratio of 0.020 , we arrive at a minimum age estimate of 600 Ma for this source. The same $\varepsilon H f$ value of +10 may also be obtained by subordinate contamination of MORB-type mantle by Archean granitoid crust or detritus, or pelagic sediments.

The 82.8 Ma old kyanite-bearing felsic dyke relates to a third magmatic phase derived from a MORB-type reservoir again. The dykes have crystallised at $610^{\circ} \mathrm{C}$ and $1.0-1.2$ GPa (Jan \& Karim, 1995). They probably originate from partial melting of granulitic, eclogitic or pyroxenitic lithologies from the deep base of the arc. Remelting of 20 Myr old mantle-derived materials is not expected to produce a discernible isotopic contrast in the Lu-Hf isotopic system. Therefore, we consider the eHf value of +14 to be inherited from the initial stage of partial melting of a MORB-type mantle source. The data also indicate that the felsic rocks, including granodiorite intrusions in the Metadioritic Sequence as well as felsic dykes, are not derived from post-obduction melting of the overridden Indian plate but form part of the arc magmatic sequence.

### 4.1.6 Geodynamic implications for the Kohistan Arc

The $\mathrm{U}-\mathrm{Pb}$ and Hf isotopic data from zircons of the Kohistan Arc Complex delineate a complex magmatic evolution from the initiation of an island arc to its mature stage. Since Lu-Hf is a very conservative isotopic system with respect to contamination by released slab fluids, changes in Hf isotopic characteristics probably reflect changes of source compositions within the mantle wedge. Therefore, these isotopic values indicate that mantle sources change in composition during the evolution of an island arc system.

Any tectonic models for the evolution of the arc system have to be consistent with these geochemical arguments.

The precise $\mathrm{U}-\mathrm{Pb}$ age determinations of zircons clearly separate the inititation of the Kohistan Arc arc from pre-100 Ma to at least 90 Ma , when melts were derived from an asthenospheric MORB-type mantle, from intra-arc rifting starting at around 86 Ma when melts possessed a less radiogenic Hf signature.

We envisage four potential hypotheses for the origin of this lithospheric component in the Chilas intrusion (Fig. 4.3):
(1) Sediments or a sliver of continental Archean to Proterozoic crust within the oceanic lithosphere of the subducted slab; this would infer the existence of continental relics or detritus in the Tethyan oceanic domain.
(2) Subduction and melting of an old fertile mantle portion attached to a sliver of continental crust within the Tethys oceanic domain.
(3) Contamination by pelagic sediments from the oceanic basement of the arc edifice.
(4) Dragging of $>600 \mathrm{Ma}$-old mantle with an intrinsic fertility into the hangingwall of the subduction zone - for instance a piece of Gondwana-derived lithosphere, from the back-arc area into the area of partial melting beyond the arc.

Trace and rare earth element compositions of the Chilas Complex are typical for arcrelated rocks and point to generation of the melts in rather shallow levels of the mantle wedge (Khan et al., 1993). The melts were contaminated by a "subduction component" containing large ion lithophile elements ( $\mathrm{K}, \mathrm{Rb}, \mathrm{Sr}, \mathrm{Ba}$ ), light rare earth elements (Khan et al., 1993) and increased Sr isotope ratios above mantle values (Mikoshiba et al., 1999). Hafnium transfer from the slab into the mantle wedge $(1,2)$ would require partial melting of slab material and destabilization of Hf-bearing phases - zircon in the case of felsic crustal material, titanate phases (rutile and ilmenite) for more mafic lithologies. This conflicts with the depletion of $\mathrm{Zr}, \mathrm{Hf}$, and Ti in Chilas samples (Khan et al., 1993), and is at odds with thermal models for subducting oceanic lithosphere. Contamination by pelagic sediments of the oceanic sub-arc crust (3, Fig. 3b) would probably produce lateral isotopic variations within the Chilas body, a hypothesis yet to be tested. Hypothesis (4) invokes dragging through mantle corner flow of a piece of old
metasomatically enriched mantle from the the hangingwall of the subduction into the region of partial melting in the mantle wedge. This process would introduce a new fertile mantle component into the system, which could provide large volumes of evolved melts necessary to produce the huge Chilas intrusion. Neither of these hypotheses can, however, be substantiated with the present data at the present time.


Fig. 4.3: Magmatic sources in a tectonic model for the initial arc buildup at $>100-90$ Ma (a) and for intra-arc rifting at 85-80 Ma (b). Hf isotopic values of the two phases of magmatism are indicated. Numbers 1 to 4 delineate hypothetical magma sources and are discussed in the text.

### 4.1.7 Conclusions

The Kohistan Arc Complex evolved within some 20 Myr from initial arc buildup (prior to $100-90 \mathrm{Ma}$ ) to a mature stage with back-arc spreading and magmatism (starting at 85 and lasting to at least 80 Ma ago). This mature stage is characterized by a shift in the Hf isotopic compositions of the magmatic liquids from MORB-type values $(\varepsilon H f=+14)$ to less radiogenic compositions $(\varepsilon \mathrm{Hf}=+10)$. The change of Hf isotopic composition of arc melts coincides with a drastic change in tectonic style, at the onset of intra-arc rifting and basin opening. The Hf isotope evidence demonstrates that the Kohistan Arc Complex is an intraoceanic island arc with minimal continental influence. Subducted
components derived by melting of the slab itself were negligible for the generation of the melts.

## Acknowledgements

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### 4.2 Proterozoic intrusion in the northern Indian Plate

Immediately south of the northward dipping Indus Suture, the Indian plate is comprised of strongly foliated and folded metasediments, quartzo-feldspathic gneisses and sodic quartzo-feldspathic gneisses. The gneisses belong to the Besham Group (Fletcher et al., 1986), which contains coarse grained metagranodiorites and metadiorites. The metadiorite body at Duber Bazar (Fig. 4.4) was dated by the $\mathrm{U}-\mathrm{Pb}$ zircon technique with an age of $1858.8 \pm 7.2 \mathrm{Ma}\left({ }^{206} \mathrm{~Pb} /{ }^{238} \mathrm{U}\right.$ age Tab. 4.2; Fig. 4.5, sample description in Appendix 4) and shows a similar age as the Shang granodiorite gneiss (ca. 17 km S of Duber Bazar.) with $1864 \pm 4 \mathrm{Ma}$ (DiPietro et al., 2001). The gneisses of the Besham Group formed 2120 to 2200 Ma ago and re-equilibrated during 1950 Ma metamorphism (Shah et al. 1992). Ar-Ar ages on hornblende determined by Baig (1991) and Baig et al. (1989) have been interpreted as one or more metamorphic episodes in the range of $2031 \pm 6$ to $1865 \pm 3$ Ma. Treloar et al. (1989a) obtained an Ar-Ar age on hornblende of $1920 \pm 24 \mathrm{Ma}$ for an amphibolite sheet within a granite at Duber Bazar.


Fig. 4.4: Geological map of the southern part of the Kohistan Arc Complex with dated samples from the Metaplutonic Complex, Chilas Complex and the Indian plate diorite at Duber Bazar.

| Nr. | Description | Weight [mg] | nr. of grains | Concentrations |  |  | $\mathrm{Th} / \mathrm{U}$ | Atomic ratios |  |  |  |  |  |  | Apparent ages |  |  | Error corr. |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  |  |  | U | $\begin{gathered} \mathrm{Pb} \\ \text { rad. } \\ {[\mathrm{ppm}]} \end{gathered}$ | Pb nonrad. [pg] |  | 206/204 <br> c) | $206 / 238$ <br> d) | $\begin{gathered} \hline \text { Error } \\ 2 \sigma[\%] \end{gathered}$ | 207/235 <br> d) | $\begin{gathered} \hline \text { Error } \\ 2 \sigma[\%] \end{gathered}$ | 207/206 <br> d) | $\begin{gathered} \hline \text { Error } \\ 2 \sigma[\%] \end{gathered}$ | 206/238 | 207/235 | 207/206 |  |
| Diorite Indian plate, Swat Valley |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| 18 | prism tips | 0.0316 | 4 | 376 | 131.45 | 2.4 | 0.43 | 104052 | 0.32727 | 0.33 | 5.1298 | 0.38 | 0.11368 | 0.10 | 1825.1 | 1841.1 | 1859.1 | 0.97 |
| 19 | small spr | 0.0156 | 6 | 232 | 78.47 | 2.9 | 0.45 | 24992 | 0.31502 | 0.34 | 4.9295 | 0.38 | 0.11349 | 0.11 | 1765.4 | 1807.3 | 1856.1 | 0.96 |
| 20 | frags | 0.0209 | 5 | 273 | 91.62 | 3.9 | 0.42 | 28915 | 0.31528 | 0.33 | 4.9383 | 0.38 | 0.11360 | 0.11 | 1766.6 | 1808.8 | 1857.8 | 0.96 |
| 21 | spr | 0.0329 | 5 | 210 | 67.80 | 15.2 | 0.41 | 8739 | 0.30394 | 0.33 | 4.7610 | 0.38 | 0.11361 | 0.11 | 1710.8 | 1778.0 | 1857.9 | 0.96 |

a) all zircon fractions abraded; prism $=$ prismatic; $\mathrm{spr}=$ short-prismatic
b) Calculated on the basis of radiogenic $\mathrm{Pb} 208 / \mathrm{Pb} 206$ ratios, assuming concordance
c) Corrected for fractionation and spike
d) Corrected for fractionation, spike, blank and common lead (Stacey \& Kramers, 1975)

Table 4.2: U-Pb isotopic data of zircons from the Indian plate diorite.


Fig. 4.5: U-Pb concordia diagram with age determination for the Indian plate metadiorite. Ellipses denote analytical uncertainty at the 2 olevel of individual analyses.

Ductile deformation is concentrated at the borders of the metadiorites. Thin (ca. 30-80 cm thick) pegmatites are generally undeformed and cut the weak, northeastward dipping foliation. This implies that the metadiorite, located ca. 300 m beneath the Indus Suture, was not or only little affected by the India-Kohistan collision. Furthermore, the $\mathrm{Ar}-\mathrm{Ar}$ Hornblende age of ca. 1920 Ma (Treloar et al., 1989a) was not reset during Himalayan collision, indicating that temperatures did not exceed the closure temperature of $530 \pm 50$ ${ }^{\circ} \mathrm{C}$ (McDougall \& Harrison, 1988). However, Himalayan metamorphism is recorded in the Swat area with Ar-Ar hornblende and muscovite ages ranging in their majority from 36 - 30 Ma (Anczkiewicz, 1998). The emplacement of the metadiorite is part of extensive intraplate plutonism around 1858 Ma in the Indian Plate as indicated by geochemical data (chapter 3.2).

### 4.3 EXhumation across the ISZ, antiform growth or back sliding of the

 HANGING WALL?
### 4.3.1 Introduction

Studying the exhumation history of the Kohistan Arc Complex is a key for postcollisional tectonic interpretation. Zircon and apatite fission-track (FT) ages were used to constrain cooling through ca. $250{ }^{\circ} \mathrm{C}$ and $100{ }^{\circ} \mathrm{C}$, respectively. Especially the comparison of these ages across faults gives insight in relative movements and also provide times, when the differential movement across the faults terminated. An extensive study was carried out by Zeitler (1985), providing fission track ages over the whole KAC (Fig. 4.6).


Fig. 4.6: Contoured zircon (upper map) and apatite (lower map) fission track ages of base-level samples (Zeitler, 1985). Area of this study is squared.

Prior to 15 Ma the cooling histories of both the KAC and the Indian Plate are different with the older ages in the hanging wall of the Indus Suture. Apatite ages are more similar on both sides of the Suture, leading to the interpretation that since 15 Ma no differential movement occurred across the Indus Suture (Zeitler, 1985). However, the contouring used by Zeitler (1985) smoothens the age pattern across major faults, which were not taken into account as discrete movement planes. Therefore, Zeitler's interpretation (Zeitler, 1985) is appropriate on a large scale but a more detailed study on distinct parts of the sutures is necessary. The following zircon and apatite fission-track ages were obtained in the Indus valley.

### 4.3.2 Results

Zircon and apatite fission-track ages were done on samples collected along the Karakoram Highway covering a ca. 35 km long section starting in the SW at Duber Bazar and ending in the NE 6 km SW of Dasu. Sample preparation followed the routine technique described by Seward (1989). Counting was carried out by Diane Seward at the ETH fission track lab. Etching of the apatite grains was done with $7 \% \mathrm{HNO}_{3}$ at 21 ${ }^{\circ} \mathrm{C}$ for 55 seconds. Zircon grains were etched in a eutectic mixture of KOH and NaOH at $220^{\circ} \mathrm{C}$ for between 20 and 50 hours. Irradiation was carried out at the ANSTO facility, Lucas Heights, Australia. Microscopic analysis was completed using an optical microscope with a computer driven stage ("Langstage" software from Dumitru (1995)). All ages were determined using the zeta approach (Hurford \& Green, 1983) with a zeta value of $355 \pm 5$ for CN5 and $120 \pm 5$ for CN1. They are reported as central ages (Galbraith \& Laslett, 1993) with a $2 \sigma$ error (Table 4.3). Where possible, 20 crystals of each sample were counted for age determination. The magnification used was x1250 for apatite and x1600 (oil) for zircon. Horizontal confined track lengths were measured at x1250.

Several samples from the Jijal Complex and southern Metaplutonic Complex failed FT dating due to the lack of zircon and/or apatite or due to an extremely low $U$ content in apatite. This prevented some ages being determined on several apatite samples in particular because the density of spontaneous tracks was so very low. Any determination would yield ages with such large errors that they would have not been useful in comparing them with neighbouring samples. In the Indian Plate a deformed
metadiorite from ca. 100 m beneath the ISZ close to Jijal (sample 00/30/12), representing the footwall of the ISZ was dated and yields zircon and apatite FT ages of $22.6 \pm 2.6(2 \sigma) \mathrm{Ma}$ and $3.7 \pm 1 \mathrm{Ma}$ respectively (Fig. 4.7 for sample location, table 4.3 and appendix 4 for details). In the southern Metaplutonic Complex, representing the hanging wall of the ISZ four samples could be dated. The oldest dated zircon FT age is yielded by a coarse grained pegmatite cutting the dioritic host rocks ca. 5 km NE of Patan with $41.8 \pm 8.8 \mathrm{Ma}$ (sample 56/02/01). A coarse-grained hornblende diorite from ca. 1 km S of Kiru yields a zircon FT age of $27.9 \pm 5.2 \mathrm{Ma}$ (sample 91/07/01). The northernmost dated sample was collected ca. 200 m S of the bridge where the Karakoram Highway and the Indus River bend sharply towards the E . This coarse grained granite yields a zircon FT age of $23.6 \pm 5.6 \mathrm{Ma}$ (sample 00/10/01). Only one apatite FT age was obtained in the Metaplutonic Complex. The sample was taken ca. 3 km E of Patan in an undeformed part of the Sarangar gabbro and yields an apatite FT age of 10.6 $\pm 2.4$ Ma (sample 16/02/01).


Fig. 4.7: Zircon fission track samples from the Metaplutonic Complex and the Indian plate diorite. Inset: Map of lower Swat region with two ages on the Indian Plate (Anczkiewicz, 1998 and D. Seward, unpubl.).

New Geochronological constraints...

| Sample <br> number | Mineral | Number <br> of <br> grains | Standard track <br> density $\times 10^{4} \mathrm{~cm}^{-2}$ <br> (counted) | $\rho_{\mathrm{S}} \times 10^{4} \mathrm{~cm}^{-2}$ <br> (counted) | $\rho_{\mathrm{i}} \times 10^{4} \mathrm{~cm}^{-2}$ <br> $($ counted) | U conc. <br> ppm | $\mathrm{P}\left(\chi^{2}\right)$ <br> (Variation $\%)$ | Central Age <br> $\pm 2 \sigma(\mathrm{Ma})$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $95-196$ | A | 16 | $110(2439)$ | $12.9(111)$ | $225(1929)$ | 26 | $97(0)$ | $11.2 \pm 2.2$ |
| $95-198$ | A | 20 | $122(2439)$ | $10.8(82)$ | $156(1185)$ | 16 | $83(0.3)$ | $15.0 \pm 3.4$ |
| $16 / 02 / 01$ | A | 20 | $121(2808)$ | $0.73(126)$ | $14.5(2509)$ | 2 | $12(25)$ | $10.6 \pm 2.4$ |
| $00 / 30 / 12$ | A | 20 | $130(2978)$ | $3.64(60)$ | $24.3(4007)$ | 23 | $8(25)$ | $3.7 \pm 1.0$ |
| $95-196$ | Z | 12 | $49.1(2376)$ | $287(378)$ | $414(546)$ | 338 | $8(18)$ | $20.5 \pm 3.6$ |
| $95-198$ | Z | 8 | $56.2(2376)$ | $473(885)$ | $563(1054)$ | 400 | $11(9)$ | $28.2 \pm 3.4$ |
| $56 / 02 / 01$ | Z | 20 | $73.3(2543)$ | $14.6(180)$ | $15.3(189)$ | 8 | $100(0)$ | $41.8 \pm 8.8$ |
| $00 / 10 / 01$ | Z | 4 | $64.8(2543)$ | $223(151)$ | $372(252)$ | 230 | $22(9)$ | $23.6 \pm 5.6$ |
| $91 / 07 / 01$ | Z | 20 | $71.6(2543)$ | $39.2(217)$ | $60.7(336)$ | 33.9 | $66(10)$ | $27.9 \pm 5.2$ |
| $00 / 30 / 12$ | Z | 20 | $69.3(2543)$ | $532(1726)$ | $1002(3253)$ | 578 | $0(19)$ | $22.6 \pm 2.6$ |

Table 4.3: Zircon fission track results from the Metaplutonic Complex, the Indian plate diorite and two ages of lower Swat region (Anczkiewicz, D. Seward, unpubl.). $A=$ apatite, $Z=$ zircon. $\rho s$ and $\rho i$ represent sample spontaneous and induced track densities; $P\left(\chi^{2}\right)$ is the probability of $X^{2}$ for $v$ degrees of freedom where $v=n o$. of crystals $-1 . \lambda_{D}=1.55125 \times 10^{-10}$. Zeta $=355 \pm 5$ for apatite and CN5, $120 \pm 5$ for zircon and CN1. Samples were irradiated at the ANSTO facility, Australia. Ages are reported as central ages (Galbraith, 1981). Detailed table with sample coordinates is provided in Table Appendix 4.2.

### 4.3.3 Tectonic Interpretation

The FT ages yield low temperature tectonothermal information. There is clearly a younging across the Indus Suture from north to south. Within the Metaplutonic Complex ages show an opposite younging direction. The difference in zircon cooling ages could represent tilting effects, or discrete events across faults. If the latter is the case across the ISZ then there is evidence for a relative back sliding of the hanging wall (KAC) along the ISZ. The jump across the ISZ is significant for both, zircon and apatite FT data. The difference between the zircon FT ages representing footwall and hanging wall of today's ISZ geometry of ca. 20 Ma points to an onset of normal faulting during this time span. The apatite FT ages point to a maximum exhumation rate of ca. 0.8 $\mathrm{mm} / \mathrm{a}$ in the footwall over the last 3.7 Ma and of ca. $0.3 \mathrm{~mm} / \mathrm{a}$ for the last 11 Ma in the hanging wall. The young apatite FT ages in the Indian Plate ( 3 to 5 Ma ) may reflect a still growing Besham antiform (Coward, 1986; Treloar et al., 1989c). To the W, in the Swat valley, apatite FT ages in the Indian Plate are older. Two samples collected by R. Anczkiewicz ca. 4 km NE of Mingora (sample 95-196) and ca. 5 km ENE of Manglaur (sample $95-198$ ) yield ages of $11.2 \pm 2.2 \mathrm{Ma}$ and $15.0 \pm 3.4 \mathrm{Ma}$ respectively (D. Seward, unpubl.; Fig. 4.7, inset), indicating a faster and/or younger exhumation in the Indus area. As a preliminary interpretation the zircon and apatite FT data indicate normal faulting along this specific part of the Indus Suture from 23 Ma to at least 3 Ma . This
contradicts the termination of differential movement at around 16 Ma on the Indus Suture (Zeitler, 1985; Vince \& Treloar, 1996). The northward younging direction of ages within the KAC may be interpreted with rotations and the subsequent relative faster exhumations of the northern part against the southern part of the KAC (Fig. 4.8).

The difference of ca. 14 Ma between the 41.8 Ma zircon FT age 6 km NE of Patan and the 27.9 Ma zircon FT age 1 km S of Kiru is significant. An E-W trending, ca. $70-80^{\circ}$ northward dipping reverse fault is the prominent fault between the sampled rocks. Rollover geometries are the result of half grabens formed during crustal extension that is accommodated by listric shaped normal faults, causing collapse of the hanging wall (Hamblin, 1965). Rollovers have been observed in rift basins worldwide (e.g. Bally, 1983; Nunns, 1991) and geometric and physical models of hanging wall collapse are described (e.g., Gibbs, 1983; Rowan \& Kligfield, 1989; Groshong, 1990; White \& Yielding, 1991). In the investigated area no basin developed in a rollover setting during extension is recorded. This implies the following possibilities: 1) the subsidence in the southern edge (hanging wall located north of footwall) of the hanging wall was not large enough to create a sediment filled basin along the Indus Suture, 2) the theoretic rollover surface was always above sedimentation level and 3) the whole geometry was exhumed faster relative to subsidence in the hanging wall and sediment filled basins trapped north of the Indian Plate were eroded subsequently. The development of rollover geometry is characterised by a rotation of the southern part of the hanging wall. As shown in the tectonic sketch presented in Figure 4.8 the rotation brings the early antithetic normal fault into the position of an apparent reverse fault.

Backsliding of the hanging wall is reported on several places of the ISZ: the Naran region (Burg et al., 1996), the Indus region (Vince \& Treloar, 1996) and the Swat valley (Anczkiewicz, 1998). The earliest onset of extension is estimated by Burg et al. (1996) at 47 Ma , based on dated peak metamorphism of $47 \pm 3 \mathrm{Ma}$ from SHRIMP U-Pb analysis of zircon rims (Smith et al., 1994). Extension was completed at ca. 20 Ma (Burg et al., 1996; Vince \& Treloar, 1996) based on ages from Zeitler (1985). A comparison of the cooling ages from the northern, middle, southern Metaplutonic Complex and the Indian Plate along the same SW-NE section as in Figure 4.8 shows that the relative normal movement along the Indus Suture outlasted the 20 Ma (Fig. 4.9).


Fig. 4.8: Tectonic sketch integrating available FT-ages on a SW-NE profile along the Indus River. A rollover effect in the hanging wall rotates an early antithetic normal fault to an apparent reverse fault. Ages in Ma are given for samples when passing the $250^{\circ}$ and $100{ }^{\circ} \mathrm{C}$ isograde. The time steps of $40,28,20,10,5 \mathrm{Ma}$ and recent reflect midpoints of certain time ranges. Samples: $\backslash$ : apatite, $\downarrow$ : zircon, $\downarrow$ : both; PAZ: partial annealing zone; 79I-samples from Zeitler, (1985). Simple depth temperature correlation is used and does not reflect real conditions.

The zircon FT ages of 41.8 Ma for the southern Metaplutonic Complex and the apatite FT age of 3.7 Ma for the Indian Plate are in agreement with the FT ages obtained by Zeitler (1985) ( $41.4 \pm 2.6 \mathrm{Ma}$ and $5.6 \pm 2.3 \mathrm{Ma}$ ). Two explanations for the young differential movement across the ISZ in the Indus region are left: 1) The KAC was backsliding along the ISZ as a block until ca. 3 Ma or 2) the recent growth of the Besham antiform is decoupled from the overlying KAC (here mainly the Jijal Complex) and the ISZ is accommodating the differential "growth" with normal faulting (Fig. 4.10).


Fig. 4.9: Temperature - time plot showing the cooling history of four different blocks as shown in Fig. 4.8. They are from $S$ to N: Indian Plate, southern, middle and northern Metaplutonic Complex. Cooling paths are rough estimations and not modelled. PAZ: partial annealing zone; 79I-samples from Zeitler, (1985).

The first explanation is unlikely because apatite FT ages younger than 5 Ma are not observed within the Indian Plate along the whole length south of the contact to the KAC. Instead in the Swat valley apatite FT ages on both sites of the Indus Suture are comparable with ca. 11-15 Ma in the south and 15-17 Ma (Zeitler, 1985) in the north, indicating no or little relative movement across the Indus Suture later than 15 Ma ago. The second explanation is consistent with the young apatite FT ages in the Indus area S of the ISZ and older ages E and W of the Besham antiform (Fig. 4.7). Therefore, following conclusion is drawn: the young differential movement across the ISZ in the Indus area is a local phenomena caused by the growth of the Besham antiform and the ISZ is accommodating the differential "growth" with normal faulting.

The duration and the onset of the Besham antiform growth are not directly dated. Assuming the Besham antiform represents a small-scale equivalent to the Nanga Parbat Syntaxis (Coward, 1986; Treloar et al., 1989c) with close generic relation to the Nanga Parbat Syntaxis, located ca. 100 km E of the Besham antiform, the timing can be estimated. The rapid denudation during growth of the Nanga Parbat lasted some 4 Ma (Burg \& Podladchikov, 2000). It started not earlier than ca. 5 Ma ago (Zeitler et al. 1989, 1993). This implies for the Besham antiform an onset of denudation during growth younger than 5 Ma with a probably faster exhumation at the beginning of the
buckling which may be represented by the 3-4 Ma apatite FT age. A recent growth is noticeable in topographic features. The Indus River excavated a steep gorge S of Thakot (ca. 16 km S of Besham). However, earthquakes or seismicity does not report recent normal faulting along the Indus Suture. Therefore it is not known, if normal faulting along the Indus Suture N of Besham is still an active process.


Fig. 4.10: Simplified block diagram showing the Besham antiform descending towards north beneath the Kohistan Arc Complex along the ISZ (MC: Metaplutonic Complex, JC: Jijal Complex).

## 5. ANASTOMOSING SHEAR ZONES IN THE LOWER KOHISTAN COMPLEX

Chapter 5.1 is "Pre-collisional anastomosing shear zones in the Kohistan arc, NW Pakistan" by L. Arbaret, J.P. Burg, G. Zeilinger, N. Chaudhry, S. Hussain and H. Dawood (Arbaret et al., 2000).

### 5.1 Pre-collisional anastomosing shear zones in the Kohistan arc, NW Pakistan

## Abstract

Ductile strain localisation commonly forms a pattern of shear zones anastomosing around lenses of less deformed rock. Initiation and development history of anastomosing shear zones, which are little documented, are studied through the description of the structures and the deformation history of plutonic rocks that form the lower crust of the Kohistan Complex. Structures and textures developed in these rocks result from primary magmatic to solid state regional strain overprinted by anastomosing shear zones. The primary strain was mainly acquired during magmatic emplacement at 100-90 Ma. Strain localisation took place continuously from magmatic emplacement to solid state deformation during cooling of the plutons and formed 3 successive sets of shear zones. Set 1 is composed of associated discrete Riedel and thrust shear zones developed above solidus conditions during SW-ward thrusting. Continuous deformation from solidus to amphibolite facies conditions between 100 and 83 Ma formed the second set of shear zones. The lower amphibolite facies set 3 shear zones are differentiated by larger strain recorded in the thicker mylonitic zones and enlargement of the spacing between shear zones during cooling. The anastomosing pattern of shear zones described here probably represents arc-related deformation during subduction of the Tethys oceanic lithosphere below the Kohistan Arc Complex.

### 5.1.1 Introduction

Ductile strain is commonly concentrated into narrow shear zones, anastomosing around lenses of lower strain, the symmetry of the anastomosing pattern reflecting the symmetry of the bulk deformation history (Gapais et al., 1987). Anastomosing patterns are commonly reported (e.g. Mitra, 1978; Ramsay, 1980; Harris \& Cobbold, 1984;

Gapais et al., 1987; Lafrance et al., 1998), but with limited information on their geometric detail and growth history. Therefore, the way such anastomosing patterns form and evolve remains an open question. In particular, the relationships between regionally distributed strain and strain localisation in shear zones is not clear.

We aim to provide preliminary answers in describing the structures and deformation history of the plutonic rocks of the lower Kohistan Complex. Some of these plutonic rocks display an early fabric that defines the regional pattern, but an important part of the deformation resulted in strain localisation into shear zones anastomosing around lenses of less, or undeformed rock (Treloar et al., 1990). We argue that the regional background strain is composite and was mainly acquired during magmatic emplacement of the various plutons. The later anastomosing shear zones result from the nucleation and growth of three successive sets of shear zones sharing similar directions. Our petrologic information does not improve previous P-T estimates, which will be referred to in the following paragraphs. Therefore we will not duplicate here the geothermobarometry of these rocks. Correlation between structures and the P-T-time history provides a detailed tectonic evolution of the lower part of the Kohistan Complex.

### 5.1.2 Geological setting

The Kohistan Complex was an island arc within the Tethys ocean during the Late Mesozoic (Tahirkheli et al., 1979; Bard, 1983; Coward et al., 1986). The tectonic history involves the 102 to 75 Ma old accretion to the Asian plate, to the north (Petterson \& Windley, 1985; Treloar et al., 1989a) and closure of the Tethys ocean, to the south, at $c .55 \mathrm{Ma}$, followed by obduction of the Kohistan Complex onto India (Coward et al., 1986). Our structural analysis was focused on the southern, i.e. lower, part of the arc sequence (Fig. 5.1). The study area is comprised of meta-gabbros, hornblende-gabbros and diorites intruded by, and intruding, hornblendites (the southern part of the Kamila Amphibolite Belt, Treloar et al., 1990; Khan et al., 1993). The southernmost gabbro is granulitic (it is the upper part of the Jijal Complex, Jan \& Howie, 1981; Bard, 1983; Yamamoto \& Yoshino, 1998). The structurally upper gabbro, named in this study Sarangar gabbro, forms the lower part of the Patan Complex of Miller et al. (1991). The detailed petrography of these rocks has been presented by Jan
\& Howie (1981), Treloar et al. (1990), Miller et al. (1991), Yamamoto (1993), Yoshino et al. (1998) and Ringuette et al. (1999). All agree that they represent calc-alkaline magmas emplaced during the arc activity (Khan et al., 1993). Ductile deformation related to the Kohistan / India fault contact (the so-called Indus Suture or Main Mantle Thrust, Tahirkheli et al., 1979) is limited to a few hundred meters above this major fault (Burg et al., 1998). It was therefore important to know why ductile deformation in shear zones (Treloar et al., 1990) was appearing several thousand meters higher in the Kohistan sequence, and to understand the kinematics involved in this deformation event. We will argue that it is an arc related shear deformation synthetic to the subduction zone above which the arc has grown.


Fig. 5.1: Simplified geological map of the working area. Inset: localisation in the Kohistan Complex (modified from Bard, 1983 and Burg et al., 1998).

### 5.1.3 Regional strain

The regional strain is defined by heterogeneously distributed foliations with associated mineral lineations. Two-dimensional mineral shape anisotropy has been measured in sections parallel to the lineation and perpendicular to the foliation (XZ planes of rocks) to estimate the orientation intensity of the mineral fabric (Fig. 5.2). We applied the
intercept method (Rink, 1976; Launeau et al., 1990; Launeau \& Robin, 1996) to feldspars because they are both the most common and the weakest mineral phase in all the studied rocks. Measurements close to the shear zones were not taken into consideration in order to avoid later strain effects.


Fig. 5.2: Example of image analysis performed in an XZ section of the magmatic assemblage (left) showing the centimetre-scale variation of the fabric intensity Is of plagioclase between 1.05 and 1.2 (right). Note that the foliation trace (black double arrows) remains constant all over the image, independently of the fabric intensity (Sarangar gabbro, location in Fig. 5.14).

Two intensity fabrics were recognised. The weak one (1.05 to 1.2, Fig. 5.2) is recorded in the core of the two structurally lower gabbros (Fig. 5.3). The weak intensity fabric corresponds to the shape preferred orientation of the magmatic assemblages characterised by cumulate textures. The cumulate assemblage of the lowermost granulitic gabbro is composed of hypidiomorphic pyroxene, garnet, plagioclase and amphibole in textural equilibrium (Ringuette et al., 1999). Fabrics of this assemblage describe the $c .40^{\circ} \mathrm{NW}$-dipping foliation and the $c .35^{\circ} \mathrm{NW}$-plunging lineation (Fig. 5.3 and 5.4). In the coarse grained Sarangar gabbro that lies to the north of, and above the granulitic gabbro (Fig. 5.1), weak intensity fabrics correspond to the coarse grained magmatic cumulate, garnet-free assemblage of pyroxene and plagioclase (Fig. 5.5a). These intensity fabrics describe E-W trending vertical foliations bearing a W-plunging lineation (Fig. 5.3 and 5.4).

The highest intensity fabrics ( 1.2 to $>1.6$ ) were measured in three main locations. First, in the granulitic gabbro, the high intensity fabrics progressively replace the magmatic, low intensity fabrics, in particular toward the basal contact with ultramafic rocks (Fig.
5.3). Higher intensity fabrics mark a transition from magmatic to solid state deformation by plastic deformation of plagioclase and rigid rotation of pyroxene and other mineral phases towards the pluton boundaries. Local S-C structures and the angular relationship between the mineral fabric and magmatic layering indicate a bulk SW-ward shear (Fig. 5.5b).


Fig. 5.3: Trajectories of magmatic and solid state foliations and lineations. Black ellipses indicate direction and intensity (percentage of anisotropy) of fabric measured in $X Z$ planes.

foliations $=62$, lineations $=56$
hornblende-gabbro and diorite laccoliths: tectonic fabric


Fig. 5.4: Lower hemisphere equal area projections of poles to foliations and lineations. Starkey density contours: 2, 4, 6 and $8 \%$.

The second location of high intensity fabrics is the contact between the Granulitic and the Sarangar gabbros. The extension of this fabric on both sides of the contact has a maximum width of about 2000 m . In the Sarangar gabbro, this high intensity fabric of oriented pyroxene and deformed plagioclase decreases and progressively gives way to a magmatic fabric towards the core of the pluton. Consistent foliation, lineation and kinematic directions and this geographic distribution suggest that the periplutonic highintensity fabric was solid state, ductile deformation at the boundary with the granulitic gabbro, during the Sarangar gabbro emplacement. In the granulitic gabbro, the high
intensity fabric has overprinted the magmatic fabric (Fig. 5.3). These structural relationships indicate that the Sarangar gabbro is younger and has intruded the cooler than solidus granulitic gabbro.


Fig. 5.5: (a) Magmatic texture in the Sarangar gabbro in an XZ section, trace of the foliation is vertical. (b) Crosscutting relationships between the magmatic layering and the trace of the tectonic foliation indicating $S W$-ward shear in the granulitic gabbro. Coin size $=23 \mathrm{~mm}$. Locations in Fig. 5.14.

Finally, strong and homogeneous orientation intensity fabrics occur throughout the hornblende-gabbro and diorite laccoliths that have intruded the Sarangar gabbro to the north (Fig. 5.3). Attitudes have an average $38^{\circ} \mathrm{N}$-dip foliation with a NE-striking lineation (Fig. 5.3 and 5.4). There is no preserved magmatic fabric in these rocks whose deformation is fundamentally solid state with rotated porphyroclasts indicating SWward sense of shear, as recognised also in the northern parts of the Kamila belt (Treloar et al., 1990).

The metamorphic history of the granulitic gabbro includes sub-solidus isobaric cooling of the magmatic paragenesis within granulite facies conditions ( $\mathrm{T} \sim 750{ }^{\circ} \mathrm{C}$ and $\mathrm{P} \sim 1.8$ GPa, Ringuette et al., 1999). A Sm-Nd cooling age of 91 Ma (Yamamoto \& Nakamura,
1996) is slightly younger than $\mathrm{Sm}-\mathrm{Nd}$ and $\mathrm{Rb}-\mathrm{Sr}$ ages of 96 and 101 Ma , respectively, measured by Anczkiewicz \& Vance (2000). Preserved plagioclase-clinopyroxene assemblages indicate crystallisation at $800^{\circ} \mathrm{C}$ and $0.8-1.1 \mathrm{GPa}$ in the Sarangar gabbro (Yoshino et al., 1998), whose age is not available. However, it should be bracketed by the 100-90 Ma old granulitic gabbro and the regional hydration phase at about 83 Ma (Ar-Ar age of amphibole, Treloar et al., 1990).

Therefore, the regional fabric fundamentally reflects magmatic emplacement of mafic, calc-alkaline magmas during the pre-collisional arc formation.

### 5.1.4 Shear strain localisation

Shear strain localisation is responsible for the anastomosing pattern of shear zones in the Kamila Shear Zone (Treloar et al., 1990) from which we have separated three successive sets.

### 5.1.4.1 Set 1

Set 1 is composed of shear zones less than 1 m long and a few centimetres wide (Fig. 5.6 and 5.7). They bend the earlier magmatic foliation and layering, where present, without noticeable change in grain size and mineral assemblage. These shear zones apparently formed at conditions above the solidus because the magmatic paragenesis remained stable in their most deformed centres. Set 1 shear zones were found along structural and compositional discontinuities and within the bulk rock mass where they may be cut by millimetre-thick feldspar-rich magmatic joints (Fig. 5.7). Structural discontinuities are discrete, feldspar-rich magmatic joints, hornblendite boundaries and few millimetres thick veins of quartz, feldspar and amphibole (Fig. 5.8), which supports the interpretation of set 1 shear zones being above solidus features.

Most of set 1 shear zones have a normal sense in the field and are organised in two orientation populations (Fig. 5.9). The dominant one corresponds to shallowly Sdipping to horizontal shear zones with a SE- to SW-vergence. The second, apparently conjugate population, is formed of homogeneously NE-verging, N -dipping shear zones. These conjugate populations could indicate some sub-vertical flattening. However, field observation shows that many of the NE-verging shear zones are localised along upper and lower borders of hornblendite bodies that have rotated consistently with the bulk SE- to SW-ward shear (Fig. 5.6). Reverse set 1 shear zones developed along more
homogeneously NE-dipping planes with the same lineation direction as the normal shear zones. Thus, reverse and normal set 1 shear zones are kinematically consistent. Differences in orientation fit Riedel orientations known in movement zones (Riedel, 1929). The most common population of normal SW-ward shear zones would represent Riedel shears R (Fig. 5.6). The reverse N -ward shear zones have an attitude consistent with R' Riedel shears (Fig. 5.6). We contend that the NE-dipping reverse shear zones were the closest to the main shear plane (primary shear bands D or, rather, thrust shears P, Tchalenko, 1968). Indeed, they are nearly parallel to the regional lithological boundaries of intensely sheared rocks (the regional reference plane), which were nearly horizontal before tilting of the Kohistan Complex subsequent to its obduction over India (Tahirkheli et al., 1979; Coward et al., 1986).


Fig. 5.6: Synthetic sketch of the three successive sets of shear zones. Set 1, centimetre-scale shear zones are developed along Riedel orientations consistent with a general, nearly horizontal $S W$-ward shear during the late magmatic emplacement of the gabbros. Upper amphibolite facies set 2 shear zones developed later along the same orientations as set 1 shear zones. Lengthwise propagation of set 1 shear zones formed a metre-scale, anastomosing pattern in which partial melting yielded garnet bearing quartzo-feldspathic veins. $N E$-ward normal set 2 shear zones developed essentially in response to rotation of hornblendite bodies and split lenses. The third set of shear zones, characterised by a larger spacing, developed in lower amphibolite facies conditions along and within similarly oriented set 2 shear zones.

All Set 1 shear zones are therefore attributed to a general SW-ward shear. Set 1 shear zones have the same orientation in both the granulitic and the Sarangar gabbros, independent of the magmatic fabrics. We conclude that set 1 shear zones are responses to regional stresses. Limited grain size reduction in their centres and measurements of the passively curved foliation indicate shear strains $\gamma<5$ (March 1932; Ramsay, 1980). The geographical distribution of these shear zones is undetermined because solid state deformation in the other plutons and Set 2 shear zones may have overprinted them.


Fig. 5.7: Set 1 shear zones (location in Fig. 5.14). (a) Normal $W$-ward Riedel shear zone crosscut by a plagioclase-rich magmatic joint (arrows). (b) Lengthwise terminating set 1 shear zone. Arrow as in (a).

### 5.1.4.2 Set 2

Set 2 is the most spectacular anastomosing pattern of mylonitic shear zones that wrap around lenses of less deformed rock with their long axis plunging to the north-east (Fig. 5.6 and 5.10). Bordering continuous strain gradients include progressively curved rockfoliations whose fabric intensities increase to 1.9. The replacement of diopside and hypersthene by ferroan pargasitic to tschermakitic hornblendes (Treloar et al., 1990) is correlated with the increasing intensity of the fabric in the strain gradients. However, pyroxene porphyroclasts exist in the mylonite, which suggests that metamorphic conditions were dry enough to prevent full retrogression.


Fig. 5.8: Synthetic sketch of crosscutting relationships between structural features in the Sarangar gabbro based on two successive outcrops (rigth).
$66 \%$ of set 2 shear zones are N -dipping to flat-lying, SW-verging reverse shear zones bearing a $c .30^{\circ}$ NE-plunging lineation (Fig. 5.11). They branch into horizontal to Wdipping, SW-verging normal shear zones bearing a $c .14^{\circ}$ SW-plunging lineation. The change in attitude accommodates the shape of the preserved rock lenses. $11 \%$ of set 2 shear zones are NE-verging normal shear zones that developed, as some of the set 1
zones, around the tips of rotated hornblendite bodies (Fig. 5.6). The other set 2 shear zones correspond to reactivated, similarly oriented SE -verging set 1 shear zones. The thickness of the mylonitic centre of set 2 shear zones, defined by mylonitic foliations parallel to the shear zone borders, varies along a single shear zone from less than 5 cm to more than 4 m (Fig. 5.12). Centimetre-wide splay shears may occur within lenses in response to rapid direction changes of set 2 shear zones or between rotating parts of split lenses (Fig. 5.6 and 5.10b).

## set 1 shear zones


set 1 normal shear zones

set 1 reverse shear zones


Fig. 5.9: Lower hemisphere equal area projections of poles to planes and stretching lineations of set 1 shear zones, arrows pointing to the shear direction. Starkey density contours: 2, 4, 6 and $8 \%$..

Long axes, a, and perpendicular short axes, c, of rock lenses were measured on 60 regularly shaped lenses of less-deformed gabbro observed in XZ outcrops with respect to the stretching lineation and foliation of the bounding set 2 shear zones. Measurements were limited for sake of reliability to less than 5 m long lenses. Axis a versus axis $\mathbf{c}$ define a linear relation at any scale (Fig. 5.13). Similar asymmetric lens shapes at all scales and strongly dominant SW-vergence indicate that set 2 shear zones developed during distributed non coaxial deformation (Gapais et al., 1987).


Fig. 5.10: Set 2 shear zones pattern (location in Fig. 5.14). (a) Shear zone wrapping around a lens of undeformed gabbro. (b) Rapid change in orientation of a shear zone inducing nucleation of centimetre-wide splay shears in the lens (parallel to the white arrows). (c) and (d) Syn-mylonitic quartz-felspar garnet-bearing veins in set 2 shear zones with (black arrow in d) undeformed segment of vein in a transtensional zone in a neck region of a lens of the Sarangar gabbro.

Set 2 shear zones occur over the whole area (Fig. 5.14). In the Sarangar gabbro, they form a pervasive pattern with an average spacing of 8-10 m between shear zones. This regular spacing indicates periodicity in their distribution. The anastomosing pattern becomes dense, forming metric lenses of less deformed rock along the contact with the underlying granulitic gabbro in which the set 2 pattern has a metric distribution above the hornblendite-rich area. In the lower part of the granulitic gabbro, the spacing
between set 2 shear zones increases in the hornblendite rich area. No shear zone was found in the lowest 1000 m . In the hornblende-gabbros and diorite laccoliths, to the north, i.e. above the Sarangar gabbro, spacing between set 2 shear zones is about 50 m .

## set 2-3 shear zones


top to north-east; set 2-3 normal shear zones


Fig. 5.11: Lower hemisphere equal area projections of poles to planes and stretching lineations of set 2-3 shear zones, arrows pointing to the shear direction. Sense of shear determination is based on the curvature direction of foliations and the geometry of feldspar and garnet porphyroclasts in mylonites. Starkey density contours: 2, 4, 6, 8 and $10 \%$.
top to south-west; set 2-3 shear zones


73 reverse shear zones (white poles and arrows)
32 normal shear zones (black poles and arrows)

Metamorphic conditions associated with set 2 shear zones were estimated at $550-650^{\circ} \mathrm{C}$ under pressures of 0.9-1.0 GPa (Bard, 1983; Treloar et al., 1990). Partial melting in set 2 mylonites generated quartz-plagioclase-garnet bearing segregation veins, some of which having been deformed by subsequent shearing (Fig. 5.10c) while others remained almost undeformed in transtensional zones between lenses (Fig. 5.10d). Partial melting points to temperatures exceeding $650^{\circ} \mathrm{C}$ under pressures of $>0.8 \mathrm{GPa}$ (Burnham, 1979) for the initiation of set 2 shear zones, which is consistent with the upper amphibolite facies crystal plasticity described by Treloar et al. (1990).

The minimum age for set 2 shear zones is inferred from the 87 Ma amphibole-epidoteparagonite assemblage that has partially overprinted the granulitic assemblage in the granulitic gabbro (Anczkiewicz, 1998).


Fig. 5.12: Thickness distribution of mylonitic central zones of set 2-3 shear zones (average thickness $=1.27 \mathrm{~m}$; median value $=50 \mathrm{~cm}$ ) and feldspar porphyroclast-rich set 3 shear zones (average thickness $=2.8 \mathrm{~m}$ and median value $=1.9 \mathrm{~m}$ ).

### 5.1.4.3 Set 3

Set 3 corresponds to several, tens metres long, shallow S-dipping to $30^{\circ}$ NE-dipping shear zones with a wavelength of over 50 m (Fig. 5.6 and 5.15). Thickness distribution of the set 3 ductile mylonite zones presents a unimodal distribution (Fig. 5.12). Mylonitic zones display strongly foliated and fine-grained matrix containing plagioclase with brittle deformation and relicts of amphiboles. Grain size reduction and rotated porphyroclasts of garnet and plagioclase derived from stretched pegmatite veins indicate shear strain $\gamma>10$. Curvature of the foliation along with the hornblendite and felsic veins within the shear zone boundaries indicate SW-ward shearing in set 3 shear zones (Fig. 5.15b). Set 3 shear zones, which contain metric lenses of undeformed rock,
developed along and within, similarly oriented set 2 shear zones. The mylonitic assemblage is composed of quartz, plagioclase, amphibole and porphyroclastic garnet. This assemblage points to lower amphibolite facies conditions, which are not dated but are probably around the $83 \mathrm{Ma} \mathrm{Ar}-\mathrm{Ar}$ age proposed for the end of the regional amphibolite facies metamorphism (Treloar et al., 1989a; Wartho et al., 1996). Set 3 shear zones distribution is limited to the Sarangar gabbro and the hornblende-gabbro and diorite laccoliths with an average spacing of over 50 m (Fig. 5.14).


Fig. 5.13: Major axis a versus short axis c calculated from lenses of less-deformed gabbro observed in outcrop sections perpendicular to the foliation plane and parallel to the stretching lineation of the surrounding shear zones. Long axes of lenses vary between 35 cm to 420 cm and perpendicular short axes range from 8 to 180 cm .

### 5.1.5 Discussion

The geometry of the anastomosing pattern and kinematic analysis of strain localisation from regionally distributed fabrics provide a preliminary answer to the deformation history of the plutonic rocks forming the lower part of the Kohistan arc.

The composite regional fabric, developed above solidus conditions, fundamentally reflects magmatic emplacement of mafic, calc-alkaline plutons around $100-90 \mathrm{Ma}$. The onset of shear strain localisation is marked by the occurrence of set 1 shear zones whose Riedel-type attitudes were controlled by regional stresses. The continuous evolution from magmatic fabrics to solid state deformation indicates that strain
localisation began above or at solidus conditions of the mafic magmas. Similarity of attitudes suggests that the asymmetric anastomosing pattern of set 2 and set 3 shear zones formed by the lengthwise propagation of the set 1 shear zones. This cooling evolution is marked by a progressive increase in spacing of the array of shear zones.

Treloar et al. (1990) regarded shearing in the Kamila Amphibolite Belt at temperatures greater than $500{ }^{\circ} \mathrm{C}$ until $80 \pm 5 \mathrm{Ma}$ as the major expression of arc thickening produced by collision with Asia, to the north, between 102 and 75 Ma . Treloar et al. (1996) revisited this scenario and suggested that heating metamorphism and melting of the lower arc were rather related to magmatic underplating and intra arc-rifting (Khan et al., 1989, 1993).


Fig. 5.14: Distribution and density of the set 2 and 3 shear zone patterns. Squared numbers give location of photographic plates.

Age constraints indicate that SW-ward thrusting responsible for the three successive sets of anastomosing shear zones was active between 100 and 80 Ma at the latest. This supports the interpretation in terms of progressive development of the anastomosing pattern during magmatic, quasi isobaric cooling and subsequent retrogression through the hornblende blocking temperature of the calc-alkaline plutons.


Fig. 5.15: Set 3 shear zones (locations in Fig. 5.14). (a) Feldspar-rich normal shear zone (scale hammer is squared). (b) Hornblendite and felsic veins curved within the lower boundary of a $S W$-ward normal shear zone.

This time span range includes shearing in the Kamila Amphibolitic Belt between 100 and 83 Ma (Treloar et al., 1990) but also the extensional emplacement of the Chilas Complex (Burg et al., 1998) at c. 84 Ma (U-Pb age, Zeitler et al., 1981). Because of apparent age discrepancy and the structural presence of an extensional zone between the southern Kohistan and the Northern Suture, with Asia, we do not follow Treloar et al. (1996) in their collisional interpretation of the studied shear zones and shear fabrics. Instead, we attribute the shear zones anastomosed in the lower part of the Kohistan, to arc-related deformation during northward subduction of the Tethys lithosphere below the Kohistan Complex. The sense of subduction has imparted the bulk, syn-magmatic sense of shear.

### 5.1.6 Conclusion

The calc-alkaline plutons forming the lower part of the Kohistan Complex have recorded the dynamics during their emplacement and subsequent solid state deformation. Regionally distributed magmatic-related fabrics and set 1 shear zones developed under solidus conditions. Set 2 and 3 anastomosing shear zones were superposed onto set 1 shear zones, with increasingly localised deformation along Riedel shear directions. The anastomosing pattern has grown through lengthwise propagation of initially discrete set 1 shear zones while thickening of the propagating shear zones remained minor. These events occurred successively between 100 and 83 Ma during the continuous SW-ward shearing, probably initiated by subduction accretion of the Tethys lithosphere below the Kohistan arc.

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### 5.2 SPATIAL SHEAR ZONE DISTRIBUTION

Chapter 5.2 provides additional supportive data for chapter 5.1 and extends the investigated area to the north.

### 5.2.1 Approach of data analysis

Close to 450 ductile shear zones were measured in the mapped area. A statistical approach groups these data and shows mean values. Before calculating mean values, a grouping of the data was done according to geological aspects. Geological criteria were:
a) lithologies, b) structural similarity and c) direction of tectonic transport. Six lithologies (respectively areas) were chosen based on the results of detailed mapping and structural appearance: 1) Indian Plate, 2) Granulitic gabbro and Sarangar gabbro (Patan area), 3) Northern sheared Gabbros and Diorites, 4) Kiru Amphibolites, 5) Kamila Sequence and 6) Northern Kamila Sequence. The Indian Plate rocks are structurally decoupled from the Kohistan Arc Complex along the Indus Suture and therefore treated as a distinct area. The Granulitic gabbro and Sarangar gabbro are two large intrusive bodies with similar ductile deformation structures without a decoupling horizon in between. In the Northern sheared Gabbros and Diorites lithologies alternate in laccolitic intrusions from gabbros, hornblendite gabbros, diorites, norites and tonalites. A dense pattern of anastomosing shear zones is observed in some of the coarse grained gabbros which alternate with interlayered coarse grained and fine grained diorites. This difference to the appearance of the Sarangar gabbro in the south and to the Kiru Amphibolites in the north suggested grouping the Northern sheared Gabbros and Diorites as a single area. In the Kiru Amphibolites shear plane orientations are dominantly E-W striking and the pronounced anastomosing pattern observed towards the south disappeared. The area comprises sills of dominantly fine grained rocks of gabbroic to dioritic composition, interlayered gabbros and diorites with more pegmatite and granitic sills towards the north and a medium grained hornblende-rich diorite with locally strong foliation forming spaced anastomosing shear zones. The Kamila Sequence is build of several lithologies like fine banded amphibolites, granite, metagabbros and hornblendites. The sheet wise granitic intrusions are characteristic. Ductile shearing is concentrated in the finer grained amphibolites and gabbros, whereas the bulkier granite sheets are internally less affected. The Northern Kamila Sequence comprises in its southern part coarse grained amphibolites alternating with fine grained
amphibolites and tonalitic intrusions and in its northern part a coarse grained metagabbro with quartzite sheets and xenoliths of meta-carbonates. Ductile shearing is weaker than in the Kamila Sequence and more heterogeneous distributed in its northern part. This above briefly described lithologies/areas and their distribution and appearance of the ductile shear zones led to the applied separation.


Fig. 5.16: Equal area projection, lower hemisphere, of 446 measured shear zones within the investigated area. Plots on the left side show pole of planes, plots on the right side show lineations. Contouring is done with small circle count and contour interval 0.5. The overall NE-SW trend is clearly visible with contours.

Beside a regional separation data were grouped according to their orientation of stretching lineation and sense of shear. This was done with following reasons: 1) a main tectonic transport towards the SW occurred in the Lower Kohistan Arc Complex as recognised in the study of Treloar et al. (1990) and in the results shown in chapter 5.1, 2) top to the NE shear zone account for ca. $20 \%$ of the data, therefore they have statistical importance and possibly reflect a tectonic event different to the top to the SW shearing and 3) shear zones not representative for top to SW or top to NE shearing were treated separately to identify other regional preferred orientations. Separated segments are: 1) for transport to SW (Lineation $0-90^{\circ}$ and $180-270^{\circ}$ ), 2) transport to NE (Lineation $0-90^{\circ}$ and $180-270^{\circ}$ ) and 3) SE-NW transport (Lineations in SE and NW quadrants).

Mean foliations and lineations were calculated using a grid (spacing ca. 800 meter) and calculating the mean direction for each grid cell. The actual number of data for one cell can range from 2 to ca. 28 with an average of ca. 10 measurements for a counted grid cell (cells with 1 measurement were rejected). The mean direction (foliation or lineation) for each grid cell containing more than two measurements is plotted with one symbol. For each of the six lithologies/areas, a lower hemisphere equal area projection provides the measurements (see also appendix for data).

### 5.2.2 Results

The shear zone data presented here (Fig. 5.16) show a similar pattern in the lower hemisphere projection as set 2 and set 3 shear zones defined in chapter 5.1 (Fig. 5.11). However, a direct analogous to set 2 and set 3 shear zones cannot be done, especially for the northern parts, where shear zones are of less anastomosed character. The statistical maximum of the shear plane poles plunges ca. 45 towards SW. The lineations plunge with ca. $35^{\circ}$ towards NE. A separated analysis of shear plane poles for top to NE shear zones and top to SW shear zones provided the same maximum contouring orientation of ca. 210/55. Lineations belonging to top to NE shear zones generally dip towards NE and are similar to the orientation of the top to SW shear zone lineation, but with opposite sense of shearing.

The regional distribution of shear plane poles is shown in Fig. 5.17. In the Patan area the poles scatter due to the geometry of anastomosed shear zones, but still define a
maximum at ca. 230/60. Towards the N , in the northern sheared gabbros and diorites the poles are more concentrated in the SW quadrant. Most striking is the pronounced E-W strike of shear zone planes within the Kiru amphibolites with an average dip of $50^{\circ}$ towards N. This reflects a transitional (?) change in the shear zone appearance from an anastomosing pattern in the Sarangar gabbro to a more parallel pattern in the Kamila Amphibolites.

Shear zone planes: Spatial Averaging


Fig. 5.17: Spatial Distribution of shear zone planes and lower hemisphere equal area projections of pole to planes. Foliations shown on the map were calculated using a grid (spacing ca. 800 meter) and calculating the mean direction for each grid cell. Therefore the actual number of data for one cell can range from 2 to ca. 28 with an average of ca. 10 measurements for a counted grid cell (cells with 1 measurement were rejected). Lower hemisphere equal area projections of pole to planes account for each of the six areas and provide the measurements. Contouring is done with small circle count and contour interval 1.

Shear zone planes within the Kamila sequence are more scattered, defining a maximum of ca. 215/50 for the poles. The mean orientations are similar for normal and reverse sense shear zones.

Top to SW Shear zone lineation: Spatial Averaging


Fig. 5.18: Spatial Distribution of lineations representing top to $S W$ shear zones and lower hemisphere equal area projections of lineations of set 2-3 shear zones. Lineations shown on the map were calculated using similar procedure and settings as described in Fig. 5.17 except for top to SW normal shear zones, where cells with 1 measurement were not rejected. Lower hemisphere equal area projections of lineations and pole of planes belonging to them provide the measurements. Contouring is done only for lineations.

Looking at the orientations of lineations related to southwest-ward transport direction, a difference between the southern and northern parts can be observed (Fig. 5.18). In the Patan area the orientations of the lineation varies within the NE quadrant and a significant number plots also in the SW quadrant. Also the lineations plotted on the map show that the SW plunging lineations (normal sense) and NE plunging lineations (reverse sense) are not concentrated in specific locations. At outcrop scale they are coexisting, which is explained with the bending of the shear zones around the undeformed pods (see Fig. 5.6), with the vast majority being reverse sense shear zones. In the northern sheared gabbros and diorites the orientation of the lineations is similar to that of the Patan area. Within the Kiru amphibolites the lineations show a consistency in orientation, slightly oblique to the dip direction of the shear zone planes. Only in the more western part lineations are a more westward directed than along the Indus valley, but they still keep their relative orientation to the shear plane consistent. The regional metamorphic foliation is bended there as well, suggesting that the orientation of shear zones follow the direction predefined by regional foliation, which is almost parallel to the lithological boundaries. The Kamila sequence displays a very homogenous orientation for the lineations at ca. 70/35. Lineations are oblique to the shear planes, suggesting a sinistral strike slip component during shearing. The lineations measured within Precambrian gneisses directly in the footwall of the ISZ show SSW thrusting, the direction in which the Kohistan Arc Complex was thrust over the Indian Plate.

The top to NE shear zones are frequent in the Patan area and the northern sheared gabbros and diorites (Fig. 5.19) and are northeast-directed normal set 2 shear zones (for definition refer to chapter 5.1.4) developed essentially in response to rotation of hornblendite bodies and split lenses. The lineations show a preferred orientation of ca. 60/40 and 40/35. In the Kiru amphibolites the normal shear zones are concentrated in the northern part. The lineations cluster around a mean value of $50 / 60$ which is almost $20^{\circ}$ steeper than the lineations in the southern ares. The shear planes are also generally steep dipping towards N . Within the Kamila sequence the top to NE shear zones play a less significant role.

Shear zones not fitting the top to the NE or top to SW relative transport criteria are summarized in Fig. 5.20. The variation in lineation orientation in the Patan area is
strong and is most likely due to the anastomosing geometry of the shear zones. To accommodate deformation around lens-shaped undeformed pods, especially at the tips of the lens, shear direction varies from the dominant NE-SW direction. In the Kamila sequence the SE-NW directed lineations are noticeable. Their orientation relative to the shear planes point to sinistral strike slip component.

Top to NE Shear zone lineation: Spatial Averaging


Fig. 5.19: Spatial Distribution of lineations representing top to NE shear zones and lower hemisphere equal area projections of lineations of set 2-3 shear zones. Lineations shown on the map were calculated using similar procedure and settings as described in Fig. 5.17. Lower hemisphere equal area projections of lineations and pole of planes belonging to them provide the measurements. Contouring is done only for lineations.

SE-NW Shear zone lineation: Spatial Averaging


Fig. 5.20: Spatial Distribution of lineations representing shear zones with NW and SE directed lineations and lower hemisphere equal area projections of these lineations. Lineations shown on the map were calculated using similar procedure and settings as described in Fig. 5.17. Lower hemisphere equal area projections of lineations and pole of planes belonging to them provide the measurements. Contouring is done only for lineations.

### 5.2.3 Interpretation and Discussion

The orientations of shear planes, lineations and a bulk sense of shear towards SW for amphibolite facies shear zones in the Patan area and for the northern sheared gabbros and diorites is similar to the shear zones described in chapter 5.1. Most of the data presented here were obtained on the same shear zones used for defining set 2 and set 3
shear zones (chapter 5.1). They are explained by the development of anastomosing shear zones as described in chapter 5.1.

Shear zones in the Kiru amphibolites are heterogeneously distributed. In the coarse grained diorites and gabbros shear zones can be observed as discrete zones, dominantly mylonitic developed with a thickness of centimeter to half a meter and an average spacing of 50 to 100 m . The fine grained amphibolites are distinguished in banded and unbanded (Treloar et al., 1990). The banded amphibolites are characterised by alternating hornblende-rich and plagioclase rich layers commonly a few cm wide. Most of these fine grained, banded amphibolites are intensely sheared plutonics and metavolcanics and/or meta-sediments (Treloar et al., 1990). The mylonitic assemblage is composed of quartz, plagioclase, amphibole and porphyroclastic garnet. This assemblage points to amphibolite facies conditions, which are not dated but lasted probably until the $83 \mathrm{Ma} \mathrm{Ar}-\mathrm{Ar}$ age proposed for the end of the regional amphibolite facies metamorphism (Treloar et al., 1989a; Wartho et al., 1996). The steep to north dipping shear planes are parallel to regional foliation and are well exposed along the contacts between coarse grained and banded amphibolites (Fig. 5.21b). Within the banded amphibolites spacing is variable. The shear zones in the Kiru amphibolites are more parallel without a pronounced anastomosing pattern. Pods enclosed by anastomosed shear zones are larger than tens of meter, leading to a more parallel pattern in the outcrop scale. The more gently dipping shear planes crosscut the regional foliation (Fig. 5.21a) and are interpreted as being a later shearing event in the continuous SW-ward shearing. Spacing between the flat shear zones is around 50 to 100 m . In the northern part of the Kiru amphibolites several N -side down shear zones are observed in a zone, where lithology alternates sheet-wise between coarse grained diorites, gabbros, fine grained banded and unbanded amphibolites within tens to hundreds of meter. The mylonitic assemblage is composed of quartz, plagioclase, amphibole, epidote and porphyroclastic garnet which points to epidote-amphibolite facies conditions. These shear zones may reflect an extensional phase. The timing of the N -side down shear event remains unclear but should be prior to the end of the regional amphibolite facies metamorphism. A suitable linkage for the extensional phase is an intra arc-rifting event proposed by Khan et al. $(1989,1993)$ with the extensional emplacement of the Chilas Complex (Burg et al., 1998) at about 85 Ma. Possible
mechanism for the extension are: 1) Retreating subduction zone as proposed by Treloar et al. (1996), based on a scenario described by Royden (1993) and 2) gravity collapse of an abnormal thickened crust.

Shear zones in the Kamila sequence are observed within fine to coarse grained metagabbros and meta-diorites. Two large granite sheets are not pervasively sheared, probably reacting in a more rigid manner. Between the two granite sheets pegmatiteswarm are stretched and sheared (Fig. 5.21c). Within this zone the top to SW shear zones comprises a sinistral dip-slip component within the frame of dominantly SW directed relative transport, indicated by a top to the W-WSW shearing direction oblique to the roughly NE dipping shear planes.


Fig. 5.21: Different shear zone pattern in the central Metaplutonic Complex: a) flat dipping top to SW shear zone close to Kiru; b) steep to $N$ dipping top to SW shear zone within metadiorites close to Kiru; c) gently northward dipping shear zone in amphibolites with sheared pegmatite lenses between the two big granite bodies.

### 5.2.4 Conclusions

The presented shear zones in the Patan area and in the northern sheared gabbros and diorites reflecting with their varying orientations the shape of the anastomosing shear zone pattern are interpreted as set 2 and set 3 shear zones (as defined in chapter 5.1). Towards the N shear zones are getting more parallel, especially where lithology alternates between coarse grained meta-plutonics, fine grained amphibolites and banded amphibolites, reflecting a strong layer parallel strain localisation. The anastomosing pattern is less developed and expressed dominantly by flat lying to gently N dipping shear zones, cutting the layer parallel shear zones (Fig. 5.21a). In the Kamila sequence shearing is dominantly SW directed.

The time span range for shearing in the Kamila Amphibolitic Belt between 100 and 83 Ma (Treloar et al., 1990). The continuous SW-ward shearing, probably initiated by subduction accretion of the Tethys lithosphere below the Kohistan arc, produced the anastomosed shear zones in the lower part of the Kohistan arc. However, there is evidence for an extensional phase expressed by lower amphibolite facies N -side down shear zones in the northern Kiru amphibolites. They are probably related to an intra arcrifting event (Khan et al., 1989, 1993) contemporaneous to the extensional emplacement of the Chilas Complex (Burg et al., 1998) at about 85 Ma .

## 6. FAULT SYSTEMS AND PALEO-STRESS TENSORS IN THE INDUS SUTURE ZONE

Chapter 6 is "Fault systems and Paleo-stress tensors in the Indus Suture Zone" by G. Zeilinger, J.P. Burg, N. Chaudhry, H. Dawood and S. Hussain (Zeilinger et al., 2000).

## Abstract

Analysis of fault-striations measured in the Kohistan part of the Indus Suture Zone (NW Himalaya, Pakistan) has been carried out to document dynamic evolution during the brittle stage of the collision of India and Asia. Processing of the data with a direct inversion method identified four stress fields which were chronologically ordered from field evidence as SSE-NNW compression, E-W compression, radial extension and SSW-NNE compression. The last corresponds to the present-day stress field defined from seismic activity. The earlier stress fields are related to times during the Miocene, when convergence-related stresses were disturbed by the formation of the nearby Nanga Parbat and Indus syntaxes.

### 6.1. Introduction

In NW Pakistan, the Kohistan Complex developed as an island arc above the northward subducted Tethyan lithosphere during the Mesozoic (Bard, 1983; Coward et al., 1986). The southern boundary of the Kohistan Complex is the Indus Suture (the Main Mantle Thrust $=$ MMT of Tahirkheli et al., 1979), which is a crustal-scale, 35-50 northward dipping fault contact (Malinconico, 1989), along which the Kohistan Complex has been thrust over India (Fig. 6.1). Collision between India and Kohistan began at about 65 Ma (Beck et al., 1996), and continued with subsequent obduction onto India (Coward et al., 1987).


Fig. 6.1: Studied area (squared) located in the structural framework of Northern Pakistan, with the Kohistan Complex, main sutures and syntaxes.

The syn- and late-collisional history of the Indus Suture includes southward directed thrusting and extension (Treloar et al., 1991a; Burg et al., 1996). Fission track ages (Zeitler et al., 1982) record the termination of significant differential movement across the Indus Suture at 15 Ma .

Ductile structures have been investigated by several authors (e.g. Coward \& Butler, 1985; Coward et al., 1986, 1987; Treloar et al., 1990; Arbaret et al., 2000) in order to understand early collisional processes. However, faulting is an important part of the long-lived deformation history that has produced the present day structure of the suture zone (Fig. 6.2a). Some areas are seismically active (Seeber \& Armbruster, 1979) and it is evident in the field that, despite the lack of surface rupturing during particular earthquakes (e.g. Pennington, 1979), rocks have been fractured. This study concentrates on the analysis of fault-striations (Fig. 6.2b and 6.2c) in order to document the dynamics of this part of the Indus Suture Zone during hypercollision.


Fig. 6.2: a) fault zone in meta-gabbro, 4 km NE of Duber Kale (locality in Fig. 6.4) (circled cow as scale). Slip plane: 305/36. This fault shows sinistral movement and is attributed to population 4. b) normal fault (slip plane 129/88) with serpentine fibres (046/58) indicating the downward relative movement of the hangingwall block ( $4 \mathrm{~km} N$ of Duber Bazaar located in Fig. 6.4). c) same fault plane as $1 b$ with several fibre orientations: 046/58 and 050/21; fibre continuity indicates a relatively short change in movement direction. For fault population distinction only the principal fibre direction was used, fitting population 3. d) superposed striations on one fault plane (slip plane 290/18), the older one (lower arrow, 274/17, population 2) is cut by the younger one (upper arrow, 011/03, population 4).

The investigated area straddles the Indus Suture and comprises three main units that are from S to N, i.e. from the structurally lowest (Kazmi \& Jan, 1997; Treloar et al., 1996):

- The Indian unit that includes granodiorite and intensely foliated and folded gneisses with a marked SW trending stretching lineation (Coward et al., 1988; Treloar et al., 1989b).
- The lower Kohistan unit comprising ultramafic and mafic rocks at the bottom, and mainly gabbros, hornblende-gabbros, diorites, norites, amphibolites and small bodies of hornblendite (the Kamila Amphibolite Belt of Jan, 1988) towards the top.
- The Chilas Complex which is composed of gabbro-norite and subordinate diorites (Jan et al., 1984).


### 6.2. Method

Ca. 650 measurements of fault/striation pairs from 30 sites form our data set. Outcrops defined as sites are sections not longer than 100-200 m. The size of faults, quality of criteria for sense of movement and faulting style (for example discrete plane versus gouge zone) were also noted in order to characterize homogeneous fault populations. In particular, we distinguished faults with new crystallization (chlorite, apatite etc.) from "dry" faults with striated slip planes and without new crystallization, indicating that "dry" faults were formed at lower temperature, closer to the surface, and therefore later than the crystal coated planes. Conjugate fault sets were used in the field for a first estimate of the bulk shortening (compression) direction giving evidence of separate faulting events. Superposed striations on single fault planes further supported the chronological classification (Fig. 6.2d). These field observations, which yielded homogeneous data sets, were the key for computer aided fault distinction and stress tensor calculations. A first computer analysis of single-site data consistently required more than one stress tensor to explain the field measurements. Consequently, fault striation data have been separated for every site to determine successive stress tensors (Fig. 6.3). Data processing was carried out with Software FSA written by B. Célérier, which is based on a direct inversion algorithm (Bott, 1959; Compton, 1966; Etchecopar et al., 1981). Assumptions of the method are: 1) Direction of the resolved shear stress on the fault is parallel to the measured striation and records the slip direction on the
fault; 2) Slip on each fault is independent from other new or reactivated faults and 3) the stress directions are homogenous on the scale of a site.


Fig. 6.3: Data processing illustrated with data set 29/08 ( 6 km NE of Patan along the Karakorum Highway, length of section 150 m ). This site is part of the "regional data set", a combined set of measurements belonging to an area where faulting can be assumed to be homogeneous. Once the successive, preliminary regional tensors have been calculated from the "regional data set", each site was filtered through these stress tensors (misfit angle $\leq 40^{\circ}$ ). In addition, data sets from each site were separated without employing the regional stress tensors. Both resulting subsets were compared and rearranged to populations with coherent ol directions. These separated faults were processed again with a random tensor search to obtain local deviations in terms of shape and orientation of the local stress tensor. Populations of less than 5 faults and unexplained data were considered to be not significant. Therefore they are not shown in Fig. 6.5 to 6.8 .

To ascertain regional stress fields, 248 measurements within a single rock unit, a metagabbro of Lower Kohistan, near Patan, were processed as one data set (Fig. 6.4). The aim was to eliminate small, local deviations and to integrate over a large area the bulk brittle tectonics. In the first processing step, tensors of maximum ( $\sigma 1$ ), intermediate ( $\sigma 2$ ) and minimum ( $\sigma 3$ ) principal stresses (positive in compression) best fitting $30 \%$ of the measurements were randomly calculated. This random tensor search is a Monte Carlo approach to search a stress tensor that best explain the slip directions. This mainly follows the method proposed in Etchecopar et al. (1981) and Etchecopar (1984). In a first step of the random tensor search, the stress tensor data are generated by using a random variable, so that the orientations are uniformly distributed in space. In a second step, for each tensor, the angular misfit between the predicted and observed slip direction for each fault slip data is computed. In a third step, for each tensor, an average angular misfit, is computed. An adjustable percentage of the fault slip data must be explained. The angular misfit is the average of the angular misfit for the chosen percentage of the data that have the lowest angular misfit. In a fourth step the tensors are ranked by increasing value of the angular misfit and only the first tensors are retained. These stress tensors are the most compatible with the fault data set.

The best fitting tensor integrated the highest number of misfit angles $<30^{\circ}$ and as close as possible to $0^{\circ}$ between the optimal tensor for each fault/striation pair and the calculated ones. After extracting the corresponding fault data and checking their consistency as a population compared with field determination, computation was performed on the extracted population with a stepwise increase of the best fit for 40,50 , 60 and finally $70 \%$ of the measurements. The stepwise increase is justified by the fact that every extraction decreases the number of possible faults in the extracted data set. This procedure yielded 4 bulk stress tensors that explain $70 \%$ of four separate populations embracing most of the 248 measurements. These 4 tensors were taken as the preliminary regional, yet smoothed, reference because they approximately fit compression/extension directions estimated from conjugate sets in the field. About $10 \%$ of the measured faults were not consistent with any of the four reference tensors.


The unexplained fault planes may be due to local accommodation between adjacent, rotating blocks, thus recording local deformation unrelated to the regional deformation. Treated together the un-explained faults did not yield any acceptable stress tensor.

Data from every site were then separated by successively imposing the 4 regional stress tensors and calculating the misfit angle of each fault/striation pair. Pre-existing anisotropy in the rocks caused by foliation and earlier fault planes justified allowing a misfit of $40^{\circ}$ as the separation criterion for fault data fitting each tensor.

The obtained data subsets were compared to populations identified from geological observation. Results overlapped by more than $75 \%$. We rearranged the subsets by comparing faults best fitting the tensors obtained through single site separation with the fault population obtained by imposing the regional tensors. The main criterion for rearranging the subsets was the similarity of the mean stress direction independently from the misfit angle.

In order to obtain the local orientation of the stress tensors, each subset underwent a further random tensor search accounting for $70 \%$ of the subset. Each local stress tensor has specific orientation and R -ratio, which is defined as $\mathrm{R}=(\sigma 1-\sigma 2) /(\sigma 1-\sigma 3)$ and describes the shape of the stress ellipsoid (Célérier, 1988; Célérier, 1995). An R-value close to 1 indicates small differences between $\sigma 2$ and $\sigma 3$. R close to 0 indicates that $\sigma 1$ and $\sigma 2$ are almost equal.

The four regional tensors were not obtained at every site. Some tensors may have escaped calculation either because the relevant fault population was locally too small, or because the site area did not deform during the particular brittle event. A few fault data could be attributed to one or other of two different stress tensors. Omission of these data did not affect either calculation. Regional correlation from site to site relies on both the orientation and the shape of stress tensors and the chronological sequence based on field relationships between identified fault and striation sets.

Fig. 6.4 (left page): Unseparated data from 30 sites. Sites used for calculation of the regional stress tensors are not separated by major faults and they are clustered in one, reasonably small geographic area (marked by dashed line) in a single rock unit (a massive metagabbro). Fault/striation plotted in stereo-net, lower hemisphere (fault plane: great circle, striation: arrow, pointing towards movement direction of the hanging block).

A question is whether the direct inversion method is a valid tool for separating multiphase data sets. We might get only mixed stress tensors which are geologically meaningless (Nemcok \& Lisle, 1995; Lisle \& Vandycke, 1996). The latter assertion is correct when calculating stress tensors on inhomogeneous data sets without taking geological observations such as fault style and conjugate faults into account. The computed results presented here were verified by information independent of our computation to avoid mixed stress tensors. Although our attempt is not free from uncertainties inherent to areas of complex-deformation, the results give a plausible overview of the development and the succession of stress fields in this particular area of the Himalayas.

### 6.3. RESULTS

The distribution of the calculated stress tensors depicts 4 regional stress fields.

### 6.3.1. Population 1: SSE - NNW directed ol + ENE - WSW o3

Population 1 faults are dominantly steep N-S (sinistral) and (dextral) WNW-ESE strike-slip faults (Fig. 6.5). They are sharp planes with chlorite crystallization and/or a few centimetres to half a meter thick gouge zones. The R-values vary from site to site from 0.31 to 0.99 pointing to a nearly uniaxial SSE-NNW compression at some places. Site deviation from the regional stress field is mainly expressed by swapping $\sigma 2$ and $\sigma 3$, which is consistent with $\mathrm{R} \cong 1$. Ca . $20^{\circ}$ plunges of some $\sigma 1$ axes are attributed to block rotations due to later faulting and/or to local variations in stress field orientation.

### 6.3.2. Population 2: $E-W$ directed $\sigma 1+$ subvertical $\sigma 3$

Population 2 faults are dominated by $30^{\circ}$ to $55^{\circ}$ dipping thrust faults and secondary, steep strike-slip faults well defined N of Dasu and N of Duber Kale (Fig. 6.6) where sinistral reverse movement is notable in a broad serpentinized zone, which is the northern continuation of the Alpurai fault (Anczkiewicz et al. 1998). Fault thickness ranges from one centimeter to few decameter. Faults often display gouge zones and/or tectonic breccias. Local plunges $\left(<15^{\circ}\right)$ of $\sigma 1$ are attributed to late block tilting and/or to local variations in stress field orientation.

Fig. 6.5 (right page): Population 1, dominantly strike-slip faults fitting SSE-NNW directed or. $R$-values and direction of stress axes are plotted for each site (lower hemisphere, 1: o1, 2: o2, 3: o3). Orientation of regional stress axes plotted as "beachball" (white: compression, black: dilatation, arrows: expected movement on intersection between compressional and dilatative quadrants). Distribution of misfit angle is given in a histogram (n: number of faults, $x$-axis: misfit angle) and the stress tensor aspect ratio $R$ is represented in a Mohr circle diagram ( $\tau$ : shear stress, on: normal stress).

Fault systems and Paleo-stress tensors in the Indus Suture Zone



Fig. 6.6: Population 2, strike-slip and thrust faults that were activated by E-W compression. Legend as Fig. 6.5.

Highest R-values (0.75-0.90) are within the Kohistan Complex. Smaller R-values to the S of the Indus suture and towards the N of Patan imply a different shape of the stress ellipsoid in the footwall ( $\sigma 2>\sigma 3, \sigma 3$ subvertical) and immediate hanging-wall ( $\sigma 2 \approx$ $\sigma 3, \sigma 2$ subvertical) of the Suture. This suggests a direct influence of the Indus Suture on the stress distribution.

### 6.3.3. Population 3: $W N W-E S E$ extension (subvertical $\sigma 1$ )

Normal faults, occurring as slicken-sided fault planes and as wider gouge zones can be observed in many places. The calculated stress tensors are characterised by $\mathrm{N}-\mathrm{S}$ and $\mathrm{E}-$ W extensional directions (Fig. 6.7). R-values dominantly close to 1 indicate little difference between $\sigma 2$ and $\sigma 3$, i.e. radial extension. Sites with $\mathrm{R}<0.6$ have preferred E-W directed $\sigma 3$. N of Dasu and in the Kandiah valley (K1 and K2, Fig. 4), normal faults do not play a significant role whereas closer to Chilas (C1, Fig. 6.7) normal faulting is better expressed. This seems to indicate that brittle extension was more localised than earlier brittle deformation.

### 6.3.4. Population 4: SSW - NNE directed ol + WNW - ESE o3

Population 4 is defined by a set of dry $\mathrm{N}-\mathrm{S}$ dextral and conjugate $\mathrm{SW}-\mathrm{NE}$ sinistral strike-slip and subordinate thrust faults (Fig. 6.8). Fault planes are neat fractures without crystallization on their surfaces. Movement indicators are Riedel shears and striae on weakly polished planes. This fault population is dense around Patan but little represented within India, which suggests some brittle "decoupling" between the Indian footwall and the allochthonous Kohistan Complex. R ranges from 0.25 to 0.98 with a majority of ratios $\geq 0.8$. The interchangeable orientation of $\sigma 2$ and $\sigma 3$ points to faulting under close to surface conditions (small burial loads). Single sites with low R-ratio have either $\sigma 2$ or $\sigma 3$ subvertical which means that this difference reflects very local effects and depends whether thrust faults ( $\sigma 3$ steep) are more abundant than strike-slip faults ( $\sigma 2$ steep).


Fig. 6.7: Population 3, normal faulting. Legend as Fig. 6.5


Fig. 6.8: Population 4, SSW-NNE directed ol. Legend as Fig. 6.5
$80 \%$ of the measurements from Kandiah valley (K1 and K2) and W of Chilas (C1) can be explained by the regional stress tensors. In these far distant sites, field observation and computer processing yielded a stress history similar to that obtained in the Patan region but with slightly different orientations. The orientation changes indicate that the regional stress fields have orientation variations not covered in this work.

### 6.4. InTERPRETATION AND DISCUSSION

The analysis of the regional fault-framework provides evidence for four stress fields in the Patan region and probably over a large part of the Kohistan Complex. The idea of using a regional data set for separating subsets belonging to specific stress tensors is an appropriate way to depart from block rotations and noisy, background faulting. Block rotation in this area of complex faulting history was not large enough to seriously affect our separation procedure and is reflected in local, small plunges of the oldest stress axes. Supporting evidence is that, no significant, sudden change in attitude of ca. 90 Ma old shear zones and magmatic fabrics (Treloar et al., 1990; Arbaret et al., 2000) pointing to large block rotations has been observed.

The relative timing of the stress fields derived by crosscutting striations and fault planes that are unequivocally and separately attributable to one specific stress tensor yields the following sequence: population 1 is older than population 3 , population 2 is older than population 3 and population 4 is younger than population 2. Relative age constraints for population 1 and 2 and population 3 and 4 were not found so that we were left with four possibilities: the adopted 1234 sequence or eventually 2134,2143 and 1243 history. The proposed timing was inferred from the following observations:
-Fault planes containing chlorite fibres are the highest temperature faults and are thus likely to be the oldest. They essentially belong to population 1 which seems, therefore, older than population 2. The indicated SSE-NNW compression fits the convergence direction expected in this Himalayan region (e.g. Molnar \& Tapponnier, 1975). Apatite fission track ages are $7.3 \pm 2 \mathrm{Ma} 20 \mathrm{~km}$ N of Dasu and $5.6 \pm 2.3 \mathrm{Ma}$ close to Besham (sample 79I-9 and 79I-12 in Zeitler, 1985). Assuming that chlorite, apatite and fibrous mineralisation found on population 1 fault planes were formed not much over 100$140^{\circ} \mathrm{C}$, which is the range for the apatite fission track closure, a maximum age for these faults can be estimated at around 9 Ma (Fig. 6.9). This suggests, that the four brittle
events recognised in this paper are younger than the normal faulting in ductile-brittle conditions along the Indus Suture (Burg et al., 1996; Vince \& Treloar, 1996) which has been dated as older than 18 Ma (Vince \& Treloar, 1996).

We relate the E-W compression recorded in population 2 faults to the formation of $\mathrm{N}-\mathrm{S}$ trending crustal antiforms such as the nearby Nanga Parbat and Indus syntaxes (Coward, 1985; DiPietro et al. 1999) (Fig. 6.1). They are crustal scale NS buckle folds that should reflect E-W compression sometime within the last 5 Ma (Zeitler, 1985). Indeed, reconnaissance work has revealed that the importance of $\mathrm{E}-\mathrm{W}$ compression increases eastward towards the main syntaxes. The studied area is structurally the hangingwall of the syntaxes.


Fig. 6.9: Proposed time sequence and orientation of regional stress tensors.

Population 3 extension took place after the E-W compression. We tentatively relate this extension to lateral collapse of the hanging-wall of the crustal anticlines, in particular the Nanga Parbat structure.
-Fault characteristics of population 4 indicate cold, subsurface conditions, as expected for recent fracturing. The $\sigma 1$ direction calculated from population 4 faults corresponds to the present-day compression (Fig. 6.10), which produced the Patan earthquake in December 1974 (Ambraseys et al., 1975). Therefore, we deduce that population 4
documents the ongoing southward thrusting of the Kohistan Complex onto the Indian Plate (Pennington, 1979) and is younger than population 3.


Fig. 6.10: Recent stress field in Northern Pakistan and adjacent countries (World stress map, release 1997/1). Additional: Focal mechanism solution for the Patan earthquake, lower hemisphere, indicating nearly pure thrusting in a NNE-SSW direction, along either of the two nodal planes (Pennington, 1979). Population 4 is in accordance with this record.

### 6.5. CONCLUSIONS

Faulting in and near the Indus Suture Zone results from the existence of four successive stress fields that point to a complex tectonic evolution. Early faults formed during southward thrusting of the Kohistan Complex, probably in Miocene times. Population 2 faults occurred during E-W compression related to the formation of N-S striking buckle folds such as the Nanga Parbat and Indus syntaxes. That compression disturbed the stress field related to N-S convergence between India and Eurasia. We speculate that population 3 developed as lateral collapse structures in the hanging-wall of the Nanga Parbat crustal anticline. Population 4 documents a present day return to the stress field controlled by the ongoing southward thrusting of the Kohistan Complex as documented by the focal mechanism solution of the Patan earthquake. Faults subjected to potential reactivation under these stress conditions can be identified with adequate computation.

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## 7. SUMMARY AND CONCLUSIONS

The KAC was installed on the oceanic crust somewhere in the Tethys around the equatorial zone and close to the southern marginal terrane (Karakoram block) of the Eurasian continent (e.g. Yoshida et al., 1996). The Indian plate was, during this time, in the southern hemisphere (Klootwijk et al., 1986a,b). The northward dipping subduction formed during Mid Cretaceous times produced a volcanic arc (Fig. 7.1, >100 Ma). The Sarangar gabbro emplaced 99 Ma ago at pressures of ca. 1.0 GPa. It shows a geochemical signature typical for island arcs. At this time the arc had already a thickened crust (Fig. 7.1, $100-90 \mathrm{Ma}$ ). Consequently, the initiation of subduction must have started before 100 Ma . Remnants of the oceanic crust on which the arc was built are most likely preserved as enclaves in the lower to mid crustal level of the arc (Jan, 1988; Khan et. al., 1993; Treloar et. al., 1996). PT conditions of emplacement, retrograde metamorphism and the continuous section with mantle rocks (Jijal Complex ultramafic sequence) at the base confirm that the Metaplutonic Complex is representing the lower - middle level of the arc. In the Indus area the Metaplutonic Complex is built up mainly by a) Sheared gabbro-diorites (Patan region) comprising retrograde metagabbros with locally preserved igneous layering; b) Kiru sequence consisting of more than hundred meters thick imbricate intrusions of gabbroic to dioritic composition; c) Kamila sequence, an amphibolitised sequence composed of gabbroic, dioritic, granitic and tonalitic rocks intruded in a crustal sequence, whose remnants are represented by fine layered amphibolites with epidote-rich lenses, enclaves of carbonates and quartzite bands.

The KAC evolved quickly from early arc buildup (Fig. 7.1, $>100 \mathrm{Ma}$ ) to a mature stage with back-arc spreading and magmatism (starting at 85 and lasting to at least 80 Ma ago; Fig. 7.1, $90-80 \mathrm{Ma}$ ). As shown in Figure $7.1(100-90 \mathrm{Ma})$ the buildup is represented by a suite of calc-alkaline to tholeiitic plutons intruded in the lower to mid crustal levels of the arc (the Metaplutonic Complex is symbolized in the lower central and southern part of the arc). Some of these plutons yield intrusion ages of $98.9 \pm 0.4$, $97.1 \pm 0.2$ and $91.8 \pm 1.4 \mathrm{Ma}$ (Fig. 7.2, $\mathrm{U} / \mathrm{Pb}$ ages) and represent parts of multiple, fast differentiation cycles from gabbros to granites.
$>100 \mathrm{Ma}$


Fig. 7.1: Proposed model for the Kohistan Arc Complex build-up during northwarddirected subduction of the Tethyan ocean and subsequent collision with Eurasia and India. The Chilas complex and Kohistan batholith III are drawn for simplicity as huge magma chambers, which are unlikely to have existed at this extend.

Kohistan Arc Complex Timeline


Fig. 7.2: Kohistan Arc Complex history showing dated and interpreted events covering the magmatic emplacement, metamorphic overprint, cooling, deformation events and the regional implication. Data from: Zeitler et al. (1981), Zeitler (1985), Peterson \& Windley (1985), Treloar et al. (1989a), George et al. (1993), Yamamoto \& Nakamura (1996, 2000), Anczkiewicz (1998), Anczkiewicz \& Vance (2000) and this work. Extension and compression regime after Treloar et al. (1996).

The early gabbro-dioritic suite of the Kohistan batholith, drawn on a more northern position in the arc (Fig. 7.1, $100-90 \mathrm{Ma}$ ) was also formed within this island arc setting between 110 - 85 Ma (Petterson \& Windley, 1991; Petterson et al., 1991b). The Granulitic and Sarangar gabbros most likely intruded at the base of a thickened arcat pressures of ca. $1.0-1.2 \mathrm{GPa}$ (equivalent to a depth of $\sim 40 \mathrm{~km}$ ) or higher as it is suggested by Ringuette et al. (1999). The mantle - crust transition is then supposed to have been at depths of $40-50 \mathrm{~km}$ around $100-90 \mathrm{Ma}$ (Fig. 7.1). The Hf isotopes give further evidence that the Kohistan Arc Complex is an intraoceanic island arc with minimal continental influence. Subducted components derived by melting of the slab itself were negligible for the generation of the melts. In the highly dynamic setting above a subduction zone the intruding plutons recorded the deformation at the base of the island arc during their emplacement and subsequent solid state deformation. This is symbolized in Figure $7.1(100-90 \mathrm{Ma})$ by the anastomosing pattern. The Metaplutonic Complex and the Granulitic gabbro (the unit just above the in dark grey drawn Ultramafic sequence at the base of the arc in Figure 7.1) preserved an intense, dominantly SW directed shear deformation.

To illustrate the early arc history a temperature, pressure and time relation (Fig. 7.3) is sketched where pressure is assumed to be overburden depth. The emplacement of gabbroic magmas occurred under high temperature $>1000{ }^{\circ} \mathrm{C}$ (Yoshino et al., 1998) and pressures above or around 1.5 GPa at ca. 99 Ma (Sarangar gabbro) or earlier followed by a subsequent granulite facies metamorphism under intrusion conditions or with a prograde path of $\Delta \mathrm{P} 0.3 \mathrm{GPa}$ and $\Delta \mathrm{T} 80^{\circ} \mathrm{C}$. The temperature, pressure and time relation suggests for both, the Granulitic and Sarangar gabbros cooling and decompression from 100 to 97 Ma and nearly "isothermal" decompression from 97 to 87 Ma , where the exhumation path follows the $600^{\circ} \mathrm{C}$ isotherm (Fig. 7.3). Metamorphic conditions associated with set 2 shear zones (chapter 5.1) were estimated at $550-650$ ${ }^{\circ} \mathrm{C}$ under pressures of $0.9-1.0 \mathrm{GPa}$ (Bard, 1983; Treloar et al., 1990).


Fig. 7.3: Simplified geothermal and geobarometrical evolution of the lowest Kohistan Arc Complex (Jijal and Metaplutonic Complex). Pressure gradient is assumed constant with depth. PT values are estimates: at 100 Ma crystallisation of Sarangar gabbro, pressure of ca. 1.5 GPa in granulitic gabbro; at 95 Ma cooling below $700{ }^{\circ} \mathrm{C}$ (Sm-Nd age, Anczkiewicz \& Vance, 2000); at 83 Ma cooling below $500{ }^{\circ} \mathrm{C}$ (Ar-Ar age of hornblende, Treloar et al., 1990); the stippled line represent the cooling path obtained by FT-dating (see Chapter 4.3). Pressures were taken from Fig. 3.11. Since 40 Ma a geothermal gradient of $30{ }^{\circ} \mathrm{C} / \mathrm{km}$ was assumed. The strong "upwelling" of the isotherms coincides with the intrusion of the Chilas Complex at ca. 85 Ma .

Results presented earlier (chapter 3.1) imply that the onset of shearing happened under granulite facies conditions at ca. $800^{\circ} \mathrm{C}$ and $\sim 1.1 \mathrm{GPa}$ and shearing was ongoing under amphibolite facies to conditions of $600^{\circ} \mathrm{C}$ and 0.8 GPa . This range is in agreement with the conditions of the $97-87$ Ma part of the history. Higher temperatures mentioned earlier are supported by the fact that partial melting in set 2 mylonites generated quartz-plagioclase-garnet bearing segregation veins and by the upper amphibolite facies crystal plasticity reported by Treloar et al. (1990). The minimum age for set 2 shear zones is inferred from the $86.7 \pm 4.6 \mathrm{Ma}$ amphibole-epidote-paragonite assemblage that has partially overprinted the granulitic assemblage in the granulitic gabbro (Anczkiewicz, 1998, Ar-Ar, paragonite). Set 2 shear zones were then successively replaced by the later set 3 shear zones (chapter 5.1) which show lower amphibolite facies mylonitic assemblages. The set 2 mylonite analysed in the Sarangar gabbro recorded temperatures around $540{ }^{\circ} \mathrm{C}$ and pressures around 0.6 GPa (chapter 3.1). Therefore it cannot be excluded, that some of the set 2 shear zones were still active when most of the deformation was accommodated by flat lying and less anastomosed set 3 shear zones under lower amphibolite facies conditions.

At ca. 85 Ma a strong "upwelling" of isotherms occurred, coinciding with a drastic change in tectonic style. This was the onset of intra-arc rifting and the intrusion of the Chilas Complex $85.73 \pm 0.15 \mathrm{Ma}$ ago (Fig. 7.1, $90-80 \mathrm{Ma}$; Fig. 7.2, U/Pb ages). The intrusive Chilas Complex is constituted of several intrusions of noritic to dioritic character with ${ }^{87} \mathrm{Sr} /{ }^{86} \mathrm{Sr}$ ratio typical for island arc magma (Miksohiba et al., 1999). The relatively shallow isotherms at this time (Fig. 7.3) are most likely the result of the extensional emplacement of the Chilas Complex (Burg et al., 1998). This stage is characterized by a shift in the Hf isotopic compositions of the magmatic liquids from MORB-type values $(\varepsilon H f=+14)$ to less radiogenic compositions $(\varepsilon H f=+10)$. The Chilas melts $\varepsilon H f$ value is significantly lower than the older magmatic rocks of the Metaplutonic Complex, but well within the range of oceanic basalts. This suggests that an old mantle component with a lower time-integrated $\mathrm{Lu} / \mathrm{Hf}$ ratio than MORB, or even sedimentary or continental crustal contaminants, were involved in the generation of the Chilas parental melts.

The termination of shearing in the Set 3 shear zones is constrained to be under amphibolite facies conditions at ca. $550^{\circ} \mathrm{C}$ and $\sim 0.4 \mathrm{GPa}$, probably between 83 and 76 (Fig. 7.2, Ar-Ar ages) proposed for the end of the regional amphibolite facies metamorphism (Treloar et al., 1989a; Wartho et al., 1996). After ca. 83 Ma further cooling and decompression (exhumation) took place under apparently moderate gradients (Fig. 7.3) with a temperature decrease of $200^{\circ} \mathrm{C}$ and only a slight pressure drop until 40 Ma .

What were the regional tectonic events for rapid decompression and cooling? An important hint for this question is provided by several N -side down epidote-amphibolite to amphibolite facies shear zones described by Treloar et al. (1990) and in Chapter 5.2 in the northern part of the Kiru sequence, which reflect extension. The metamorphic conditions recorded in the N -side down shear zones suggest, that they were not active later than the end of the regional amphibolite facies metamorphism ( $83-76 \mathrm{Ma}, \mathrm{Ar}-\mathrm{Ar}$ age on hornblende, Treloar et al., 1989a, Wartho et al., 1996). The proposed intra arcrifting at ca. 85 Ma (Khan et al., 1989, 1993) could explain extension and resulting development of epidote-amphibolite to amphibolite facies normal sense shear zones. These shear zones may have supported a faster exhumation of the Metaplutonic

Complex. An essential question for the tectonic interpretation is the timing for collision of the Kohistan Arc with the Karakoram plate (Asia). This collision changed the system north of the Indian plate from an island arc to an active continental margin. The closure of the Northern Suture is bracketed between $102 \pm 12$ and $85-75 \mathrm{Ma}$ (Petterson \& Windley, 1985; Treloar et al., 1989a). Treloar et al. (1996) suggest an emplacement of the Chilas Complex "after suturing" of the Kohistan Arc to the Karakoram Plate due to the lower degree of deformation within the Chilas Complex compared to the Metaplutonic Complex. If this assertion is appropriate and the $\mathrm{Rb}-\mathrm{Sr}, \mathrm{Ar}-\mathrm{Ar}$ and $\mathrm{K}-\mathrm{Ar}$ dating is giving precise values, the closure of the Northern Suture was prior to the emplacement of the Chilas Complex or at the same time (Fig. 7.4a).

b) emplacement of Chilas Complex prior to suturing, retreating subduction zone

c) emplacement of Chilas Complex prior to suturing, gravitative collapse


Fig. 7.4: Discussed settings during emplacement of the Chilas Complex at 85 Ma: a) post- to syn- collisional (Kohistan arc - Karakoram) emplacement; the reason of extension remains unclear; b) prior to the closure of the SSZ with extension induced in the upper plate by a retreating subduction zone; c) prior to closure of the SSZ with extension driven by gravity.

In this scenario, the extension would occur in a system, where general compression between the arc and the continental margin is expected. The emplacement of the Chilas Complex could have been also prior to closure of the Northern Suture. In this scenario following processes are possible: 1) Change in subduction angle/speed with retreating subduction zone as proposed by Treloar et al. (1996), based on a scenario described by Royden (1993) (Fig. 7.4b) and 2) gravity collapse of an abnormally thickened arc crust (Fig. 7.4c). Changes in plate tectonics movement pattern or combinations of the processes mentioned above cannot be excluded as potential reason for extension.

The mantle sources were different for the $99-92$ Ma magmatism of the arc and the extensional magmatism at ca. 85 Ma . This reveals a change in tectonic setting. The early magmatism corresponds to a calc-alkaline to tholeiitic differentiation trend. Gabbros, diorites, tonalites and granites were emplaced in a island (volcanic-) arc setting. The related gabbros and diorites preserved a common geochemical signature from the melts and/or melt-source. The overall geochemical signature is similar to volcanics of the Tonga, New Hebrides and Mariana arc systems, clearly showing the arc-type character of the analysed gabbros and diorites. They are different to the high Ti- amphibolites (Khan et al., 1993) and therefore do not represent a remnant ocean floor, which is suggested to be also present in the "Kamila Amphibolite Belt". Diorites intruded ca. 1865 Ma ago in the Indian Plate (at Duber Bazar) are within plate plutons with a clear discrepancy in the geochemical signature compared to the diorites and the arc rocks.

A dominantly basaltic juvenile arc crust is likely to evolve with time towards more acidic and HIE-enriched compositions similar to continental crust. The HIE depletion of the analysed gabbros and diorites tends to "balance" the enrichment of the continental crust. Different processes are likely involved in the arc differentiation and granite formation. Based on the preliminary dataset the cumulitic characters of the gabbros are relatively subtle and these gabbros cannot be considered as proper "cumulates", therefore fractional crystallisation seems to be a minor process. This also indicates that there was no common large magma chamber involved in the formation of the analysed gabbros and diorites. A dehydration of hornblendites possibly triggered incipient partial melting of the lower arc section along with pervasive upward migration of small melt
fraction. The formation of granites is most likely the result of partial melting at the base of the arc. This process would efficiently remove granitic components from the base of the arc and transport them up to shallower levels.

The main bulk of the Kohistan batholith comprises a gabbro-diorite suite at 85 to 60 Ma (Fig. 7.1, $80-70 \mathrm{Ma}$ ) and a granitic suite, emplaced at $60-40 \mathrm{Ma}$ (Petterson \& Windley, 1991; Petterson et al., 1991b). A postcollisional magmatism with leucogranitic sills and dikes occurred between 40 and 26 Ma (Petterson \& Windley, 1985).

Extension took place after the collision between the Kohistan Complex and the passive margin of the Indian plate at ca. 55 Ma , and subsequent thrusting onto the Indian Plate. It is indicated by NE vergent folds, interpreted as collapse structures related to the backsliding of the Kohistan Arc (Burg et al., 1996; Vince \& Treloar, 1996; Anczkiewicz, 1998). The earliest onset of extension is estimated at 47 Ma (Burg et al., 1996). Anczkiewicz et al. (2001) interpret the 29 Ma emplacement age of a pegmatite in the Swat valley, which is deformed by folds related to normal faulting, as expression of an already ongoing extension or as maximum age for reactivation of the ISZ as normal fault. The difference between the zircon FT ages representing footwall ( $\sim 23 \mathrm{Ma}$ ) and hanging wall ( $\sim 42 \mathrm{Ma}$ ) of today's ISZ geometry at the Indus valley (Fig. 7.1, $<10 \mathrm{Ma}$ ) is ca. 20 Ma . The onset of normal faulting occurred latest during this time span. Consequently normal faulting along the investigated part of the Indus Suture started at least $\sim 23 \mathrm{Ma}$ ago. Apatite FT ages are $\sim 3.7 \mathrm{Ma}$ for the footwall and $\sim 10 \mathrm{Ma}$ for the hanging wall. This suggests that normal faulting outlasted the termination of extension of other ISZ segments, where it is reported between ca. 20 Ma (Burg et al., 1996; Vince \& Treloar, 1996) and 15 Ma (Swat valley, Anczkiewicz et al., 2001) based on FT ages from Zeitler (1985). The young differential movement across the ISZ in the Indus area is here a local phenomenon caused by the growth of the Besham antiform, a possible small-scale equivalent to the Nanga Parbat "Syntaxis" (Coward, 1986; Treloar et al., 1989c).

Based on the paleostress analysis, four stress fields could be identified. They were chronologically ordered in SSE-NNW compression, E-W compression, radial extension and SSW-NNE compression (Fig. 7.2, deformation/regional events). The
dynamic evolution during the brittle stage of the collision of India and Asia is interpreted by the four stress fields and their relative timing. The earlier stress fields were active during the Miocene, when convergence-related N-S directed stresses were dominant. That compression was disturbed by an E-W compression which is related to the formation of the nearby Nanga Parbat "Syntaxis" at ca. $10-5 \mathrm{Ma}$ (e.g. Butler \& Prior 1988). The radial extension is probably the expression of a lateral collapse in the hanging-wall of the Nanga Parbat crustal anticline. The last compression, related to $\mathrm{N}-\mathrm{S}$ convergence between India and Eurasia fits the present-day stress field defined from seismic activity.

## Appendix

## APPENDIX Chapter 3

- Results of electron microprobe analysis.


## Sample location:

## Sarangar Gabbro GbS (16/01/01) E 7301'23" N 35 ${ }^{\circ} 07^{\prime} 16^{\prime \prime}$

The sample was collected ca. 2 km ENE of Patan along the KKH . The mylonitic part covered by the sample belongs to a ca. 8 cm thick shear zone with a foliation dipping $20^{\circ}$ towards N and a top to S shear sense.

## Sarangar Gabbro GbS (90/02/01) E $73^{\circ} 01{ }^{\prime} 37^{\prime \prime} \mathrm{N} 35^{\circ} 07^{\prime} 01^{\prime \prime}$

The sample was collected ca. 2 km E of Patan along the Jeep-road Patan - Chor Nala close to Shalkanabad. The gently NE dipping shear zone is ca. 50 cm thick and contains stretched plagioclase veins and numerous rotated plagioclase porphyroclasts. Sense of shear is towards SW

## Granulitic Gabbro GtGrlt (13/01/01) E 725 ${ }^{\circ}$ '56" N 35º6'44"

The sample locality is ca. 200 m SW of the bridge crossing the Patan Banil Khwar at Patan at KKH. The shear zone is ca. $50^{\circ} \mathrm{W}$ dipping and yields a top to E sense of shear.

## Granulitic Gabbro GtGrlt (94/04/01) E 7259'57" N 35 ${ }^{\circ} 066^{\prime \prime} 43^{\prime \prime}$

Ca .40 m SSE of sample 13/01/01 in stronger retrogressed part of the granulitic Gabbro. Covered shear zone is ca. 20 cm thick. The shear zone is ca. $20^{\circ} \mathrm{NE}$ dipping and yields a top to SW shear sense.

## Kiru Amphibolite AmKi (46/02/01) E 7306'22" N 35 ${ }^{\circ} 11^{\prime} 06^{\prime \prime}$

The sample was collected ca. 150 mW of the metal bridge close to Kiru along the KKH within fine banded ( cm scale) amphibolites. Foliation is steeply $>70^{\circ} \mathrm{N}$ dipping and sense of shear is top to SSW. Protolith was a diorite.

Table Appendix 3.1: Representative chemical composition of plagioclase.

| Mineral: plagioclase |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Granulitic Gabbro undeformed 13/01/01 |  |  |  |  |  |  |  |
| $\begin{gathered} \text { p02s03pl3 } \\ \text { s03 } \end{gathered}$ | $\begin{gathered} \text { po2s03pl1 } \\ \text { s03 } \end{gathered}$ | $\begin{gathered} \text { p02s03pl1 } \\ \text { s03 } \end{gathered}$ | $\begin{gathered} \text { p02s03pl1 } \\ \text { s03 } \end{gathered}$ | $\begin{gathered} \text { p02s03p11 } \\ \text { s03 } \end{gathered}$ | $\begin{gathered} \text { p02s03p11 } \\ \text { s03 } \end{gathered}$ | $\begin{gathered} \text { p02s03pl2 } \\ \text { s03 } \end{gathered}$ | $\begin{gathered} \text { p02s03p11 } \\ \text { s03 } \end{gathered}$ |
| in contact to amp |  |  |  |  |  |  |  |
| 58.44 | 61.35 | 59.05 | 58.81 | 57.26 | 61.54 | 57.78 | 58.45 |
| 0.00 | 0.01 | 0.00 | 0.00 | 0.02 | 0.04 | 0.00 | 0.00 |
| 0.01 | 0.07 | 0.02 | 0.00 | 0.07 | 0.04 | 0.04 | 0.00 |
| 27.25 | 25.02 | 25.71 | 26.92 | 28.70 | 24.63 | 26.43 | 26.55 |
| 0.00 | 0.05 | 0.03 | 0.02 | 0.00 | 0.04 | 0.00 | 0.00 |
| 0.04 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.10 | 0.06 |
| 0.00 | 0.05 | 0.02 | 0.02 | 0.00 | 0.04 | 0.00 | 0.02 |
| 0.00 | 0.00 | 0.03 | 0.00 | 0.00 | 0.00 | 0.00 | 0.01 |
| 8.57 | 5.56 | 6.13 | 8.33 | 6.57 | 4.80 | 8.28 | 6.74 |
| 6.75 | 8.17 | 7.42 | 6.99 | 6.05 | 8.34 | 6.55 | 6.89 |
| 0.16 | 0.48 | 0.74 | 0.06 | 1.88 | 0.64 | 0.17 | 0.77 |
| 101.22 | 100.76 | 99.17 | 101.16 | 100.55 | 100.09 | 99.33 | 99.49 |
| 2.584 | 2.705 | 2.651 | 2.599 | 2.549 | 2.728 | 2.607 | 2.624 |
| 0.000 | 0.000 | 0.000 | 0.000 | 0.001 | 0.001 | 0.000 | 0.000 |
| 0.001 | 0.002 | 0.001 | 0.000 | 0.002 | 0.001 | 0.001 | 0.000 |
| 1.420 | 1.300 | 1.361 | 1.402 | 1.506 | 1.286 | 1.405 | 1.405 |
| 0.000 | 0.002 | 0.001 | 0.001 | 0.000 | 0.001 | 0.000 | 0.000 |
| 0.002 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.004 | 0.002 |
| 0.000 | 0.002 | 0.001 | 0.001 | 0.000 | 0.001 | 0.000 | 0.001 |
| 0.000 | 0.000 | 0.002 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 |
| 0.406 | 0.263 | 0.295 | 0.395 | 0.313 | 0.228 | 0.400 | 0.324 |
| 0.579 | 0.699 | 0.646 | 0.599 | 0.522 | 0.717 | 0.573 | 0.599 |
| 0.009 | 0.027 | 0.042 | 0.003 | 0.107 | 0.036 | 0.010 | 0.044 |
| 5.000 | 5.000 | 5.000 | 5.000 | 5.000 | 5.000 | 5.000 | 5.000 |
| 0.582 | 0.707 | 0.657 | 0.601 | 0.554 | 0.731 | 0.583 | 0.619 |
| 0.408 | 0.266 | 0.300 | 0.396 | 0.333 | 0.232 | 0.407 | 0.335 |
| 0.009 | 0.027 | 0.043 | 0.003 | 0.113 | 0.037 | 0.010 | 0.045 |
| calculated with | "Norm" pros | Pam by P | Imer using: F | PAR |  |  |  |


| Mineral: plagioclase |  |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Granulitic Gabbro mylonitic 13/01/01 |  |  |  |  |  |  |  |  |
| Sample | p01s05p11 | p01s05pl2 | p01s05pl3 | p01s05pl4 | p01s05p15 | p01s06p11 | p01s03in3 | p01s03gt6 |
| Site | s05 | s05 | s05 | s05 | s05 | s06 | s03 | s03 |
| nr. | 24 | 25 | 26 | 27 | 28 | 29 | 38 | 43 |
| note |  |  | in defor | ed part |  |  | \| incl. in gt | adj. to gt |
| $\mathrm{SiO}_{2}$ | 58.22 | 57.98 | 57.80 | 57.85 | 57.98 | 58.35 | 60.12 | 63.48 |
| $\mathrm{TiO}_{2}$ | 0.00 | 0.00 | 0.09 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 |
| $\mathrm{Cr}_{2} \mathrm{O}_{3}$ | 0.01 | 0.01 | 0.00 | 0.02 | 0.00 | 0.00 | 0.00 | 0.00 |
| $\mathrm{Al}_{2} \mathrm{O}_{3}$ | 26.89 | 27.02 | 27.04 | 27.28 | 27.05 | 27.14 | 26.20 | 24.13 |
| $\mathrm{Fe}_{2} \mathrm{O}_{3}$ | 0.00 | 0.00 | 0.03 | 0.00 | 0.04 | 0.00 | 0.00 | 0.32 |
| FeO | 0.06 | 0.04 | 0.00 | 0.02 | 0.00 | 0.00 | 0.16 | 0.10 |
| MnO | 0.00 | 0.02 | 0.00 | 0.05 | 0.00 | 0.03 | 0.05 | 0.03 |
| MgO | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 |
| NiO | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 |
| CaO | 8.35 | 8.74 | 8.68 | 8.57 | 8.46 | 8.31 | 7.37 | 4.86 |
| $\mathrm{Na}_{2} \mathrm{O}$ | 6.79 | 6.66 | 6.72 | 6.70 | 6.87 | 6.92 | 7.44 | 9.02 |
| $\mathrm{K}_{2} \mathrm{O}$ | 0.06 | 0.08 | 0.07 | 0.05 | 0.07 | 0.07 | 0.05 | 0.08 |
| Total | 100.37 | 100.55 | 100.43 | 100.53 | 100.46 | 100.82 | 101.41 | 102.01 |
| Si | 2.595 | 2.582 | 2.576 | 2.575 | 2.580 | 2.587 | 2.647 | 2.759 |
| Ti | 0.000 | 0.000 | 0.003 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 |
| Cr | 0.000 | 0.001 | 0.000 | 0.001 | 0.000 | 0.000 | 0.000 | 0.000 |
| Al | 1.413 | 1.418 | 1.420 | 1.431 | 1.419 | 1.418 | 1.360 | 1.236 |
| Fe3 | 0.000 | 0.000 | 0.001 | 0.000 | 0.001 | 0.000 | 0.000 | 0.011 |
| Fe2 | 0.002 | 0.002 | 0.000 | 0.001 | 0.000 | 0.000 | 0.006 | 0.004 |
| Mn | 0.000 | 0.001 | 0.000 | 0.002 | 0.000 | 0.001 | 0.002 | 0.001 |
| Mg | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 |
| Ni | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 |
| Ca | 0.399 | 0.417 | 0.414 | 0.409 | 0.403 | 0.395 | 0.348 | 0.226 |
| Na | 0.587 | 0.575 | 0.581 | 0.578 | 0.593 | 0.595 | 0.635 | 0.760 |
| K | 0.004 | 0.005 | 0.004 | 0.003 | 0.004 | 0.004 | 0.003 | 0.004 |
| $\Sigma$ Cations | 5.000 | 5.000 | 5.000 | 5.000 | 5.000 | 5.000 | 5.000 | 5.000 |
| Albite | 0.593 | 0.577 | 0.581 | 0.584 | 0.593 | 0.599 | 0.644 | 0.767 |
| Anorthite | 0.403 | 0.419 | 0.415 | 0.413 | 0.403 | 0.397 | 0.353 | 0.228 |
| Orthoclase | 0.004 | 0.005 | 0.004 | 0.003 | 0.004 | 0.004 | 0.003 | 0.004 |




| Mineral: plagioclase |  |  |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Sarangar Gabbro undeformed 90/02/01 |  |  |  |  |  |  |  |  |  |
| p10s01p07 | p10s01p08 | p10s01p14 | p10s01p16 | p10s02p06 | p10s03p14 | p10s03p15 | p10s03p24 | p10s03p27 | p10s03p28 |
| s01 | s01 | s01 | s01 | s02 | s03 | s03 | s03 | s03 | s03 |
| 7 | 8 | 14 | 2 | 22 | 61 | 62 | 71 | 74 | 75 |
| in contact to amp |  | in contact to amp \|incl. in cpx | in contact to amp rim |  |  |  |  | in contact to amp rim |  |  |
| 53.82 | 54.18 | 54.45 | 54.27 | 54.65 | 52.24 | 52.24 | 53.31 | 53.35 | 54.15 |
| 0.02 | 0.01 | 0.00 | 0.01 | 0.02 | 0.03 | 0.00 | 0.00 | 0.03 | 0.01 |
| 0.00 | 0.00 | 0.00 | 0.03 | 0.00 | 0.00 | 0.02 | 0.00 | 0.04 | 0.01 |
| 29.40 | 29.27 | 29.71 | 28.76 | 28.30 | 28.65 | 28.70 | 28.46 | 28.67 | 28.22 |
| 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.13 | 0.08 | 0.10 | 0.13 | 0.09 |
| 0.18 | 0.07 | 0.17 | 0.18 | 0.22 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 |
| 0.00 | 0.03 | 0.00 | 0.01 | 0.01 | 0.01 | 0.00 | 0.04 | 0.00 | 0.00 |
| 0.00 | 0.00 | 0.00 | 0.11 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 |
| 0.04 | 0.00 | 0.00 | 0.02 | 0.00 | 0.00 | 0.05 | 0.00 | 0.02 | 0.04 |
| 12.19 | 12.09 | 12.00 | 11.92 | 11.36 | 12.57 | 12.47 | 12.12 | 11.90 | 11.78 |
| 4.09 | 4.12 | 4.18 | 4.73 | 5.05 | 4.43 | 4.60 | 4.78 | 4.86 | 5.09 |
| 0.09 | 0.09 | 0.09 | 0.08 | 0.08 | 0.11 | 0.10 | 0.15 | 0.14 | 0.16 |
| 99.83 | 99.87 | 100.61 | 100.12 | 99.70 | 98.19 | 98.27 | 98.97 | 99.14 | 99.56 |
| 2.451 | 2.466 | 2.459 | 2.454 | 2.478 | 2.411 | 2.405 | 2.436 | 2.433 | 2.457 |
| 0.001 | 0.000 | 0.000 | 0.000 | 0.001 | 0.001 | 0.000 | 0.000 | 0.001 | 0.000 |
| 0.000 | 0.000 | 0.000 | 0.001 | 0.000 | 0.000 | 0.001 | 0.000 | 0.001 | 0.001 |
| 1.578 | 1.570 | 1.582 | 1.533 | 1.512 | 1.558 | 1.557 | 1.533 | 1.541 | 1.509 |
| 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.004 | 0.003 | 0.003 | 0.005 | 0.003 |
| 0.007 | 0.003 | 0.006 | 0.007 | 0.009 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 |
| 0.000 | 0.001 | 0.000 | 0.001 | 0.001 | 0.001 | 0.000 | 0.001 | 0.000 | 0.000 |
| 0.000 | 0.000 | 0.000 | 0.007 | 0.000 | 0.001 | 0.000 | 0.001 | 0.000 | 0.000 |
| 0.001 | 0.000 | 0.000 | 0.001 | 0.000 | 0.000 | 0.002 | 0.000 | 0.001 | 0.002 |
| 0.595 | 0.590 | 0.581 | 0.578 | 0.552 | 0.621 | 0.615 | 0.593 | 0.582 | 0.572 |
| 0.361 | 0.364 | 0.366 | 0.415 | 0.444 | 0.397 | 0.411 | 0.424 | 0.429 | 0.447 |
| 0.005 | 0.006 | 0.005 | 0.005 | 0.005 | 0.007 | 0.006 | 0.009 | 0.008 | 0.009 |
| 5.000 | 5.000 | 5.000 | 5.000 | 5.000 | 5.000 | 5.000 | 5.000 | 5.000 | 5.000 |
| 0.375 | 0.379 | 0.385 | 0.416 | 0.444 | 0.387 | 0.398 | 0.413 | 0.421 | 0.435 |
| 0.619 | 0.615 | 0.610 | 0.579 | 0.551 | 0.606 | 0.596 | 0.578 | 0.571 | 0.556 |
| 0.005 | 0.006 | 0.005 | 0.005 | 0.005 | 0.006 | 0.006 | 0.009 | 0.008 | 0.009 |
|  |  |  |  |  |  |  |  |  |  |


| Mineral: plagioclase |  |  |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Sarangar Gabbro mylonitic 90/02/01 |  |  |  |  |  |  |  |  |  |
| Sample | p07s05p17 | p07s05p25 | p07s05p29 | p07s06p17 | p07s06p18 | p07s06p19 | p07s06p20 | p07s06p23 | p07s06p25 |
| Site | s05 | s05 | s05 | s06 | s06 | s06 | s06 | s06 | s06 |
| nr. | 87 | 95 | 99 | 117 | 118 | 119 | 120 | 123 | 125 |
| note |  | at amp | at gt |  | at gt, am | mp, epi |  | at am | p, epi |
| $\mathrm{SiO}_{2}$ | 59.43 | 60.41 | 59.10 | 59.37 | 60.16 | 59.46 | 59.88 | 60.02 | 59.89 |
| $\mathrm{TiO}_{2}$ | 0.00 | 0.00 | 0.00 | 0.02 | 0.03 | 0.02 | 0.00 | 0.04 | 0.00 |
| $\mathrm{Cr}_{2} \mathrm{O}_{3}$ | 0.00 | 0.02 | 0.01 | 0.00 | 0.00 | 0.01 | 0.01 | 0.02 | 0.00 |
| $\mathrm{Al}_{2} \mathrm{O}_{3}$ | 25.74 | 25.63 | 25.27 | 25.41 | 25.63 | 25.79 | 25.74 | 25.61 | 25.78 |
| $\mathrm{Fe}_{2} \mathrm{O}_{3}$ | 0.17 | 0.00 | 0.12 | 0.39 | 0.11 | 0.07 | 0.10 | 0.13 | 0.10 |
| FeO | 0.00 | 0.16 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 |
| MnO | 0.05 | 0.00 | 0.00 | 0.02 | 0.04 | 0.05 | 0.02 | 0.03 | 0.00 |
| MgO | 0.00 | 0.02 | 0.00 | 0.00 | 0.00 | 0.00 | 0.03 | 0.00 | 0.00 |
| NiO | 0.08 | 0.00 | 0.03 | 0.06 | 0.00 | 0.00 | 0.00 | 0.08 | 0.01 |
| CaO | 7.42 | 6.78 | 7.19 | 7.01 | 7.19 | 7.48 | 7.25 | 7.60 | 7.72 |
| $\mathrm{Na}_{2} \mathrm{O}$ | 7.78 | 7.65 | 7.70 | 7.88 | 7.80 | 7.66 | 7.59 | 7.76 | 7.57 |
| $\mathrm{K}_{2} \mathrm{O}$ | 0.06 | 0.08 | 0.09 | 0.05 | 0.07 | 0.10 | 0.08 | 0.07 | 0.09 |
| Total | 100.75 | 100.75 | 99.52 | 100.20 | 101.03 | 100.65 | 100.72 | 101.36 | 101.16 |
| Si | 2.627 | 2.673 | 2.644 | 2.638 | 2.652 | 2.632 | 2.651 | 2.639 | 2.640 |
| Ti | 0.000 | 0.000 | 0.000 | 0.001 | 0.001 | 0.001 | 0.000 | 0.001 | 0.000 |
| Cr | 0.000 | 0.001 | 0.001 | 0.000 | 0.000 | 0.001 | 0.001 | 0.001 | 0.000 |
| Al | 1.341 | 1.337 | 1.333 | 1.331 | 1.332 | 1.345 | 1.343 | 1.327 | 1.339 |
| Fe3 | 0.006 | 0.000 | 0.004 | 0.013 | 0.004 | 0.002 | 0.004 | 0.004 | 0.004 |
| Fe 2 | 0.000 | 0.006 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 |
| Mn | 0.002 | 0.000 | 0.000 | 0.001 | 0.002 | 0.002 | 0.001 | 0.001 | 0.000 |
| Mg | 0.001 | 0.001 | 0.000 | 0.000 | 0.000 | 0.000 | 0.002 | 0.000 | 0.000 |
| Ni | 0.003 | 0.000 | 0.001 | 0.002 | 0.000 | 0.000 | 0.000 | 0.003 | 0.000 |
| Ca | 0.351 | 0.321 | 0.345 | 0.334 | 0.340 | 0.355 | 0.344 | 0.358 | 0.365 |
| Na | 0.667 | 0.656 | 0.668 | 0.679 | 0.667 | 0.657 | 0.651 | 0.662 | 0.647 |
| K | 0.004 | 0.005 | 0.005 | 0.003 | 0.004 | 0.006 | 0.005 | 0.004 | 0.005 |
| $\Sigma$ Cations | 5.000 | 5.000 | 5.000 | 5.000 | 5.000 | 5.000 | 5.000 | 5.000 | 5.000 |
| Albite | 0.653 | 0.668 | 0.656 | 0.669 | 0.660 | 0.646 | 0.651 | 0.647 | 0.636 |
| Anorthite | 0.344 | 0.327 | 0.339 | 0.329 | 0.336 | 0.348 | 0.344 | 0.350 | 0.359 |
| Orthoclase | 0.004 | 0.005 | 0.005 | 0.003 | 0.004 | 0.006 | 0.005 | 0.004 | 0.005 |


| Mineral: plagioclase |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: |
| Sarangar Gabbro undeformed 16/01/01 |  |  |  |  |  |
| p13s1cpl3 | p13s1cpl4 | p13s3apl1 | p13s3apl2 | p13s3bpl2 | p13s3bpl3 |
| s01 | s01 | s03 | s03 | s03 | s03 |
| around amp |  | in contact to amp |  | in contact to amp |  |
| 54.80 | 54.81 | 52.39 | 52.56 | 53.68 | 54.03 |
| 0.02 | 0.00 | 0.00 | 0.03 | 0.02 | 0.01 |
| 0.00 | 0.02 | 0.00 | 0.00 | 0.00 | 0.00 |
| 30.14 | 30.16 | 31.30 | 31.00 | 30.69 | 30.39 |
| 0.00 | 0.00 | 0.00 | 0.00 | 0.10 | 0.00 |
| 0.08 | 0.06 | 0.07 | 0.11 | 0.00 | 0.02 |
| 0.06 | 0.00 | 0.03 | 0.00 | 0.00 | 0.00 |
| 0.00 | 0.00 | 0.00 | 0.02 | 0.00 | 0.00 |
| 11.65 | 11.51 | 13.21 | 13.50 | 12.27 | 12.20 |
| 5.00 | 5.01 | 4.00 | 3.59 | 4.69 | 4.70 |
| 0.08 | 0.07 | 0.13 | 0.20 | 0.09 | 0.08 |
| 101.81 | 101.63 | 101.12 | 101.00 | 101.53 | 101.44 |
| 2.431 | 2.435 | 2.351 | 2.370 | 2.390 | 2.408 |
| 0.001 | 0.000 | 0.000 | 0.001 | 0.001 | 0.000 |
| 0.000 | 0.001 | 0.000 | 0.000 | 0.000 | 0.000 |
| 1.576 | 1.579 | 1.655 | 1.647 | 1.611 | 1.597 |
| 0.000 | 0.000 | 0.000 | 0.000 | 0.003 | 0.000 |
| 0.003 | 0.002 | 0.003 | 0.004 | 0.000 | 0.001 |
| 0.002 | 0.000 | 0.001 | 0.000 | 0.000 | 0.000 |
| 0.000 | 0.000 | 0.000 | 0.001 | 0.000 | 0.000 |
| 0.554 | 0.548 | 0.635 | 0.652 | 0.586 | 0.583 |
| 0.430 | 0.432 | 0.348 | 0.314 | 0.405 | 0.407 |
| 0.005 | 0.004 | 0.007 | 0.011 | 0.005 | 0.005 |
| 5.000 | 5.000 | 5.000 | 5.000 | 5.000 | 5.000 |
| 0.435 | 0.439 | 0.351 | 0.321 | 0.407 | 0.409 |
| 0.560 | 0.557 | 0.642 | 0.667 | 0.588 | 0.586 |
| 0.005 | 0.004 | 0.007 | 0.012 | 0.005 | 0.005 |


| Mineral: plagioclase |  |  |  |  |  |  |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Sarangar Gabbro mylonitic 16/01/01 |  |  |  |  |  |  |  |  |  |  |  |  |  |
| Sample | p11s08pl1 | p11s08pl2 | p11s08pl4 | p11s08p15 | p11s07p11 | p11s04p\|1 | p11s04p\|2 | p11s04pl4 | p11s04p15 | p11s04pl6 | p11s09pl1 | p11s09pl2 | p11109pl3 |
| Site | s08 | s08 | s08 | s08 | s07 | s04 | s04 | s04 | s04 | s04 | s09 | s09 | s09 |
| note | around mica |  |  |  | at clast | in contact to epi (plg layer) |  |  |  |  | in contact to amp |  |  |
| $\mathrm{SiO}_{2}$ | 61.23 | 60.86 | 58.37 | 57.26 | 60.35 | 58.07 | 58.06 | 58.66 | 60.02 | 59.80 | 58.94 | 58.05 | 58.27 |
| $\mathrm{TiO}_{2}$ | 0.00 | 0.01 | 0.02 | 0.04 | 0.02 | 0.02 | 0.01 | 0.00 | 0.00 | 0.03 | 0.00 | 0.00 | 0.00 |
| $\mathrm{Cr}_{2} \mathrm{O}_{3}$ | 0.00 | 0.00 | 0.01 | 0.01 | 0.03 | 0.02 | 0.00 | 0.00 | 0.00 | 0.00 | 0.02 | 0.00 | 0.00 |
| $\mathrm{Al}_{2} \mathrm{O}_{3}$ | 25.33 | 25.86 | 27.09 | 27.53 | 25.64 | 26.31 | 26.26 | 25.86 | 26.47 | 26.31 | 27.39 | 27.01 | 27.71 |
| $\mathrm{Fe}_{2} \mathrm{O}_{3}$ | 0.00 | 0.00 | 0.00 | 0.05 | 0.00 | 0.03 | 0.04 | 0.04 | 0.00 | 0.03 | 0.07 | 0.00 | 0.01 |
| FeO | 0.03 | 0.04 | 0.05 | 0.00 | 0.04 | 0.00 | 0.00 | 0.00 | 0.03 | 0.00 | 0.00 | 0.03 | 0.05 |
| MnO | 0.01 | 0.01 | 0.02 | 0.00 | 0.02 | 0.02 | 0.02 | 0.01 | 0.03 | 0.01 | 0.00 | 0.00 | 0.02 |
| MgO | 0.00 | 0.00 | 0.01 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 |
| NiO |  |  |  |  |  |  |  |  |  |  |  |  |  |
| CaO | 6.00 | 6.70 | 8.73 | 9.39 | 7.06 | 6.99 | 7.14 | 6.64 | 7.32 | 7.33 | 8.17 | 8.34 | 8.75 |
| $\mathrm{Na}_{2} \mathrm{O}$ | 8.09 | 7.82 | 6.58 | 6.39 | 7.61 | 7.73 | 7.62 | 7.94 | 7.43 | 7.54 | 7.14 | 6.77 | 6.68 |
| $\mathrm{K}_{2} \mathrm{O}$ | 0.05 | 0.05 | 0.08 | 0.10 | 0.07 | 0.06 | 0.07 | 0.06 | 0.09 | 0.09 | 0.05 | 0.07 | 0.09 |
| Total | 100.72 | 101.36 | 100.96 | 100.77 | 100.83 | 99.26 | 99.22 | 99.22 | 101.39 | 101.15 | 101.77 | 100.27 | 101.59 |
| Si | 2.703 | 2.674 | 2.592 | 2.548 | 2.669 | 2.599 | 2.602 | 2.624 | 2.641 | 2.635 | 2.586 | 2.590 | 2.568 |
| Ti | 0.000 | 0.000 | 0.001 | 0.001 | 0.001 | 0.001 | 0.000 | 0.000 | 0.000 | 0.001 | 0.000 | 0.000 | 0.000 |
| Cr | 0.000 | 0.000 | 0.000 | 0.000 | 0.001 | 0.001 | 0.000 | 0.000 | 0.000 | 0.000 | 0.001 | 0.000 | 0.000 |
| Al | 1.318 | 1.339 | 1.418 | 1.444 | 1.336 | 1.388 | 1.387 | 1.364 | 1.373 | 1.367 | 1.417 | 1.421 | 1.440 |
| Fe3 | 0.000 | 0.000 | 0.000 | 0.002 | 0.000 | 0.001 | 0.001 | 0.002 | 0.000 | 0.001 | 0.003 | 0.000 | 0.000 |
| Fe 2 | 0.001 | 0.001 | 0.002 | 0.000 | 0.002 | 0.000 | 0.000 | 0.000 | 0.001 | 0.000 | 0.000 | 0.001 | 0.002 |
| Mn | 0.000 | 0.001 | 0.001 | 0.000 | 0.001 | 0.001 | 0.001 | 0.000 | 0.001 | 0.000 | 0.000 | 0.000 | 0.001 |
| Mg | 0.000 | 0.000 | 0.001 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 |
| Ni |  |  |  |  |  |  |  |  |  |  |  |  |  |
| Ca | 0.284 | 0.316 | 0.416 | 0.448 | 0.335 | 0.335 | 0.343 | 0.318 | 0.345 | 0.346 | 0.384 | 0.399 | 0.413 |
| Na | 0.692 | 0.667 | 0.566 | 0.552 | 0.652 | 0.671 | 0.662 | 0.689 | 0.634 | 0.645 | 0.607 | 0.586 | 0.571 |
| K | 0.003 | 0.003 | 0.004 | 0.006 | 0.004 | 0.004 | 0.004 | 0.003 | 0.005 | 0.005 | 0.003 | 0.004 | 0.005 |
| $\Sigma$ Cations | 5.000 | 5.000 | 5.000 | 5.000 | 5.000 | 5.000 | 5.000 | 5.000 | 5.000 | 5.000 | 5.000 | 5.000 | 5.000 |
| Albite | 0.707 | 0.677 | 0.574 | 0.549 | 0.658 | 0.664 | 0.656 | 0.682 | 0.644 | 0.647 | 0.611 | 0.592 | 0.577 |
| Anorthite | 0.290 | 0.320 | 0.421 | 0.446 | 0.338 | 0.332 | 0.340 | 0.315 | 0.351 | 0.348 | 0.386 | 0.404 | 0.418 |
| Orthoclase | 0.003 | 0.003 | 0.004 | 0.005 | 0.004 | 0.003 | 0.004 | 0.003 | 0.005 | 0.005 | 0.003 | 0.004 | 0.005 |



Table Appendix 3.2: Representative chemical composition of garnet.

| Mineral: garnet |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Granulitic Gabbro mylonitic 13/01/01 |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| Sample Site | $\begin{gathered} \text { p01s01gt1 } \\ \text { s01 } \end{gathered}$ | $\begin{gathered} \text { p01s01gt2 } \\ \text { s01 } \end{gathered}$ | $\begin{gathered} \text { p01s01gt3 } \\ \text { s01 } \end{gathered}$ | $\begin{gathered} \text { p01s01gt4 } \\ \text { s01 } \end{gathered}$ | $\begin{gathered} \text { p01s01gt5 } \\ \text { s01 } \end{gathered}$ | $\begin{gathered} \text { p01s01 gt6 } \\ \text { s01 } \end{gathered}$ | $\begin{gathered} \text { p01s01gt7 } \\ \text { s01 } \end{gathered}$ | $\begin{gathered} \text { p01so2gt2 } \\ \text { s02 } \end{gathered}$ | $\begin{gathered} \text { p01so2gt1 } \\ \text { s02 } \end{gathered}$ | $\begin{gathered} \text { p01 s03gt4 } \\ \text { s03 } \end{gathered}$ | $\begin{gathered} \text { p01s03gt6 } \\ \text { s03 } \end{gathered}$ | $\begin{gathered} \text { p01s03gt5 } \\ \text { s03 } \end{gathered}$ | $\begin{gathered} \text { p01s03gt3 } \\ \text { s03 } \end{gathered}$ | $\begin{gathered} \text { p01so4gt5 } \\ \text { s04 } \end{gathered}$ | $\begin{gathered} \text { p01so4gt3 } \\ \text { s04 } \end{gathered}$ | $\begin{gathered} \text { p01so4gt2 } \\ \text { s04 } \end{gathered}$ | $\begin{gathered} \text { p01s04gt1 } \\ \text { s04 } \end{gathered}$ | $\begin{gathered} \text { p01s09gt3 } \\ \text { s09 } \end{gathered}$ | $\begin{gathered} \text { p01s09gt2 } \\ \text { s09 } \end{gathered}$ |
| note | profile across porphyroclast |  |  |  |  |  |  | gt with inclusion S02 |  | gt porpyroclast with inclusion |  |  |  | big gt porphyroclast in less def. part |  |  |  | gt in contact to amp |  |
| $\mathrm{SiO}_{2}$ | 39.78 | 39.88 | 39.71 | 39.86 | 39.80 | 39.81 | 40.00 | 39.81 | 39.66 | 39.98 | 39.64 | 39.97 | 40.00 | 39.66 | 39.59 | 39.76 | 39.81 | 39.40 | 39.19 |
| $\mathrm{TiO}_{2}$ | 0.10 | 0.10 | 0.11 | 0.09 | 0.12 | 0.09 | 0.12 | 0.15 | 0.14 | 0.11 | 0.10 | 0.09 | 0.10 | 0.06 | 0.05 | 0.05 | 0.05 | 0.15 | 0.26 |
| $\mathrm{Cr}_{2} \mathrm{O}_{3}$ | 0.00 | 0.06 | 0.07 | 0.09 | 0.09 | 0.06 | 0.06 | 0.04 | 0.13 | 0.01 | 0.00 | 0.00 | 0.01 | 0.05 | 0.04 | 0.01 | 0.00 | 0.00 | 0.00 |
| $\mathrm{Al}_{2} \mathrm{O}_{3}$ | 22.07 | 22.00 | 21.99 | 22.03 | 22.00 | 21.97 | 22.01 | 22.11 | 21.87 | 22.34 | 22.17 | 22.53 | 22.27 | 22.45 | 22.49 | 22.46 | 22.51 | 21.86 | 21.95 |
| $\mathrm{Fe}_{2} \mathrm{O}_{3}$ | 1.10 | 0.91 | 0.76 | 1.35 | 1.64 | 1.17 | 0.10 | 0.77 | 1.24 | 0.14 | 0.48 | 0.64 | 0.66 | 0.57 | 1.12 | 0.01 | 0.69 | 1.13 | 0.93 |
| FeO | 20.71 | 20.75 | 20.79 | 20.61 | 20.49 | 20.28 | 21.06 | 20.74 | 20.31 | 21.79 | 22.76 | 21.31 | 21.38 | 20.20 | 19.50 | 20.57 | 20.28 | 21.39 | 21.37 |
| MnO | 0.57 | 0.61 | 0.60 | 0.62 | 0.63 | 0.60 | 0.57 | 0.53 | 0.50 | 0.47 | 0.71 | 0.54 | 0.54 | 0.51 | 0.52 | 0.56 | 0.56 | 0.48 | 0.44 |
| MgO | 9.55 | 9.58 | 9.56 | 9.60 | 9.80 | 9.73 | 9.36 | 9.76 | 9.65 | 9.69 | 8.27 | 9.58 | 9.62 | 10.02 | 9.69 | 9.85 | 9.60 | 9.17 | 9.20 |
| NiO |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| CaO | 7.17 | 7.10 | 7.02 | 7.17 | 6.93 | 7.25 | 7.36 | 6.94 | 7.31 | 6.39 | 7.14 | 6.92 | 6.79 | 6.81 | 7.62 | 6.78 | 7.43 | 7.03 | 6.84 |
| $\mathrm{Na}_{2} \mathrm{O}$ | 0.02 | 0.03 | 0.04 | 0.03 | 0.04 | 0.01 | 0.03 | 0.04 | 0.03 | 0.03 | 0.03 | 0.01 | 0.03 | 0.03 | 0.05 | 0.05 | 0.02 | 0.01 | 0.02 |
| $\mathrm{K}_{2} \mathrm{O}$ | 0.03 | 0.03 | 0.00 | 0.02 | 0.02 | 0.04 | 0.01 | 0.00 | 0.01 | 0.02 | 0.00 | 0.00 | 0.01 | 0.01 | 0.03 | 0.00 | 0.03 | 0.00 | 0.00 |
| Total | 101.10 | 101.06 | 100.64 | 101.46 | 101.55 | 101.01 | 100.68 | 100.88 | 100.84 | 100.98 | 101.30 | 101.58 | 101.42 | 100.37 | 100.70 | 100.09 | 100.98 | 100.60 | 100.20 |
| Si | 2.989 | 2.996 | 2.995 | 2.985 | 2.978 | 2.990 | 3.014 | 2.992 | 2.986 | 3.003 | 2.995 | 2.986 | 2.995 | 2.986 | 2.974 | 3.001 | 2.986 | 2.984 | 2.978 |
| Ti | 0.006 | 0.006 | 0.006 | 0.005 | 0.007 | 0.005 | 0.007 | 0.008 | 0.008 | 0.006 | 0.006 | 0.005 | 0.005 | 0.003 | 0.003 | 0.003 | 0.003 | 0.008 | 0.015 |
| Cr | 0.000 | 0.004 | 0.004 | 0.005 | 0.005 | 0.004 | 0.004 | 0.002 | 0.008 | 0.001 | 0.000 | 0.000 | 0.001 | 0.003 | 0.003 | 0.001 | 0.000 | 0.000 | 0.000 |
| Al | 1.954 | 1.948 | 1.955 | 1.945 | 1.940 | 1.945 | 1.955 | 1.959 | 1.940 | 1.978 | 1.975 | 1.984 | 1.966 | 1.992 | 1.991 | 1.998 | 1.989 | 1.951 | 1.966 |
| $\mathrm{Fe}^{3}$ | 0.062 | 0.052 | 0.043 | 0.076 | 0.092 | 0.066 | 0.006 | 0.043 | 0.070 | 0.008 | 0.027 | 0.036 | 0.038 | 0.032 | 0.063 | 0.000 | 0.039 | 0.064 | 0.053 |
| $\mathrm{Fe}^{2}$ | 1.301 | 1.304 | 1.312 | 1.291 | 1.282 | 1.274 | 1.327 | 1.304 | 1.279 | 1.369 | 1.438 | 1.332 | 1.339 | 1.272 | 1.225 | 1.298 | 1.272 | 1.355 | 1.358 |
| Mn | 0.036 | 0.039 | 0.038 | 0.039 | 0.040 | 0.038 | 0.037 | 0.034 | 0.032 | 0.030 | 0.046 | 0.034 | 0.034 | 0.033 | 0.033 | 0.036 | 0.036 | 0.031 | 0.028 |
| Mg | 1.069 | 1.073 | 1.075 | 1.071 | 1.093 | 1.089 | 1.052 | 1.093 | 1.083 | 1.085 | 0.931 | 1.067 | 1.073 | 1.125 | 1.085 | 1.108 | 1.073 | 1.035 | 1.042 |
| Ni |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| Ca | 0.577 | 0.572 | 0.567 | 0.576 | 0.555 | 0.584 | 0.594 | 0.559 | 0.590 | 0.515 | 0.578 | 0.554 | 0.545 | 0.550 | 0.614 | 0.548 | 0.597 | 0.571 | 0.557 |
| Na | 0.003 | 0.005 | 0.005 | 0.005 | 0.005 | 0.002 | 0.004 | 0.006 | 0.004 | 0.004 | 0.004 | 0.002 | 0.004 | 0.004 | 0.008 | 0.007 | 0.003 | 0.001 | 0.003 |
| K | 0.002 | 0.003 | 0.000 | 0.002 | 0.002 | 0.004 | 0.001 | 0.000 | 0.001 | 0.002 | 0.000 | 0.000 | 0.001 | 0.001 | 0.002 | 0.000 | 0.003 | 0.000 | 0.001 |
| $\Sigma$ Cations | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 |
| Grossular | 0.158 | 0.159 | 0.162 | 0.149 | 0.133 | 0.157 | 0.188 | 0.159 | 0.154 | 0.163 | 0.176 | 0.165 | 0.159 | 0.164 | 0.172 | 0.181 | 0.179 | 0.154 | 0.152 |
| Almandine | 0.435 | 0.436 | 0.438 | 0.433 | 0.431 | 0.426 | 0.440 | 0.435 | 0.428 | 0.456 | 0.480 | 0.445 | 0.447 | 0.427 | 0.414 | 0.434 | 0.427 | 0.453 | 0.454 |
| Pyrope | 0.358 | 0.358 | 0.359 | 0.359 | 0.367 | 0.364 | 0.349 | 0.365 | 0.362 | 0.361 | 0.311 | 0.357 | 0.358 | 0.377 | 0.367 | 0.370 | 0.360 | 0.346 | 0.349 |
| Spessartite | 0.012 | 0.013 | 0.013 | 0.013 | 0.013 | 0.013 | 0.012 | 0.011 | 0.011 | 0.010 | 0.015 | 0.012 | 0.011 | 0.011 | 0.011 | 0.012 | 0.012 | 0.010 | 0.009 |
| Andradite | 0.031 | 0.026 | 0.021 | 0.038 | 0.046 | 0.033 | 0.003 | 0.022 | 0.035 | 0.004 | 0.014 | 0.018 | 0.019 | 0.016 | 0.032 | 0.000 | 0.020 | 0.032 | 0.027 |






|  |  | gt porphyroclast |  |  |  |
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| $\Sigma$ Cations | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 |
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| Grossular | 0.228 | 0.237 | 0.223 | 0.199 | 0.222 | 0.223 | 0.241 | 0.177 | 0.177 | 0.186 | 0.184 | 0.178 | 0.185 | 0.154 | 0.175 | 0.172 | 0.186 | 0.215 | 0.237 |
| Almandine | 0.454 | 0.468 | 0.466 | 0.466 | 0.469 | 0.458 | 0.462 | 0.467 | 0.433 | 0.432 | 0.461 | 0.438 | 0.436 | 0.439 | 0.45 | 0.42 | 0.4 | 0.437 | 0.435 |
| Pyrope | 0.284 | 0.280 | 0.295 | 0.299 | 0.296 | 0.286 | 0.275 | 0.280 | 0.324 | 0.328 | 0.291 | 0.329 | 0.330 | 0.356 | 0.31 | 0.34 | 0.326 | 0.31 | 0.316 |
| Spessartite | 0.009 | 0.010 | 0.010 | 0.011 | 0.009 | 0.010 | 0.010 | 0.015 | 0.012 | 0.012 | 0.016 | 0.012 | 0.012 | 0.012 | 0.014 | 0.011 | 0.011 | 0.010 | 0.010 |
| Andradite | 0.022 | 0.000 | 0.002 | 0.019 | 0.000 | 0.017 | 0.010 | 0.055 | 0.051 | 0.038 | 0.043 | 0.041 | 0.035 | 0.038 | 0.041 | 0.050 | 0.045 | 0.020 | 0.000 |



Grossular
Almandine
Pyrope
Spessartite
Andradite

|  | Mineral: ga | rnet |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
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|  | Granulitic G | abbro unde | formed 94 | /04/01 |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| Sample | p06s01p06 | p06s01p07 | p06s01p08 | p06s01p16 | p06s01p17 | p06s01p18 | p06s01p25 | p06s02p10 | p06s02p11 | p06s02p12 | p06s02p13 | p06s02p14 | p06s02p15 | p06s02p16 | p06s02p17 | p06s02p28 | p06s02p29 | p06s02p30 | p06s02p31 |
| Site | s01 | s01 | s01 | s01 | s01 | s01 | s01 | s02 | s02 | s02 | s02 | s02 | s02 | s02 | s02 | s02 | s02 | s02 | s02 |
| nr. | 6 | 7 | 8 | 16 | 17 | 18 | 25 | 35 | 1 | 2 | 3 | 4 | 5 | 6 | 7 | 19 | 20 | 21 | 22 |
| note | gt in | contact to |  | gt in | contact to |  | \| gt to cpx | |  |  |  | profile th | rough gt |  |  |  |  | profile th | rough gt |  |
| $\mathrm{SiO}_{2}$ | 39.63 | 39.38 | 39.59 | 39.30 | 39.52 | 39.32 | 38.98 | 38.58 | 38.17 | 37.61 | 38.16 | 38.54 | 39.47 | 39.53 | 38.73 | 39.60 | 39.44 | 39.65 | 39.63 |
| $\mathrm{TiO}_{2}$ | 0.07 | 0.05 | 0.06 | 0.06 | 0.07 | 0.07 | 0.08 | 0.06 | 0.07 | 0.07 | 0.04 | 0.04 | 0.05 | 0.09 | 0.05 | 0.05 | 0.08 | 0.09 | 0.06 |
| $\mathrm{Cr}_{2} \mathrm{O}_{3}$ | 0.00 | 0.00 | 0.01 | 0.00 | 0.00 | 0.02 | 0.05 | 0.01 | 0.00 | 0.00 | 0.01 | 0.00 | 0.01 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 |
| $\mathrm{Al}_{2} \mathrm{O}_{3}$ | 21.81 | 21.74 | 21.82 | 21.94 | 22.18 | 21.81 | 21.74 | 22.20 | 21.59 | 21.27 | 22.16 | 21.95 | 21.60 | 21.58 | 21.78 | 21.85 | 21.33 | 21.73 | 22.06 |
| $\mathrm{Fe}_{2} \mathrm{O}_{3}$ | 0.00 | 0.62 | 0.95 | 0.56 | 1.16 | 0.19 | 1.20 | 0.94 | 2.18 | 2.70 | 2.66 | 2.15 | 0.23 | 1.18 | 1.52 | 0.36 | 0.00 | 0.47 | 0.75 |
| FeO | 22.14 | 21.95 | 21.63 | 23.04 | 20.83 | 21.30 | 20.48 | 20.20 | 19.70 | 21.09 | 19.03 | 19.51 | 20.73 | 20.97 | 20.69 | 21.63 | 22.40 | 21.43 | 20.81 |
| MnO | 0.43 | 0.47 | 0.40 | 0.62 | 0.39 | 0.36 | 0.38 | 0.47 | 0.46 | 0.59 | 0.50 | 0.50 | 0.46 | 0.49 | 0.52 | 0.47 | 0.45 | 0.50 | 0.46 |
| MgO | 7.95 | 7.92 | 7.86 | 5.97 | 7.66 | 7.70 | 8.15 | 7.75 | 7.81 | 5.65 | 7.70 | 7.70 | 7.14 | 7.72 | 7.85 | 8.10 | 6.11 | 8.05 | 7.75 |
| NiO | 0.00 | 0.00 | 0.00 | 0.06 | 0.00 | 0.00 | 0.00 | 0.03 | 0.03 | 0.03 | 0.00 | 0.00 | 0.00 | 0.08 | 0.00 | 0.00 | 0.00 | 0.05 | 0.00 |
| CaO | 8.10 | 8.27 | 8.77 | 9.87 | 9.59 | 9.10 | 8.71 | 8.99 | 9.01 | 10.32 | 9.64 | 9.68 | 10.37 | 9.15 | 8.48 | 8.42 | 9.70 | 8.70 | 9.63 |
| $\mathrm{Na}_{2} \mathrm{O}$ | 0.05 | 0.00 | 0.02 | 0.00 | 0.02 | 0.00 | 0.02 | 0.02 | 0.01 | 0.00 | 0.00 | 0.00 | 0.00 | 0.06 | 0.05 | 0.02 | 0.04 | 0.00 | 0.00 |
| $\mathrm{K}_{2} \mathrm{O}$ | 0.00 |  | 0.00 | 0.00 | 0.01 | 0.01 | 0.02 | 0.02 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.01 | 0.02 | 0.00 | 0.02 | 0.00 | 0.00 |
| Total | 100.17 | 100.40 | 101.12 | 101.44 | 101.44 | 99.89 | 99.79 | 99.28 | 99.04 | 99.33 | 99.92 | 100.06 | 100.07 | 100.85 | 99.70 | 100.51 | 99.56 | 100.68 | 101.17 |
| Si | 3.023 | 3.003 | 2.997 | 2.996 | 2.980 | 3.007 | 2.982 | 2.965 | 2.950 | 2.938 | 2.922 | 2.947 | 3.018 | 3.002 | 2.973 | 3.009 | 3.053 | 3.010 | 2.993 |
| Ti | 0.004 | 0.003 | 0.003 | 0.003 | 0.004 | 0.004 | 0.005 | 0.004 | 0.004 | 0.004 | 0.002 | 0.002 | 0.003 | 0.005 | 0.003 | 0.003 | 0.005 | 0.005 | 0.004 |
| Cr | 0.000 | 0.000 | 0.001 | 0.000 | 0.000 | 0.001 | 0.003 | 0.001 | 0.000 | 0.000 | 0.001 | 0.000 | 0.001 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 |
| Al | 1.961 | 1.954 | 1.947 | 1.971 | 1.971 | 1.966 | 1.960 | 2.011 | 1.967 | 1.958 | 2.000 | 1.978 | 1.947 | 1.931 | 1.970 | 1.957 | 1.946 | 1.944 | 1.964 |
| $\mathrm{Fe}^{3}$ | 0.000 | 0.035 | 0.054 | 0.032 | 0.066 | 0.011 | 0.069 | 0.055 | 0.127 | 0.159 | 0.153 | 0.124 | 0.013 | 0.067 | 0.088 | 0.020 | 0.000 | 0.027 | 0.043 |
| $\mathrm{Fe}^{2}$ | 1.412 | 1.400 | 1.370 | 1.469 | 1.313 | 1.363 | 1.310 | 1.299 | 1.273 | 1.378 | 1.219 | 1.247 | 1.325 | 1.331 | 1.328 | 1.375 | 1.450 | 1.361 | 1.315 |
| Mn | 0.028 | 0.030 | 0.026 | 0.040 | 0.025 | 0.024 | 0.024 | 0.031 | 0.030 | 0.039 | 0.032 | 0.032 | 0.030 | 0.032 | 0.034 | 0.030 | 0.030 | 0.032 | 0.029 |
| Mg | 0.904 | 0.900 | 0.887 | 0.679 | 0.861 | 0.878 | 0.929 | 0.888 | 0.899 | 0.658 | 0.878 | 0.877 | 0.814 | 0.873 | 0.898 | 0.917 | 0.704 | 0.910 | 0.873 |
| Ni | 0.000 | 0.000 | 0.000 | 0.004 | 0.000 | 0.000 | 0.000 | 0.002 | 0.002 | 0.002 | 0.000 | 0.000 | 0.000 | 0.005 | 0.001 | 0.000 | 0.000 | 0.003 | 0.000 |
| Ca | 0.662 | 0.675 | 0.712 | 0.806 | 0.775 | 0.746 | 0.714 | 0.741 | 0.746 | 0.864 | 0.791 | 0.793 | 0.849 | 0.744 | 0.697 | 0.685 | 0.804 | 0.708 | 0.779 |
| Na | 0.008 | 0.000 | 0.003 | 0.000 | 0.003 | 0.000 | 0.003 | 0.002 | 0.002 | 0.000 | 0.001 | 0.000 | 0.000 | 0.010 | 0.007 | 0.003 | 0.005 | 0.000 | 0.000 |
| K | 0.000 | 0.000 | 0.000 | 0.001 | 0.001 | 0.001 | 0.002 | 0.002 | 0.000 | 0.001 | 0.001 | 0.000 | 0.001 | 0.001 | 0.002 | 0.000 | 0.002 | 0.000 | 0.001 |
| $\Sigma$ Cations | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 |
| Grossular | 0.217 | 0.205 | 0.208 | 0.251 | 0.224 | 0.239 | 0.200 | 0.220 | 0.186 | 0.211 | 0.190 | 0.205 | 0.273 | 0.212 | 0.189 | 0.215 | 0.266 | 0.219 | 0.237 |
| Almandine | 0.469 | 0.466 | 0.457 | 0.491 | 0.441 | 0.453 | 0.439 | 0.438 | 0.432 | 0.469 | 0.417 | 0.423 | 0.439 | 0.446 | 0.449 | 0.457 | 0.484 | 0.452 | 0.439 |
| Pyrope | 0.300 | 0.299 | 0.296 | 0.227 | . 289 | 0.292 | 0.312 | 0.300 | 0.305 | 0.224 | 0.301 | 0.297 | 0.270 | 0.293 | 0.303 | 0.30 | 0.23 | 0.302 | 0.291 |
| Spessartite | 0.009 | 0.010 | 0.009 | 0.013 | . 008 | 0.008 | 0.008 | . 010 | 0.010 | 0.013 | . 011 | 0.011 | 0.010 | 0.011 | . 011 | 0.010 | 0.010 | 0.011 | . 010 |
| Andradite | 0.000 | 0.018 | 0.027 | 0.016 | 0.033 | 0.005 | 0.035 | 0.028 | 0.064 | 0.081 | 0.079 | 0.063 | 0.006 | 0.034 | 0.044 | 0.010 | 0.000 | 0.013 | 0.021 |



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| 00.52 | 100.63 | 99.86 | 100.17 | 101.00 | 100.52 | 99.28 | 99.33 | 99.71 | 100.02 | 99.55 |
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| Si | 2.992 | 2.994 | 3.004 | 2.960 | 2.971 | 2.963 | 2.981 | 2.935 | 2.973 | 2.963 | 2.952 | 2.943 | 2.987 | 2.965 | 2.995 | 2.975 | 2.982 |
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| Ti | 0.006 | 0.001 | 0.003 | 0.002 | 0.005 | 0.002 | 0.002 | 0.004 | 0.000 | 0.002 | 0.003 | 0.003 | 0.006 | 0.009 | 0.003 | 0.004 | 0.003 |
| Cr | 0.005 | 0.001 | 0.002 | 0.001 | 0.001 | 0.000 | 0.001 | 0.001 | 0.001 | 0.003 | 0.002 | 0.001 | 0.001 | 0.001 | 0.000 | 0.000 | 0.002 |
| AI | 1.984 | 2.014 | 2.012 | 2.010 | 1.990 | 2.006 | 1.989 | 1.999 | 1.999 | 1.994 | 1.983 | 1.992 | 2.001 | 2.006 | 2.005 | 1.999 | 2.010 |
| $\mathrm{Fe}^{3}$ | 0.025 | 0.002 | 0.000 | 0.065 | 0.069 | 0.076 | 0.052 | 0.124 | 0.062 | 0.075 | 0.113 | 0.121 | 0.019 | 0.052 | 0.009 | 0.043 | 0.021 |
| $\mathrm{Fe}^{2}$ | 1.699 | 1.711 | 1.730 | 1.664 | 1.666 | 1.631 | 1.690 | 1.628 | 1.665 | 1.647 | 1.633 | 1.643 | 1.698 | 1.677 | 1.703 | 1.692 | 1.720 |
| Mn | 0.154 | 0.162 | 0.139 | 0.152 | 0.119 | 0.125 | 0.117 | 0.156 | 0.150 | 0.160 | 0.158 | 0.153 | 0.151 | 0.141 | 0.147 | 0.148 | 0.144 |
| Mg | 0.490 | 0.481 | 0.460 | 0.490 | 0.435 | 0.463 | 0.486 | 0.475 | 0.492 | 0.492 | 0.492 | 0.466 | 0.476 | 0.497 | 0.469 | 0.483 | 0.466 |
| Ni | 0.000 | 0.004 | 0.002 | 0.000 | 0.002 | 0.000 | 0.001 | 0.003 | 0.000 | 0.000 | 0.000 | 0.000 | 0.001 | 0.000 | 0.003 | 0.000 | 0.003 |
| Ca | 0.639 | 0.625 | 0.646 | 0.655 | 0.732 | 0.724 | 0.671 | 0.673 | 0.649 | 0.664 | 0.659 | 0.674 | 0.653 | 0.643 | 0.656 | 0.655 | 0.645 |
| Na | 0.008 | 0.005 | 0.001 | 0.000 | 0.011 | 0.010 | 0.009 | 0.002 | 0.008 | 0.001 | 0.006 | 0.004 | 0.007 | 0.008 | 0.010 | 0.000 | 0.004 |
| k | 0.001 | 0.000 | 0.001 | 0.001 | 0.001 | 0.001 | 0.000 | 0.000 | 0.000 | 0.001 | 0.000 | 0.001 | 0.001 | 0.000 | 0.000 | 0.001 | 0.000 |
| $\Sigma$ Cations | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 |
| Grossular | 0.196 | 0.208 | 0.214 | 0.186 | 0.209 | 0.206 | 0.198 | 0.163 | 0.188 | 0.183 | 0.164 | 0.165 | 0.204 | 0.184 | 0.213 | 0.196 | 0.203 |
| Almandine | 0.569 | 0.574 | 0.581 | 0.562 | 0.564 | 0.554 | 0.570 | 0.555 | 0.563 | 0.556 | 0.555 | 0.559 | 0.569 | 0.565 | 0.572 | 0.568 | 0.578 |
| Pyrope | 0.164 | 0.161 | 0.155 | 0.165 | 0.147 | 0.157 | 0.164 | 0.162 | 0.166 | 0.166 | 0.167 | 0.159 | 0.159 | 0.168 | 0.157 | 0.162 | 0.157 |
| Spessartite | 0.051 | 0.054 | 0.047 | 0.051 | 0.040 | 0.042 | 0.039 | 0.053 | 0.051 | 0.054 | 0.054 | 0.052 | 0.051 | 0.047 | 0.049 | 0.050 | 0.048 |
| Andradite | 0.012 | 0.000 | 0.000 | 0.033 | 0.035 | 0.039 | 0.026 | 0.063 | 0.031 | 0.038 | 0.057 | 0.062 | 0.010 | 0.026 | 0.004 | 0.022 | 0.011 | calculated with＂Norm＂program by P．Ulmer using：GARNET N

Normalization on the basis of 7 CATIONS and 24 CHARGES

| Mineral: garnet |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
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| Sarangar Gabbro mylonitic 16/01/01 |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| Sample <br> Site nr. <br> note | $\begin{gathered} \text { p11s01gt1 } \\ \text { s01 } \end{gathered}$ | $\begin{gathered} \text { p11s01gt2 } \\ \text { s01 } \end{gathered}$ <br> profile | p11 $\mathrm{c} 019 \mathrm{gt3}$ s 01 through ha |  | $\begin{gathered} \text { p11s01 gtt6 } \\ \text { s01 } \end{gathered}$ | $\begin{gathered} \text { p11s01in1 } \\ \text { s01 } \end{gathered}$ | $\begin{gathered} \text { p11s01pr1 } \\ \text { s01 } \end{gathered}$ | $\begin{gathered} \text { p11s01pr1 } \\ \text { s01 } \end{gathered}$ | $\begin{gathered} \text { p11s01pr1 } \\ \text { s01 } \end{gathered}$ | $\begin{gathered} \text { p11s01pr1 } \\ \text { s01 } \end{gathered}$ | $\begin{gathered} \text { p11s01pr2 } \\ \text { s01 } \end{gathered}$ | $\begin{gathered} \text { p11s01pr2 } \\ \text { s01 } \end{gathered}$ | $\begin{gathered} \text { p11s01pr2 } \\ \text { s01 } \end{gathered}$ | $\begin{gathered} \text { p11s01pr2 } \\ \text { s01 } \end{gathered}$ | $\begin{gathered} \text { p11s01pr2 } \\ \text { s01 } \end{gathered}$ | $\begin{gathered} \text { p11s01pr2 } \\ \text { s01 } \end{gathered}$ | $\begin{gathered} \text { p11s01pr2 } \\ \text { s01 } \end{gathered}$ | $\begin{gathered} \text { p11s03in3 } \\ \text { s03 } \end{gathered}$ | $\begin{aligned} & \text { p11s03in2 } \\ & \text { s03 } \end{aligned}$ |
| $\mathrm{SiO}_{2}$ | 37.25 | 38.58 | 38.63 | 38.75 | 38.79 | 38.34 | 38.73 | 38.67 | 38.04 | 38.80 | 39.10 | 39.21 | 38.68 | 38.80 | 38.72 | 38.68 | 38.73 | 38.95 | 38.83 |
| $\mathrm{TiO}_{2}$ | 0.04 | 0.04 | 0.12 | 0.12 | 0.04 | 0.08 | 0.04 | 0.08 | 0.05 | 0.09 | 0.07 | 0.07 | 0.03 | 0.04 | 0.08 | 0.00 | 0.08 | 0.09 | 0.02 |
| $\mathrm{Cr}_{2} \mathrm{O}_{3}$ | 0.00 | 0.03 | 0.02 | 0.04 | 0.00 | 0.03 | 0.01 | 0.00 | 0.02 | 0.05 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 |
| $\mathrm{Al}_{2} \mathrm{O}_{3}$ | 21.30 | 21.85 | 21.60 | 21.71 | 21.87 | 21.80 | 21.91 | 21.76 | 21.62 | 21.74 | 21.82 | 21.83 | 21.48 | 21.89 | 21.86 | 21.76 | 21.68 | 21.71 | 21.81 |
| $\mathrm{Fe}_{2} \mathrm{O}_{3}$ | 2.78 | 0.22 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.85 | 0.00 | 0.00 | 0.00 | 0.00 | 0.40 | 0.01 | 0.00 | 0.00 | 0.00 | 0.00 |
| FeO | 24.96 | 25.63 | 25.02 | 25.66 | 25.67 | 25.47 | 26.46 | 25.61 | 24.95 | 25.34 | 25.93 | 25.91 | 25.57 | 25.83 | 25.84 | 26.04 | 25.45 | 25.35 | 25.99 |
| MnO | 3.54 | 3.55 | 3.08 | 3.19 | 3.32 | 3.63 | 3.69 | 3.56 | 3.49 | 3.35 | 3.32 | 3.26 | 3.27 | 3.52 | 3.75 | 3.99 | 3.28 | 3.31 | 3.84 |
| MgO | 4.39 | 4.50 | 4.68 | 4.82 | 4.89 | 4.45 | 4.48 | 4.32 | 4.25 | 4.58 | 4.84 | 4.98 | 4.83 | 4.82 | 4.28 | 4.30 | 4.77 | 4.70 | 4.49 |
| NiO |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| CaO | 6.20 | 6.83 | 7.32 | 6.78 | 6.65 | 6.78 | 6.07 | 6.89 | 7.27 | 7.20 | 6.63 | 6.59 | 6.68 | 6.43 | 7.02 | 6.35 | 6.75 | 6.73 | 6.56 |
| $\mathrm{Na}_{2} \mathrm{O}$ | 0.06 | 0.03 | 0.01 | 0.02 | 0.02 | 0.01 | 0.02 | 0.01 | 0.03 | 0.03 | 0.02 | 0.00 | 0.03 | 0.03 | 0.01 | 0.00 | 0.02 | 0.04 | 0.02 |
| $\mathrm{K}_{2} \mathrm{O}$ | 0.00 | 0.01 | 0.01 | 0.00 | 0.00 | 0.00 | 0.00 | 0.02 | 0.01 | 0.00 | 0.02 | 0.00 | 0.01 | 0.02 | 0.02 | 0.00 | 0.00 | 0.00 | 0.01 |
| Total | 100.52 | 101.27 | 100.50 | 101.11 | 101.25 | 100.61 | 101.42 | 100.91 | 100.60 | 101.19 | 101.75 | 101.86 | 100.58 | 101.78 | 101.59 | 101.13 | 100.76 | 100.89 | 101.57 |
| Si | 2.932 | 2.994 | 3.012 | 3.006 | 3.003 | 2.995 | 3.006 | 3.013 | 2.977 | 3.009 | 3.015 | 3.019 | 3.016 | 2.994 | 2.999 | 3.013 | 3.014 | 3.028 | 3.007 |
| Ti | 0.002 | 0.002 | 0.007 | 0.007 | 0.002 | 0.005 | 0.003 | 0.004 | 0.003 | 0.005 | 0.004 | 0.004 | 0.002 | 0.002 | 0.005 | 0.000 | 0.004 | 0.006 | 0.001 |
| Cr | 0.000 | 0.002 | 0.001 | 0.002 | 0.000 | 0.002 | 0.000 | 0.000 | 0.001 | 0.003 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 |
| Al | 1.976 | 1.998 | 1.985 | 1.985 | 1.996 | 2.007 | 2.004 | 1.998 | 1.994 | 1.987 | 1.983 | 1.981 | 1.974 | 1.991 | 1.996 | 1.997 | 1.989 | 1.989 | 1.991 |
| $\mathrm{Fe}^{3}$ | 0.165 | 0.013 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.050 | 0.000 | 0.000 | 0.000 | 0.000 | 0.023 | 0.001 | 0.000 | 0.000 | 0.000 | 0.000 |
| $\mathrm{Fe}^{2}$ | 1.643 | 1.663 | 1.632 | 1.665 | 1.662 | 1.664 | 1.718 | 1.669 | 1.632 | 1.643 | 1.672 | 1.668 | 1.667 | 1.667 | 1.674 | 1.696 | 1.657 | 1.648 | 1.683 |
| Mn | 0.236 | 0.233 | 0.204 | 0.210 | 0.218 | 0.241 | 0.243 | 0.235 | 0.232 | 0.220 | 0.217 | 0.213 | 0.216 | 0.230 | 0.246 | 0.264 | 0.216 | 0.218 | 0.252 |
| Mg | 0.515 | 0.521 | 0.545 | 0.557 | 0.564 | 0.519 | 0.519 | 0.502 | 0.496 | 0.530 | 0.556 | 0.572 | 0.562 | 0.554 | 0.494 | 0.500 | 0.553 | 0.545 | 0.518 |
| Ni |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| Ca | 0.522 | 0.568 | 0.612 | 0.564 | 0.552 | 0.567 | 0.505 | 0.576 | 0.610 | 0.599 | 0.548 | 0.544 | 0.558 | 0.531 | 0.583 | 0.530 | 0.563 | 0.560 | 0.544 |
| Na | 0.008 | 0.004 | 0.001 | 0.003 | 0.003 | 0.002 | 0.003 | 0.001 | 0.004 | 0.004 | 0.003 | 0.000 | 0.004 | 0.005 | 0.001 | 0.001 | 0.003 | 0.007 | 0.002 |
| K | 0.000 | 0.001 | 0.001 | 0.001 | 0.000 | 0.000 | 0.000 | 0.002 | 0.001 | 0.000 | 0.002 | 0.000 | 0.001 | 0.002 | 0.002 | 0.000 | 0.000 | 0.000 | 0.001 |
| $\Sigma$ Cations | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 | 8.000 |
| Grossular | 0.093 | 0.181 | 0.200 | 0.183 | 0.183 | 0.186 | 0.167 | 0.190 | 0.177 | 0.195 | 0.180 | 0.179 | 0.184 | 0.165 | 0.191 | 0.177 | 0.185 | 0.185 | 0.181 |
| Almandine | 0.563 | 0.557 | 0.545 | 0.555 | 0.554 | 0.556 | 0.575 | 0.559 | 0.549 | 0.549 | 0.558 | 0.557 | 0.555 | 0.558 | 0.558 | 0.568 | 0.554 | 0.554 | 0.561 |
| Pyrope | 0.176 | 0.174 | 0.182 | 0.186 | 0.188 | 0.173 | 0.174 | 0.168 | 0.167 | 0.177 | 0.186 | 0.191 | 0.187 | 0.186 | 0.165 | 0.167 | 0.185 | 0.183 | 0.173 |
| Spessartite | 0.081 | 0.078 | 0.068 | 0.070 | 0.073 | 0.080 | 0.081 | 0.079 | 0.078 | 0.073 | 0.072 | 0.071 | 0.072 | 0.077 | 0.082 | 0.088 | 0.072 | 0.073 | 0.084 |
| Andradite | 0.085 | 0.007 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.025 | 0.000 | 0.000 | 0.000 | 0.000 | 0.012 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 |

Table Appendix 3.3: Representative chemical composition of pyroxene.


| Mineral: clinopyroxene |  |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Granulitic Gabbro mylonitic 13/01/01 |  |  |  |  |  |  |  |  |
| Sample | p01s11px9 | p01s11px8 | p01s11px6 | p01s11px5 | p01s 11px4 | p01s12am4 | p01s12am3 | p01s12am2 |
| Site | s11 | s11 | s11 | s11 | s11 | s12 | s12 | s12 |
| note | cpx relicts |  |  |  |  |  | cpx in contact to gt |  |
| $\mathrm{SiO}_{2}$ | 50.22 | 50.77 | 50.78 | 50.02 | 50.99 | 51.27 | 51.48 | 51.38 |
| $\mathrm{TiO}_{2}$ | 0.55 | 0.60 | 0.54 | 0.63 | 0.61 | 0.52 | 0.49 | 0.47 |
| $\mathrm{Cr}_{2} \mathrm{O}_{3}$ | 0.05 | 0.13 | 0.07 | 0.12 | 0.08 | 0.04 | 0.07 | 0.04 |
| $\mathrm{Al}_{2} \mathrm{O}_{3}$ | 6.08 | 6.20 | 6.40 | 6.52 | 6.37 | 5.55 | 5.87 | 5.45 |
| $\mathrm{Fe}_{2} \mathrm{O}_{3}$ | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 |
| FeO | 6.73 | 7.42 | 6.23 | 7.48 | 7.28 | 6.74 | 6.75 | 6.64 |
| MnO | 0.09 | 0.06 | 0.04 | 0.06 | 0.07 | 0.02 | 0.06 | 0.11 |
| MgO | 11.66 | 11.67 | 11.61 | 11.46 | 12.00 | 12.67 | 12.62 | 12.76 |
| NiO |  |  |  |  |  |  |  |  |
| CaO | 21.18 | 20.80 | 20.76 | 20.59 | 21.20 | 21.44 | 21.29 | 21.81 |
| $\mathrm{Na}_{2} \mathrm{O}$ | 1.32 | 1.37 | 1.50 | 1.30 | 1.23 | 1.18 | 1.25 | 1.07 |
| $\mathrm{K}_{2} \mathrm{O}$ | 0.06 | 0.02 | 0.02 | 0.04 | 0.05 | 0.04 | 0.08 | 0.07 |
| Total | 97.94 | 99.04 | 97.94 | 98.23 | 99.89 | 99.47 | 99.95 | 99.80 |
| Si | 1.893 | 1.894 | 1.904 | 1.883 | 1.886 | 1.901 | 1.898 | 1.900 |
| Ti | 0.016 | 0.017 | 0.015 | 0.018 | 0.017 | 0.015 | 0.014 | 0.013 |
| Cr | 0.002 | 0.004 | 0.002 | 0.004 | 0.002 | 0.001 | 0.002 | 0.001 |
| Al | 0.270 | 0.273 | 0.283 | 0.289 | 0.278 | 0.243 | 0.255 | 0.238 |
| $\mathrm{Fe}^{3}$ | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 |
| $\mathrm{Fe}^{2}$ | 0.212 | 0.232 | 0.195 | 0.235 | 0.225 | 0.209 | 0.208 | 0.205 |
| Mn | 0.003 | 0.002 | 0.001 | 0.002 | 0.002 | 0.001 | 0.002 | 0.003 |
| Mg | 0.655 | 0.649 | 0.649 | 0.643 | 0.662 | 0.700 | 0.694 | 0.703 |
| Ni |  |  |  |  |  |  |  |  |
| Ca | 0.856 | 0.831 | 0.834 | 0.830 | 0.840 | 0.852 | 0.841 | 0.864 |
| Na | 0.096 | 0.099 | 0.109 | 0.095 | 0.088 | 0.085 | 0.089 | 0.077 |
| K | 0.003 | 0.001 | 0.001 | 0.002 | 0.002 | 0.002 | 0.004 | 0.004 |
| $\Sigma$ Cations | 4.005 | 4.001 | 3.993 | 4.002 | 4.002 | 4.006 | 4.006 | 4.008 |
| xMg Fe(tot) | 0.755 | 0.737 | 0.769 | 0.732 | 0.746 | 0.770 | 0.769 | 0.774 |
| Wollastonite | 0.383 | 0.371 | 0.376 | 0.366 | 0.372 | 0.384 | 0.377 | 0.389 |
| Enstatite | 0.326 | 0.324 | 0.327 | 0.321 | 0.330 | 0.348 | 0.345 | 0.349 |
| Ferrosilite | 0.106 | 0.116 | 0.098 | 0.118 | 0.112 | 0.104 | 0.103 | 0.102 |
| cpx calculated with "Norm" program by P. Ulmer using: CPX Norm (Cawthom \& Collerson) cpx normalization on the basis of 12 CHARGES and $\mathrm{Fe}^{3+}=\mathrm{Acmite}$ |  |  |  |  |  |  |  |  |



| Mineral: clinopyroxene |  |  |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Granulitic Gabbro mylonitic 94/04/01 |  |  |  |  |  |  |  |  |  |
| Sample | p04s01p19 | p04s01p20 | p04s01p21 | p04s01p22 | p04s01p23 | p04s05p01 | p04s05p02 | p04s05p16 | p04s05p17 |
| Site | s01 | s01 | s01 | s01 | s01 | s05 | s05 | s05 | s05 |
| nr . | 19 | 20 | 21 | 22 | 23 | 1 | 2 | 16 | 17 |
| note | fractured cpx in contact to amp |  |  |  |  | big cpx clast in contact to gt |  |  |  |
| $\mathrm{SiO}_{2}$ | 50.82 | 49.36 | 50.45 | 50.38 | 50.56 | 50.23 | 50.82 | 52.31 | 52.65 |
| $\mathrm{TiO}_{2}$ | 0.54 | 1.03 | 0.53 | 0.47 | 0.58 | 0.55 | 0.44 | 0.35 | 0.37 |
| $\mathrm{Cr}_{2} \mathrm{O}_{3}$ | 0.00 | 0.03 | 0.01 | 0.00 | 0.00 | 0.00 | 0.04 | 0.00 | 0.00 |
| $\mathrm{Al}_{2} \mathrm{O}_{3}$ | 5.34 | 4.69 | 5.23 | 5.21 | 5.37 | 5.16 | 4.91 | 3.10 | 3.13 |
| $\mathrm{Fe}_{2} \mathrm{O}_{3}$ | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 |
| FeO | 8.15 | 9.75 | 8.37 | 8.62 | 8.48 | 8.76 | 9.31 | 7.80 | 7.42 |
| MnO | 0.06 | 0.07 | 0.07 | 0.07 | 0.06 | 0.07 | 0.09 | 0.08 | 0.09 |
| MgO | 11.75 | 11.76 | 11.57 | 11.75 | 11.57 | 11.64 | 12.36 | 12.95 | 13.11 |
| NiO | 0.03 | 0.05 | 0.00 | 0.03 | 0.00 | 0.00 | 0.00 | 0.00 | 0.02 |
| CaO | 22.36 | 21.55 | 21.81 | 21.48 | 22.11 | 21.28 | 20.26 | 22.73 | 22.65 |
| $\mathrm{Na}_{2} \mathrm{O}$ | 1.23 | 1.00 | 1.30 | 1.25 | 1.31 | 1.27 | 1.21 | 1.04 | 1.09 |
| $\mathrm{K}_{2} \mathrm{O}$ | 0.02 | 0.09 | 0.01 | 0.03 | 0.02 | 0.03 | 0.00 | 0.02 | 0.01 |
| Total | 100.30 | 99.38 | 99.37 | 99.30 | 100.06 | 98.98 | 99.44 | 100.39 | 100.54 |
| Si | 1.889 | 1.869 | 1.893 | 1.893 | 1.886 | 1.893 | 1.904 | 1.939 | 1.943 |
| Ti | 0.015 | 0.029 | 0.015 | 0.013 | 0.016 | 0.016 | 0.012 | 0.010 | 0.010 |
| Cr | 0.000 | 0.001 | 0.000 | 0.000 | 0.000 | 0.000 | 0.001 | 0.000 | 0.000 |
| Al | 0.234 | 0.209 | 0.231 | 0.231 | 0.236 | 0.229 | 0.217 | 0.135 | 0.136 |
| $\mathrm{Fe}^{3}$ | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 |
| $\mathrm{Fe}^{2}$ | 0.253 | 0.309 | 0.263 | 0.271 | 0.265 | 0.276 | 0.292 | 0.242 | 0.229 |
| Mn | 0.002 | 0.002 | 0.002 | 0.002 | 0.002 | 0.002 | 0.003 | 0.002 | 0.003 |
| Mg | 0.651 | 0.663 | 0.647 | 0.658 | 0.643 | 0.654 | 0.690 | 0.715 | 0.721 |
| Ni | 0.001 | 0.001 | 0.000 | 0.001 | 0.000 | 0.000 | 0.000 | 0.000 | 0.001 |
| Ca | 0.890 | 0.874 | 0.877 | 0.864 | 0.884 | 0.859 | 0.813 | 0.903 | 0.896 |
| Na | 0.089 | 0.074 | 0.094 | 0.091 | 0.095 | 0.093 | 0.088 | 0.075 | 0.078 |
| K | 0.001 | 0.005 | 0.001 | 0.001 | 0.001 | 0.001 | 0.001 | 0.001 | 0.001 |
| $\Sigma$ Cations | 4.024 | 4.036 | 4.024 | 4.025 | 4.027 | 4.024 | 4.019 | 4.021 | 4.018 |
| xMg Fe(tot) | 0.720 | 0.682 | 0.711 | 0.708 | 0.709 | 0.703 | 0.703 | 0.748 | 0.759 |
| Woll astonite | 0.399 | 0.390 | 0.395 | 0.388 | 0.396 | 0.387 | 0.367 | 0.427 | 0.426 |
| Enstatite | 0.318 | 0.321 | 0.316 | 0.321 | 0.313 | 0.320 | 0.338 | 0.350 | 0.355 |
| Ferrosilite | 0.124 | 0.149 | 0.128 | 0.132 | 0.129 | 0.135 | 0.143 | 0.118 | 0.113 |


| Mineral: orthopyroxene |  |  |  |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Sarangar Gabbro undeformed 90/02/01 |  |  |  |  |  |  |  |  |  |  |
| p10s02p09 | p10s02p10 | p10s02p11 | p10s02p12 | p10s03p03 | p10s03p04 | p10s03p05 | p10s03p06 | p10s03p07 | p10s03p08 | p10s03p09 |
| s02 | s02 | s02 | s02 | s03 | s03 | s03 | s03 | s03 | s03 | s03 |
| 25 | 26 | 27 | 28 | 50 | 51 | 52 | 53 | 54 | 55 | 56 |
|  | in contact to opx |  |  | opx rimed by amp |  |  |  |  |  |  |
| 49.26 | 49.34 | 49.07 | 49.12 | 48.17 | 49.07 | 48.54 | 48.81 | 49.22 | 48.71 | 48.44 |
| 0.17 | 0.11 | 0.03 | 0.13 | 0.09 | 0.08 | 0.12 | 0.11 | 0.09 | 0.05 | 0.06 |
| 0.03 | 0.05 | 0.04 | 0.02 | 0.02 | 0.04 | 0.02 | 0.04 | 0.04 | 0.01 | 0.00 |
| 2.77 | 2.59 | 2.54 | 2.54 | 2.57 | 2.54 | 2.66 | 2.63 | 2.75 | 2.54 | 2.91 |
| 1.20 | 1.02 | 1.23 | 1.60 | 1.86 | 0.50 | 0.93 | 1.43 | 0.14 | 1.89 | 1.52 |
| 27.99 | 28.59 | 27.85 | 27.93 | 28.38 | 28.62 | 27.84 | 27.12 | 25.56 | 27.52 | 26.89 |
| 0.82 | 0.82 | 0.77 | 0.74 | 0.66 | 0.67 | 0.66 | 0.73 | 0.60 | 0.72 | 0.71 |
| 16.27 | 16.10 | 16.18 | 16.47 | 15.34 | 16.12 | 16.24 | 16.47 | 15.73 | 16.53 | 16.03 |
| 0.08 | 0.00 | 0.00 | 0.00 | 0.04 | 0.02 | 0.00 | 0.00 | 0.06 | 0.03 | 0.00 |
| 0.75 | 0.71 | 0.75 | 0.53 | 0.64 | 0.55 | 0.44 | 0.74 | 3.18 | 0.40 | 1.11 |
| 0.03 | 0.01 | 0.06 | 0.03 | 0.08 | 0.00 | 0.03 | 0.06 | 0.13 | 0.00 | 0.07 |
| 0.00 | 0.00 | 0.00 | 0.00 | 0.01 | 0.00 | 0.00 | 0.01 | 0.00 | 0.00 | 0.02 |
| 99.38 | 99.35 | 98.53 | 99.11 | 97.87 | 98.20 | 97.48 | 98.15 | 97.50 | 98.41 | 97.78 |
| 1.915 | 1.922 | 1.924 | 1.915 | 1.913 | 1.931 | 1.921 | 1.917 | 1.936 | 1.912 | 1.911 |
| 0.005 | 0.003 | 0.001 | 0.004 | 0.003 | 0.003 | 0.004 | 0.003 | 0.003 | 0.001 | 0.002 |
| 0.001 | 0.002 | 0.001 | 0.001 | 0.001 | 0.001 | 0.001 | 0.001 | 0.001 | 0.001 | 0.000 |
| 0.127 | 0.119 | 0.117 | 0.117 | 0.120 | 0.118 | 0.124 | 0.122 | 0.127 | 0.118 | 0.135 |
| 0.035 | 0.030 | 0.036 | 0.047 | 0.056 | 0.015 | 0.028 | 0.042 | 0.004 | 0.056 | 0.045 |
| 0.910 | 0.932 | 0.913 | 0.911 | 0.943 | 0.942 | 0.922 | 0.891 | 0.841 | 0.903 | 0.887 |
| 0.027 | 0.027 | 0.025 | 0.024 | 0.022 | 0.022 | 0.022 | 0.024 | 0.020 | 0.024 | 0.024 |
| 0.943 | 0.935 | 0.946 | 0.957 | 0.908 | 0.946 | 0.958 | 0.964 | 0.922 | 0.967 | 0.943 |
| 0.003 | 0.000 | 0.000 | 0.000 | 0.001 | 0.001 | 0.000 | 0.000 | 0.002 | 0.001 | 0.000 |
| 0.031 | 0.030 | 0.032 | 0.022 | 0.027 | 0.023 | 0.019 | 0.031 | 0.134 | 0.017 | 0.047 |
| 0.003 | 0.001 | 0.004 | 0.002 | 0.007 | 0.000 | 0.002 | 0.004 | 0.010 | 0.000 | 0.005 |
| 0.001 | 0.001 | 0.000 | 0.000 | 0.001 | 0.000 | 0.000 | 0.001 | 0.000 | 0.000 | 0.001 |
| 4.000 | 4.000 | 4.000 | 4.000 | 4.000 | 4.000 | 4.000 | 4.000 | 4.000 | 4.000 | 4.000 |
| 0.499 | 0.493 | 0.499 | 0.500 | 0.476 | 0.497 | 0.502 | 0.508 | 0.522 | 0.502 | 0.503 |
| 0.016 | 0.015 | 0.016 | 0.011 | 0.014 | 0.012 | 0.009 | 0.015 | 0.067 | 0.008 | 0.023 |
| 0.453 | 0.449 | 0.454 | 0.458 | 0.434 | 0.457 | 0.460 | 0.461 | 0.446 | 0.462 | 0.449 |
| 0.436 | 0.447 | 0.438 | 0.436 | 0.450 | 0.454 | 0.443 | 0.426 | 0.406 | 0.431 | 0.423 |

[^1]| Mineral: clinopyroxene |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Sarangar Gabbro undeformed 90/02/01 |  |  |  |  |  |  |  |
| Sample | p10s02p01 | p10s02p02 | p10s02p03 | p10s02p04 | p10s02p05 | p10s02p07 | p10s02p08 |
| Site | s02 | s02 | s02 | s02 | s02 | s02 | s02 |
| nr . | 17 | 18 | 19 | 20 | 21 | 23 | 24 |
| note | big cpx with ilm incl. and plg incl. |  |  |  |  | smaller cpx |  |
| $\mathrm{SiO}_{2}$ | 49.76 | 49.92 | 49.98 | 49.75 | 49.83 | 50.00 | 49.62 |
| $\mathrm{TiO}_{2}$ | 0.35 | 0.34 | 0.38 | 0.30 | 0.37 | 0.36 | 0.34 |
| $\mathrm{Cr}_{2} \mathrm{O}_{3}$ | 0.16 | 0.16 | 0.18 | 0.14 | 0.15 | 0.09 | 0.06 |
| $\mathrm{Al}_{2} \mathrm{O}_{3}$ | 3.78 | 3.73 | 3.79 | 3.54 | 3.37 | 3.98 | 3.82 |
| $\mathrm{Fe}_{2} \mathrm{O}_{3}$ | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 |
| FeO | 11.67 | 11.72 | 11.44 | 11.60 | 11.62 | 11.77 | 11.69 |
| MnO | 0.35 | 0.35 | 0.32 | 0.34 | 0.38 | 0.33 | 0.31 |
| MgO | 10.60 | 10.70 | 10.76 | 10.74 | 10.92 | 10.67 | 10.78 |
| NiO | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.06 | 0.00 |
| CaO | 21.34 | 21.27 | 21.43 | 21.44 | 21.49 | 21.24 | 21.00 |
| $\mathrm{Na}_{2} \mathrm{O}$ | 0.67 | 0.67 | 0.73 | 0.66 | 0.72 | 0.70 | 0.70 |
| $\mathrm{K}_{2} \mathrm{O}$ | 0.02 | 0.00 | 0.00 | 0.02 | 0.00 | 0.02 | 0.00 |
| Total | 98.69 | 98.87 | 99.02 | 98.51 | 98.85 | 99.21 | 98.34 |
| Si | 1.910 | 1.912 | 1.910 | 1.914 | 1.912 | 1.908 | 1.910 |
| Ti | 0.010 | 0.010 | 0.011 | 0.009 | 0.011 | 0.010 | 0.010 |
| Cr | 0.005 | 0.005 | 0.005 | 0.004 | 0.005 | 0.003 | 0.002 |
| Al | 0.171 | 0.168 | 0.171 | 0.160 | 0.152 | 0.179 | 0.174 |
| $\mathrm{Fe}^{3}$ | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 |
| $\mathrm{Fe}^{2}$ | 0.375 | 0.376 | 0.366 | 0.373 | 0.373 | 0.376 | 0.376 |
| Mn | 0.011 | 0.011 | 0.011 | 0.011 | 0.012 | 0.011 | 0.010 |
| Mg | 0.607 | 0.611 | 0.613 | 0.616 | 0.625 | 0.607 | 0.619 |
| Ni | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.002 | 0.000 |
| Ca | 0.878 | 0.873 | 0.878 | 0.884 | 0.884 | 0.869 | 0.866 |
| Na | 0.050 | 0.050 | 0.054 | 0.049 | 0.053 | 0.052 | 0.052 |
| K | 0.001 | 0.000 | 0.000 | 0.001 | 0.000 | 0.001 | 0.000 |
| $\Sigma$ Cations | 4.017 | 4.016 | 4.018 | 4.020 | 4.026 | 4.017 | 4.019 |
| $\times \mathrm{Mg} \mathrm{Fe}$ (tot) | 0.618 | 0.619 | 0.627 | 0.623 | 0.626 | 0.618 | 0.622 |
| Wollastonite | 0.401 | 0.399 | 0.401 | 0.405 | 0.405 | 0.395 | 0.395 |
| Enstatite | 0.298 | 0.300 | 0.301 | 0.302 | 0.304 | 0.299 | 0.304 |
| Ferrosilite | 0.184 | 0.185 | 0.179 | 0.183 | 0.182 | 0.185 | 0.185 |


| Mineral: orthopyroxene |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Sarangar Gabbro undeformed 16/01/01 |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| Sample | ps04am9 | ps04am 13 | ps04am14 | ps04am15 | ps04am16 | ps05am6 | ps 05 am 8 | ps05am9 | ps05am10 | ps05am 11 | ps05am12 | ps05am13 | ps05am14 | ps05am15 | ps05am16 | ps05am 16 | ps05am 16 | ps05am 16 | ps05am16 |
| Site | s04 | s04 | s04 | s04 | s04 | s05 | s05 | s05 | s05 | s05 | s05 | s05 | s05 | s05 | s05 | s05 | s05 | s05 | s05 |
| nr. | 21 | 25 | 26 | 27 | 28 | 59 | 60 | 61 | 62 | 63 | 64 | 65 | 66 | 67 | 83 | 84 | 85 | 86 | 87 |
| note | opx in center of amp with small qtz at rim |  |  |  |  | opx in amp, section |  |  |  |  |  |  |  |  | opx in amp, profile |  |  |  |  |
| $\mathrm{SiO}_{2}$ | 53.66 | 55.01 | 54.42 | 54.86 | 54.36 | 54.37 | 53.37 | 54.00 | 54.01 | 55.14 | 53.41 | 53.10 | 53.84 | 51.58 | 54.19 | 53.59 | 50.70 | 52.68 | 52.89 |
| $\mathrm{TiO}_{2}$ | 0.08 | 0.01 | 0.05 | 0.02 | 0.05 | 0.05 | 0.06 | 0.07 | 0.09 | 0.09 | 0.10 | 0.07 | 0.02 | 0.12 | 0.08 | 0.04 | 0.16 | 0.08 | 0.09 |
| $\mathrm{Cr}_{2} \mathrm{O}_{3}$ | 0.00 | 0.00 | 0.00 | 0.03 | 0.00 | 0.00 | 0.00 | 0.02 | 0.03 | 0.00 | 0.01 | 0.00 | 0.00 | 0.02 | 0.00 | 0.04 | 0.00 | 0.01 | 0.00 |
| $\mathrm{Al}_{2} \mathrm{O}_{3}$ | 3.70 | 2.24 | 2.51 | 1.68 | 2.35 | 1.68 | 2.69 | 2.37 | 2.02 | 2.63 | 2.72 | 3.26 | 2.34 | 5.72 | 2.39 | 3.11 | 7.31 | 3.99 | 4.03 |
| $\mathrm{Fe}_{2} \mathrm{O}_{3}$ | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 |
| FeO | 19.47 | 19.81 | 20.18 | 20.75 | 19.82 | 21.82 | 21.85 | 21.68 | 22.01 | 21.60 | 21.24 | 21.50 | 21.69 | 19.49 | 21.13 | 20.66 | 18.49 | 20.36 | 20.34 |
| MnO | 0.76 | 0.93 | 0.88 | 0.89 | 0.83 | 0.81 | 0.79 | 0.79 | 0.80 | 0.78 | 0.83 | 0.82 | 0.84 | 0.69 | 0.81 | 0.77 | 0.59 | 0.79 | 0.77 |
| Mgo | 17.38 | 17.76 | 17.76 | 18.33 | 18.10 | 17.49 | 16.95 | 17.25 | 17.41 | 18.01 | 16.90 | 16.82 | 17.37 | 15.57 | 17.87 | 16.98 | 14.70 | 16.61 | 16.56 |
| Nio | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 |
| CaO | 3.65 | 2.53 | 2.89 | 1.94 | 2.54 | 1.58 | 2.19 | 2.02 | 1.71 | 2.21 | 2.34 | 2.56 | 2.04 | 4.81 | 1.97 | 2.69 | 6.13 | 3.41 | 3.37 |
| $\mathrm{Na}_{2} \mathrm{O}$ | 0.32 | 0.26 | 0.29 | 0.19 | 0.30 | 0.29 | 0.31 | 0.26 | 0.31 | 0.35 | 0.34 | 0.42 | 0.28 | 0.61 | 0.29 | 0.40 | 0.71 | 0.40 | 0.43 |
| $\mathrm{K}_{2} \mathrm{O}$ | 0.12 | 0.10 | 0.06 | 0.05 | 0.11 | 0.17 | 0.09 | 0.09 | 0.05 | 0.09 | 0.13 | 0.13 | 0.07 | 0.18 | 0.11 | 0.09 | 0.18 | 0.11 | 0.10 |
| Total | 99.13 | 98.66 | 99.04 | 98.73 | 98.46 | 98.26 | 98.31 | 98.55 | 98.43 | 100.90 | 98.03 | 98.69 | 98.48 | 98.79 | 98.83 | 98.37 | 98.98 | 98.45 | 98.57 |
| Si | 1.999 | 2.052 | 2.030 | 2.051 | 2.035 | 2.053 | 2.020 | 2.034 | 2.039 | 2.025 | 2.024 | 2.003 | 2.031 | 1.941 | 2.030 | 2.018 | 1.903 | 1.986 | 1.990 |
| Ti | 0.002 | 0.000 | 0.002 | 0.001 | 0.001 | 0.002 | 0.002 | 0.002 | 0.002 | 0.003 | 0.003 | 0.002 | 0.001 | 0.003 | 0.002 | 0.001 | 0.005 | 0.002 | 0.002 |
| Cr | 0.000 | 0.000 | 0.000 | 0.001 | 0.000 | 0.000 | 0.000 | 0.001 | 0.001 | 0.000 | 0.000 | 0.000 | 0.000 | 0.001 | 0.000 | 0.001 | 0.000 | 0.000 | 0.000 |
| Al | 0.162 | 0.098 | 0.110 | 0.074 | 0.104 | 0.075 | 0.120 | 0.105 | 0.090 | 0.114 | 0.121 | 0.145 | 0.104 | 0.254 | 0.105 | 0.138 | 0.323 | 0.178 | 0.179 |
| $\mathrm{Fe}^{3}$ | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 |
| $\mathrm{Fe}^{2}$ | 0.607 | 0.618 | 0.629 | 0.649 | 0.621 | 0.689 | 0.692 | 0.683 | 0.695 | 0.663 | 0.673 | 0.678 | 0.684 | 0.613 | 0.662 | 0.651 | 0.581 | 0.642 | 0.640 |
| Mn | 0.024 | 0.030 | 0.028 | 0.028 | 0.026 | 0.026 | 0.025 | 0.025 | 0.026 | 0.024 | 0.027 | 0.026 | 0.027 | 0.022 | 0.026 | 0.025 | 0.019 | 0.025 | 0.025 |
| Mg | 0.965 | 0.988 | 0.988 | 1.022 | 1.010 | 0.984 | 0.956 | 0.968 | 0.979 | 0.986 | 0.954 | 0.946 | 0.976 | 0.874 | 0.998 | 0.953 | 0.823 | 0.934 | 0.929 |
| Ni | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 |
| Ca | 0.146 | 0.101 | 0.115 | 0.078 | 0.102 | 0.064 | 0.089 | 0.081 | 0.069 | 0.087 | 0.095 | 0.104 | 0.082 | 0.194 | 0.079 | 0.109 | 0.247 | 0.138 | 0.136 |
| Na | 0.023 | 0.019 | 0.021 | 0.014 | 0.022 | 0.022 | 0.023 | 0.019 | 0.023 | 0.025 | 0.025 | 0.031 | 0.021 | 0.045 | 0.021 | 0.029 | 0.052 | 0.029 | 0.031 |
| K | 0.006 | 0.005 | 0.003 | 0.002 | 0.005 | 0.008 | 0.004 | 0.005 | 0.003 | 0.004 | 0.006 | 0.006 | 0.003 | 0.009 | 0.005 | 0.005 | 0.009 | 0.005 | 0.005 |
| $\Sigma$ Cations | 3.933 | 3.910 | 3.925 | 3.919 | 3.925 | 3.922 | 3.931 | 3.923 | 3.926 | 3.931 | 3.929 | 3.941 | 3.929 | 3.955 | 3.928 | 3.928 | 3.961 | 3.939 | 3.936 |
| xMg Fe(tot) | 0.614 | 0.615 | 0.611 | 0.612 | 0.619 | 0.588 | 0.580 | 0.587 | 0.585 | 0.598 | 0.586 | 0.582 | 0.588 | 0.588 | 0.601 | 0.594 | 0.586 | 0.592 | 0.592 |
| Wollastonite | 0.073 | 0.050 | 0.058 | 0.039 | 0.051 | 0.032 | 0.044 | 0.041 | 0.035 | 0.043 | 0.047 | 0.052 | 0.041 | 0.097 | 0.040 | 0.054 | 0.123 | 0.069 | 0.068 |
| Enstatite | 0.482 | 0.494 | 0.494 | 0.511 | 0.505 | 0.492 | 0.478 | 0.484 | 0.490 | 0.493 | 0.477 | 0.473 | 0.488 | 0.421 | 0.499 | 0.476 | 0.385 | 0.463 | 0.462 |
| Ferrosilite | 0.303 | 0.309 | 0.315 | 0.324 | 0.310 | 0.345 | 0.346 | 0.341 | 0.347 | 0.332 | 0.336 | 0.339 | 0.342 | 0.295 | 0.331 | 0.325 | 0.271 | 0.319 | 0.319 |

Table Appendix 3.4: Representative chemical composition of amphiboles.

| Mineral: amphibole |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Granulitic Gabbro mylonitic 13/01/01 |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| Sample | p01s07mt4 | p01s07m16 | p01s07mt8 | p01s07mt9 | p01s07mt1 | p01s07mt1 | p01s07mt2 | p01s09hb3 | p01s09hb1 | p01s09hb4 | p01s09hb6 | p01s09hb7 | p01s09hb8 | p01s09hb2 | p01s13ri1 | p01s13hb1 | p01s13nb2 | p01s13hb3 | p01s 13hb4 |
| Site | s07 | s07 | s07 | s07 | s07 | s07 | s07 | s09 | s09 | s09 | s09 | s09 | s09 | s09 | s13 | s13 | s13 | s13 | s13 |
| nr. | 71 | 73 | 75 | 76 | 77 | 79 | 69 | 83 | 80 | 82 | 91 | 92 | 93 | 81 | 98 | 94 | 95 | 96 | 97 |
| note | amphibole in matrix |  |  |  |  |  |  | amp in contact to gt |  |  |  |  |  |  |  |  |  |  |  |
| $\mathrm{SiO}_{2}$ | 43.59 | 44.17 | 49.83 | 45.17 | 48.49 | 44.23 | 44.22 | 44.02 | 43.02 | 42.53 | 41.44 | 43.08 | 42.82 | 43.65 | 49.70 | 42.19 | 42.51 | 42.73 | 42.93 |
| $\mathrm{TiO}_{2}$ | 0.50 | 0.48 | 0.46 | 0.43 | 0.44 | 0.44 | 0.50 | 0.34 | 0.38 | 0.39 | 0.39 | 0.43 | 0.41 | 0.31 | 0.31 | 0.20 | 0.21 | 0.21 | 0.21 |
| $\mathrm{Cr}_{2} \mathrm{O}_{3}$ | 0.00 | 0.05 | 0.07 | 0.00 | 0.08 | 0.03 | 0.00 | 0.02 | 0.03 | 0.00 | 0.00 | 0.05 | 0.00 | 0.05 | 0.02 | 0.03 | 0.00 | 0.05 | 0.00 |
| $\mathrm{Al}_{2} \mathrm{O}_{3}$ | 14.95 | 14.65 | 14.16 | 14.21 | 12.54 | 15.15 | 14.93 | 15.31 | 15.67 | 16.59 | 18.10 | 14.48 | 16.36 | 15.55 | 15.07 | 16.78 | 16.86 | 16.22 | 16.42 |
| $\mathrm{Fe}_{2} \mathrm{O}_{3}$ | 3.38 | 4.32 | 0.00 | 4.68 | 0.83 | 4.31 | 4.43 | 5.87 | 4.41 | 5.59 | 4.91 | 3.24 | 5.15 | 5.39 | 0.00 | 6.31 | 5.48 | 6.10 | 5.00 |
| FeO | 8.94 | 8.23 | 11.20 | 7.87 | 10.22 | 8.41 | 8.03 | 7.89 | 8.85 | 8.79 | 8.97 | 9.34 | 8.45 | 8.17 | 11.35 | 7.73 | 8.34 | 7.41 | 7.91 |
| MnO | 0.16 | 0.13 | 0.12 | 0.16 | 0.18 | 0.17 | 0.09 | 0.13 | 0.20 | 0.12 | 0.13 | 0.15 | 0.13 | 0.18 | 0.10 | 0.15 | 0.11 | 0.13 | 0.11 |
| MgO | 11.50 | 11.81 | 10.07 | 12.36 | 11.99 | 11.69 | 11.85 | 11.33 | 10.90 | 10.44 | 10.11 | 11.24 | 10.79 | 11.19 | 8.86 | 10.73 | 10.78 | 11.27 | 11.38 |
| Nio | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 |
| CaO | 11.45 | 11.26 | 10.19 | 11.46 | 10.76 | 11.37 | 11.33 | 10.93 | 11.27 | 10.96 | 11.23 | 11.23 | 11.11 | 10.98 | 9.92 | 11.06 | 11.23 | 11.14 | 11.35 |
| $\mathrm{Na}_{2} \mathrm{O}$ | 1.82 | 1.85 | 1.80 | 1.79 | 1.61 | 1.90 | 1.77 | 2.04 | 1.87 | 2.16 | 2.11 | 1.94 | 1.98 | 2.10 | 1.74 | 2.07 | 2.09 | 2.05 | 2.05 |
| $\mathrm{K}_{2} \mathrm{O}$ | 0.46 | 0.52 | 0.51 | 0.44 | 0.39 | 0.47 | 0.46 | 0.46 | 0.50 | 0.58 | 0.75 | 0.51 | 0.54 | 0.48 | 0.54 | 0.56 | 0.56 | 0.52 | 0.53 |
| Cl | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 |
| $\mathrm{H}_{2} \mathrm{O}$ | 2.05 | 2.07 | 2.11 | 2.10 | 2.10 | 2.08 | 2.08 | 2.08 | 2.05 | 2.07 | 2.06 | 2.02 | 2.07 | 2.08 | 2.09 | 2.06 | 2.07 | 2.07 | 2.07 |
| Total | 98.80 | 99.53 | 100.52 | 100.66 | 99.61 | 100.26 | 99.70 | 100.44 | 99.14 | 100.20 | 100.21 | 97.71 | 99.81 | 100.12 | 99.69 | 99.87 | 100.25 | 99.91 | 99.96 |
| Si | 6.375 | 6.402 | 7.088 | 6.460 | 6.938 | 6.368 | 6.388 | 6.333 | 6.289 | 6.172 | 6.023 | 6.391 | 6.217 | 6.308 | 7.140 | 6.130 | 6.154 | 6.190 | 6.210 |
| Ti | 0.054 | 0.052 | 0.049 | 0.046 | 0.047 | 0.048 | 0.054 | 0.037 | 0.042 | 0.042 | 0.043 | 0.048 | 0.045 | 0.034 | 0.034 | 0.022 | 0.023 | 0.023 | 0.023 |
| Cr | 0.001 | 0.006 | 0.007 | 0.000 | 0.009 | 0.003 | 0.001 | 0.003 | 0.003 | 0.000 | 0.000 | 0.006 | 0.000 | 0.005 | 0.003 | 0.003 | 0.000 | 0.006 | 0.000 |
| Al | 2.577 | 2.502 | 2.374 | 2.396 | 2.115 | 2.572 | 2.543 | 2.597 | 2.699 | 2.837 | 3.100 | 2.531 | 2.800 | 2.648 | 2.551 | 2.873 | 2.877 | 2.769 | 2.799 |
| $\mathrm{Fe}^{3}$ | 0.372 | 0.471 | 0.000 | 0.504 | 0.089 | 0.467 | 0.481 | 0.636 | 0.485 | 0.611 | 0.536 | 0.362 | 0.562 | 0.586 | 0.000 | 0.690 | 0.597 | 0.665 | 0.545 |
| $\mathrm{Fe}^{2}$ | 1.093 | 0.998 | 1.333 | 0.941 | 1.223 | 1.012 | 0.970 | 0.949 | 1.083 | 1.067 | 1.091 | 1.159 | 1.026 | 0.987 | 1.364 | 0.940 | 1.010 | 0.897 | 0.957 |
| Mn | 0.019 | 0.016 | 0.014 | 0.019 | 0.021 | 0.020 | 0.011 | 0.016 | 0.025 | 0.014 | 0.016 | 0.019 | 0.016 | 0.022 | 0.012 | 0.018 | 0.014 | 0.016 | 0.013 |
| Mg | 2.508 | 2.552 | 2.135 | 2.634 | 2.558 | 2.509 | 2.551 | 2.430 | 2.374 | 2.258 | 2.191 | 2.485 | 2.334 | 2.409 | 1.897 | 2.324 | 2.326 | 2.433 | 2.453 |
| Ni | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 |
| Ca | 1.794 | 1.748 | 1.554 | 1.756 | 1.650 | 1.754 | 1.754 | 1.685 | 1.765 | 1.704 | 1.749 | 1.785 | 1.729 | 1.700 | 1.527 | 1.721 | 1.742 | 1.730 | 1.759 |
| Na | 0.516 | 0.521 | 0.497 | 0.497 | 0.446 | 0.532 | 0.496 | 0.570 | 0.529 | 0.608 | 0.595 | 0.557 | 0.557 | 0.588 | 0.484 | 0.583 | 0.585 | 0.577 | 0.575 |
| K | 0.086 | 0.095 | 0.093 | 0.080 | 0.071 | 0.086 | 0.085 | 0.084 | 0.093 | 0.107 | 0.139 | 0.097 | 0.100 | 0.089 | 0.098 | 0.105 | 0.104 | 0.097 | 0.098 |
| OH | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 |
| $\Sigma$ Cations | 15.397 | 15.364 | 15.143 | 15.333 | 15.167 | 15.372 | 15.335 | 15.340 | 15.387 | 15.420 | 15.483 | 15.439 | 15.385 | 15.377 | 15.108 | 15.408 | 15.431 | 15.404 | 15.432 |
| xMg Fe(tot) | 0.631 | 0.635 | 0.616 | 0.646 | 0.661 | 0.629 | 0.637 | 0.605 | 0.602 | 0.574 | 0.574 | 0.620 | 0.595 | 0.605 | 0.582 | 0.588 | 0.592 | 0.609 | 0.620 |
| Tschermak | 1.434 | 1.486 | 1.216 | 1.451 | 1.245 | 1.507 | 1.522 | 1.642 | 1.560 | 1.704 | 1.745 | 1.385 | 1.669 | 1.616 | 1.225 | 1.740 | 1.673 | 1.676 | 1.600 |
| Edenite | 0.397 | 0.364 | 0.143 | 0.333 | 0.167 | 0.372 | 0.335 | 0.340 | 0.387 | 0.420 | 0.483 | 0.439 | 0.385 | 0.377 | 0.108 | 0.408 | 0.431 | 0.404 | 0.432 |
| Plagioclase | 0.206 | 0.252 | 0.446 | 0.244 | 0.350 | 0.246 | 0.246 | 0.315 | 0.235 | 0.296 | 0.251 | 0.215 | 0.271 | 0.300 | 0.473 | 0.279 | 0.258 | 0.270 | 0.241 |


| Mineral: amphibole |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Granulitic Gabbro undeformed |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| Sample | p02s03p11 | p02s03p11 | p02s03p11 | p02s03pl1 | p02s03p11 | p02s03ri1 | p02s03ri2 | p02s03hb1 | p02s03nb2 | p02s07am1 | p02s07am2 | p02s07am3 | p02s07am4 | p02s07am6 | p02s07am7 | p02s07am8 |
| Site | s03 | s03 | s03 | s03 | s03 | s03 | s03 | s03 | s03 | s07 | s07 | s07 | s07 | s07 | s07 | s07 |
| nr . | 30 | 33 | 35 | 52 | 55 | 59 | 60 | 67 | 68 | 1 | 2 | 3 | 4 | 14 | 15 | 16 |
| note | amp in contact to plg |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| $\mathrm{SiO}_{2}$ | 49.58 | 51.07 | 52.13 | 48.05 | 51.05 | 50.43 | 52.50 | 50.95 | 49.73 | 55.67 | 50.39 | 53.16 | 56.63 | 51.79 | 54.83 | 57.16 |
| $\mathrm{TiO}_{2}$ | 0.47 | 0.17 | 0.12 | 0.61 | 0.29 | 0.39 | 0.43 | 0.29 | 0.49 | 0.49 | 0.54 | 0.51 | 2.02 | 0.60 | 0.63 | 0.51 |
| $\mathrm{Cr}_{2} \mathrm{O}_{3}$ | 0.08 | 0.05 | 0.10 | 0.08 | 0.09 | 0.07 | 0.14 | 0.04 | 0.01 | 0.05 | 0.07 | 0.09 | 0.07 | 0.08 | 0.07 | 0.05 |
| $\mathrm{Al}_{2} \mathrm{O}_{3}$ | 6.43 | 7.17 | 6.26 | 11.03 | 6.26 | 7.92 | 6.43 | 7.07 | 7.53 | 6.29 | 7.26 | 6.19 | 6.07 | 6.46 | 6.36 | 6.18 |
| $\mathrm{Fe}_{2} \mathrm{O}_{3}$ | 0.00 | 0.67 | 2.06 | 1.44 | 3.38 | 0.40 | 0.16 | 1.74 | 3.16 | 0.00 | 1.94 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 |
| FeO | 8.51 | 10.20 | 7.19 | 7.12 | 9.16 | 10.85 | 9.47 | 8.30 | 9.69 | 9.77 | 8.97 | 9.33 | 8.69 | 11.16 | 10.49 | 10.50 |
| MnO | 0.13 | 0.22 | 0.20 | 0.10 | 0.32 | 0.27 | 0.23 | 0.20 | 0.24 | 0.22 | 0.17 | 0.20 | 0.14 | 0.26 | 0.28 | 0.28 |
| MgO | 12.93 | 14.81 | 16.56 | 15.39 | 14.63 | 14.21 | 15.53 | 15.53 | 13.83 | 13.55 | 15.28 | 12.95 | 11.40 | 13.54 | 13.53 | 12.43 |
| NiO | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 |
| CaO | 18.93 | 12.55 | 12.47 | 12.29 | 11.99 | 12.58 | 12.27 | 12.28 | 11.85 | 11.43 | 12.48 | 15.51 | 10.87 | 11.84 | 11.62 | 10.90 |
| $\mathrm{Na}_{2} \mathrm{O}$ | 1.01 | 0.74 | 0.66 | 1.18 | 0.78 | 0.75 | 0.73 | 0.76 | 0.92 | 0.67 | 0.81 | 0.95 | 0.62 | 0.78 | 0.73 | 0.64 |
| $\mathrm{K}_{2} \mathrm{O}$ | 0.13 | 0.19 | 0.18 | 0.93 | 0.21 | 0.21 | 0.17 | 0.27 | 0.19 | 0.14 | 0.37 | 0.12 | 0.21 | 0.19 | 0.18 | 0.13 |
| Cl | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 |
| $\mathrm{H}_{2} \mathrm{O}$ | 1.94 | 2.10 | 2.12 | 2.11 | 2.10 | 2.10 | 2.12 | 2.10 | 2.08 | 2.12 | 2.11 | 2.03 | 2.07 | 2.07 | 2.12 | 2.13 |


| Total | 100.13 | 99.94 | 100.06 | 100.32 | 100.26 | 100.18 | 100.19 | 99.54 | 99.73 | 100.40 | 100.39 | 101.04 | 98.79 | 98.75 | 100.84 | 100.90 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Si | 7.666 | 7.298 | 7.358 | 6.817 | 7.293 | 7.213 | 7.426 | 7.271 | 7.161 | 7.863 | 7.175 | 7.837 | 8.204 | 7.510 | 7.744 | 8.041 |
| Ti | 0.055 | 0.018 | 0.012 | 0.065 | 0.031 | 0.042 | 0.046 | 0.032 | 0.053 | 0.053 | 0.057 | 0.057 | 0.220 | 0.065 | 0.067 | 0.054 |
| Cr | 0.009 | 0.005 | 0.011 | 0.009 | 0.011 | 0.008 | 0.016 | 0.004 | 0.001 | 0.006 | 0.008 | 0.010 | 0.008 | 0.010 | 0.008 | 0.006 |
| Al | 1.172 | 1.208 | 1.042 | 1.845 | 1.054 | 1.335 | 1.072 | 1.190 | 1.277 | 1.047 | 1.218 | 1.076 | 1.036 | 1.103 | 1.059 | 1.024 |
| $\mathrm{Fe}^{3}$ | 0.000 | 0.072 | 0.219 | 0.154 | 0.363 | 0.043 | 0.018 | 0.187 | 0.342 | 0.000 | 0.208 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 |
| $\mathrm{Fe}^{2}$ | 1.100 | 1.219 | 0.848 | 0.844 | 1.094 | 1.297 | 1.120 | 0.990 | 1.167 | 1.154 | 1.068 | 1.150 | 1.053 | 1.353 | 1.239 | 1.236 |
| Mn | 0.017 | 0.027 | 0.024 | 0.012 | 0.039 | 0.032 | 0.028 | 0.025 | 0.029 | 0.027 | 0.021 | 0.025 | 0.017 | 0.031 | 0.033 | 0.034 |
| Mg | 2.980 | 3.154 | 3.485 | 3.254 | 3.116 | 3.029 | 3.275 | 3.302 | 2.969 | 2.852 | 3.244 | 2.845 | 2.461 | 2.927 | 2.848 | 2.606 |
| Ni | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 |
| Ca | 3.137 | 1.922 | 1.887 | 1.868 | 1.835 | 1.928 | 1.860 | 1.878 | 1.829 | 1.729 | 1.904 | 2.449 | 1.687 | 1.839 | 1.758 | 1.644 |
| Na | 0.302 | 0.205 | 0.180 | 0.324 | 0.217 | 0.209 | 0.201 | 0.210 | 0.257 | 0.182 | 0.225 | 0.272 | 0.175 | 0.218 | 0.199 | 0.174 |
| K | 0.026 | 0.035 | 0.033 | 0.168 | 0.038 | 0.039 | 0.030 | 0.050 | 0.036 | 0.025 | 0.066 | 0.023 | 0.038 | 0.035 | 0.032 | 0.023 |
| OH | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 |
| $\Sigma$ Cations | 16.465 | 15.162 | 15.100 | 15.360 | 15.090 | 15.175 | 15.091 | 15.137 | 15.122 | 14.936 | 15.195 | 15.743 | 14.901 | 15.092 | 14.989 | 14.840 |


| Total | 100.13 | 99.94 | 100.06 | 100.32 | 100.26 | 100.18 | 100.19 | 99.54 | 99.73 | 100.40 | 100.39 | 101.04 | 98.79 | 98.75 | 100.84 | 100.90 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Si | 7.666 | 7.298 | 7.358 | 6.817 | 7.293 | 7.213 | 7.426 | 7.271 | 7.161 | 7.863 | 7.175 | 7.837 | 8.204 | 7.510 | 7.744 | 8.041 |
| Ti | 0.055 | 0.018 | 0.012 | 0.065 | 0.031 | 0.042 | 0.046 | 0.032 | 0.053 | 0.053 | 0.057 | 0.057 | 0.220 | 0.065 | 0.067 | 0.054 |
| Cr | 0.009 | 0.005 | 0.011 | 0.009 | 0.011 | 0.008 | 0.016 | 0.004 | 0.001 | 0.006 | 0.008 | 0.010 | 0.008 | 0.010 | 0.008 | 0.006 |
| Al | 1.172 | 1.208 | 1.042 | 1.845 | 1.054 | 1.335 | 1.072 | 1.190 | 1.277 | 1.047 | 1.218 | 1.076 | 1.036 | 1.103 | 1.059 | 1.024 |
| $\mathrm{Fe}^{3}$ | 0.000 | 0.072 | 0.219 | 0.154 | 0.363 | 0.043 | 0.018 | 0.187 | 0.342 | 0.000 | 0.208 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 |
| $\mathrm{Fe}^{2}$ | 1.100 | 1.219 | 0.848 | 0.844 | 1.094 | 1.297 | 1.120 | 0.990 | 1.167 | 1.154 | 1.068 | 1.150 | 1.053 | 1.353 | 1.239 | 1.236 |
| Mn | 0.017 | 0.027 | 0.024 | 0.012 | 0.039 | 0.032 | 0.028 | 0.025 | 0.029 | 0.027 | 0.021 | 0.025 | 0.017 | 0.031 | 0.033 | 0.034 |
| Mg | 2.980 | 3.154 | 3.485 | 3.254 | 3.116 | 3.029 | 3.275 | 3.302 | 2.969 | 2.852 | 3.244 | 2.845 | 2.461 | 2.927 | 2.848 | 2.606 |
| Ni | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 |
| Ca | 3.137 | 1.922 | 1.887 | 1.868 | 1.835 | 1.928 | 1.860 | 1.878 | 1.829 | 1.729 | 1.904 | 2.449 | 1.687 | 1.839 | 1.758 | 1.644 |
| Na | 0.302 | 0.205 | 0.180 | 0.324 | 0.217 | 0.209 | 0.201 | 0.210 | 0.257 | 0.182 | 0.225 | 0.272 | 0.175 | 0.218 | 0.199 | 0.174 |
| K | 0.026 | 0.035 | 0.033 | 0.168 | 0.038 | 0.039 | 0.030 | 0.050 | 0.036 | 0.025 | 0.066 | 0.023 | 0.038 | 0.035 | 0.032 | 0.023 |
| OH | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 |
| $\Sigma$ Cations | 16.465 | 15.162 | 15.100 | 15.360 | 15.090 | 15.175 | 15.091 | 15.137 | 15.122 | 14.936 | 15.195 | 15.743 | 14.901 | 15.092 | 14.989 | 14.840 |
| xMg Fe(tot) | 0.730 | 0.710 | 0.766 | 0.765 | 0.681 | 0.693 | 0.742 | 0.737 | 0.663 | 0.712 | 0.718 | 0.712 | 0.700 | 0.684 | 0.697 | 0.678 |
| Tschermak | 0.006 | 0.619 | 0.655 | 0.955 | 0.783 | 0.684 | 0.623 | 0.715 | 0.888 | 0.345 | 0.725 | 0.000 | 0.010 | 0.558 | 0.487 | 0.155 |
| Edenite | 0.328 | 0.162 | 0.100 | 0.360 | 0.090 | 0.175 | 0.091 | 0.137 | 0.122 | 0.000 | 0.195 | 0.163 | 0.000 | 0.092 | 0.000 | 0.000 |
| Plagioclase | 0.000 | 0.078 | 0.114 | 0.132 | 0.165 | 0.072 | 0.140 | 0.122 | 0.171 | 0.207 | 0.096 | 0.000 | 0.214 | 0.161 | 0.231 | 0.196 |


| Mineral: amphibole |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
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| Granulitic Gabbro mylonitic |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| Sample | p04s02p04 | p04s02p05 | p04s02p09 | p04s02p10 | p04s02p14 | p04s02p16 | p04s02p17 | p04s03p11 | p04s03p 12 | p04s03p16 | p04s04pro | p04s04pro | p04s04pro | p04s04pro | p04s04pro | p04504p06 | p04s05p08 |
| Site | s02 | s02 | s02 | s02 | s02 | s02 | s02 | s03 | s03 | s03 | s04 | s04 | s04 | s04 | s04 | s04 | s05 |
| nr . | 4 | 5 | 9 | 10 | 14 | 16 | 17 | 36 | 37 | 41 | 4 | 14 | 15 | 16 | 17 | 51 | 8 |
| note | amp matrix in contact with gt |  |  |  |  |  |  | amp in contact to gt and plg |  |  |  |  | amp matrix |  |  | to gt |  |
| $\mathrm{SiO}_{2}$ | 42.57 | 41.63 | 40.20 | 39.68 | 44.02 | 50.44 | 50.97 | 41.07 | 43.63 | 47.94 | 49.60 | 44.77 | 43.11 | 42.12 | 49.24 | 42.93 | 41.11 |
| $\mathrm{TiO}_{2}$ | 0.06 | 0.12 | 0.14 | 0.10 | 0.23 | 0.22 | 0.21 | 0.11 | 0.08 | 0.06 | 0.37 | 0.34 | 0.29 | 0.73 | 0.44 | 0.33 | 0.12 |
| $\mathrm{Cr}_{2} \mathrm{O}_{3}$ | 0.00 | 0.00 | 0.00 | 0.03 | 0.02 | 0.00 | 0.01 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.04 | 0.00 | 0.04 | 0.00 | 0.00 |
| $\mathrm{Al}_{2} \mathrm{O}_{3}$ | 17.39 | 16.47 | 18.08 | 18.95 | 14.30 | 7.40 | 7.03 | 17.61 | 15.47 | 10.66 | 19.14 | 12.54 | 15.75 | 13.57 | 8.62 | 15.38 | 16.70 |
| $\mathrm{Fe}_{2} \mathrm{O}_{3}$ | 6.45 | 6.84 | 8.86 | 8.79 | 3.29 | 5.37 | 5.91 | 6.55 | 7.43 | 5.60 | 0.00 | 6.39 | 4.33 | 6.46 | 7.07 | 3.36 | 6.46 |
| FeO | 6.57 | 6.22 | 6.03 | 7.42 | 9.31 | 5.43 | 4.94 | 6.94 | 5.54 | 5.80 | 8.04 | 7.36 | 9.86 | 8.03 | 5.76 | 8.73 | 7.46 |
| MnO | 0.05 | 0.05 | 0.09 | 0.07 | 0.09 | 0.05 | 0.06 | 0.02 | 0.02 | 0.05 | 0.05 | 0.08 | 0.10 | 0.06 | 0.08 | 0.00 | 0.05 |
| Mgo | 11.92 | 11.71 | 11.00 | 10.08 | 11.79 | 16.06 | 16.23 | 11.36 | 12.64 | 14.42 | 7.92 | 12.41 | 10.65 | 11.56 | 14.63 | 11.41 | 11.07 |
| Nio | 0.00 | 0.00 | 0.00 | 0.00 | 0.02 | 0.00 | 0.01 | 0.00 | 0.02 | 0.04 | 0.00 | 0.00 | 0.03 | 0.00 | 0.01 | 0.00 | 0.01 |
| CaO | 11.21 | 11.14 | 11.05 | 11.08 | 11.04 | 11.44 | 11.29 | 11.29 | 11.10 | 11.18 | 8.39 | 10.77 | 10.86 | 11.03 | 10.92 | 10.84 | 11.06 |
| $\mathrm{Na}_{2} \mathrm{O}$ | 2.54 | 2.27 | 2.53 | 2.72 | 2.52 | 1.36 | 1.29 | 2.25 | 2.24 | 1.61 | 3.87 | 2.15 | 2.58 | 2.10 | 1.46 | 2.40 | 2.38 |
| $\mathrm{K}_{2} \mathrm{O}$ | 0.70 | 0.49 | 0.76 | 0.77 | 0.51 | 0.18 | 0.18 | 1.02 | 0.53 | 0.35 | 0.55 | 0.41 | 0.61 | 0.60 | 0.26 | 0.55 | 0.74 |
| Cl | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 |
| $\mathrm{H}_{2} \mathrm{O}$ | 2.11 | 2.05 | 2.07 | 2.08 | 2.05 | 2.12 | 2.13 | 2.07 | 2.10 | 2.10 | 2.10 | 2.06 | 2.06 | 2.02 | 2.12 | 2.03 | 2.04 |
| Total | 101.57 | 99.00 | 100.81 | 101.77 | 99.19 | 100.05 | 100.25 | 100.31 | 100.80 | 99.82 | 100.04 | 99.30 | 100.28 | 98.28 | 100.66 | 97.97 | 99.20 |
| Si | 6.063 | 6.083 | 5.820 | 5.730 | 6.430 | 7.137 | 7.184 | 5.958 | 6.232 | 6.835 | 7.084 | 6.516 | 6.268 | 6.256 | 6.974 | 6.332 | 6.037 |
| Ti | 0.007 | 0.013 | 0.015 | 0.011 | 0.026 | 0.023 | 0.022 | 0.012 | 0.008 | 0.007 | 0.040 | 0.037 | 0.032 | 0.081 | 0.047 | 0.037 | 0.014 |
| Cr | 0.000 | 0.000 | 0.000 | 0.003 | 0.003 | 0.000 | 0.001 | 0.000 | 0.000 | 0.000 | 0.000 | 0.001 | 0.004 | 0.000 | 0.004 | 0.000 | 0.000 |
| Al | 2.919 | 2.835 | 3.086 | 3.226 | 2.462 | 1.234 | 1.167 | 3.011 | 2.605 | 1.792 | 3.222 | 2.150 | 2.699 | 2.376 | 1.440 | 2.674 | 2.890 |
| $\mathrm{Fe}^{3}$ | 0.692 | 0.752 | 0.965 | 0.955 | 0.362 | 0.571 | 0.627 | 0.715 | 0.799 | 0.601 | 0.000 | 0.699 | 0.473 | 0.722 | 0.754 | 0.373 | 0.714 |
| $\mathrm{Fe}^{2}$ | 0.783 | 0.761 | 0.730 | 0.896 | 1.137 | 0.642 | 0.582 | 0.842 | 0.661 | 0.691 | 0.961 | 0.896 | 1.199 | 0.998 | 0.682 | 1.076 | 0.916 |
| Mn | 0.006 | 0.007 | 0.010 | 0.009 | 0.011 | 0.006 | 0.007 | 0.002 | 0.002 | 0.006 | 0.006 | 0.009 | 0.012 | 0.008 | 0.010 | 0.000 | 0.006 |
| Mg | 2.531 | 2.551 | 2.374 | 2.170 | 2.567 | 3.386 | 3.409 | 2.458 | 2.690 | 3.064 | 1.687 | 2.691 | 2.309 | 2.560 | 3.089 | 2.508 | 2.423 |
| Ni | 0.000 | 0.000 | 0.000 | 0.000 | 0.002 | 0.000 | 0.002 | 0.001 | 0.002 | 0.005 | 0.001 | 0.001 | 0.004 | 0.000 | 0.002 | 0.000 | 0.001 |
| Ca | 1.711 | 1.743 | 1.714 | 1.715 | 1.727 | 1.734 | 1.705 | 1.755 | 1.699 | 1.708 | 1.284 | 1.680 | 1.691 | 1.755 | 1.657 | 1.712 | 1.739 |
| Na | 0.701 | 0.644 | 0.711 | 0.761 | 0.712 | 0.373 | 0.351 | 0.634 | 0.620 | 0.445 | 1.071 | 0.608 | 0.728 | 0.605 | 0.401 | 0.688 | 0.679 |
| K | 0.127 | 0.091 | 0.141 | 0.143 | 0.095 | 0.033 | 0.032 | 0.189 | 0.097 | 0.064 | 0.101 | 0.077 | 0.114 | 0.114 | 0.047 | 0.104 | 0.138 |
| OH | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 |
| $\Sigma$ Cations | 15.539 | 15.478 | 15.566 | 15.619 | 15.535 | 15.140 | 15.088 | 15.577 | 15.416 | 15.216 | 15.455 | 15.364 | 15.532 | 15.474 | 15.105 | 15.504 | 15.556 |
| xMg Fe(tot) | 0.632 | 0.628 | 0.583 | 0.540 | 0.631 | 0.736 | 0.738 | 0.612 | 0.648 | 0.703 | 0.637 | 0.628 | 0.580 | 0.598 | 0.683 | 0.634 | 0.598 |
| Tschermak | 1.687 | 1.696 | 1.900 | 1.936 | 1.307 | 0.989 | 1.023 | 1.709 | 1.653 | 1.241 | 1.176 | 1.441 | 1.508 | 1.516 | 1.264 | 1.452 | 1.668 |
| Edenite | 0.539 | 0.478 | ${ }^{0.566}$ | 0.619 | ${ }^{0.535}$ | 0.140 | ${ }^{0.088}$ | ${ }^{0.577}$ | 0.416 | 0.216 | 0.455 | ${ }^{0.364}$ | ${ }^{0.532}$ | 0.474 | ${ }^{0.105}$ | ${ }^{0.504}$ | 0.556 |
| Plagioclase | 0.289 | 0.257 | 0.286 | 0.285 | 0.273 | 0.266 | 0.295 | 0.245 | 0.301 | 0.292 | 0.716 | 0.321 | 0.309 | 0.245 | 0.343 | 0.288 | 0.261 |


| Mineral: amphibole |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
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| Sarangar Gabbro mylonitic 90/02/01 |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| Sample | p07a01p05 | p07a01p07 | p07a01p11 | p07a01p15 | p07a01p16 | p07a01p17 | p07a03p10 | p07a03p11 | p07a04p02 | p07a04p05 | p07a04p07 | p07a04p11 | p07a04p12 | p07a04p16 | p07a04p20 | p07a04p21 | p07a05p13 | p07a08p05 | p07a08p07 |
| Site | a01 | a01 | a01 | a01 | a01 | a01 | a03 | a03 | a04 | a04 | a04 | a04 | a04 | a04 | a04 | a04 | a05 | a08 | a08 |
| $n \mathrm{n}$. | 5 | 7 | 11 | 15 | 16 | 17 | 36 | 37 | 53 | 56 | 58 | 62 | 63 | 67 | 71 | 72 | 89 | 105 | 107 |
| note | amp at epi |  |  | amp at epi and plg |  |  | to epi and gt |  | porphyroclast in contact to epi \| |  |  | porphyroclast |  | amp in plg rich layer |  |  | $\mid \mathrm{amp}$ at gt\| | in gt crack |  |
| $\mathrm{SiO}_{2}$ | 40.77 | 41.33 | 40.43 | 40.98 | 41.13 | 40.46 | 40.53 | 40.67 | 40.36 | 40.59 | 40.57 | 40.34 | 40.43 | 40.88 | 40.71 | 40.75 | 40.71 | 39.96 | 40.97 |
| $\mathrm{TiO}_{2}$ | 0.64 | 0.63 | 0.62 | 0.71 | 0.67 | 0.75 | 0.68 | 0.74 | 0.71 | 0.79 | 0.71 | 0.58 | 0.67 | 0.71 | 0.66 | 0.75 | 0.62 | 0.55 | 0.71 |
| $\mathrm{Cr}_{2} \mathrm{O}_{3}$ | 0.00 | 0.02 | 0.05 | 0.02 | 0.04 | 0.01 | 0.00 | 0.03 | 0.04 | 0.03 | 0.00 | 0.03 | 0.04 | 0.03 | 0.03 | 0.04 | 0.03 | 0.02 | 0.04 |
| $\mathrm{Al}_{2} \mathrm{O}_{3}$ | 16.91 | 16.60 | 16.52 | 16.36 | 17.35 | 16.30 | 16.01 | 15.89 | 16.12 | 15.47 | 15.15 | 15.44 | 16.23 | 16.14 | 15.83 | 15.72 | 16.44 | 16.31 | 16.38 |
| $\mathrm{Fe}_{2} \mathrm{O}_{3}$ | 5.63 | 4.08 | 3.46 | 4.27 | 5.07 | 5.57 | 4.65 | 4.10 | 4.11 | 5.85 | 5.85 | 5.08 | 5.43 | 5.49 | 4.00 | 3.74 | 5.50 | 5.09 | 4.11 |
| FeO | 10.68 | 11.75 | 12.09 | 11.46 | 9.84 | 10.30 | 11.35 | 11.84 | 11.95 | 10.81 | 11.09 | 11.82 | 10.94 | 11.01 | 11.78 | 12.15 | 11.17 | 10.91 | 10.30 |
| MnO | 0.08 | 0.16 | 0.08 | 0.05 | 0.13 | 0.14 | 0.04 | 0.16 | 0.11 | 0.11 | 0.17 | 0.14 | 0.13 | 0.09 | 0.09 | 0.10 | 0.16 | 0.09 | 0.05 |
| MgO | 9.18 | 9.17 | 8.83 | 9.53 | 9.87 | 9.44 | 9.24 | 9.25 | 8.88 | 9.35 | 9.32 | 8.86 | 9.16 | 9.26 | 9.10 | 9.19 | 8.99 | 9.20 | 9.78 |
| Nio | 0.00 | 0.02 | 0.03 | 0.00 | 0.02 | 0.00 | 0.05 | 0.00 | 0.02 | 0.01 | 0.05 | 0.04 | 0.03 | 0.00 | 0.09 | 0.00 | 0.04 | 0.07 | 0.02 |
| CaO | 11.36 | 11.62 | 11.56 | 11.70 | 11.60 | 11.34 | 11.53 | 11.69 | 11.48 | 11.41 | 11.43 | 11.52 | 11.42 | 11.48 | 11.47 | 11.66 | 11.43 | 11.55 | 11.32 |
| $\mathrm{Na}_{2} \mathrm{O}$ | 1.92 | 1.83 | 1.73 | 1.91 | 1.89 | 1.87 | 1.83 | 1.83 | 1.84 | 1.82 | 1.90 | 1.71 | 1.88 | 1.83 | 1.81 | 1.79 | 1.84 | 1.83 | 1.88 |
| $\mathrm{K}_{2} \mathrm{O}$ | 0.64 | 0.65 | 0.67 | 0.66 | 0.65 | 0.66 | 0.59 | 0.62 | 0.62 | 0.57 | 0.58 | 0.63 | 0.64 | 0.58 | 0.66 | 0.64 | 0.66 | 0.64 | 0.62 |
| Cl | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 |
| $\mathrm{H}_{2} \mathrm{O}$ | 2.03 | 2.03 | 1.99 | 2.03 | 2.05 | 2.01 | 2.00 | 2.00 | 1.99 | 2.00 | 2.00 | 1.99 | 2.01 | 2.02 | 1.99 | 2.00 | 2.02 | 1.99 | 2.01 |
| Total | 99.86 | 99.88 | 98.06 | 99.67 | 100.32 | 98.88 | 98.50 | 98.82 | 98.23 | 98.82 | 98.81 | 98.18 | 99.03 | 99.52 | 98.22 | 98.53 | 99.61 | 98.19 | 98.19 |
| Si | 6.015 | 6.099 | 6.085 | 6.064 | 6.008 | 6.028 | 6.074 | 6.086 | 6.076 | 6.071 | 6.083 | 6.094 | 6.032 | 6.062 | 6.120 | 6.115 | 6.040 | 6.013 | 6.109 |
| Ti | 0.071 | 0.070 | 0.070 | 0.079 | 0.074 | 0.085 | 0.077 | 0.083 | 0.081 | 0.089 | 0.080 | 0.066 | 0.075 | 0.080 | 0.075 | 0.085 | 0.070 | 0.062 | 0.080 |
| Cr | 0.001 | 0.003 | 0.006 | 0.002 | 0.004 | 0.002 | 0.001 | 0.003 | 0.004 | 0.003 | 0.000 | 0.004 | 0.005 | 0.003 | 0.004 | 0.005 | 0.004 | 0.002 | 0.004 |
| Al | 2.941 | 2.887 | 2.931 | 2.854 | 2.988 | 2.863 | 2.828 | 2.802 | 2.860 | 2.726 | 2.678 | 2.749 | 2.854 | 2.820 | 2.806 | 2.781 | 2.874 | 2.892 | 2.878 |
| $\mathrm{Fe}^{3}$ | 0.625 | 0.453 | 0.392 | 0.475 | 0.557 | 0.624 | 0.524 | 0.462 | 0.465 | 0.658 | 0.660 | 0.577 | 0.610 | 0.612 | 0.453 | 0.422 | 0.614 | 0.576 | 0.462 |
| $\mathrm{Fe}^{2}$ | 1.318 | 1.450 | 1.522 | 1.418 | 1.202 | 1.284 | 1.423 | 1.481 | 1.505 | 1.353 | 1.390 | 1.493 | 1.366 | 1.365 | 1.481 | 1.525 | 1.386 | 1.372 | 1.285 |
| Mn | 0.010 | 0.020 | 0.011 | 0.006 | 0.016 | 0.018 | 0.006 | 0.020 | 0.014 | 0.014 | 0.021 | 0.018 | 0.017 | 0.012 | 0.012 | 0.012 | 0.021 | 0.011 | 0.007 |
| Mg | 2.019 | 2.017 | 1.980 | 2.101 | 2.149 | 2.097 | 2.063 | 2.063 | 1.993 | 2.085 | 2.082 | 1.995 | 2.038 | 2.046 | 2.040 | 2.055 | 1.988 | 2.064 | 2.173 |
| Ni | 0.001 | 0.002 | 0.003 | 0.000 | 0.002 | 0.000 | 0.006 | 0.000 | 0.003 | 0.002 | 0.006 | 0.005 | 0.004 | 0.000 | 0.011 | 0.001 | 0.005 | 0.008 | 0.002 |
| Ca | 1.796 | 1.838 | 1.864 | 1.854 | 1.815 | 1.810 | 1.851 | 1.874 | 1.851 | 1.828 | 1.836 | 1.864 | 1.826 | 1.824 | 1.847 | 1.875 | 1.817 | 1.862 | 1.808 |
| Na | 0.550 | 0.523 | 0.505 | 0.548 | 0.537 | 0.541 | 0.532 | 0.530 | 0.536 | 0.528 | 0.554 | 0.501 | 0.545 | 0.526 | 0.529 | 0.520 | 0.531 | 0.532 | 0.544 |
| K | 0.121 | 0.122 | 0.128 | 0.125 | 0.120 | 0.125 | 0.113 | 0.119 | 0.120 | 0.109 | 0.112 | 0.122 | 0.121 | 0.109 | 0.126 | 0.123 | 0.125 | 0.123 | 0.117 |
| OH | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 |
| $\Sigma$ Cations | 15.466 | 15.482 | 15.497 | 15.527 | 15.472 | 15.476 | 15.496 | 15.522 | 15.507 | 15.465 | 15.501 | 15.487 | 15.492 | 15.458 | 15.501 | 15.518 | 15.472 | 15.518 | 15.469 |
| xMg Fe(tot) | 0.510 | 0.514 | 0.509 | 0.526 | 0.550 | 0.524 | 0.515 | 0.515 | 0.503 | 0.509 | 0.504 | 0.491 | 0.508 | 0.508 | 0.513 | 0.513 | 0.499 | 0.514 | 0.555 |
| Tschermak | 1.724 | 1.581 | 1.554 | 1.554 | 1.704 | 1.686 | 1.580 | 1.519 | 1.566 | 1.636 | 1.580 | 1.555 | 1.651 | 1.657 | 1.531 | 1.492 | 1.671 | 1.607 | 1.613 |
| Edenite | 0.466 | 0.482 | 0.497 | 0.527 | 0.472 | 0.476 | 0.496 | 0.522 | 0.507 | 0.465 | 0.501 | 0.487 | 0.492 | 0.458 | 0.501 | 0.518 | 0.472 | 0.518 | 0.469 |
| Plagioclase | 0.204 | 0.162 | 0.136 | 0.146 | 0.185 | 0.191 | 0.149 | 0.127 | 0.149 | 0.172 | 0.164 | 0.136 | 0.174 | 0.177 | 0.153 | 0.125 | 0.183 | 0.138 | 0.192 |




|  <br>  | $\stackrel{\text { ¢\% }}{\text { ¢ }}$ |
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calculated with "Norm" program by P. Ulmer using: Amphibole Norm NAMP
CATIONS calculated on the bases of 23 oxygens and 13 cations $+\mathrm{K}+\mathrm{Na}+\mathrm{Ca}$
Byyyyyyyyyyy

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\stackrel{\curvearrowleft}{\square} \underset{0}{\text { Na }}
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둥 문


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| :--- |
|  |
| $\times M g$ Fe(tot) |
| Tschermak |
| Edenite |
| Plagioclase |



| Mineral: amphibole |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $\begin{aligned} & \text { Sample } \\ & \text { Site } \\ & \text { nr. } \\ & \text { note } \end{aligned}$ | Sarangar Gabbro undeformed 16/01/01 |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
|  | p13s05am17 | cso5am18 | 3s05am19 | so5am20 | p13s1auk 1 | p13s 1auk5 | p13sichb1 | p13sten | p13s3abh7 |  |  |  | ps3abb1 | 3ahb | ns3ab4 | pssbam 1 | ps3bam2 | ps3bams |
|  |  |  | s05 | s05 |  | s1a |  | s1c |  | s3b | s3c | s3c | s3a | s3a | s3a |  |  |  |
|  |  |  | 91 | 92 |  | 10 | 34 | 40 | 99 |  |  | 132 | 93 |  |  |  |  |  |
|  | amp around opx, second sit |  |  |  | 10 |  | amp around ilm |  |  |  |  |  |  |  |  | amp at symplectite (epi) |  |  |
|  |  |  |  |  | \| close to symplectite | amp at symplectite |  | amp at symplectite (epi) |  |  |  |  |  |  |  |
| $\mathrm{SiO}_{2}$ | 42.44 | 42.81 | 42.80 | 40.42 |  |  | 40.40 | 41.52 | 40.15 | 39.90 | 44.01 | 43.87 | 51.77 | 51.45 | 51.42 | 53.63 | 51.12 | 52.49 | 51.26 | 51.60 |
| $\mathrm{TiO}_{2}$ | 0.26 | 0.24 | 0.32 | 0.08 | 0.52 | 0.60 |  |  | 0.52 | 0.55 | 0.41 | 0.42 | 0.25 | 0.24 | 0.35 | 0.25 | 0.28 | 0.27 | 0.26 | 0.27 |
| $\mathrm{Cr}_{2} \mathrm{O}_{3}$ | 0.00 | 0.00 | 0.04 | 0.02 | 0.04 | 0.02 | 0.02 | 0.04 | 0.05 | 0.00 | 0.02 | 0.03 | 0.06 | 0.03 | 0.00 | 0.00 | 0.04 | 0.04 |
| $\mathrm{Al}_{2} \mathrm{O}_{3}$ | 17.05 | 16.53 | 16.92 | 20.13 | 18.89 | 17.48 | 18.97 | 18.53 | 14.05 | 14.36 | 5.91 | 5.88 | 5.85 | 4.80 | 6.32 | 5.03 | 6.25 | 6.21 |
| $\mathrm{Fe}_{2} \mathrm{O}_{3}$ | 5.86 | 5.89 | 5.53 | 5.96 | 7.67 | 5.27 | 4.74 | 5.35 | 5.51 | 5.55 | 1.88 | 1.83 | 1.86 | 0.96 | 2.12 | 1.81 | 2.36 | 1.78 |
| FeO | 10.21 | 10.20 | 10.45 | 10.91 | 10.45 | 11.69 | 14.52 | 14.55 | 9.29 | 10.66 | 11.97 | 11.55 | 10.82 | 10.71 | 10.83 | 11.20 | 10.96 | 10.74 |
| MnO | 0.20 | 0.28 | 0.29 | 0.32 | 0.30 | 0.22 | 0.31 | 0.26 | 0.28 | 0.27 | 0.30 | 0.24 | 0.27 | 0.27 | 0.28 | 0.21 | 0.28 | 0.26 |
| Mgo | 8.96 | 8.98 | 8.98 | 7.39 | 7.34 | 8.01 | 5.57 | 5.62 | 10.57 | 9.91 | 14.65 | 14.61 | 14.85 | 16.30 | 14.99 | 15.23 | 14.61 | 15.04 |
| NiO | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 |
| CaO | 10.94 | 10.88 | 11.01 | 10.80 | 10.45 | 10.85 | 10.84 | 10.69 | 11.08 | 11.23 | 11.03 | 11.10 | 11.60 | 11.55 | 11.06 | 11.55 | 11.35 | 11.32 |
| $\mathrm{Na}_{2} \mathrm{O}$ | 1.69 | 1.64 | 1.61 | 1.79 | 1.73 | 1.75 | 1.49 | 1.62 | 1.45 | 1.57 | 0.66 | 0.69 | 0.65 | 0.45 | 0.68 | 0.67 | 0.81 | 0.70 |
| $\mathrm{K}_{2} \mathrm{O}$ | 0.57 | 0.49 | 0.64 | 1.02 | 0.94 | 0.71 | 0.81 | 0.96 | 0.58 | 0.51 | 0.13 | 0.14 | 0.22 | 0.13 | 0.16 | 0.16 | 0.22 | 0.27 |
| Cl | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 |
| $\mathrm{H}_{2} \mathrm{O}$ | 2.06 | 2.05 | 2.07 | 2.06 | 2.05 | 2.04 | 2.01 | 2.01 | 2.05 | 2.06 | 2.10 | 2.09 | 2.09 | 2.13 | 2.09 | 2.11 | 2.10 | 2.10 |


|  | 100.24 | 100.03 | 100.66 | 100.90 | 100.76 | 100.16 | 99.96 | 100.08 | 99.33 | 100.41 | 100.67 | 99.84 | 100.03 | 101.21 | 99.93 | 100.73 | 100.50 | 100.33 |
| :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- |


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| Mineral：amphibole |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Kiru amphibolite 46／02／01 |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| Sample | p19s01p05 | p19s01p06 | p19s01p07 | p19s02p11 | p19s02p13 | p19s03p01 | p19s03p02 | p19s03p04 | p19s03p09 | p19s03p12 | p19s03p17 | p19s03p18 | p19s04p08 | p19s04p10 | p19s04p17 | p19s04p18 | p19s05p04 | p19s05p05 | p19s05p08 |
| Site | s01 | s01 | s01 | s02 | s02 | s03 | s03 | s03 | s03 | s03 | s03 | s03 | s04 | s04 | s04 | s04 | s05 | s05 | s05 |
| nr ． | 5 | 6 | 7 | 22 | 24 | 31 | 32 | 34 | 39 | 42 | 47 | 48 | 68 | 70 | 77 | 78 | 84 | 85 | 89 |
| note | amp in contact to epi and plg |  |  | in contact to plg |  | amp in contact to chl |  |  | in contact to chl |  | in contact to plg |  | in contact to sph |  | in contact to plg |  | amp in contact to plg |  |  |
| $\mathrm{SiO}_{2}$ | 42.51 | 42.74 | 42.05 | 43.21 | 43.43 | 42.19 | 42.61 | 42.09 | 43.19 | 42.76 | 43.35 | 44.29 | 43.11 | 44.15 | 42.77 | 42.80 | 44.21 | 43.19 | 44.08 |
| $\mathrm{TiO}_{2}$ | 0.68 | 0.68 | 0.63 | 0.61 | 0.70 | 0.75 | 0.68 | 0.80 | 0.53 | 0.76 | 0.58 | 0.59 | 0.53 | 0.66 | 0.78 | 0.75 | 0.55 | 0.59 | 0.64 |
| $\mathrm{Cr}_{2} \mathrm{O}_{3}$ | 0.02 | 0.04 | 0.00 | 0.03 | 0.02 | 0.04 | 0.01 | 0.03 | 0.00 | 0.03 | 0.05 | 0.05 | 0.03 | 0.04 | 0.04 | 0.03 | 0.02 | 0.03 | 0.00 |
| $\mathrm{Al}_{2} \mathrm{O}_{3}$ | 14.13 | 14.21 | 13.68 | 13.88 | 13.21 | 13.50 | 13.69 | 13.88 | 13.81 | 14.02 | 13.60 | 12.78 | 13.76 | 13.06 | 13.58 | 13.62 | 12.98 | 13.61 | 12.38 |
| $\mathrm{Fe}_{2} \mathrm{O}_{3}$ | 3.26 | 2.26 | 4.89 | 2.41 | 4.02 | 2.74 | 3.79 | 2.29 | 2.97 | 3.38 | 2.84 | 2.86 | 2.57 | 4.92 | 2.44 | 3.13 | 2.43 | 2.40 | 3.26 |
| FeO | 11.37 | 12.01 | 9.97 | 11.31 | 10.25 | 11.85 | 10.70 | 11.93 | 11.04 | 11.13 | 10.92 | 10.51 | 11.41 | 9.12 | 11.70 | 11.30 | 10.58 | 10.85 | 10.61 |
| MnO | 0.35 | 0.35 | 0.34 | 0.32 | 0.33 | 0.32 | 0.36 | 0.38 | 0.35 | 0.31 | 0.37 | 0.36 | 0.36 | 0.33 | 0.35 | 0.38 | 0.33 | 0.35 | 0.35 |
| MgO | 10.78 | 10.62 | 10.89 | 11.01 | 11.38 | 10.65 | 10.87 | 10.41 | 10.94 | 10.74 | 11.28 | 11.67 | 10.70 | 11.36 | 10.73 | 10.79 | 11.52 | 11.20 | 11.75 |
| NiO | 0.00 | 0.00 | 0.03 | 0.03 | 0.00 | 0.00 | 0.00 | 0.02 | 0.00 | 0.01 | 0.00 | 0.06 | 0.00 | 0.00 | 0.05 | 0.05 | 0.01 | 0.02 | 0.07 |
| CaO | 12.03 | 12.09 | 11.74 | 11.92 | 11.88 | 11.85 | 11.77 | 11.91 | 11.72 | 11.76 | 12.07 | 11.91 | 11.76 | 11.10 | 11.90 | 11.89 | 11.88 | 11.87 | 12.00 |
| $\mathrm{Na}_{2} \mathrm{O}$ | 1.60 | 1.58 | 1.46 | 1.66 | 1.38 | 1.72 | 1.50 | 1.46 | 1.59 | 1.57 | 1.63 | 1.48 | 1.68 | 1.50 | 1.52 | 1.50 | 1.37 | 1.63 | 1.44 |
| $\mathrm{K}_{2} \mathrm{O}$ | 1.05 | 1.01 | 1.05 | 0.89 | 1.03 | 1.01 | 1.07 | 1.13 | 1.10 | 1.10 | 0.77 | 0.84 | 0.73 | 0.96 | 1.05 | 1.06 | 0.80 | 0.81 | 0.96 |
| Cl | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 |
| $\mathrm{H}_{2} \mathrm{O}$ | 2.03 | 2.03 | 2.01 | 2.03 | 2.04 | 2.00 | 2.02 | 2.00 | 2.03 | 2.03 | 2.04 | 2.04 | 2.02 | 2.04 | 2.01 | 2.02 | 2.03 | 2.02 | 2.04 |
| Total | 99.80 | 99.63 | 98.74 | 99.31 | 99.67 | 98.64 | 99.08 | 98.32 | 99.27 | 99.59 | 99.51 | 99.43 | 98.65 | 99.24 | 98.92 | 99.32 | 98.71 | 98.56 | 99.57 |
| Si | 6.280 | 6.320 | 6.268 | 6.381 | 6.389 | 6.318 | 6.325 | 6.317 | 6.386 | 6.317 | 6.385 | 6.504 | 6.407 | 6.477 | 6.366 | 6.345 | 6.525 | 6.411 | 6.487 |
| Ti | 0.075 | 0.075 | 0.071 | 0.067 | 0.077 | 0.085 | 0.076 | 0.090 | 0.059 | 0.084 | 0.065 | 0.065 | 0.059 | 0.072 | 0.087 | 0.084 | 0.061 | 0.066 | 0.071 |
| Cr | 0.002 | 0.004 | 0.001 | 0.003 | 0.003 | 0.005 | 0.001 | 0.003 | 0.001 | 0.003 | 0.006 | 0.006 | 0.003 | 0.004 | 0.004 | 0.004 | 0.002 | 0.003 | 0.001 |
| AI | 2.460 | 2.477 | 2.403 | 2.416 | 2.290 | 2.382 | 2.395 | 2.454 | 2.406 | 2.441 | 2.361 | 2.213 | 2.410 | 2.259 | 2.383 | 2.381 | 2.259 | 2.381 | 2.147 |
| $\mathrm{Fe}^{3}$ | 0.363 | 0.252 | 0.548 | 0.268 | 0.445 | 0.309 | 0.423 | 0.258 | 0.330 | 0.376 | 0.315 | 0.316 | 0.288 | 0.543 | 0.273 | 0.349 | 0.270 | 0.268 | 0.361 |
| $\mathrm{Fe}^{2}$ | 1.404 | 1.486 | 1.242 | 1.397 | 1.260 | 1.484 | 1.329 | 1.497 | 1.365 | 1.375 | 1.345 | 1.290 | 1.418 | 1.119 | 1.457 | 1.401 | 1.306 | 1.347 | 1.306 |
| Mn | 0.044 | 0.044 | 0.043 | 0.040 | 0.042 | 0.041 | 0.045 | 0.049 | 0.044 | 0.039 | 0.046 | 0.045 | 0.045 | 0.042 | 0.045 | 0.047 | 0.042 | 0.044 | 0.043 |
| Mg | 2.373 | 2.341 | 2.420 | 2.424 | 2.495 | 2.377 | 2.405 | 2.329 | 2.410 | 2.365 | 2.477 | 2.554 | 2.370 | 2.485 | 2.380 | 2.384 | 2.535 | 2.478 | 2.577 |
| Ni | 0.000 | 0.000 | 0.004 | 0.003 | 0.000 | 0.000 | 0.000 | 0.003 | 0.000 | 0.001 | 0.000 | 0.007 | 0.000 | 0.000 | 0.005 | 0.006 | 0.002 | 0.002 | 0.008 |
| Ca | 1.905 | 1.916 | 1.874 | 1.886 | 1.872 | 1.902 | 1.872 | 1.916 | 1.856 | 1.861 | 1.904 | 1.875 | 1.872 | 1.744 | 1.897 | 1.889 | 1.878 | 1.887 | 1.892 |
| Na | 0.458 | 0.452 | 0.422 | 0.476 | 0.394 | 0.501 | 0.431 | 0.424 | 0.455 | 0.449 | 0.466 | 0.421 | 0.484 | 0.427 | 0.439 | 0.431 | 0.393 | 0.468 | 0.411 |
| K | 0.198 | 0.191 | 0.199 | 0.168 | 0.194 | 0.194 | 0.203 | 0.216 | 0.208 | 0.208 | 0.145 | 0.157 | 0.139 | 0.180 | 0.200 | 0.201 | 0.150 | 0.152 | 0.180 |
| OH | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 |
| $\Sigma$ Cations | 15.561 | 15.560 | 15.496 | 15.530 | 15.460 | 15.596 | 15.506 | 15.555 | 15.518 | 15.518 | 15.514 | 15.453 | 15.494 | 15.351 | 15.536 | 15.520 | 15.420 | 15.508 | 15.484 |
| xMg Fe（tot） | 0.573 | 0.574 | 0.575 | 0.593 | 0.594 | 0.570 | 0.579 | 0.570 | 0.587 | 0.575 | 0.599 | 0.614 | 0.582 | 0.599 | 0.579 | 0.577 | 0.617 | 0.605 | 0.607 |
| Tschermak | 1.254 | 1.204 | 1.362 | 1.203 | 1.280 | 1.184 | 1.298 | 1.213 | 1.240 | 1.304 | 1.197 | 1.168 | 1.226 | 1.428 | 1.200 | 1.246 | 1.177 | 1.194 | 1.137 |
| Edenite | 0.561 | 0.560 | 0.496 | 0.530 | 0.460 | 0.596 | 0.506 | 0.555 | 0.518 | 0.518 | 0.514 | 0.453 | 0.494 | 0.351 | 0.536 | 0.520 | 0.420 | 0.508 | 0.484 |
| Plagioclase | 0.095 | 0.084 | 0.126 | 0.114 | 0.128 | 0.098 | 0.128 | 0.084 | 0.144 | 0.139 | 0.096 | 0.125 | 0.128 | 0.256 | 0.103 | 0.111 | 0.122 | 0.113 | 0.108 |




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 | Tschermak |
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| $\begin{array}{l}\text { Edenite } \\ \text { Plagioclase }\end{array}$ |

Table Appendix 3．5：Representative chemical composition of biotite，muscovite， chlorite．

|  | 守○○ |  | $\left\|\begin{array}{l} \infty \\ \stackrel{\infty}{\Omega} \\ \hline \Omega \end{array}\right\|$ |  | $\overline{\mathrm{O}} \mathrm{~B}$ | － | $\stackrel{\circ}{\infty}$ | ＋ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $\propto$ |  |  | $\left\|\begin{array}{l} \infty \\ \dot{\infty} \\ \dot{\Omega} \end{array}\right\|$ |  | 응웅웅 | $\begin{aligned} & \mathrm{O} \\ & \mathrm{o} \end{aligned}$ | ¢ | 8080 |
|  | 8 甘 8 ले 888 GOO ¢ NOO－ |  | $\left\|\begin{array}{c} 8 \\ 0 \\ 0 \\ 0 \end{array}\right\|$ |  | 응 쿵 Nㅓㅇ | $\begin{aligned} & \mathrm{O} \\ & \mathrm{o} \end{aligned}$ | $\stackrel{N}{\infty}$ | ＋ |
|  |  <br>  | $\stackrel{\varrho}{\circ}$ | $\left\|\begin{array}{c} \infty \\ 0 \\ 0 \\ 0 \end{array}\right\|$ |  <br>  | $\stackrel{0}{\circ} \stackrel{\infty}{\circ} \stackrel{\leftrightarrow}{\circ} \stackrel{L}{\circ}$ | $\begin{aligned} & \mathrm{O} \\ & \mathrm{i} \end{aligned}$ | $\begin{aligned} & \text { 毋 } \\ & \text { O } \\ & \text { ón } \end{aligned}$ | $\begin{aligned} & \infty \\ & \stackrel{\circ}{\mathrm{o}} \end{aligned}$ |
|  | $\text { foo No } 0$ |  | ｜ | os o o in Mo No No No OO | No 웅 | $\begin{aligned} & \mathrm{O} \\ & \text { ì } \end{aligned}$ | $\stackrel{\sim}{0}$ | ＋ |


| Mineral：biotite |  |  |  | Mineral：biotite |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Granulitic Gabbro mylonitic 13／01／0 |  |  |  | Sarangar Gabbro undeformed 16／01／01 |  |  |  |  |  |
| Sample | p01s10bt2 | p01s10bt1 | p01s10bt3 | p13s02bt1 | p13s02bt6 | p13s02bt2 | p13s02bt3 | p13s02bt5 | p13s02bt4 |
| Site | s10 | s10 | s10 | s02 | s02 | s02 | s02 | s02 | s02 |
| note | bt in gt－fracture |  |  | bt in contact to amp |  |  |  |  |  |
| $\mathrm{SiO}_{2}$ | 33.45 | 33.12 | 34.56 | 30.18 | 31.25 | 33.34 | 31.82 | 34.73 | 29.30 |
| $\mathrm{TiO}_{2}$ | 1.80 | 1.59 | 1.76 | 1.59 | 4.26 | 2.25 | 2.02 | 2.12 | 2.08 |
| $\mathrm{Cr}_{2} \mathrm{O}_{3}$ | 0.58 | 0.70 | 0.60 | 0.12 | 0.17 | 0.18 | 0.14 | 0.23 | 0.20 |
| $\mathrm{Al}_{2} \mathrm{O}_{3}$ | 17.22 | 16.37 | 17.44 | 18.15 | 16.68 | 17.87 | 18.51 | 18.56 | 18.59 |
| $\mathrm{Fe}_{2} \mathrm{O}_{3}$ | 3.08 | 0.00 | 0.00 | 25.95 | 10.84 | 12.67 | 17.92 | 0.80 | 24.43 |
| FeO | 14.57 | 15.72 | 16.54 | 0.00 | 10.42 | 9.13 | 5.08 | 16.68 | 0.00 |
| MnO | 0.04 | 0.05 | 0.02 | 0.21 | 0.16 | 0.19 | 0.12 | 0.11 | 0.17 |
| MgO | 12.71 | 11.92 | 12.74 | 14.20 | 13.07 | 13.03 | 13.62 | 12.57 | 14.86 |
| NiO |  |  |  |  |  |  |  |  |  |
| CaO | 0.14 | 0.25 | 0.09 | 0.54 | 2.91 | 0.50 | 0.34 | 0.09 | 0.96 |
| $\mathrm{Na}_{2} \mathrm{O}$ | 0.18 | 0.21 | 0.25 | 0.08 | 0.08 | 0.11 | 0.08 | 0.27 | 0.11 |
| $\mathrm{K}_{2} \mathrm{O}$ | 6.62 | 6.76 | 7.81 | 1.43 | 0.36 | 3.29 | 2.97 | 6.94 | 0.40 |
| $\mathrm{H}_{2} \mathrm{O}$ | 3.81 | 3.65 | 3.84 | 4.02 | 3.89 | 3.98 | 4.00 | 3.93 | 4.01 |
| Total | 94.21 | 90.33 | 95.66 | 96.47 | 94.07 | 96.54 | 96.62 | 97.02 | 95.11 |
| Si | 2.629 | 2.724 | 2.695 | 2.254 | 2.411 | 2.510 | 2.385 | 2.649 | 2.191 |
| Ti | 0.107 | 0.098 | 0.103 | 0.090 | 0.247 | 0.127 | 0.114 | 0.122 | 0.117 |
| Cr | 0.036 | 0.045 | 0.037 | 0.007 | 0.010 | 0.011 | 0.009 | 0.014 | 0.012 |
| Al | 1.596 | 1.587 | 1.603 | 1.597 | 1.516 | 1.586 | 1.635 | 1.669 | 1.638 |
| $\mathrm{Fe}^{3}$ | 0.182 | 0.000 | 0.000 | 1.458 | 0.629 | 0.718 | 1.011 | 0.046 | 1.375 |
| $\mathrm{Fe}^{2}$ | 0.958 | 1.081 | 1.079 | 0.000 | 0.672 | 0.575 | 0.318 | 1.064 | 0.000 |
| Mn | 0.003 | 0.004 | 0.002 | 0.014 | 0.010 | 0.012 | 0.008 | 0.007 | 0.011 |
| Mg | 1.490 | 1.461 | 1.481 | 1.581 | 1.503 | 1.462 | 1.522 | 1.429 | 1.656 |
| Ni |  |  |  |  |  |  |  |  |  |
| Ca | 0.012 | 0.022 | 0.008 | 0.043 | 0.241 | 0.040 | 0.027 | 0.007 | 0.077 |
| Na | 0.028 | 0.033 | 0.037 | 0.011 | 0.012 | 0.016 | 0.011 | 0.040 | 0.015 |
| K | 0.664 | 0.709 | 0.777 | 0.136 | 0.035 | 0.316 | 0.284 | 0.675 | 0.038 |
| OH |  |  |  |  |  |  |  |  |  |
| H | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 | 2.000 |
| $\Sigma$ Cations | 7.703 | 7.764 | 7.822 | 7.190 | 7.287 | 7.372 | 7.322 | 7.722 | 7.130 |
| xMg Fe（tot） | 0.566 | 0.575 | 0.579 | 0.520 | 0.536 | 0.531 | 0.534 | 0.563 | 0.546 |
| biotite calculated with＂Norm＂program by P．Ulmer using：Mica Norm BIO（Standard） CATIONS calculated on the bases of 11 oxygens 7 cations $+\mathrm{Na}+\mathrm{K}+\mathrm{Ca}$ muscovite calculated with＂Norm＂program by P．Ulmer using：Mica Norm MUS（Standard） CATIONS calculated on the bases of 11 oxygens and 6 cations $+\mathrm{Na}+\mathrm{K}+\mathrm{Ca}$ chlorite calculated with＂Norm＂program by P．Ulmer using：General OXIDE Norm all Fe（II＋） CATIONS calculated on the basis of 18 ．Oxygens |  |  |  |  |  |  |  |  |  |
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N $\begin{array}{rr}29.88 & 29.73 \\ 0.02 & 0.17\end{array}$ $\begin{array}{rr}0.02 & 0.17 \\ 0.05 & 0.00 \\ 12.73 & 17.51\end{array}$ $\begin{array}{rr}0.05 & 17.51 \\ 12.73 & 17.51 \\ 0.00 & 0.00\end{array}$ $\begin{array}{rr}0.00 & \\ 28.96 & 23.43 \\ 0.14 & 0.23\end{array}$ $\begin{array}{rr}0.14 & 0.23 \\ 13.16 & 15.45 \\ 0.00 & 0.04\end{array}$ $\begin{array}{ll}0.00 & 0.04 \\ 0.31 & 0.31\end{array}$ $\begin{array}{ll}0.31 & 0.31 \\ 0.03 & 0.02 \\ 0.03 & 0.37 \\ 2.23 & 2.37\end{array}$ $\begin{array}{ll}0.03 & 0.37 \\ 2.23 & 2.37\end{array}$ ع9'68 ${ }^{\text {tG L8 }}$ 87.54 biotite calculated with "Norm" program by P. Ulimer using: Mica Norm BIO (Standard)
CATIONS calculated on the bases of 11 oxygens 7 cations $+\mathrm{Na}+\mathrm{K}+\mathrm{Ca}$
 chlorite calculated with "Norm" program by P. UImer using: General OXIDE Norm all Fe (ll+) chlorite calculated with "Norm" program oy
CATIONS calculated on the basis of 18 . Oxygens
$\begin{array}{ll}4.010 & 3.768 \\ 0.002 & 0.016 \\ 0.005 & 0.000\end{array}$ $\begin{array}{ll}0.002 & \\ 0.005 & 0.000 \\ 2.014 & 2.616\end{array}$ $\begin{array}{ll}2.014 & 0.60 \\ 0.000 & 0.000 \\ 3.250 & 2.484\end{array}$ $\begin{array}{ll}0.0250 & 2.484 \\ 0.025\end{array}$ $\begin{array}{ll}0.016 & \\ 2.633 & 2.919 \\ 0.000 & 0.004\end{array}$ $\begin{array}{ll}0.000 & 0.004 \\ 0.044 & 0.042 \\ & 0.005\end{array}$ 0
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 $\begin{array}{lrrrr} & & & & \\ \mathrm{SiO}_{2} & 24.61 & 23.59 & 23.93 & 24.97 \\ \mathrm{TiO}_{2} & 0.09 & 0.06 & 0.03 & 0.06 \\ \mathrm{Cr}_{2} \mathrm{O}_{3} & 0.00 & 0.05 & 0.00 & 0.02 \\ \mathrm{Al}_{2} \mathrm{O}_{3} & 20.88 & 20.10 & 21.03 & 21.33 \\ \mathrm{Fe}_{2} \mathrm{O}_{3} & 0.00 & 0.00 & 0.00 & 0.00 \\ \mathrm{FeO} & 25.49 & 25.56 & 26.31 & 26.00 \\ \mathrm{MnO} & 0.28 & 0.19 & 0.27 & 0.26 \\ \mathrm{MgO} & 14.85 & 14.31 & 14.34 & 15.12 \\ \mathrm{NiO} & 0.00 & 0.00 & 0.03 & 0.00 \\ \mathrm{CaO} & 0.13 & 0.11 & 0.09 & 0.06 \\ \mathrm{Na}_{2} \mathrm{O} & 0.02 & 0.05 & 0.07 & 0.06 \\ \mathrm{~K}_{2} \mathrm{O} & 0.03 & 0.03 & 0.04 & 0.00 \\ \mathrm{H}_{2} \mathrm{O} & 2.30 & 2.22 & 2.27 & 2.34 \\ & & & & \end{array}$ | 88.68 | 86.26 | 88.41 | 90.23 |
| :--- | :--- | :--- | :--- | :--- |

 $\begin{array}{lllll}\Sigma \text { Cations } & 14.174 & 14.217 & 14.218 & 14.180\end{array}$
$\Sigma$ Cations
$\times \mathrm{xMg}$ Fe(tot)

Table Appendix 3．6：Representative chemical composition of epidote and sphene．

|  |  <br>  <br>  <br>  <br>  <br>  <br>  | － |  mo no <br>  <br>  <br>  <br>  लö No <br>  Noonoo <br>  |  <br>  <br>  <br>  <br>  00 －0．o． <br>  0 <br>  |  |  |
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| Mineral：epidote |  |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Granulitic Gabbro mylonitic 94／04／01 |  |  |  |  |  |  |  |  |
| Sample | p04501p07 | p04802206 | p04s03918 | p04s04pro | p04s04pro | po4s09pro | 55p12 | po4s50520 |
| Site | s01 | s02 | s03 | s04 | 504 | 504 | s05 |  |
| nr． | 7 |  | 43 | profile through matrix |  |  | 12 | 22 |
| note | to plg | to gt | epiats03 |  |  |  | ｜epi in contact togt |  |
| $\mathrm{SiO}_{2}$ | 38.64 | 38.60 | 38.11 | 38.47 | 37.60 | 38.82 | 38.46 | 38.49 |
| $\mathrm{TiO}_{2}$ | 0.07 | 0.04 | 0.07 | 0.17 | 0.09 | 0.02 | 0.04 | 0.05 |
| $\mathrm{Cr}_{2} \mathrm{O}_{3}$ | 0.00 | 0.00 | 0.00 | 0.03 | 0.03 | 0.00 | 0.00 | 0.00 |
| $\mathrm{Al}_{2} \mathrm{O}_{3}$ | 26.35 | 26.61 | 26.83 | 29.00 | 28.11 | 28.42 | 27.07 | 26.01 |
| $\mathrm{Fe}_{2} \mathrm{O}_{3}$ | 9.26 | 9.28 | 9.19 | 6.07 | 7.42 | 6.90 | 8.97 | 10.25 |
| $\begin{array}{llllllllll}\text { FeO } & 0.00 & 0.00 & 0.00 & 0.00 & 0.00 & 0.00 & 0.00 & 0.00 \\ \mathrm{MnO} & & & \end{array}$ |  |  |  |  |  |  |  |  |
|  |  |  |  |  |  |  |  |  |  |  |  |
| $\mathrm{Mn}_{2} \mathrm{O}_{3}$ | 0.02 | 0.08 | 0.11 | 0.07 | 0.09 | 0.03 | 0.04 | 0.07 |
| Mgo | 0.07 | 0.12 | 0.11 | 0.15 | 0.10 | 0.10 | 0.00 | 0.15 |
| Nio | 0.04 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.01 | 0.00 |
| CaO | 23.68 | 23.24 | 23.48 | 23.10 | 23.31 | 23.41 | 23.44 | 22.67 |
| $\mathrm{Na}_{2} \mathrm{O}$ | 0.01 | 0.02 | 0.03 | 0.02 | 0.04 | 0.06 | 0.00 | 0.03 |
| $\mathrm{K}_{2} \mathrm{O}$ | 0.04 | 0.02 | 0.00 | 0.02 | 0.01 | 0.01 | 0.01 | 0.00 |
| $\mathrm{H}_{2} \mathrm{O}$ | 1.92 | 1.92 | 1.91 | 1.92 | 1.90 | 1.93 | 1.92 | 1.91 |
| Total | 100.10 | 99.93 | 99.83 | 99.02 | 98.69 | 99.71 | 99.97 | 99.64 |
| Si | 3.019 | 3.016 | 2.985 | 2.997 | 2.962 | 3.014 | 3.002 | 3.022 |
| Ti | 0.004 | 0.003 | 0.004 | 0.010 | 0.005 | 0.001 | 0.002 | 0.003 |
| Cr | 0.000 | 0.000 | 0.000 | 0.002 | 0.002 | 0.000 | 0.000 | 0.000 |
| Al | 2.426 | 2.450 | 2.476 | 2.663 | 2.609 | 2.600 | 2.490 | 2.406 |
| $\mathrm{Fe}^{3}$ | 0.545 | 0.545 | 0.542 | 0.356 | 0.440 | 0.403 | 0.527 | 0.606 |
| $\mathrm{Fe}^{2}$ | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 |
| Mn |  |  |  |  |  |  |  |  |
| Mn ${ }^{3+}$ | 0.001 | 0.005 | 0.006 | 0.004 | 0.006 | 0.002 | 0.003 | 0.005 |
| Mg | 0.008 | 0.014 | 0.013 | 0.018 | 0.012 | 0.011 | 0.000 | 0.018 |
| Ni | 0.003 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.001 | 0.000 |
| Ca | 1.982 | 1.945 | 1.970 | 1.928 | 1.967 | 1.947 | 1.960 | 1.906 |
| Na | 0.002 | 0.003 | 0.005 | 0.004 | 0.006 | 0.010 | 0.001 | 0.004 |
| K | 0.004 | 0.002 | 0.000 | 0.002 | 0.001 | 0.001 | 0.001 | 0.001 |
| О H | 1.000 | 1.000 | 1.000 | 1.000 | 1.000 | 1.000 | 1.000 | 1.000 |
| $\Sigma$ Cations | 18.006 | 18.016 | 17.999 | 18.017 | 17.991 | 18.012 | 18.013 | 18.030 |
| ZoisiseEpidote | 0.430 | 0.453 | 0.465 | 0.672 | 0.578 | 0.601 | 0.492 | 0.409 |
|  | 0.545 | 0.545 | 0.535 | ${ }^{0.328}$ | 0.422 | 0.399 | 0.508 | 0.591 |
|  |  |  |  | Ining | e Norm |  |  |  |


| Mineral: epidote |  |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Sarangar Gabbro undeformed 16/01/01 |  |  |  |  |  |  |  |  |
| p13s3ahb8 | p13s3asy1 | p13s3asy | p13s3asy6 | p13s3asy7 | p13s3asy9 | p13s3cuk2 | p13s3csy1 | p13s3csy2 |
| s3a | s3a | s3a | s3a | s3a | s3a | s3c | s3c | s3c |
| 100 | 101 | 103 | 106 | 107 | 109 | 127 | 128 | 129 |
| epi between amp and plg (sympl) in contact to qtz. |  |  |  |  | \|epi between amp and plg (sympl) |  |  |  |
| 37.89 | 39.04 | 39.47 | 39.13 | 39.19 | 39.42 | 39.40 | 39.19 | 39.39 |
| 0.07 | 0.06 | 0.03 | 0.06 | 0.06 | 0.03 | 0.01 | 0.00 | 0.00 |
| 0.00 | 0.05 | 0.00 | 0.01 | 0.02 | 0.04 | 0.02 | 0.00 | 0.02 |
| 27.42 | 27.67 | 28.32 | 28.35 | 28.44 | 28.70 | 29.41 | 28.44 | 28.02 |
| 8.60 | 8.33 | 7.21 | 7.23 | 7.37 | 6.66 | 6.30 | 7.47 | 7.71 |
| 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 |
| 0.23 | 0.25 | 0.25 | 0.19 | 0.27 | 0.24 | 0.23 | 0.18 | 0.16 |
| 0.08 | 0.08 | 0.06 | 0.06 | 0.06 | 0.05 | 0.05 | 0.05 | 0.05 |
| 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 |
| 22.63 | 23.13 | 23.25 | 23.28 | 23.16 | 23.06 | 23.49 | 23.16 | 23.00 |
| 0.04 | 0.06 | 0.04 | 0.04 | 0.09 | 0.08 | 0.01 | 0.00 | 0.06 |
| 0.03 | 0.02 | 0.04 | 0.04 | 0.07 | 0.04 | 0.01 | 0.00 | 0.04 |
| 1.90 | 1.94 | 1.95 | 1.94 | 1.95 | 1.95 | 1.96 | 1.95 | 1.94 |
| 98.86 | 100.61 | 100.59 | 100.31 | 100.65 | 100.23 | 100.89 | 100.42 | 100.39 |
| 2.983 | 3.016 | 3.036 | 3.020 | 3.015 | 3.036 | 3.013 | 3.021 | 3.039 |
| 0.004 | 0.004 | 0.002 | 0.003 | 0.004 | 0.002 | 0.001 | 0.000 | 0.000 |
| 0.000 | 0.003 | 0.000 | 0.001 | 0.001 | 0.002 | 0.001 | 0.000 | 0.001 |
| 2.545 | 2.520 | 2.567 | 2.579 | 2.580 | 2.605 | 2.650 | 2.584 | 2.548 |
| 0.509 | 0.484 | 0.417 | 0.420 | 0.427 | 0.386 | 0.363 | 0.433 | 0.448 |
| 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 |
| 0.014 | 0.015 | 0.015 | 0.011 | 0.016 | 0.014 | 0.014 | 0.010 | 0.010 |
| 0.009 | 0.009 | 0.007 | 0.006 | 0.007 | 0.006 | 0.006 | 0.005 | 0.006 |
| 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 |
| 1.909 | 1.914 | 1.916 | 1.925 | 1.909 | 1.903 | 1.924 | 1.912 | 1.901 |
| 0.007 | 0.009 | 0.005 | 0.006 | 0.013 | 0.011 | 0.002 | 0.001 | 0.009 |
| 0.003 | 0.002 | 0.003 | 0.004 | 0.007 | 0.004 | 0.001 | 0.000 | 0.004 |
| 1.000 | 1.000 | 1.000 | 1.000 | 1.000 | 1.000 | 1.000 | 1.000 | 1.000 |
| 18.017 | 18.025 | 18.032 | 18.024 | 18.021 | 18.033 | 18.026 | 18.034 | 18.036 |
| 0.533 | 0.526 | 0.569 | 0.583 | 0.584 | 0.608 | 0.652 | 0.584 | 0.549 |
| 0.467 | 0.474 | 0.417 | 0.417 | 0.416 | 0.386 | 0.348 | 0.416 | 0.448 |


| Mineral: epidote |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Sarangar Gabbro mylonitic 90/02/01 |  |  |  |  |  |  |
| Sample | p07a01p08 | p07a01p09 | p07a01p10 | p07a02p15 | p07a03p08 | p07a03p23 |
| Site | a01 | a01 | a01 | a02 | a03 | a03 |
| nr. | 8 | 9 | 10 | 55 | 34 | 49 |
| note | in contact to plg and amp |  |  | \|epi at amp| | to amp |  |
| $\mathrm{SiO}_{2}$ | 37.61 | 37.51 | 37.99 | 36.27 | 36.99 | 37.61 |
| $\mathrm{TiO}_{2}$ | 0.17 | 0.16 | 0.22 | 0.15 | 0.13 | 0.16 |
| $\mathrm{Cr}_{2} \mathrm{O}_{3}$ | 0.05 | 0.04 | 0.06 | 0.03 | 0.02 | 0.00 |
| $\mathrm{Al}_{2} \mathrm{O}_{3}$ | 27.51 | 26.14 | 27.39 | 25.46 | 27.29 | 26.49 |
| $\mathrm{Fe}_{2} \mathrm{O}_{3}$ | 8.78 | 8.38 | 8.09 | 8.24 | 8.50 | 8.73 |
| FeO | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 |
| MnO |  |  |  |  |  |  |
| $\mathrm{Mn}_{2} \mathrm{O}_{3}$ | 0.09 | 0.21 | 0.17 | 0.17 | 0.15 | 0.14 |
| MgO | 0.10 | 0.10 | 0.12 | 0.09 | 0.09 | 0.11 |
| NiO | 0.00 | 0.00 | 0.00 | 0.01 | 0.00 | 0.04 |
| CaO | 24.32 | 24.18 | 24.72 | 23.19 | 23.82 | 24.18 |
| $\mathrm{Na}_{2} \mathrm{O}$ | 0.02 | 0.00 | 0.03 | 0.00 | 0.00 | 0.00 |
| $\mathrm{K}_{2} \mathrm{O}$ | 0.00 | 0.00 | 0.01 | 0.00 | 0.02 | 0.01 |
| $\mathrm{H}_{2} \mathrm{O}$ | 1.92 | 1.89 | 1.93 | 1.83 | 1.89 | 1.90 |
| Total | 100.57 | 98.60 | 100.70 | 95.43 | 98.90 | 99.36 |
| Si | 2.931 | 2.981 | 2.953 | 2.976 | 2.929 | 2.968 |
| Ti | 0.010 | 0.010 | 0.013 | 0.009 | 0.008 | 0.010 |
| Cr | 0.003 | 0.002 | 0.004 | 0.002 | 0.001 | 0.000 |
| Al | 2.527 | 2.449 | 2.509 | 2.463 | 2.546 | 2.463 |
| $\mathrm{Fe}^{3}$ | 0.515 | 0.501 | 0.473 | 0.509 | 0.507 | 0.518 |
| $\mathrm{Fe}^{2}$ | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 |
| Mn |  |  |  |  |  |  |
| $\mathrm{Mn}^{3+}$ | 0.006 | 0.013 | 0.010 | 0.011 | 0.009 | 0.009 |
| Mg | 0.012 | 0.012 | 0.014 | 0.011 | 0.011 | 0.013 |
| Ni | 0.000 | 0.000 | 0.000 | 0.001 | 0.000 | 0.002 |
| Ca | 2.030 | 2.059 | 2.058 | 2.040 | 2.021 | 2.044 |
| Na | 0.004 | 0.000 | 0.005 | 0.001 | 0.001 | 0.000 |
| K | 0.000 | 0.001 | 0.001 | 0.000 | 0.002 | 0.001 |
| OH | 1.000 | 1.000 | 1.000 | 1.000 | 1.000 | 1.000 |
| $\Sigma$ Cations | 17.964 | 17.973 | 17.961 | 17.978 | 17.967 | 17.972 |
| Zoisite | 0.471 | 0.442 | 0.478 | 0.451 | 0.484 | 0.441 |
| Epidote | 0.515 | 0.501 | 0.473 | 0.509 | 0.507 | 0.518 |


| Mineral: sphene |  |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Kiru amphibolite 46/02/01 |  |  |  |  |  |  |  |  |
| Sample | p19s02p08 | p19s02p09 | p19s02p10 | p19s04p01 | p19s04p02 | p19s04p03 | p19s04p04 | p19s04p05 |
| Site | s02 | s02 | s02 | s04 | s04 | s04 | s04 | s04 |
| nr. | 19 | 20 | 21 | 61 | 62 | 63 | 64 | 65 |
| note | new grown? |  |  | sph in contact to amp |  |  |  |  |
| $\mathrm{SiO}_{2}$ | 29.75 | 29.93 | 29.68 | 29.55 | 29.84 | 29.72 | 28.90 | 29.14 |
| $\mathrm{TiO}_{2}$ | 35.81 | 36.04 | 36.60 | 35.08 | 34.76 | 35.55 | 35.05 | 36.18 |
| $\mathrm{Cr}_{2} \mathrm{O}_{3}$ | 0.03 | 0.03 | 0.02 | 0.07 | 0.05 | 0.02 | 0.01 | 0.02 |
| $\mathrm{Al}_{2} \mathrm{O}_{3}$ | 1.04 | 1.17 | 1.22 | 0.66 | 0.83 | 0.93 | 0.69 | 0.66 |
| $\mathrm{Fe}_{2} \mathrm{O}_{3}$ | 0.96 | 0.66 | 0.86 | 0.55 | 0.39 | 0.66 | 0.51 | 0.60 |
| FeO | 0.00 | 0.00 | 0.00 | 0.02 | 0.17 | 0.00 | 0.00 | 0.00 |
| MnO | 0.10 | 0.07 | 0.08 | 0.04 | 0.10 | 0.07 | 0.11 | 0.11 |
| $\mathrm{Mn}_{2} \mathrm{O}_{3}$ |  |  |  |  |  |  |  |  |
| MgO | 0.03 | 0.02 | 0.00 | 0.00 | 0.00 | 0.22 | 0.02 | 0.00 |
| NiO | 0.02 | 0.03 | 0.02 | 0.00 | 0.00 | 0.01 | 0.04 | 0.04 |
| CaO | 28.25 | 28.32 | 28.45 | 26.42 | 26.33 | 26.88 | 26.06 | 28.23 |
| $\mathrm{Na}_{2} \mathrm{O}$ | 0.02 | 0.00 | 0.02 | 0.10 | 0.07 | 0.09 | 0.12 | 0.39 |
| $\mathrm{K}_{2} \mathrm{O}$ | 0.00 | 0.02 | 0.00 | 0.03 | 0.06 | 0.03 | 0.05 | 0.07 |
| $\mathrm{H}_{2} \mathrm{O}$ | 0.30 | 0.29 | 0.32 | 0.22 | 0.23 | 0.27 | 0.23 | 0.31 |
| Total | 96.31 | 96.59 | 97.28 | 92.73 | 92.83 | 94.47 | 91.81 | 95.77 |
| Si | 1.002 | 1.004 | 0.990 | 1.034 | 1.042 | 1.019 | 1.022 | 0.984 |
| Ti | 0.907 | 0.909 | 0.918 | 0.923 | 0.913 | 0.917 | 0.932 | 0.919 |
| Cr | 0.001 | 0.001 | 0.001 | 0.002 | 0.001 | 0.001 | 0.000 | 0.001 |
| Al | 0.041 | 0.047 | 0.048 | 0.027 | 0.034 | 0.038 | 0.029 | 0.026 |
| $\mathrm{Fe}^{3}$ | 0.024 | 0.017 | 0.022 | 0.015 | 0.010 | 0.017 | 0.014 | 0.015 |
| $\mathrm{Fe}^{2}$ | 0.000 | 0.000 | 0.000 | 0.001 | 0.005 | 0.000 | 0.000 | 0.000 |
| Mn | 0.003 | 0.002 | 0.002 | 0.001 | 0.003 | 0.002 | 0.003 | 0.003 |
| $\mathrm{Mn}^{3+}$ |  |  |  |  |  |  |  |  |
| Mg | 0.002 | 0.001 | 0.000 | 0.000 | 0.000 | 0.012 | 0.001 | 0.000 |
| Ni | 0.001 | 0.001 | 0.001 | 0.000 | 0.000 | 0.000 | 0.001 | 0.001 |
| Ca | 1.019 | 1.018 | 1.017 | 0.990 | 0.985 | 0.988 | 0.987 | 1.022 |
| Na | 0.001 | 0.000 | 0.001 | 0.007 | 0.005 | 0.006 | 0.009 | 0.026 |
| K | 0.000 | 0.001 | 0.000 | 0.001 | 0.003 | 0.001 | 0.002 | 0.003 |
| OH | 0.068 | 0.065 | 0.072 | 0.052 | 0.053 | 0.063 | 0.053 | 0.071 |
| $\Sigma$ Cations | 3.068 | 3.065 | 3.072 | 3.052 | 3.053 | 3.063 | 3.054 | 3.071 |

[^2]| Mineral: epidote |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Kiru amphibolite 46/02/01 |  |  |  |  |  |  |  |
| Sample | p19s01p01 | p19s01p02 | p19s01p03 | p19s02p01 | p19s02p02 | p19s02p04 | p19s02p05 |
| Site | s01 | s01 | s01 | s02 | s02 | s02 | s02 |
| nr. | 1 | 2 | 3 | 12 | 13 | 15 | 16 |
| note | epi in contact to amp |  |  | epi in contact to plg and qtz |  |  |  |
| $\mathrm{SiO}_{2}$ | 37.58 | 37.47 | 37.75 | 38.16 | 37.87 | 38.21 | 38.11 |
| $\mathrm{TiO}_{2}$ | 0.11 | 0.14 | 0.14 | 0.05 | 0.18 | 0.17 | 0.15 |
| $\mathrm{Cr}_{2} \mathrm{O}_{3}$ | 0.02 | 0.01 | 0.03 | 0.02 | 0.00 | 0.02 | 0.02 |
| $\mathrm{Al}_{2} \mathrm{O}_{3}$ | 27.57 | 27.64 | 27.60 | 27.75 | 27.24 | 27.76 | 27.59 |
| $\mathrm{Fe}_{2} \mathrm{O}_{3}$ | 8.18 | 8.05 | 8.18 | 8.16 | 8.45 | 8.36 | 8.15 |
| FeO | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 |
| MnO |  |  |  |  |  |  |  |
| $\mathrm{Mn}_{2} \mathrm{O}_{3}$ | 0.23 | 0.28 | 0.25 | 0.19 | 0.17 | 0.23 | 0.26 |
| MgO | 0.07 | 0.07 | 0.06 | 0.02 | 0.09 | 0.06 | 0.07 |
| NiO | 0.11 | 0.00 | 0.01 | 0.02 | 0.00 | 0.00 | 0.05 |
| CaO | 24.39 | 24.27 | 24.48 | 23.96 | 24.04 | 24.25 | 23.77 |
| $\mathrm{Na}_{2} \mathrm{O}$ | 0.04 | 0.04 | 0.05 | 0.00 | 0.00 | 0.02 | 0.00 |
| $\mathrm{K}_{2} \mathrm{O}$ | 0.02 | 0.05 | 0.03 | 0.00 | 0.00 | 0.00 | 0.00 |
| $\mathrm{H}_{2} \mathrm{O}$ | 1.92 | 1.91 | 1.92 | 1.93 | 1.92 | 1.94 | 1.92 |
| Total | 100.22 | 99.90 | 100.48 | 100.25 | 99.95 | 100.99 | 100.08 |
| Si | 2.937 | 2.935 | 2.941 | 2.969 | 2.962 | 2.957 | 2.971 |
| Ti | 0.007 | 0.008 | 0.008 | 0.003 | 0.011 | 0.010 | 0.009 |
| Cr | 0.001 | 0.001 | 0.002 | 0.001 | 0.000 | 0.001 | 0.001 |
| Al | 2.539 | 2.551 | 2.534 | 2.545 | 2.511 | 2.532 | 2.534 |
| $\mathrm{Fe}^{3}$ | 0.481 | 0.474 | 0.479 | 0.478 | 0.497 | 0.486 | 0.478 |
| $\mathrm{Fe}^{2}$ | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 |
| Mn |  |  |  |  |  |  |  |
| $\mathrm{Mn}^{3+}$ | 0.014 | 0.017 | 0.015 | 0.011 | 0.010 | 0.014 | 0.016 |
| Mg | 0.009 | 0.009 | 0.007 | 0.003 | 0.011 | 0.007 | 0.008 |
| Ni | 0.007 | 0.000 | 0.001 | 0.001 | 0.000 | 0.000 | 0.003 |
| Ca | 2.042 | 2.036 | 2.044 | 1.998 | 2.015 | 2.010 | 1.985 |
| Na | 0.006 | 0.006 | 0.007 | 0.001 | 0.000 | 0.003 | 0.000 |
| K | 0.002 | 0.005 | 0.003 | 0.000 | 0.000 | 0.001 | 0.001 |
| OH | 1.000 | 1.000 | 1.000 | 1.000 | 1.000 | 1.000 | 1.000 |
| $\Sigma$ Cations | 17.957 | 17.959 | 17.959 | 17.990 | 17.982 | 17.981 | 17.994 |
| Zoisite | 0.484 | 0.495 | 0.485 | 0.519 | 0.484 | 0.499 | 0.515 |
| Epidote | 0.481 | 0.474 | 0.479 | 0.478 | 0.497 | 0.486 | 0.478 |

- ICP-MS and XRF analysis.


## Sample location:

## Sarangar Gabbro, undeformed, GbS (16/01/01) E 73 ${ }^{\circ} 01$ '23' N $35^{\circ} 07{ }^{\prime} 16^{\prime \prime}$

The sample was collected 1.95 km NE of Patan (road distance) along the KKH. The mylonitic part covered by the sample belongs to a ca. 8 cm thick shear zone with a foliation dipping $20^{\circ}$ towards N and a top to S shear sense.

Sarangar Gabbro mylonite, GbS, (90/02/01) E 73 ${ }^{\circ} 01^{\prime} 37^{\prime \prime}$ N $35^{\circ} 07^{\prime} 01^{\prime \prime}$
The sample was collected 4.33 km E of Patan (road distance form Indus bridge) along the Jeep-road Patan - Chor Nala close to Shalkanabad. The gently NE dipping shear zone is ca. 50 cm thick and contains stretched plagioclase veins and numerous rotated plagioclase porphyroclasts. Sense of shear is towards SW.

## Gabbro GbDrHbl (91/03/02) E 7302'49' N 35º7'44'

Sample locality is 5.25 km NE of Patan (road distance). As shown in Fig. 3.16, this gabbro is intruded into hornblendites and was later intruded by the brighter gabbro $(91 / 03 / 03)$. Successive magma injections produced the magmatic breccia.

Gabbro GbDrHbl (91/03/03) E $73^{\circ} 02^{\prime} 49^{\prime \prime}$ N $35^{\circ} 07{ }^{\prime} 44^{\prime \prime}$
See (91/03/02).
Meta-gabbro AmO (71/03/02) E 73 ${ }^{\circ} 12^{\prime} 14^{\prime \prime}$ N 35 ${ }^{\circ} 17^{\prime} 46^{\prime \prime}$
This well foliated, fine graine amphibole rich gabbro was collected 4.8 km N of Dasu (road distance). The metamorphic foliation at the sample locality dips with $60^{\circ}$ towards north.

Meta-diorite, shear zone, DrHbl (RK2) E 7303'32' N 35 ${ }^{\circ} 08^{\prime} 06^{\prime \prime}$
The sample was collected 6.63 km NE of Patan (road distance). As shown in Fig. 3.16, the sample was taken from below a top to S amphibolite facies shear zone (Foliation: $249 / 16$, Linear: $181 / 06$ ). The steep foliation $(335 / 84)$ outside the shear zone is bend by the mylonitic foliation.

## Meta-diorite, undeformed, DrHbl (RK1) E 73 ${ }^{\circ} 03^{\prime} 32^{\prime \prime}$ N 35 ${ }^{\circ} 08^{\prime} 06^{\prime \prime}$

See RK2 for locality. The sample was taken within the shear zone (Fig. 3.16).

## Hornblende diorite HbIDr (35/02/01) E 7306'11' N 35º $\mathbf{0}^{\prime}$ '05'

Sample locality is 2.38 km S of metal bridge at Kiru (road distance). It is a coarse-grained hornblende diorite.
Hornblende diorite HbIDr (91/07/01) E 73 ${ }^{\circ} 06^{\prime} 25^{\prime \prime}$ N $35^{\circ} 10^{\prime} 40^{\prime \prime}$
The sample was collected alongside the KKH 1.1 km S of metal bridge at Kiru (road distance). Detailed description in appendix 4.
Tonalite dike in Kiru Amphibolites AmKi (36/01/01) E 73º6'32' N 35 ${ }^{\circ} \mathbf{1 0}^{\prime} 55^{\prime \prime}$
The dike was sampled 550 m S of metal bridge at Kiru (road distance) (Fig. 3.16). It is a $5-10 \mathrm{~cm}$ thick undeformed tonalite dike (orientation: 292/42) cutting a meta-gabbro of the Kiru amphibolites.

## Granite Gr (00/10/01) E $73^{\circ} 10^{\prime} 44,2^{\prime \prime}$ N $35^{\circ} 14{ }^{\prime} 29,9^{\prime \prime}$

The sample locality is ca. 200 m S of the bridge where the Karakoram Highway ( KKH ) and the Indus River bend sharply towards the E. Detailed description in appendix 4.

## Granite Gr (91/08/01) E $73^{\circ} 08^{\prime} 56^{\prime \prime} \mathrm{N} 35^{\circ} 13^{\prime} 01^{\prime \prime}$

The sample was collected in a ca. 2 m thick granitic sheet 9.8 km S of Dasu (road distance).
Indian plate diorite (00/30/14) E $72^{\circ} 53 ' 51,7^{\prime \prime} \mathrm{N} 35^{\circ} 02^{\prime} 14,5^{\prime \prime}$
The sample was collected 100 m S of Duber Bazar along the KKH. Detailed description in appendix 4.
Volcanoclastic Sediment CH-VcS (92/07/01) E 7301'41' N $35^{\circ} \mathbf{2 7}^{\prime} 43^{\prime \prime}$
Sample locality is in the Kandiah valley ca. 19.3 km (road distance) W of Kandiah river - Indus confluence.

## APPENDIX Chapter 4

## - $\mathrm{U} / \mathrm{Pb}$ dating of zircons

## Analytical techniques for high-precision $\mathrm{U}-\mathrm{Pb}$ ages and initial Hf isotopic compositions

 of zircons:Zircons were air-abraded and washed in distilled acetone and water in an ultrasonic bath. Dissolution in HF-HNO3, chemical separation on anion exchange resin and mass spectrometry using ion counting on a MAT262 mass spectrometer followed standard techniques. The ion counting system was calibrated by repeated analysis of the NBS 982 standard. The procedural Pb blank was estimated at $1.5 \pm 0.75 \mathrm{pg}$; common lead in excess of this amount was corrected with depleted mantle isotopic compositions. The Hf fraction was isolated using Eichrom ${ }^{\mathrm{TM}} \mathrm{Ln}$-spec resin, and in static mode on a NuPlasma ${ }^{\mathrm{TM}}$ multi-collector ICP-MS using a MCN6000 nebulizer for sample introduction. Zircons are commonly characterized by extremely low ${ }^{176} \mathrm{Lu} /{ }^{177} \mathrm{Hf}$ of less than 0.005 ; Hf isotopic values were therfore not corrected for in-situ radiogenic ingrowth from ${ }^{176} \mathrm{Lu}$, because corrections for the 100 Ma old zircons are within the analytical uncertainty of the measured ${ }^{176} \mathrm{Hf} /{ }^{177} \mathrm{Hf}$ ratios. The Hf isotopic ratios were corrected for mass fractionation using a ${ }^{179} \mathrm{Hf} /{ }^{177} \mathrm{Hf}$ value of 0.7325 and normalized to ${ }^{176} \mathrm{Hf} /{ }^{177} \mathrm{Hf}=$ of 0.282160 of the JMC- 475 standard (BlichertToft et al., 1997). Mean ages and mean Hf isotopic values are given at the $95 \%$ confidence level.

## Sample location and description:

## Sarangar Gabbro (sample number 16/02/01) E 73 ${ }^{\circ} 01^{\prime} 41,7^{\prime \prime} \mathbf{N} 35^{\circ} 07{ }^{\prime} 15,8^{\prime \prime}$

The sample was taken ca. 2 km E of Patan next to the Karakoram Highway in an undeformed part of the Sarangar Gabbro. The gabbro shows spectacular shear zones anastomosing around undeformed and coarse grained, weakly deformed with well preserved magmatic fabric material (Arbaret et al,. 2000). The mineralogical composition within the main metamorphic fabric is dominantly pargasite and tschermakite, hypidioblastic clinopyroxene, plagioclase, rutile, garnet, epidote, quartz and zoisite. Some of the plagioclase shows anorthite richer cores as relicts of the magmatic origin. Albite twinning lamellae are gently bent. The amount of rutile is explained by the release of Ti during recrystallization of Ti-pargasite. Garnet is metamorphic intergrown with rutile along plagioclase rich domains. Large epidote grains are magmatic. Zoisite is growing as metamorphic phase within plagioclase cluster. Clinopyroxene shows breakdown in amphibole and quartz.


Fig. A4.1: Representative zircons used for U/Pb - dating of the Sarangar gabbro (16/02/01): a,c) backscattered electron image (BSE), b,d) cathodoluminescence image (CL).

## Granite (00/10/01) E 73 ${ }^{\circ} 10{ }^{\prime} \mathbf{4 4}, 2^{\prime \prime}$ N $35^{\circ} 144^{\prime} 29,9^{\prime \prime}$

The sample locality is ca. 200 m S of the bridge where the Karakoram Highway (KKH) and the Indus River bend sharply towards the E. The coarse grained granite is composed of quartz, plagioclase, muscovite, garnet, idiomorph epidote, pyrite, chlorite, apatite and zircon. Deformation is concentrated in narrow zones and at the contacts. Quartz is dynamically recrystallized. Undeformed quartz occurs as inclusions in magmatic epidote. Albite rich plagioclase shows weak normal zoning. Twinning lamellae are deformed. Orthoclase is strongly seritisized and shows pertitic exsolution.


Fig. A4.2: Representative zircons used for U/Pb - dating of a granite (00/10/01): a,c) backscattered electron image (BSE), b,d) cathodoluminescence image (CL).

## Diorite (91/07/01) E $\mathbf{7 3}^{\circ} \mathbf{0 6} 6^{\prime 25,5} 5^{\prime \prime}$ N $35^{\circ} 10^{\prime} 40,3^{\prime \prime}$

The sample was collected alongside the KKH ca. 1 km S of Kiru. It is a coarse-grained hornblende diorite composed of plagioclase, greenish hornblende, quartz and epidote, chlorite and titanite. The metamorphic fabric has overprinted the magmatic fabric. Subhedral plagioclase shows strong saussuritisation. Foliation is defined by the orientation of metamorphic hornblende. Cotectic fabric within the epidote is preserved as primary magmatic feature.


Fig. A4.3: Representative zircons used for $U / P b$ - dating of a diorite (91/07/01): a,c) backscattered electron image (BSE), b,d) cathodoluminescence image (CL).

## Kyanite bearing dyke (38/02/01) E 72 ${ }^{\circ} 27^{\prime} 08,8^{\prime \prime} \mathrm{N} 34^{\circ} 55^{\prime} 02.2^{\prime \prime}$

The sample from a 20 cm thick vein cutting diorites and gabbros was collected 2 km S of Khwazakhela, in a small quarry next to the Road Mingora-Kalam. The dyke consists of quartz, plagioclase, zoisite, scapolite, prismatic to acicular kyanite and poikiloblastic garnet. Detailed mineralogy is described by Jan \& Karim (1995).


Fig. A4.4: Representative zircons used for $U / P b$ - dating of a kyanite bearing dyke (38/02/01).

## Chilas gabbronorite (00/40/03) E 72³1'40,8' N 35 ${ }^{\circ} 07^{\prime} \mathbf{4 2}, 7^{\prime \prime}$

The sample locality is 1.5 km S of Madyan ( 55 km NNE of Mingora, Swat valley) along the road Mingora - Kalam. The gabbronorite is composed of clinopyroxene, orthopyroxene, amphibole, plagioclase, quartz, biotite and minor oxides. The magmatic fabric is in parts preserved. The well-preserved orthopyroxenes show exsolution-lamellae with beginning amphibole growth and spinel during cooling. Clinopyroxenes are less well preserved. The amphiboles are of metamorphic origin. Plagioclase is strongly saussuritized and albite twins are bended. Few bigger grains preserved anorthite rich plagioclase (An 60 ) with pericline twinning. The dominant part of plagioclase is likely to have recrystallized at temperatures of around $800^{\circ} \mathrm{C}$. Magnetite is an accessory mineral.


Fig. A4.5: Representative zircons used for U/Pb - dating of a Chilas gabbronorite (00/40/03): a,c) backscattered electron image (BSE), b,d) cathodoluminescence image (CL).

## Indian plate diorite (00/30/14) E 72 ${ }^{\circ} 53$ '51,7" N $35^{\circ} 02^{\prime} 14,5^{\prime \prime}$

The sample was collected 100 m S of Duber Bazar along the KKH. The coarse grained diorite is composed of K-feldspar, hornblende, plagioclase, quartz $\pm$ opaque oxides and traces of zircon and epidote. The magmatic fabric is generally preserved; only in strong foliated parts the fabric shows a metamorphic overprint.


Fig. A4.6: Representative zircons used for U/Pb - dating of Indian plate diorite (00/30/14): a,c) backscattered electron image (BSE), b,d) cathodoluminescence image (CL).

Table Appendix 4.1: $\mathrm{U}-\mathrm{Pb}$ isotopic data of zircons from all dated samples.

| Nr. | Description | Weight | nr . | Concentrations |  |  |  | Atomic ratios |  |  |  |  |  |  | Apparent ages |  |  | Error |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | a) | $[\mathrm{mg}]$ | of grains | U | $\begin{gathered} \mathrm{Pb} \\ \text { rad. } \\ {[\mathrm{ppm}]} \end{gathered}$ | Pb nonrad. $[\mathrm{pg}]$ | $\mathrm{Th} / \mathrm{U}$ <br> b) | $206 / 204$ <br> c) | 206/238 <br> d) | $\begin{gathered} \text { Error } \\ 2 \sigma[\%] \end{gathered}$ | 207/235 <br> d) | $\begin{gathered} \text { Error } \\ 2 \sigma[\%] \\ \hline \end{gathered}$ | 207/206 <br> d) | $\begin{gathered} \text { Error } \\ 2 \sigma[\%] \end{gathered}$ | 206/238 | 207/235 | 207/206 | corr. |
| Gabbro Sarangar |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| 1 | round | 0.0756 | 15 | 149 | 2.44 | 4.1 | 0.53 | 2733 | 0.01553 | 0.37 | 0.1030 | 0.48 | 0.04808 | 0.31 | 99.3 | 99.5 | 103.1 | 0.76 |
| 2 | round | 0.0585 | 18 | 182 | 2.99 | 2.0 | 0.58 | 5191 | 0.01544 | 0.33 | 0.1022 | 0.41 | 0.04803 | 0.19 | 98.8 | 98.8 | 98.4 | 0.89 |
| 3 | round | 0.1494 | 17 | 129 | 2.10 | 3.1 | 0.55 | 6162 | 0.01545 | 0.33 | 0.1024 | 0.40 | 0.04806 | 0.17 | 98.9 | 99.0 | 102.3 | 0.91 |
| 4 | round | 0.0171 | 21 | 189 | 3.08 | 1.9 | 0.53 | 1661 | 0.01549 | 0.39 | 0.1026 | 0.65 | 0.04805 | 0.52 | 99.1 | 99.2 | 101.9 | 0.60 |
| Granite |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| 5 | round | 0.0066 | 3 | 358 | 5.82 | 1.7 | 0.60 | 1386 | 0.01517 | 0.41 | 0.1006 | 0.81 | 0.04807 | 0.68 | 97.1 | 97.3 | 102.9 | 0.54 |
| 6 | prism incl | 0.0061 | 8 | 507 | 8.98 | 8.5 | 0.55 | 369 | 0.01517 | 0.50 | 0.1005 | 1.68 | 0.04807 | 1.55 | 97.0 | 97.3 | 102.6 | 0.40 |
| 7 | small prism | 0.0017 | 13 | 1377 | 22.25 | 2.2 | 0.57 | 1056 | 0.01512 | 0.37 | 0.1006 | 0.90 | 0.04797 | 0.78 | 97.3 | 97.3 | 100.2 | 0.51 |
| Diorite |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| 8 | prism | 0.0049 | 8 | 16.26 | 0.25 | 2.8 | 0.68 | 44 | 0.01435 | 2.12 | 0.0869 | 31.00 | 0.04392 | 29.50 | 91.8 | 84.6 | - | 0.72 |
| 9 | large oval | 0.0118 | 5 | 38.20 | 0.59 | 5.8 | 0.68 | 88 | 0.01426 | 0.86 | 0.0895 | 7.94 | 0.04550 | 7.54 | 91.3 | 87.0 | - | 0.51 |
| 10 | small oval | 0.0067 | 10 | 33.23 | 0.53 | 1.0 | 0.67 | 226 | 0.01442 | 0.71 | 0.0964 | 4.19 | 0.04848 | 3.97 | 92.3 | 93.5 | - | 0.39 |
| Swat dyke |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| 11 | acicular | 0.0119 | 12 | 6.74 | 0.08 | 1.1 | 0.05 | 77 | 0.01304 | 2.40 | 0.0802 | 16.30 | 0.04459 | 15.70 | 83.5 | 78.3 | - | 0.32 |
| 12 | prism | 0.0172 | 14 | 2.75 | 0.03 | 1.5 | 0.01 | 45 | 0.01299 | 2.16 | 0.0730 | 36.50 | 0.04077 | 35.20 | 83.2 | 71.6 | - | 0.62 |
| 13 | prism | 0.0190 | 22 | 5.29 | 0.06 | 4.3 | 0.10 | 37 | 0.01272 | 2.44 | 0.0739 | 35.44 | 0.04214 | 33.80 | 81.5 | 72.4 | - | 0.69 |
| Chilas gabbronorite |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| 14 | large round | 0.1618 | 9 | 113 | 1.75 | 5.0 | 0.91 | 3097 | 0.01337 | 0.36 | 0.0882 | 0.45 | 0.04783 | 0.24 | 85.6 | 85.8 | 90.7 | 0.85 |
| 15 | round | 0.0850 | 15 | 107 | 1.74 | 4.5 | 1.09 | 1721 | 0.01337 | 0.35 | 0.0882 | 0.52 | 0.04778 | 0.36 | 85.7 | 85.8 | 88.6 | 0.72 |
| 16 | round incl | 0.1026 | 15 | 99 | 1.62 | 8.8 | 1.12 | 993 | 0.01337 | 0.37 | 0.0884 | 0.61 | 0.04792 | 0.45 | 85.7 | 86.0 | 95.2 | 0.68 |
| 17 | prism | 0.0328 | 3 | 103 | 1.67 | 1.7 | 1.08 | 1686 | 0.01340 | 0.36 | 0.0883 | 0.64 | 0.04776 | 0.51 | 85.8 | 85.9 | 87.6 | 0.61 |
| Diorite Indian plate, Swat Valley |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| 18 | prism tips | 0.0316 | 4 | 376 | 131.45 | 2.4 | 0.43 | 104052 | 0.32727 | 0.33 | 5.1298 | 0.38 | 0.11368 | 0.10 | 1825.1 | 1841.1 | 1859.1 | 0.97 |
| 19 | small spr | 0.0156 | 6 | 232 | 78.47 | 2.9 | 0.45 | 24992 | 0.31502 | 0.34 | 4.9295 | 0.38 | 0.11349 | 0.11 | 1765.4 | 1807.3 | 1856.1 | 0.96 |
| 20 | frags | 0.0209 | 5 | 273 | 91.62 | 3.9 | 0.42 | 28915 | 0.31528 | 0.33 | 4.9383 | 0.38 | 0.11360 | 0.11 | 1766.6 | 1808.8 | 1857.8 | 0.96 |
| 21 | spr | 0.0329 | 5 | 210 | 67.80 | 15.2 | 0.41 | 8739 | 0.30394 | 0.33 | 4.7610 | 0.38 | 0.11361 | 0.11 | 1710.8 | 1778.0 | 1857.9 | 0.96 |

[^3]
## - Fission track dating

## Sample location:

## Sarangar Gabbro GbS (16/02/01) E 73 ${ }^{\circ} 01$ '31" N $35^{\circ} 07^{\prime} 16^{\prime \prime}$

The sample was taken ca. 2 km E of Patan next to the Karakoram Highway in an undeformed part of the Sarangar Gabbro. Detailed description in appendix $4 \mathrm{U} / \mathrm{Pb}$ dating samples.

## Indian Plate Diorite Dr (00/30/12) E 7254'38' N 35º ${ }^{\circ}$ '31"

The sample was collected ca. 1550 m E of Duber Bazar bridge (road distance) along KKH. The sample is similar to the diorite used for $\mathrm{U} / \mathrm{Pb}$ - dating (00/30/14).

## Pegmatite in Diorites/Hornblendites DrHbl (56/02/01) E 7302'54' N 35º7'48'

The sample locality is ca. 5000 m NE of Patan (road distance at KKH ). The coarse grained, undeformed pegmatite cuts the dioritic host rock.

## Granite (00/10/01) E 73 ${ }^{\circ} 10^{\prime} 44^{\prime \prime}$ N $35^{\circ} 144^{\prime \prime} 30^{\prime \prime}$

The sample locality is ca. 200 m S of the bridge where the Karakoram Highway ( KKH ) and the Indus River bend sharply towards the E. Detailed description in appendix $4 \mathrm{U} / \mathrm{Pb}$ dating samples.

Diorite (91/07/01) E $73^{\circ} 06^{\prime} 26^{\prime \prime}$ N $35^{\circ} 10$ '41'
The sample was collected alongside the KKH 1.1 km S of metal bridge at Kiru (road distance). Detailed description in appendix 4 $\mathrm{U} / \mathrm{Pb}$ dating samples.

Swat Granite Gneiss (95-196) E 72²4'14' N 3447'46' and (95-196) E 72²9'15' N 3449'20"
Samples collected and described by Anczkiewicz (1998). FT dating carried out by R. Anczkiewicz and D. Seward (unpublished).

Appendix Table 4.2

| Sample number | Alt. (m) | $\begin{aligned} & \text { ब्g } \\ & \text { E } \end{aligned}$ | Irradiation number | Number of grains | Standard track density $\begin{aligned} & \times 10^{4} \mathrm{~cm}^{-2} \\ & \text { (counted) } \\ & \hline \end{aligned}$ | $\rho_{\mathrm{s}} \times 10^{4} \mathrm{~cm}^{-2}$ <br> (counted) | $\begin{gathered} \rho_{\mathrm{i}} \times 10^{4} \mathrm{~cm}^{-2} \\ \text { (counted) } \\ \hline \end{gathered}$ | U conc. ppm | $\mathrm{P}\left(\chi^{2}\right)$ <br> Variation (\%) | Mean track length ( $\mu \mathrm{m}$ ) | std dev $(\mu \mathrm{m})$ | $\begin{array}{\|c} \text { Central } \\ \text { Age } \\ \pm 2 \sigma(\mathrm{Ma}) \\ \hline \end{array}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 95-196 | 800 | A | eth-86-14 | 16 | 110 (2439) | 12.9 (111) | 225 (1929) | 26 | 97 (0) | ------------ |  | $11.2 \pm 2.2$ |
| 95-198 | 1000 | A | eth-86-11 | 20 | 122 (2439) | 10.8 (82) | 156 (1185) | 16 | 83 (0.3) | ------------ |  | $15.0 \pm 3.4$ |
| 16/02/01 (GZ1) | 900 | A | eth-191-3 | 20 | 121 (2808) | 0.73 (126) | 14.5 (2509) | 2 | 12 (25) | $13.67 \pm 0.30$ | 1.99 (43) | $10.6 \pm 2.4$ |
| 00/30/12 (GZ11) | 825 | A | eth-192-3 | 20 | 130 (2978) | 3.64 (60) | 24.3 (4007) | 23 | 8 (25) | ------------ |  | $3.7 \pm 1.0$ |
| 95-196 | 800 | Z | eth-88-15 | 12 | 49.1 (2376) | 287 (378) | 414 (546) | 338 | 8 (18) |  |  | $20.5 \pm 3.6$ |
| 95-198 | 1000 | Z | eth-88-2 | 8 | 56.2 (2376) | 473 (885) | 563 (1054) | 400 | 11 (9) |  |  | $28.2 \pm 3.4$ |
| 56/02/01 (GZ3) | 1000 | Z | eth-193-4 | 20 | 73.3 (2543) | 14.6 (180) | 15.3 (189) | 8 | 100 (0) |  |  | $41.8 \pm 8.8$ |
| 00/10/01 (GZ4) | 1000 | Z | eth-193-11 | 4 | 64.8 (2543) | 223 (151) | 372 (252) | 230 | 22 (9) |  |  | $23.6 \pm 5.6$ |
| 91/07/01 (GZ8) | 960 | Z | eth-193-5 | 20 | 71.6 (2543) | 39.2 (217) | 60.7 (336) | 33.9 | 66 (10) | ------------ |  | $27.9 \pm 5.2$ |
| 00/30/12 (GZ11) | 825 | Z | eth-193-7 | 20 | 69.3 (2543) | 532 (1726) | 1002 (3253) | 578 | 0 (19) | ------ | ------------ | $22.6 \pm 2.6$ |

$\mathrm{A}=$ apatite, $\mathrm{Z}=$ zircon. $\rho_{\mathrm{S}}$ and $\rho_{\mathrm{i}}$ represent sample spontaneous and induced track densities; $\mathrm{P}\left(\chi^{2}\right)$ is the probability of $\mathrm{X}^{2}$ for v degrees of freedom where $\mathrm{v}=$ no. of crystals $-1 . \lambda_{\mathrm{D}}=1.55125 \times 10^{-10}$. Zeta $=355 \pm 5$ for apatite and CN5, $120 \pm 5$ for zircon and CN1. Samples were irradiated at the ANSTO facility, Australia. Ages are reported as central ages (Galbraith, 1981).

## APPENDIX Chapter 5

Table A5.1: Shear zone orientations data used for analysis presented in chapter 5.2.



|  | coordinates (decimal degree) |  | shear plane |  | lineation |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Outcrop | x (East) | y (North) | azimuth |  | muth | dip | sense |
| 6802 | 73.19036 | 35.24594 | 289 | 44 | 244 | 30 | normal |
| 6802 | 73.19036 | 35.24594 | 44 | 41 | 81 | 36 | reverse |
| 6802 | 73.19036 | 35.24594 | 62 | 60 | 108 | 42 | reverse (?) |
| 6803 | 73.19559 | 35.24671 | 14 | 53 | 310 | 20 | reverse (?) |
| 6803 | 73.19559 | 35.24671 | 338 | 85 | 28 | 70 | sinistral |
| 6803 | 73.19559 | 35.24671 | 18 | 60 | 72 | 50 | reverse (?) |
| 6901 | 73.20092 | 35.24607 | 60 | 38 | 76 | 35 | reverse (?) |
| 6901 | 73.20092 | 35.24607 | 70 | 24 | 80 | 22 | reverse |
| 6901 | 73.20092 | 35.24607 | 50 | 56 | 52 | 50 | reverse (?) |
| 6901 | 73.20092 | 35.24607 | 338 | 59 | 46 | 29 | reverse |
| 6901 | 73.20092 | 35.24607 | 41 | 76 | 107 | 56 | reverse |
| 6901 | 73.20092 | 35.24607 | 48 | 75 | 315 | 10 | sinistral |
| 6901 | 73.20092 | 35.24607 | 279 | 62 | 230 | 51 | reverse |
| 6902 | 73.20580 | 35.24525 | 23 | 73 | 114 | 1 | sinistral |
| 6902 | 73.20580 | 35.24525 | 20 | 39 | 4 | 15 | reverse |
| 6903 | 73.20876 | 35.24682 | 294 | 68 | 202 | 5 | sinistral |
| 6903 | 73.20876 | 35.24682 | 65 | 30 | 56 | 25 | reverse (?) |
| 6903 | 73.20876 | 35.24682 | 339 | 79 | 264 | 38 | sinistral |
| 7402 | 73.21815 | 35.27131 | 351 | 79 | 80 | 71 | normal |
| 7402 | 73.21815 | 35.27131 | 358 | 62 | 32 | 54 | normal |
| 7704 | 73.22846 | 35.21897 | 64 | 34 | 122 | 10 | normal |
| 8001 | 73.20614 | 35.22858 | 151 | 26 | 67 | 7 | reverse |
| 8001 | 73.20614 | 35.22858 | 132 | 30 | 70 | 19 | reverse (?) |
| 8002 | 73.20378 | 35.23462 | 90 | 44 | 100 | 46 | normal |
| 8202 | 73.20116 | 35.24146 | 70 | 46 | 130 | 33 | reverse |
| 8203 | 73.20471 | 35.24167 | 75 | 42 | 128 | 10 | reverse |
| 8203 | 73.20471 | 35.24167 | 69 | 86 | 340 | 9 | sinistral |
| 8302 | 73.20301 | 35.21055 | 186 | 86 | 88 | 13 | dextral |
| 8302 | 73.20301 | 35.21055 | 156 | 22 | 98 | 5 | reverse (?) |
| 8306 | 73.19836 | 35.19889 | 12 | 46 | 90 | 20 | reverse (?) |
| 8307 | 73.19525 | 35.19719 | 355 | 60 | 67 | 16 | reverse |
| 8401 | 73.15395 | 35.22531 | 140 | 24 | 68 | 8 | reverse (?) |
| 8401 | 73.15395 | 35.22531 | 290 | 24 | 248 | 16 | reverse (?) |
| 8403 | 73.15163 | 35.22490 | 130 | 27 | 65 | 10 | reverse |
| 8406 | 73.14608 | 35.22831 | 27 | 25 | 10 | 23 | normal |
| 8406 | 73.14608 | 35.22831 | 8 | 37 | 56 | 32 | normal |
| 4204 | 73.17867 | 35.24190 | 345 | 43 | 49 | 27 | normal |
| 4204 | 73.17867 | 35.24190 | 350 | 49 | 40 | 34 | normal |
| 6504 | 73.14652 | 35.21541 | 168 | 55 | 125 | 36 | reverse (?) |
| 6504 | 73.14652 | 35.21541 | 100 | 51 | 72 | 35 | sinistral |
| 6601 | 73.15185 | 35.21984 | 324 | 72 | 52 | 36 | sinistral |
| 6603 | 73.16220 | 35.22647 | 35 | 46 | 94 | 36 | reverse |
| 6701 | 73.17588 | 35.23602 | 27 | 52 | 57 | 50 | reverse |
| 6703 | 73.17762 | 35.23888 | 40 | 29 | 68 | 25 | reverse |
| 6603 | 73.16220 | 35.22647 | 28 | 57 | 75 | 45 | reverse |



|  | coordinates (decimal degree) |  | shear plane |  | lineation |  |  | coordinates(decimal degree) shear plane lineatio |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Outcrop | x (East) | y (North) a | muth | dip azi | muth | dip | sense | Outcrop | x (East) | y (North) a | nuth | dip | nuth | dip | sense |
| 6302 | 73.11785 | 35.18972 | 359 | 74 | 80 | 40 | reverse (?) | 3504 | 73.10770 | 35.17750 | 12 | 68 | 53 | 51 | reverse |
| 6302 | 73.11785 | 35.18972 | 2 | 32 | 66 | 21 | normal | 3504 | 73.10770 | 35.17750 | 10 | 44 | 48 | 36 | reverse |
| 6302 | 73.11785 | 35.18972 | 345 | 44 | 44 | 28 | reverse (?) | 3504 | 73.10770 | 35.17750 | 2 | 70 | 53 | 51 | reverse (?) |
| 6303 | 73.11877 | 35.19119 | 132 | 12 | 200 | 8 | normal | 3505 | 73.10935 | 35.18161 | 18 | 82 | 70 | 75 | reverse |
| 6303 | 73.11877 | 35.19119 | 16 | 55 | 49 | 48 | normal | 3602 | 73.11104 | 35.18378 | 350 | 84 | 74 | 36 | sinistral |
| 6303 | 73.11877 | 35.19119 | 356 | 42 | 50 | 27 | reverse (?) | 3602 | 73.11104 | 35.18378 | 335 | 28 | 55 | 5 | reverse |
| 6303 | 73.11877 | 35.19119 | 4 | 60 | 62 | 50 | reverse (?) | 3602 | 73.11104 | 35.18378 | 15 | 56 | 60 | 51 | reverse |
| 6303 | 73.11877 | 35.19119 | 50 | 22 | 40 | 21 | reverse | 3602 | 73.11104 | 35.18378 | 357 | 57 | 58 | 40 | reverse |
| 6303 | 73.11877 | 35.19119 | 3 | 25 | 38 | 25 | reverse | 3602 | 73.11104 | 35.18378 | 25 | 43 | 58 | 40 | reverse |
| 6303 | 73.11877 | 35.19119 | 25 | 47 | 60 | 36 | reverse | 2801 | 73.02574 | 35.18148 | 34 | 50 | 80 | 40 | reverse (?) |
| 6303 | 73.11877 | 35.19119 | 358 | 57 | 44 | 47 | sinistral | 2802 | 73.02639 | 35.18186 | 36 | 45 | 65 | 38 | reverse (?) |
| 6401 | 73.12173 | 35.19361 | 8 | 30 | 70 | 26 | reverse | 2806 | 73.01304 | 35.18719 | 41 | 31 | 67 | 39 | reverse |
| 6401 | 73.12173 | 35.19361 | 27 | 30 | 52 | 28 | reverse | 2905 | 73.02626 | 35.17333 | 35 | 25 | 40 | 25 | reverse |
| 6401 | 73.12173 | 35.19361 | 58 | 42 | 67 | 38 | reverse | 2906 | 73.02602 | 35.17381 | 60 | 48 | 95 | 44 | reverse (?) |
| 6401 | 73.12173 | 35.19361 | 20 | 31 | 34 | 30 | reverse | 4605 | 73.10063 | 35.15969 | 340 | 85 | 270 | 66 | reverse (?) |
| 6401 | 73.12173 | 35.19361 | 40 | 40 | 58 | 38 | reverse | 6302 | 73.11785 | 35.18972 | 348 | 40 | 48 | 34 | reverse |
| 6401 | 73.12173 | 35.19361 | 35 | 32 | 72 | 30 | reverse (?) |  |  |  |  |  |  |  |  |
| 6401 | 73.12173 | 35.19361 | 330 | 42 | 60 | 35 | reverse | Station 4: Northern sheared Gabbros and Diorites |  |  |  |  |  |  |  |
|  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
|  |  |  |  |  |  |  |  | coordinates(decimal degree) shear plane lineation |  |  |  |  |  |  |  |
| Outcrop | x (East) | y (North) azimuth |  | dip azimuth |  |  | sense |  |  |  |  |  |  |  |  |
| 6401 | 73.12173 | 35.19361 | 339 | 51 | 42 | 29 | reverse | Outcrop | x (East) | y (North) azimuth |  | dip azimuth |  | dip | sense |
| 6401 | 73.12173 | 35.19361 | 315 | 25 | 45 | 1 |  | 18026108 | 73.05584 | 35.13943 | 249 | 16 | 181 | 6 | normal |
| 6402 | 73.12371 | 35.19805 | 343 | 34 | 4 | 31 | reverse |  | 72.97931 | 35.17571 | 16 | 80 | 303 | 20 | dextral |
| 6402 | 73.12371 | 35.19805 | 13 | 36 | 37 | 35 | reverse (?) | 8904 | 72.98060 | 35.15927 | 30 | 22 | 32 | 20 | normal |
| 6402 | 73.12371 | 35.19805 | 352 | 67 | 20 | 64 | normal | 1703 | 73.04697 | 35.12888 | 340 | 70 | 54 | 40 | reverse |
| 6402 | 73.12371 | 35.19805 | 3 | 58 | 45 | 52 | reverse | 1703 | 73.04697 | 35.12888 | 358 | 56 | 359 | 52 | reverse (?) |
| 6403 | 73.12644 | 35.20133 | 205 | 88 | 94 | 74 | reverse | 1704 | 73.04795 | 35.12961 | 330 | 36 | 16 | 36 | reverse |
| 6403 | 73.12644 | 35.20133 | 346 | 34 | 18 | 30 | reverse | 1704 | 73.04795 | 35.12961 | 358 | 42 | 20 | 40 | reverse |
| 6403 | 73.12644 | 35.20133 | 36 | 67 | 60 | 64 | reverse | 1705 | 73.04989 | 35.13081 | 6 | 60 | 45 | 47 | reverse |
| 6403 | 73.12644 | 35.20133 | 2 | 32 | 60 | 20 | reverse | 1706 | 73.05073 | 35.13108 | 59 | 61 | 61 | 59 | normal |
| 6403 | 73.12644 | 35.20133 | 48 | 55 | 50 | 55 | reverse (?) | 1706 | 73.05073 | 35.13108 | 346 | 40 | 29 | 26 | normal |
| 6403 | 73.12644 | 35.20133 | 35 | 72 | 70 | 69 | normal | 1706 | 73.05073 | 35.13108 | 54 | 25 | 62 | 22 | reverse |
| 6404 | 73.12898 | 35.20352 | 11 | 51 | 345 | 45 | reverse | 1801 | 73.05722 | 35.13702 | 328 | 24 | 28 | 9 | reverse |
| 6404 | 73.12898 | 35.20352 | 24 | 74 | 312 | 50 | reverse (?) | 1801 | 73.05722 | 35.13702 | 345 | 35 | 7 | 34 | reverse |
| 6404 | 73.12898 | 35.20352 | 49 | 56 | 56 | 54 | reverse (?) | 1801 | $\begin{array}{lll}73.05722 & 35.13702\end{array}$ |  | 5 | 37 | 14 | 36 | reverse |
| 6404 | 73.12898 | 35.20352 | 357 | 72 | 80 | 20 | sinistral | 1802 | 73.0558435 .13943 |  | 344 | 76 | 50 | 64 | reverse |
| 6405 | 73.13300 | 35.20703 | 20 | 30 | 358 | 81 | reverse | 1802 | $73.05584 \quad 35.13943$ |  | 271 | 26 | 201 | 10 | normal |
| 6405 | 73.13300 | 35.20703 | 16 | 85 | 294 | 60 | normal | 1802 | $73.05584 \quad 35.13943$ |  | 348 | 70 | 34 | 53 | reverse |
| 6405 | 73.13300 | 35.20703 | 11 | 82 | 69 | 74 | normal | 1802 | $73.05584 \quad 35.13943$ |  | 345 | 60 | 51 | 40 | reverse |
| 6501 | 73.13421 | 35.20741 | 262 | 25 | 268 | 24 | normal | 1802 | 73.05584 | 35.13943 | 48 | 86 | 140 | 71 | reverse |
| 6501 | 73.13421 | 35.20741 | 359 | 76 | 88 | 2 | sinistral | 1803 | 73.05022 | 35.14135 | 48 | 29 | 24 | 36 | reverse |
| 6501 | 73.13421 | 35.20741 | 345 | 74 | 35 | 68 | reverse (?) | 1805 | 73.0574573.05745 | 35.14461 | 10 | 33 | 20 | 10 | reverse |
| 6501 | 73.13421 | 35.20741 | 2 | 74 | 42 | 66 | reverse (?) | 1805 |  | 35.14461 | 16 | 24 | 35 | 22 | reverse |
| 6502 | 73.13705 | 35.20920 | 3 | 74 | 6 | 73 | normal | 1806 | $\begin{aligned} & 73.05745 \\ & 73.05994 \end{aligned}$ | 35.14504 | 159 | 70 | 52 | 30 | reverse |
| 6502 | 73.13705 | 35.20920 | 8 | 82 | 321 | 65 | normal | 1806 | $\begin{aligned} & 73.05994 \\ & 73.05994 \end{aligned}$ | 35.14504 | 30 | 46 | 64 | 16 | reverse |
| 6502 | 73.13705 | 35.20920 | 2 | 63 | 275 | 8 | sinistral | 1902 | $\begin{aligned} & 73.05994 \\ & 73.06980 \end{aligned}$ | 35.14188 | 20 | 12 | 70 | 4 | reverse |
| 6502 | 73.13705 | 35.20920 | 5 | 55 | 70 | 20 | reverse | 1902 | $\begin{aligned} & 73.06980 \\ & 73.06980 \end{aligned}$ |  | 40 | 74 | 104 | 25 | reverse |
| 6502 | 73.13705 | 35.20920 | 20 | 67 | 2 | 42 | reverse (?) | 1902 | $73.06980$ |  | 88 | 40 | 47 | 34 | normal |
| 6502 | 73.13705 | 35.20920 | 148 | 5 | 110 | 5 | reverse | 1902 | 73.06980 | 35.14188 | 64 | 56 | 34 | 46 | normal |
| 6502 | 73.13705 | 35.20920 | 128 | 6 | 55 | 1 | reverse | 1902 | 73.06980 | 35.14188 | 96 | 40 | 49 | 34 | normal |
| 6502 | 73.13705 | 35.20920 | 92 | 60 | 92 | 60 | reverse | 1902 | $\begin{aligned} & 73.06980 \\ & 73.06980 \end{aligned}$ | 35.14188 | 27 | 50 | 60 | 40 | normal |
| 6502 | 73.13705 | 35.20920 | 248 | 28 | 72 | 26 | normal | 1902 |  | 35.14188 | 61 | 44 | 63 | 46 | normal |
| 3403 | 73.09883 | 35.15655 | 305 | 45 | 30 | 10 | sinistral | 1903 | $\begin{aligned} & 73.06980 \\ & 73.07417 \end{aligned}$ | 35.14350 | 22 | 33 | 34 | 31 | normal |
| 3403 | 73.09883 | 35.15655 | 350 | 56 | 51 | 39 | sinistral | 1903 | $73.07417$ | 35.1435035.14350 | 23 | 40 | 62 | 34 | normal |
| 3403 | 73.09883 | 35.15655 | 337 | 52 | 5 | 50 | normal | 1903 | 73.07417 |  | 352 | 30 | 50 | 15 | normal |
| 3403 | 73.09883 | 35.15655 | 325 | 70 | 30 | 64 | reverse | 1904 | 73.07792 | 35.14503 | 50 | 50 | 65 | 45 | normal |
| 3403 | 73.09883 | 35.15655 | 4 | 79 | 53 | 71 | normal | 1904 | 73.07792 | 35.14503 | 41 | 49 | 60 | 43 | normal |
| 3403 | 73.09883 | 35.15655 | 177 | 82 | 138 | 80 | reverse | 1904 | $\begin{aligned} & 73.07792 \\ & 73.07792 \end{aligned}$ | 35.14503 | 45 | 41 | 40 | 41 | reverse |
| 3403 | 73.09883 | 35.15655 | 176 | 80 | 103 | 69 | reverse (?) | 1904 |  | 35.14503 | 22 | 32 | 30 | 26 | reverse |
| 3403 | 73.09883 | 35.15655 | 254 | 15 | 188 | 6 | normal | 1904 | 73.07792 | 35.14503 | 160 | 26 | 242 | 10 | normal |
| 3404 | 73.10073 | 35.16039 | 18 | 30 | 342 | 20 | reverse | 1904 | 73.07792 | 35.14503 | 30 | 80 | 82 | 46 | normal |
| 3501 | 73.10168 | 35.16273 | 13 | 32 | 43 | 31 | reverse | 1904 | 73.07792 | 35.14503 | 68 | 39 | 58 | 36 | normal |
| 3501 | 73.10168 | 35.16273 | 322 | 71 | 10 | 62 | reverse | 1905 | 73.07929 | 35.14556 | 84 | 6 | 54 | 5 | reverse |
| 3501 | 73.10168 | 35.16273 | 40 | 50 | 26 | 46 | reverse | 1905 | 73.07929 | 35.14556 | 18 | 73 | 348 | 69 | normal |
| 3501 | 73.10168 | 35.16273 | 322 | 58 | 20 | 40 | reverse | 3205 | 73.06216 | 35.12827 | 0 | 46 | 50 | 34 | reverse (?) |
| 3502 | 73.10305 | 35.16671 | 18 | 6 | 40 | 5 | reverse | 3207 | 73.06747 | 35.13175 | 2 | 48 | 16 | 50 | reverse |
| 3502 | 73.10305 | 35.16671 | 330 | 20 | 20 | 16 | reverse | 3301 | 73.08405 | 35.13558 | 39 | 29 | 80 | 26 | reverse |
| 3502 | 73.10305 | 35.16671 | 350 | 15 | 12 | 14 | reverse | 3302 | 73.08613 | 35.13051 | 332 | 49 | 24 | 42 | normal |


|  | $\begin{gathered} \text { coordinates } \\ \text { (decimal degree) } \end{gathered}$ |  | shear plane |  | lineation |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Outcrop | x (East) | y (North) | azimuth | dip | imuth | dip | sense |
| 3302 | 73.08613 | 35.13051 | 326 | 50 | 35 | 30 | sinistral |
| 3302 | 73.08613 | 35.13051 | 335 | 56 | 34 | 41 | dextral |
| 3303 | 73.08728 | 35.13018 | 307 | 14 | 354 | 12 | reverse |
| 3303 | 73.08728 | 35.13018 | 304 | 62 | 30 | 13 | sinistral |
| 3303 | 73.08728 | 35.13018 | 250 | 13 | 230 | 11 | normal |
| 3303 | 73.08728 | 35.13018 | 266 | 25 | 200 | 9 | normal |
| 3305 | 73.08431 | 35.12575 | 224 | 26 | 214 | 25 | normal |
| 3306 | 73.08446 | 35.12149 | 275 | 54 | 208 | 15 | sinistral |
| 3306 | 73.08446 | 35.12149 | 328 | 66 | 45 | 26 | sinistral |
| 3306 | 73.08446 | 35.12149 | 358 | 27 | 40 | 21 | reverse |
| 3306 | 73.08446 | 35.12149 | 317 | 53 | 36 | 19 | sinistral |
| 3401 | 73.08600 | 35.14840 | 50 | 78 | 42 | 76 | normal |
| 3401 | 73.08600 | 35.14840 | 302 | 59 | 15 | 22 | reverse |
| 3401 | 73.08600 | 35.14840 | 38 | 56 | 2 | 55 | normal |
| 3401 | 73.08600 | 35.14840 | 10 | 42 | 26 | 37 | normal |
| 3401 | 73.08600 | 35.14840 | 57 | 28 | 29 | 24 | normal |
| 3402 | 73.08881 | 35.14965 | 26 | 28 | 42 | 26 | reverse |
| 3402 | 73.08881 | 35.14965 | 248 | 45 | 182 | 26 | normal |
| 3402 | 73.08881 | 35.14965 | 38 | 36 | 15 | 36 | normal |
| 3402 | 73.08881 | 35.14965 | 352 | 46 | 26 | 44 | reverse (?) |
| 2205 | 72.97700 | 35.14364 | 50 | 28 | 18 | 26 | reverse |
| 2505 | 72.98064 | 35.15889 | 220 | 76 | 297 | 28 | reverse (?) |
| 2509 | 72.97833 | 35.18225 | 34 | 28 | 52 | 29 | reverse (?) |
| 2510 | 72.97892 | 35.17989 | 50 | 52 | 77 | 64 | reverse (?) |
| 2510 | 72.97892 | 35.17989 | 48 | 77 | 50 | 75 | reverse (?) |
| 2511 | 72.97955 | 35.17514 | 322 | 35 | 316 | 36 | reverse (?) |
| 2901 | 73.04511 | 35.14342 | 20 | 44 | 5 | 43 | reverse |
| 2903 | 73.03230 | 35.15205 | 20 | 40 | 58 | 34 | reverse (?) |
| 4403 | 73.05253 | 35.13176 | 55 | 80 | 88 | 66 | normal |
| 4701 | 73.08569 | 35.14875 | 40 | 59 | 83 | 43 | normal |
| 8905 | 72.97778 | 35.14741 | 64 | 34 | 60 | 21 | reverse |


| Station 5: Patan Area (Sarangar Gabbro and Granulitic Gabbro) |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| coordinates (decimal degree) |  |  | shear plane |  | lineation |  |  |
| Outcrop | x (East) | y (North) | zimuth |  | muth | dip | sense |
| 1603 | 73.02804 | 35.11989 | 80 | 65 | 141 | 45 | normal |
| 5502 | 72.99897 | 35.08805 | 28 | 56 | 40 | 58 | reverse |
| 5502 | 72.99897 | 35.08805 | 50 | 32 | 15 | 28 | reverse |
| 5502 | 72.99897 | 35.08805 | 186 | 20 | 200 | 18 | normal |
| 5502 | 72.99897 | 35.08805 | 60 | 70 | 31 | 58 | reverse |
| 5901 | 73.08578 | 35.11160 | 282 | 34 | 224 | 24 | normal |
| 5902 | 73.08621 | 35.10807 | 170 | 70 | 200 | 62 | reverse |
| 5902 | 73.08621 | 35.10807 | 150 | 65 | 70 | 20 | dextral |
| 5903 | 73.09002 | 35.09814 | 290 | 29 | 9 | 1 | reverse |
| 5903 | 73.09002 | 35.09814 | 310 | 36 | 12 | 16 | reverse |
| 5903 | 73.09002 | 35.09814 | 344 | 64 | 36 | 43 | reverse |
| 5903 | 73.09002 | 35.09814 | 323 | 50 | 7 | 45 | reverse |
| 5903 | 73.09002 | 35.09814 | 324 | 70 | 359 | 69 | reverse |
| 5903 | 73.09002 | 35.09814 | 315 | 65 | 15 | 36 | reverse |
| 5903 | 73.09002 | 35.09814 | 336 | 69 | 312 | 67 | reverse |
| 5903 | 73.09002 | 35.09814 | 310 | 60 | 32 | 27 | reverse |
| 5904 | 73.09244 | 35.09359 | 332 | 85 | 40 | 80 | reverse |
| 5904 | 73.09244 | 35.09359 | 348 | 30 | 25 | 28 | reverse |
| 5904 | 73.09244 | 35.09359 | 300 | 79 | 30 | 15 | dextral |
| 5904 | 73.09244 | 35.09359 | 308 | 32 | 345 | 20 | reverse |
| 5904 | 73.09244 | 35.09359 | 316 | 72 | 36 | 34 | reverse |
| 5904 | 73.09244 | 35.09359 | 314 | 71 | 41 | 30 | reverse |
| 5904 | 73.09244 | 35.09359 | 280 | 40 | 220 | 36 | sinistral |
| 5904 | 73.09244 | 35.09359 | 346 | 80 | 49 | 66 | reverse |
| 8502 | 73.09766 | 35.07547 | 286 | 34 | 354 | 8 | reverse (?) |
| 8707 | 72.93088 | 35.10778 | 54 | 41 | 108 | 43 | normal |
| 8708 | 72.93099 | 35.11146 | 350 | 50 | 345 | 45 | reverse (?) |
| 8902 | 72.98505 | 35.12959 | 286 | 75 | 216 | 56 | reverse |
| 1210 | 73.00177 | 35.11261 | 170 | 35 | 252 | 6 | normal |
| 1210 | 73.00177 | 35.11261 | 299 | 25 | 354 | 10 | reverse |
| 1210 | 73.00177 | 35.11261 | 181 | 24 | 154 | 19 | normal |
| 1304 | 72.99934 | 35.11603 | 318 | 40 | 270 | 29 | normal |




|  | coordinates (decimal degree) |  | shear plane |  | e lineation |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Outcrop | x (East) | y (North) | azimuth |  | muth | dip | sense |
| 3106 | 73.04769 | 35.11456 | 27 | 46 | 14 | 45 | normal |
| 3106 | 73.04769 | 35.11456 | 23 | 24 | 10 | 23 | reverse |
| 3106 | 73.04769 | 35.11456 | 96 | 46 | 42 | 36 | normal |
| 3106 | 73.04769 | 35.11456 | 100 | 72 | 50 | 60 | reverse (?) |
| 3201 | 73.05086 | 35.11575 | 172 | 74 | 258 | 56 | sinistral |
| 3201 | 73.05086 | 35.11575 | 75 | 61 | 300 | 10 | reverse |
| 3202 | 73.05343 | 35.11895 | 322 | 10 | 45 | 4 | reverse |
| 3202 | 73.05343 | 35.11895 | 348 | 18 | 28 | 13 | reverse |
| 3203 | 73.05553 | 35.12295 | 330 | 16 | 26 | 7 | reverse |
| 3203 | 73.05553 | 35.12295 | 34 | 30 | 70 | 36 | reverse |
| 3203 | 73.05553 | 35.12295 | 330 | 45 | 49 | 16 | reverse |
| 3203 | 73.05553 | 35.12295 | 332 | 50 | 300 | 46 | reverse (?) |
| 2203 | 72.98422 | 35.13153 | 34 | 70 | 0 | 65 | normal |
| 2207 | 73.01397 | 35.07918 | 126 | 4 | 254 | 2 | reverse (?) |
| 2501 | 72.98090 | 35.13529 | 30 | 29 | 50 | 25 | reverse |
| 2501 | 72.98090 | 35.13529 | 54 | 60 | 20 | 39 | reverse |
| 2501 | 72.98090 | 35.13529 | 125 | 28 | 234 | 5 | reverse (?) |
| 2703 | 73.02275 | 35.06683 | 96 | 46 | 179 | 11 | reverse (?) |
| 4301 | 73.00968 | 35.11561 | 25 | 55 | 89 | 35 | normal |
| 4301 | 73.00968 | 35.11561 | 40 | 15 | 30 | 10 | reverse |
| 4304 | 73.01982 | 35.12058 | 52 | 46 | 60 | 43 | normal |
| 4305 | 73.02028 | 35.12066 | 58 | 34 | 54 | 31 | normal |


| Station 6: Indian Plate |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | $\begin{gathered} \text { coordinates } \\ \text { (decimal degree) } \end{gathered}$ |  | shear plane |  | e lineation |  |  |
| Outcrop | x (East) | y (North) | nuth |  |  | dip | sense |
| 1101 | 72.89736 | 35.03788 | 47 | 67 | 19 | 45 | reverse |
| 1101 | 72.89736 | 35.03788 | 330 | 65 | 50 | 31 | reverse |
| 1102 | 72.90066 | 35.03944 | 350 | 65 | 7 | 65 | reverse |
| 1103 | 72.91613 | 35.04110 | 104 | 59 | 36 | 12 | reverse |
| 1103 | 72.91613 | 35.04110 | 113 | 44 | 15 | 10 | reverse |
| 1104 | 72.91835 | 35.03995 | 74 | 41 | 7 | 15 | reverse |

Given coordinates represent the closest GPS point

## APPENDIX Chapter 6



Fig. A6.1: Site locality map showing the data file-ID and outcrop-ID used for paleostress analysis presented in chapter 6.

Table A6．1：Fault data used for paleo－stress analysis presented in chapter 6

| 弟 |  |  | $\begin{aligned} & \text { E. } \\ & \text { En } \\ & \text { Hen } \end{aligned}$ |  |  |  | （1） |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  | $\overline{\ddot{0}}$ | $\begin{aligned} & \text { E } \\ & \\ & H \end{aligned}$ | $\ddot{0}$ |  |  | $\begin{aligned} & \text { 己 } \\ & \text { ت̈g } \\ & 0 \\ & 0 \\ & 0 \\ & 0 \\ & 0 \\ & 000 \end{aligned}$ |  |  |
| Data－file 1106， $\mathrm{n}=27$ |  |  |  |  |  |  |  |  |  |
| ， | 102 | 70 | 29 | 39 | － | ， | 03／03 | 1 |  |
| 2 | 97 | 76 | 12 | 20 | － | 1 | A1／04 | 1 |  |
| 3 | 96 | 70 | 17 | 28 | － | 1 | A1／04 | 1 |  |
| 4 | 105 | 66 | 26 | 23 | － | 1 | A1／04 | 1 |  |
| 5 | 107 | 65 | 22 | 11 | － | 1 | A1／04 | 1 |  |
| 6 | 102 | 74 | 13 | 5 | － | 1 | A1／04 | 1 |  |
| 7 | 236 | 71 | 320 | 16 | ＋ | 1 | 11／03 | 2 |  |
| 8 | 191 | 32 | 280 | 1 | － | 1 | A1／04 | 2 |  |
| 9 | 123 | 89 | 33 | 17 | ＋ | 1 | A1／04 | 2 |  |
| 10 | 52 | 52 | 325 | 5 | － | 1 | A1／04 | 2 |  |
| 11 | 262 | 77 | 341 | 38 | ＋ | 1 | A1／04 | 2 |  |
| 12 | 62 | 82 | 347 | 63 | － | 1 | 11／03 | 3 |  |
| 13 | 230 | 67 | 291 | 49 | － | 1 | 11／03 | 3 |  |
| 14 | 48 | 73 | 62 | 73 | － | 1 | 11／06 | 3 |  |
| 15 | 249 | 86 | 186 | 81 | － | 1 | 11／06 | 3 |  |
| 16 | 37 | 57 | 13 | 55 | － | 1 | 11／06 | 3 |  |
| 17 | 80 | 70 | 96 | 69 | － | 1 | 11／06 | 3 |  |
| 18 | 70 | 80 | 358 | 61 | － | 1 | 11／06 | 3 |  |
| 19 | 17 | 50 | 17 | 50 | － | 1 | 11／06 | 3 |  |
| 20 | 244 | 86 | 314 | 79 | － | 1 | 03／03 | 3 |  |
| 21 | 62 | 69 | 54 | 69 | － | 1 | 03／03 | 3 |  |
| 22 | 66 | 49 | 66 | 49 | － | 1 | 03／03 | 3 |  |
| 23 | 56 | 61 | 71 | 60 | － | 1 | 03／03 | 3 |  |
| 24 | 61 | 72 | 63 | 72 | － | 1 | 03／03 | 3 |  |
| 25 | 79 | 50 | 17 | 29 | － | 1 | A1／04 | 3 |  |
| 26 | 228 | 85 | 143 | 45 | － | 3 | 11／04 | 6 |  |
| 27 | 131 | 81 | 204 | 61 | － | 1 | 03／03 | 6 |  |
| Data－file 1205， $\mathrm{n}=33$ |  |  |  |  |  |  |  |  |  |
| 28 | 320 | 68 | 42 | 18 | ＋ | 1 | 12／05 | 1 | reg |
| 29 | 274 | 71 | 358 | 18 | ＋ | 1 | 12／06 | 1 | reg |
| 30 | 72 | 84 | 159 | 32 | ＋ | 1 | 12／07 | 1 | reg |
| 31 | 261 | 46 | 190 | 20 | － | 1 | A0／01 | 1 | reg |
| 32 | 330 | 89 | 244 | 73 | － | 1 | A0／01 | 1 | reg |
| 33 | 303 | 84 | 225 | 62 | ＋ | 1 | 12／05 | 2 | reg |
| 34 | 231 | 59 | 312 | 15 | ＋ | 1 | 12／07 | 2 | reg |
| 35 | 301 | 80 | 213 | 13 | ＋ | 4 | 12／07 | 2 | reg |
| 36 | 200 | 36 | 236 | 30 | － | 1 | 12／03 | 3 | reg |
| 37 | 102 | 57 | 77 | 55 | － | 1 | 12／03 | 3 | reg |
| 38 | 91 | 77 | 56 | 74 | － | 1 | 12／03 | 3 | reg |
| 39 | 91 | 62 | 67 | 60 | － | 1 | 12／03 | 3 | reg |
| 40 | 40 | 70 | 349 | 60 | － | 1 | 12／05 | 3 | reg |
| 41 | 280 | 58 | 271 | 58 | － | 1 | 12／06 | 3 | reg |
| 42 | 97 | 71 | 157 | 56 | － | 1 | 12／07 | 3 | reg |
| 43 | 130 | 75 | 77 | 67 | － | 1 | 12／07 | 3 | reg |
| 44 | 253 | 74 | 203 | 66 | － | 1 | 12／07 | 3 | reg |
| 45 | 348 | 60 | 298 | 48 | － | 1 | A0／01 | 3 | reg |
| 46 | 330 | 70 | 294 | 66 |  | 1 | A0／01 | 3 | reg |
| 47 | 45 | 19 | 62 | 18 | － | 1 | A0／01 | 3 | reg |
| 48 | 65 | 69 | 112 | 61 | － | 1 | A0／01 | 3 | reg |
| 49 | 67 | 71 | 90 | 70 | － | 1 | A0／01 | 3 | reg |
| 50 | 64 | 78 | 25 | 74 | － | 1 | A0／01 | 3 | reg |
| 51 | 85 | 26 | 58 | 23 | － | 1 | A0／01 | 3 | reg |
| 52 | 71 | 46 | 36 | 41 | － | 1 | A0／01 | 3 | reg |
| 53 | 52 | 84 | 339 | 71 | － | 1 | A0／01 | 3 | reg |
| 54 | 278 | 77 | 314 | 74 | － | 1 | A0／01 | 3 | reg |
| 55 | 13 | 71 | 315 | 57 | － | 1 | A0／01 | 3 | reg |
| 56 | 238 | 11 | 247 | 11 | － | 1 | A0／01 | 3 | reg |
| 57 | 223 | 73 | 171 | 63 | ＋ | 1 | 12／07 | 4 | reg |
| 58 | 116 | 77 | 192 | 46 | － | 1 | 12／05 | 6 | reg |
| 59 | 339 | 52 | 281 | 34 | － | 1 | 12／05 | 6 | reg |
| 60 | 344 | 76 | 273 | 53 | － | 1 | 12／06 | 6 | reg |
| 61 | 4 | 56 | 295 | 28 | － | 1 | A0／01 | 6 | reg |
| Data－file 1209， $\mathrm{n}=20$ |  |  |  |  |  |  |  |  |  |
| 62 | 130 | 26 | 163 | 22 | ＋ | 1 | 12／08 | 1 | reg |
| 63 | 128 | 80 | 39 | 7 | － | 1 | 12／08 | 1 | reg |
| 64 | 106 | 82 | 17 | 6 | － | 1 | 12／08 | 1 | reg |


|  |  |  |  | $\stackrel{\tilde{O}}{\square}$ | $\begin{aligned} & \ddot{0} \\ & \stackrel{\rightharpoonup}{0} \\ & \ddot{\sim} \end{aligned}$ |  |  | $\begin{aligned} & \text { E } \\ & \text { ⿹ㅡN } \\ & \text { H } \end{aligned}$ |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 65 | 150 | 89 | 239 | 35 | ＋ | 2 | 13／06 | 1 | reg |
| 66 | 290 | 35 | 333 | 27 | ＋ | 1 | 13／07 | 1 | reg |
| 67 | 219 | 60 | 135 | 9 | ＋ | 1 | 13／07 | 1 | reg |
| 68 | 228 | 60 | 300 | 28 | ＋ | 1 | 13／07 |  | reg |
| 69 | 180 | 63 | 251 | 33 | － | 4 | 12／08 | 3 | reg |
| 70 | 190 | 46 | 162 | 42 | － | 2 | 13／06 | 3 | reg |
| 71 | 330 | 76 | 21 | 69 | － | 4 | 13／07 | 3 | reg |
| 72 | 301 | 46 | 328 | 42 | － | 1 | 13／07 | 3 | reg |
| 73 | 93 | 74 | 5 | 6 | ＋ | 1 | 12／08 | 4 | reg |
| 74 | 120 | 81 | 209 | 6 | － | 1 | 12／08 | 4 | reg |
| 75 | 268 | 80 | 179 | 4 | ＋ | 1 | 12／08 | 4 | reg |
| 76 | 188 | 84 | 101 | 21 | － | 1 | 12／09 | 4 | reg |
| 77 | 270 | 89 | 360 | 20 | ＋ | 4 | 12／10 | 4 | reg |
| 78 | 162 | 69 | 77 | 12 | － | 1 | 12／10 | 4 | reg |
| 79 | 246 | 86 | 335 | 15 | － | 1 | 12／10 | 4 | reg |
| 80 | 354 | 65 | 268 | 9 | － | 4 | 13／07 | 4 | reg |
| 81 | 78 | 48 | 152 | 17 | ＋ | 1 | 12／10 | 6 | reg |
| Data－file 1402， $\mathrm{n}=16$ |  |  |  |  |  |  |  |  |  |
| 82 | 102 | 84 | 14 | 16 | － | 1 | 14／03 | 1 | reg |
| 83 | 272 | 40 | 182 | 1 | － | 3 | 31／01 | 1 | reg |
| 84 | 100 | 40 | 81 | 39 | ＋ | 1 | 14／01 | 2 | reg |
| 85 | 255 | 50 | 236 | 48 | ＋ | 3 | 31／02 | 2 | reg |
| 86 | 78 | 63 | 32 | 53 | ＋ | 2 | 31／02 | 2 | reg |
| 87 | 235 | 47 | 285 | 34 | － | 1 | 14／01 | 3 | reg |
| 88 | 260 | 45 | 235 | 42 | － | 1 | 14／02 | 3 | reg |
| 89 | 295 | 44 | 241 | 30 | － | 1 | 14／02 | 4 | reg |
| 90 | 176 | 88 | 265 | 36 | ＋ | 1 | 14／02 | 4 | reg |
| 91 | 100 | 26 | 16 | 3 | ＋ | 1 | 14／03 | 4 | reg |
| 92 | 72 | 70 | 354 | 29 | ＋ | 2 | 31／01 | 4 | reg |
| 93 | 255 | 50 | 185 | 21 | ＋ | 2 | 31／02 | 4 | reg |
| 94 | 6 | 62 | 78 | 30 | ＋ | 1 | 31／03 | 4 | reg |
| 95 | 255 | 88 | 166 | 25 | ＋ | 3 | 31／03 | 4 | reg |
| 96 | 280 | 50 | 206 | 17 | ＋ | 3 | 31／03 | 4 | reg |
| 97 | 339 | 15 | 19 | 11 | ＋ | 4 | 31／03 | 4 | reg |
| Data－file 1501， $\mathrm{n}=29$ |  |  |  |  |  |  |  |  |  |
| 98 | 79 | 34 | 143 | 16 | － | 1 | 16／02 | 1 | reg |
| 99 | 323 | 49 | 21 | 31 | ＋ | 1 | 16／02 | 1 | reg |
| 100 | 272 | 73 | 342 | 48 | ＋ | 3 | 15／01 | 2 | reg |
| 101 | 350 | 73 | 5 | 72 | ＋ | 2 | 15／01 | 2 | reg |
| 102 | 306 | 66 | 331 | 64 | ＋ | 1 | 15／01 | 2 | reg |
| 103 | 268 | 86 | 355 | 28 | ＋ | 2 | 15／01 | 2 | reg |
| 104 | 62 | 60 | 144 | 14 | ＋ | 2 | 15／01 | 2 | reg |
| 105 | 9 | 67 | 96 | 7 | ＋ | 2 | 15／01 | 2 | reg |
| 106 | 192 | 75 | 105 | 11 | － | 4 | 15／01 | 2 | reg |
| 107 | 194 | 88 | 284 | 11 | ＋ | 3 | 15／01 | 2 | reg |
| 108 | 114 | 59 | 50 | 36 | ＋ | 1 | 16／02 | 2 | reg |
| 109 | 328 | 78 | 55 | 12 | － | 1 | 16／02 | 2 | reg |
| 110 | 59 | 31 | 70 | 31 | ＋ | 1 | 16／03 | 2 | reg |
| 111 | 326 | 79 | 346 | 78 | － | 3 | 15／01 | 3 | reg |
| 112 | 164 | 89 | 77 | 69 | － | 3 | 15／01 | 3 | reg |
| 113 | 215 | 84 | 130 | 35 | － | 3 | 15／01 | 3 | reg |
| 114 | 79 | 34 | 59 | 32 | － | 1 | 16／02 | 3 | reg |
| 115 | 328 | 78 | 259 | 59 | － | ， | 16／02 | 3 | reg |
| 116 | 281 | 45 | 343 | 26 | － | 1 | 15／01 | 4 | reg |
| 117 | 164 | 89 | 254 | 8 | － | 4 | 15／01 | 4 | reg |
| 118 | 165 | 43 | 81 | 6 | － | 4 | 15／01 | 4 | reg |
| 119 | 154 | 88 | 244 | 4 | ＋ |  | 15／01 | 4 | reg |
| 120 | 133 | 77 | 221 | 7 | ＋ | 2 | 15／01 | 4 | reg |
| 121 | 292 | 74 | 207 | 18 | － | 1 | 16／02 | 4 | reg |
| 122 | 3 | 24 | 36 | 20 | ＋ | 1 | 16／03 | 4 | reg |
| 123 | 334 | 75 | 359 | 73 | ＋ | 2 | 15／01 | 6 | reg |
| 124 | 40 | 89 | 128 | 69 | ＋ | 3 | 15／01 | 6 | reg |
| 125 | 333 | 88 | 61 | 45 | ＋ | 1 | 15／01 | 6 | reg |
| 126 | 74 | 72 | 141 | 50 | － | 1 | 15／05 | 6 | reg |
| Data－file 1805， $\mathrm{n}=29$ |  |  |  |  |  |  |  |  |  |
| 127 | 320 | 80 | 231 | 7 | － | 1 | 18／02 | ， | reg |
| 128 | 33 | 77 | 102 | 58 | － | 1 | 18／03 |  | reg |
| 129 | 47 | 72 | 124 | 33 | － | 1 | 18／05 | 1 | reg |
| 130 | 190 | 40 | 124 | 18 | ＋ | 4 | 47／07 | 1 | reg |
| 131 | 29 | 80 | 109 | 42 | － | 2 | 47／07 | ， | reg |
| 132 | 245 | 40 | 189 | 25 | － | 2 | 47／07 |  | reg |
| 133 | 280 | 70 | 10 | 1 | ＋ | 3 | 47／07 | ， | reg |
| 134 | 268 | 63 | 194 | 28 | － | 1 | 48／01 |  | reg |
| 135 | 97 | 65 | 180 | 14 | ＋ | 1 | 48／01 | 1 | reg |
| 136 | 230 | 89 | 316 | 76 | ＋ | 1 | 18／01 | 2 | reg |
| 137 | 230 | 89 | 320 | 5 | － | 1 | 18／01 | 2 | reg |
| 138 | 195 | 79 | 284 | 4 | ＋ | 2 | 47／07 | 2 | reg |
| 139 | 30 | 48 | 104 | 17 | ＋ | 1 | 48／01 | 2 | reg |


|  |  |  |  | 号 |  |  |  | $\begin{aligned} & \text { E. } \\ & \text { ⿹ㅡN } \\ & \text { H } \\ & 0 \end{aligned}$ |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 140 | 16 | 70 | 69 | 59 | － | 1 | 18／01 | 3 | reg |
| 141 | 246 | 85 | 329 | 52 | － | 1 | 18／02 | 3 | reg |
| 142 | 244 | 84 | 317 | 72 | － | 1 | 18／04 | 3 | reg |
| 143 | 68 | 26 | 68 | 26 | － | 1 | 18／05 | 3 | reg |
| 144 | 168 | 50 | 155 | 49 | － | 1 | 18／05 | 3 | reg |
| 145 | 80 | 89 | 80 | 89 | － | 2 | 47／07 | 3 | reg |
| 146 | 17 | 71 | 54 | 66 | － | 2 | 47／07 | 3 | reg |
| 147 | 338 | 46 | 341 | 46 | － | 2 | 47／07 | 3 | reg |
| 148 | 24 | 60 | 24 | 60 | － | 1 | 47／07 | 3 | reg |
| 149 | 329 | 47 | 354 | 44 | － | 1 | 47／07 | 3 | reg |
| 150 | 30 | 48 | 30 | 48 | － | 1 | 48／01 | 3 | reg |
| 151 | 168 | 50 | 238 | 22 | － | 1 | 18／05 | 6 | reg |
| 152 | 80 | 89 | 169 | 39 | － | 4 | 47／07 | 6 | reg |
| 153 | 30 | 67 | 94 | 46 | ＋ | 3 | 47／07 | 6 | reg |
| 154 | 286 | 55 | 219 | 30 | － | 1 | 48／01 | 6 | reg |
| 155 | 74 | 84 | 352 | 50 | $+$ | 1 | 48／01 | 6 | reg |
| Data－file 2110， $\mathrm{n}=26$ |  |  |  |  |  |  |  |  |  |
| 156 | 176 | 87 | 88 | 42 | ＋ | 2 | 21／10 | 1 |  |
| 157 | 149 | 76 | 65 | 24 | ＋ | 1 | 45／01 | 2 |  |
| 158 | 46 | 31 | 69 | 29 | ＋ | 1 | A1／02 | 2 |  |
| 159 | 24 | 26 | 102 | 5 | ＋ | 1 | A1／02 | 2 |  |
| 160 | 65 | 46 | 107 | 38 | ＋ | 1 | A1／02 | 2 |  |
| 161 | 105 | 24 | 88 | 23 | ＋ | 1 | A1／02 | 2 |  |
| 162 | 10 | 44 | 69 | 26 | ＋ | 1 | A1／02 | 2 |  |
| 163 | 40 | 34 | 90 | 24 | ＋ | 1 | A1／02 | 2 |  |
| 164 | 22 | 30 | 69 | 22 | ＋ | 1 | A1／02 | 2 |  |
| 165 | 20 | 26 | 72 | 16 | ＋ | 1 | A1／02 | 2 |  |
| 166 | 140 | 81 | 57 | 39 | － | 1 | 21／10 | 3 |  |
| 167 | 42 | 61 | 118 | 22 | － | 1 | 21／10 | 3 |  |
| 168 | 100 | 79 | 15 | 26 | － | 1 | 21／10 | 3 |  |
| 169 | 151 | 86 | 69 | 62 | － | 2 | 21／10 | 3 |  |
| 170 | 121 | 63 | 46 | 27 | － | 2 | 21／10 | 3 |  |
| 171 | 22 | 76 | 100 | 40 | － | 2 | 21／10 | 3 |  |
| 172 | 359 | 66 | 77 | 24 | － | 1 | 45／01 | 3 |  |
| 173 | 140 | 52 | 95 | 42 | － | 1 | 45／01 | 3 |  |
| 174 | 140 | 52 | 160 | 50 | － | 1 | 45／01 | 3 |  |
| 175 | 217 | 84 | 134 | 47 | ＋ | 1 | 21／10 | 4 |  |
| 176 | 127 | 77 | 207 | 37 | － | 2 | 21／10 | 6 |  |
| 177 | 80 | 60 | 143 | 38 | － | 1 | 21／10 | 6 |  |
| 178 | 64 | 89 | 153 | 41 | ＋ | 2 | 21／10 | 6 |  |
| 179 | 140 | 52 | 190 | 40 | － | 1 | 45／01 | 6 |  |
| 180 | 60 | 38 | 144 | 4 | ＋ | 1 | A1／02 | 6 |  |
| 181 | 342 | 28 | 277 | 13 | ＋ | 1 | A1／02 | 6 |  |
| Data－file 2111， $\mathrm{n}=25$ |  |  |  |  |  |  |  |  |  |
| 182 | 293 | 80 | 293 | 80 | ＋ | 2 | 21／11 | 2 |  |
| 183 | 70 | 87 | 346 | 62 | － | 2 | 21／11 | 2 |  |
| 184 | 58 | 54 | 68 | 54 | － | 2 | 21／11 | 3 |  |
| 185 | 122 | 68 | 42 | 24 | － | 3 | 21／11 | 3 |  |
| 186 | 82 | 76 | 82 | 76 | － | 2 | 21／11 | 3 |  |
| 187 | 62 | 45 | 68 | 45 | － | 1 | 21／11 | 3 |  |
| 188 | 110 | 58 | 68 | 49 | － | 1 | 21／11 | 3 |  |
| 189 | 316 | 28 | 275 | 22 | － | 2 | 21／11 | 3 |  |
| 190 | 65 | 62 | 10 | 48 | － | 3 | 21／11 | 3 |  |
| 191 | 200 | 53 | 266 | 28 | － | 2 | 21／11 | 3 |  |
| 192 | 75 | 70 | 135 | 54 | － | 1 | 45／02 | 3 |  |
| 193 | 94 | 80 | 20 | 55 | － | 1 | 45／02 | 3 |  |
| 194 | 100 | 64 | 142 | 56 | － | 1 | 45／02 | 3 |  |
| 195 | 80 | 60 | 84 | 60 | － | 1 | 45／02 | 3 |  |
| 196 | 73 | 70 | 75 | 70 | － | 1 | 45／02 | 3 |  |
| 197 | 45 | 56 | 114 | 27 | － | 1 | 45／02 | 3 |  |
| 198 | 109 | 65 | 159 | 54 | － | 1 | 45／02 | 3 |  |
| 199 | 320 | 80 | 234 | 20 | － | 1 | 45／02 | 4 |  |
| 200 | 308 | 80 | 220 | 13 | － | 1 | 45／02 | 4 |  |
| 201 | 40 | 62 | 339 | 42 | － | 2 | 21／11 | 6 |  |
| 202 | 240 | 14 | 255 | 14 | ＋ | 3 | 21／11 | 6 |  |
| 203 | 292 | 60 | 4 | 28 | － | 2 | 21／11 | 6 |  |
| 204 | 200 | 53 | 144 | 37 | － | 3 | 21／11 | 6 |  |
| 205 | 70 | 87 | 156 | 56 | ＋ | 2 | 21／11 | 6 |  |
| 206 | 292 | 84 | 19 | 28 | ＋ | 1 | 45／02 | 6 |  |
| Data－file 2201， $\mathrm{n}=20$ |  |  |  |  |  |  |  |  |  |
| 207 | 252 | 54 | 186 | 29 | － | 3 | 22／01 | 1 | reg |
| 208 | 209 | 86 | 120 | 11 | ＋ | 3 | 22／01 | 1 | reg |
| 209 | 100 | 89 | 190 | 9 | － | 1 | 22／01 | 1 | reg |
| 210 | 156 | 28 | 216 | 15 | － | 2 | 22／01 | 2 | reg |
| 211 | 196 | 69 | 110 | 9 | － | 2 | 22／01 | 2 | reg |
| 212 | 306 | 81 | 26 | 48 | － | 2 | 22／01 | 2 | reg |
| 213 | 345 | 56 | 38 | 42 | － | 2 | 22／01 | 2 | reg |
| 214 | 331 | 67 | 43 | 36 | － | 1 | 22／01 | 2 | reg |
| 215 | 198 | 80 | 266 | 65 | － | 3 | 22／01 | 3 | reg |
| 216 | 238 | 72 | 292 | 61 | － | 4 | 22／01 | 3 | reg |


|  |  |  | 䂞会 | 음 | $\begin{aligned} & \ddot{0} \\ & \stackrel{y}{\omega} \\ & \underset{\sim}{2} \end{aligned}$ |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 217 | 238 | 72 | 180 | 59 | － | 2 | 22／01 | 3 | reg |
| 218 | 210 | 70 | 135 | 34 | ＋ | 2 | 22／01 | 4 | reg |
| 219 | 315 | 61 | 4 | 50 | ＋ | 2 | 22／01 | 4 | reg |
| 220 | 190 | 78 | 237 | 73 | ＋ | 3 | 22／01 | 4 | reg |
| 221 | 200 | 86 | 267 | 80 | ＋ | 3 | 22／01 | 4 | reg |
| 222 | 9 | 45 | 44 | 40 | ＋ | 3 | 22／01 | 4 | reg |
| 223 | 12 | 53 | 45 | 48 | ＋ | 2 | 22／01 | 4 | reg |
| 224 | 192 | 76 | 263 | 51 | ＋ | 1 | 22／01 | 4 | reg |
| 225 | 12 | 30 | 326 | 22 | － | 3 | 22／01 | 6 | reg |
| 226 | 184 | 48 | 238 | 33 | － | 2 | 22／01 | 6 | reg |
| Data－file 2204， $\mathrm{n}=20$ |  |  |  |  |  |  |  |  |  |
| 227 | 226 | 46 | 244 | 45 | ＋ | 2 | 22／04 | 2 | reg |
| 228 | 48 | 80 | 322 | 20 | － | 1 | 22／04 | 2 | reg |
| 229 | 257 | 75 | 195 | 60 | ＋ | 2 | 22／04 | 2 | reg |
| 230 | 198 | 55 | 153 | 45 | － | 1 | 22／04 | 3 | reg |
| 231 | 220 | 51 | 182 | 45 | － | 2 | 22／04 | 3 | reg |
| 232 | 234 | 72 | 159 | 40 | － | 2 | 22／04 | 3 | reg |
| 233 | 226 | 66 | 163 | 45 | － | 2 | 22／04 | 3 | reg |
| 234 | 3 | 50 | 21 | 49 | － | 3 | 22／04 | 3 | reg |
| 235 | 218 | 52 | 170 | 41 | － | 1 | 22／04 | 3 | reg |
| 236 | 242 | 75 | 204 | 71 | － | 1 | 22／04 | 3 | reg |
| 237 | 12 | 55 | 303 | 26 | － | 2 | 22／04 | 3 | reg |
| 238 | 266 | 33 | 234 | 29 | － | 1 | 22／04 | 3 | reg |
| 239 | 258 | 70 | 197 | 53 | － | 4 | 22／04 | 3 | reg |
| 240 | 30 | 63 | 64 | 59 | ＋ | 2 | 22／04 | 4 | reg |
| 241 | 258 | 55 | 209 | 43 | ＋ | 1 | 22／04 | 4 | reg |
| 242 | 237 | 67 | 182 | 53 | ＋ | 1 | 22／04 | 4 | reg |
| 243 | 224 | 80 | 156 | 65 | ＋ | 1 | 22／04 | 4 | reg |
| 244 | 278 | 32 | 200 | 7 | ＋ | 2 | 22／04 | 4 | reg |
| 245 | 298 | 42 | 213 | 4 | － | 1 | 22／04 | 4 | reg |
| 246 | 242 | 59 | 193 | 48 | $+$ | 2 | 22／04 | 4 | reg |
| Data－file 2504，n＝ 19 |  |  |  |  |  |  |  |  |  |
| 247 | 81 | 76 | 156 | 45 | ＋ | 3 | 25／04 | 1 |  |
| 248 | 346 | 70 | 257 | 3 | ＋ |  | 25／04 | 1 |  |
| 249 | 235 | 44 | 165 | 18 | － | 1 | 25／04 | 1 |  |
| 250 | 212 | 46 | 249 | 39 | ＋ | 3 | 25／04 | 2 |  |
| 251 | 92 | 41 | 46 | 31 | ＋ | 2 | 25／04 | 2 |  |
| 252 | 70 | 65 | 95 | 63 | ＋ | 4 | 25／04 | 2 |  |
| 253 | 164 | 68 | 238 | 32 | － | 1 | 25／04 | 2 |  |
| 254 | 232 | 78 | 297 | 63 | ＋ | 2 | 25／04 | 2 |  |
| 255 | 220 | 10 | 216 | 10 | － | 1 | 25／04 | 2 |  |
| 256 | 42 | 59 | 100 | 42 | ＋ | 4 | 25／04 | 2 |  |
| 257 | 80 | 27 | 82 | 27 | ＋ | 4 | 25／04 | 2 |  |
| 258 | 248 | 63 | 306 | 45 | － | 2 | 25／04 | 3 |  |
| 259 | 235 | 44 | 274 | 37 | － | 4 | 25／04 | 3 |  |
| 260 | 78 | 56 | 78 | 56 | － | 2 | 25／04 | 3 |  |
| 261 | 238 | 72 | 220 | 71 | ＋ | 4 | 25／04 | 6 |  |
| 262 | 75 | 56 | 154 | 16 | ＋ | 2 | 25／04 | 6 |  |
| 263 | 52 | 46 | 331 | 8 | － | 2 | 25／04 | 6 |  |
| 264 | 156 | 48 | 226 | 21 | － | 1 | 25／04 | 6 |  |
| 265 | 66 | 20 | 139 | 6 | $+$ | 1 | 25／04 | 6 |  |
| Data－file 2803， $\mathrm{n}=11$ |  |  |  |  |  |  |  |  |  |
| 266 | 70 | 55 | 143 | 23 | － | 3 | 28／01 | 1 | reg |
| 267 | 18 | 56 | 74 | 40 | － | 3 | 28／01 | 1 | reg |
| 268 | 60 | 49 | 332 | 2 | ＋ | 1 | 28／01 | 1 | reg |
| 269 | 50 | 64 | 121 | 33 | － | 3 | 28／01 | 1 | reg |
| 270 | 264 | 56 | 307 | 47 | ＋ | 1 | 28／01 | 1 | reg |
| 271 | 112 | 39 | 72 | 31 | ＋ | 2 | 28／01 | 1 | reg |
| 272 | 246 | 41 | 171 | 13 | － | 2 | 28／04 | 1 | reg |
| 273 | 36 | 45 | 111 | 15 | ＋ | 1 | 28／02 | 2 | reg |
| 274 | 85 | 52 | 150 | 28 | ＋ | 2 | 28／03 | 2 | reg |
| 275 | 50 | 56 | 123 | 24 | ＋ | 3 | 28／03 | 2 | reg |
| 276 | 78 | 55 | 64 | 54 | － | 1 | 28／06 | 3 | reg |
| Data－file 2908， $\mathrm{n}=42$ |  |  |  |  |  |  |  |  |  |
| 277 | 68 | 72 | 143 | 39 | － | 3 | 29／08 | 1 | reg |
| 278 | 70 | 61 | 136 | 36 | － |  | 29／08 | 1 | reg |
| 279 | 85 | 89 | 174 | 42 | ＋ | 3 | 29／08 |  | reg |
| 280 | 81 | 86 | 352 | 10 | － | 4 | 29／08 | 1 | reg |
| 281 | 288 | 85 | 17 | 6 | ＋ | 1 | 48／03 | 1 | reg |
| 282 | 269 | 68 | 186 | 18 | － |  | 48／03 |  | reg |
| 283 | 260 | 75 | 177 | 24 | － | 1 | 48／03 | 1 | reg |
| 284 | 268 | 57 | 199 | 30 | － |  | 48／03 | 1 | reg |
| 285 | 322 | 10 | 330 | 10 | ＋ | 1 | 48／03 | 1 | reg |
| 286 | 139 | 54 | 62 | 17 | ＋ | 3 | 29／08 | 2 | reg |
| 287 | 163 | 57 | 77 | 7 | ＋ | ， | 48／03 | 2 | reg |
| 288 | 324 | 55 | 262 | 33 | ＋ | 1 | 48／03 | 2 | reg |
| 289 | 116 | 89 | 28 | 56 | － | 4 | 29／08 | 3 | reg |
| 290 | 176 | 80 | 180 | 80 | － | 1 | 29／08 | 3 | reg |
| 291 | 358 | 42 | 17 | 41 | － | 1 | 29／08 | 3 | reg |
| 292 | 358 | 61 | 11 | 60 | － | 2 | 29／08 | 3 | reg |


|  |  |  | 吾导 | : | $\begin{aligned} & \ddot{0} \\ & \stackrel{0}{5} \\ & 0 \end{aligned}$ |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 293 | 19 | 46 | 352 | 43 | － | 2 | 29／08 | 3 | reg |
| 294 | 8 | 75 | 15 | 75 | － | 2 | 29／08 | 3 | reg |
| 295 | 350 | 73 | 294 | 61 | － | 3 | 29／08 | 3 | reg |
| 296 | 320 | 62 | 293 | 59 | － | 2 | 29／08 | 3 | reg |
| 297 | 340 | 54 | 24 | 45 | － | 1 | 29／08 | 3 | reg |
| 298 | 353 | 57 | 7 | 56 | － | 3 | 29／08 | 3 | reg |
| 299 | 164 | 26 | 164 | 26 | － | 1 | 48／03 | 3 | reg |
| 300 | 76 | 66 | 356 | 22 | ＋ | 3 | 29／08 | 4 | reg |
| 301 | 356 | 83 | 270 | 27 | － | 2 | 29／08 | 4 | reg |
| 302 | 353 | 57 | 57 | 35 | ＋ | 3 | 29／08 | 4 | reg |
| 303 | 103 | 89 | 193 | 5 | ＋ | 1 | 48／03 | 4 | reg |
| 304 | 329 | 13 | 252 | 3 | － | 1 | 48／03 | 4 | reg |
| 305 | 310 | 12 | 222 | 1 | － | 1 | 48／03 | 4 | reg |
| 306 | 174 | 80 | 263 | 5 | ＋ | 1 | 48／03 | 4 | reg |
| 307 | 254 | 50 | 184 | 23 | ＋ | 1 | 48／03 | 4 | reg |
| 308 | 177 | 89 | 87 | 19 | － | 1 | 48／03 | 4 | reg |
| 309 | 42 | 47 | 42 | 47 | ＋ | 1 | 48／03 | 4 | reg |
| 310 | 176 | 80 | 99 | 53 | － | 2 | 29／08 | 6 | reg |
| 311 | 7 | 85 | 6 | 85 | ＋ | 2 | 29／08 | 6 | reg |
| 312 | 117 | 84 | 179 | 78 | ＋ | 1 | 29／08 | 6 | reg |
| 313 | 76 | 66 | 140 | 45 | － | 4 | 29／08 | 6 | reg |
| 314 | 345 | 44 | 308 | 37 | － | 1 | 29／08 | 6 | reg |
| 315 | 291 | 20 | 223 | 8 | ＋ | 1 | 48／03 | 6 | reg |
| 316 | 316 | 10 | 20 | 4 | － | 1 | 48／03 | 6 | reg |
| 317 | 335 | 58 | 259 | 20 | － | 1 | 48／03 | 6 | reg |
| 318 | 335 | 58 | 31 | 42 | － | 1 | 48／03 | 6 | reg |
| Data－file 2909， $\mathrm{n}=21$ |  |  |  |  |  |  |  |  |  |
| 319 | 212 | 89 | 301 | 36 | － | 2 | 29／09 | ， | reg |
| 320 | 172 | 86 | 259 | 34 | － | 3 | 29／09 | 1 | reg |
| 321 | 244 | 23 | 200 | 17 | ＋ | 3 | 29／09 | 1 | reg |
| 322 | 230 | 43 | 206 | 40 | － | 3 | 29／09 | 1 | reg |
| 323 | 265 | 46 | 194 | 18 | － | 1 | 48／04 | 1 | reg |
| 324 | 282 | 65 | 205 | 26 | － | 1 | 48／04 | 1 | reg |
| 325 | 324 | 72 | 234 | 1 | － | 1 | 48／04 | 1 | reg |
| 326 | 340 | 80 | 66 | 21 | － | 1 | 48／04 | 2 | reg |
| 327 | 166 | 76 | 166 | 76 | － | 2 | 29／09 | 3 | reg |
| 328 | 332 | 85 | 59 | 26 | － | 1 | 48／04 | 3 | reg |
| 329 | 252 | 65 | 199 | 53 | － | ， | 48／04 | 3 | reg |
| 330 | 172 | 50 | 204 | 45 | － | 1 | 48／04 | 3 | reg |
| 331 | 252 | 64 | 183 | 37 | － | 1 | 48／04 | 3 | reg |
| 332 | 274 | 49 | 188 | 5 | ＋ | ， | 29／09 | 4 | reg |
| 333 | 250 | 55 | 205 | 46 | ＋ | 4 | 29／09 | 4 | reg |
| 334 | 250 | 48 | 216 | 43 | ＋ | 2 | 29／09 | 4 | reg |
| 335 | 268 | 70 | 178 | 1 | ＋ | ， | 48／04 | 4 | reg |
| 336 | 216 | 21 | 127 | 1 | － | 1 | 29／09 | 6 | reg |
| 337 | 6 | 79 | 284 | 35 | ＋ | 1 | 29／09 | 6 | reg |
| 338 | 187 | 80 | 102 | 26 | － |  | 48／04 | 6 | reg |
| 339 | 300 | 80 | 210 | 1 | ＋ | 1 | 48／04 | 6 | reg |
| Data－file 3401， $\mathrm{n}=40$ |  |  |  |  |  |  |  |  |  |
| 340 | 255 | 53 | 177 | 16 | － | 4 | 34／01 | 1 |  |
| 341 | 289 | 73 | 210 | 33 | － | 1 | 34／03 | 1 |  |
| 342 | 85 | 72 | 152 | 50 | ＋ | 2 | 34／03 | 1 |  |
| 343 | 80 | 79 | 166 | 22 | ＋ | 2 | 34／03 | 1 |  |
| 344 | 215 | 56 | 133 | 10 | ＋ | 2 | 36／02 | 1 |  |
| 345 | 64 | 75 | 150 | 14 | － | 3 | 46／05 | 1 |  |
| 346 | 250 | 75 | 160 | 1 | － | 1 | 46／05 | 1 |  |
| 347 | 270 | 78 | 180 | 1 | － | 3 | 46／05 | 1 |  |
| 348 | 305 | 78 | 222 | 30 | － | 1 | 47／01 | 1 |  |
| 349 | 325 | 2 | 1 | 2 | ＋ | 1 | 47／01 | 1 |  |
| 350 | 268 | 53 | 205 | 31 | － | 1 | 47／01 | 1 |  |
| 351 | 205 | 51 | 160 | 41 | ＋ | 1 | 47／01 | 1 |  |
| 352 | 218 | 46 | 159 | 29 | ＋ | 1 | 47／01 | 1 |  |
| 353 | 300 | 75 | 216 | 20 | ＋ | 2 | 34／01 | 2 |  |
| 354 | 88 | 38 | 22 | 18 | ＋ | 2 | 46／05 | 2 |  |
| 355 | 184 | 56 | 152 | 52 | － | 2 | 34／01 | 3 |  |
| 356 | 80 | 79 | 80 | 79 | － | 1 | 34／03 | 3 |  |
| 357 | 282 | 74 | 346 | 57 | － | 4 | 36／01 | 3 |  |
| 358 | 148 | 74 | 178 | 71 | － | 1 | 36／02 | 3 |  |
| 359 | 215 | 56 | 256 | 48 | － | 4 | 36／02 | 3 |  |
| 360 | 358 | 78 | 46 | 72 | － | 2 | 36／02 | 3 |  |
| 361 | 340 | 60 | 26 | 51 | － | 3 | 36／02 | 3 |  |
| 362 | 18 | 20 | 18 | 20 | － | 3 | 46／05 | 3 |  |
| 363 | 144 | 84 | 208 | 76 | － | 2 | 46／05 | 3 |  |
| 364 | 195 | 26 | 184 | 26 | － | 1 | 46／05 | 3 |  |
| 365 | 218 | 46 | 232 | 45 | － | 1 | 47／01 | 3 |  |
| 366 | 277 | 62 | 208 | 34 | ＋ | 2 | 34／01 | 4 |  |
| 367 | 322 | 88 | 52 | 11 | ＋ | 1 | 34／03 | 4 |  |
| 368 | 2 | 72 | 74 | 43 | ＋ | 3 | 36／01 | 4 |  |
| 369 | 140 | 85 | 51 | 15 | － | 1 | 46／05 | 4 |  |
| 370 | 306 | 42 | 225 | 8 | － | 1 | 47／01 | 4 |  |


|  |  |  |  | 会 | $\begin{aligned} & \ddot{W} \\ & \stackrel{\rightharpoonup}{5} \\ & \sim \end{aligned}$ |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 371 | 328 | 51 | 36 | 24 | ＋ | 1 | 47／01 | 4 |  |
| 372 | 225 | 48 | 193 | 43 | ＋ | 1 | 47／01 | 4 |  |
| 373 | 238 | 42 | 186 | 29 | ＋ | 1 | 47／01 | 4 |  |
| 374 | 240 | 40 | 185 | 25 | ＋ | 1 | 47／01 | 4 |  |
| 375 | 315 | 56 | 8 | 42 | － | 3 | 34／01 | 6 |  |
| 376 | 110 | 74 | 23 | 10 | － | 1 | 46／05 | 6 |  |
| 377 | 77 | 74 | 156 | 33 | － | 4 | 46／05 | 6 |  |
| 378 | 350 | 18 | 24 | 15 | － | 1 | 46／05 | 6 |  |
| 379 | 334 | 55 | 48 | 21 | － | 1 | 47／01 | 6 |  |
| Data－file 6101， $\mathrm{n}=27$ |  |  |  |  |  |  |  |  |  |
| 380 | 349 | 47 | 298 | 34 | ＋ | 1 | 65／03 | 1 |  |
| 381 | 353 | 65 | 274 | 21 | ＋ | 1 | 65／03 | 1 |  |
| 382 | 286 | 64 | 210 | 27 | － | 1 | 65／03 | 1 |  |
| 383 | 20 | 56 | 304 | 19 | ＋ | 1 | 65／03 | 1 |  |
| 384 | 3 | 68 | 279 | 14 | ＋ | ， | 65／03 | 1 |  |
| 385 | 6 | 64 | 94 | 4 | － | 1 | 65／03 | 1 |  |
| 386 | 59 | 46 | 135 | 14 | ＋ | 1 | 65／03 | 2 |  |
| 387 | 10 | 62 | 88 | 20 | ＋ |  | 65／03 | 2 |  |
| 388 | 238 | 68 | 303 | 46 | ＋ | 1 | 65／03 | 2 |  |
| 389 | 97 | 75 | 8 | 3 | ＋ | 1 | 65／03 | 2 |  |
| 390 | 19 | 89 | 290 | 59 | － | 1 | 96／07 | 2 |  |
| 391 | 223 | 75 | 283 | 63 | ＋ | 1 | 96／07 | 2 |  |
| 392 | 24 | 83 | 297 | 21 | － | 1 | 96／07 | 2 |  |
| 393 | 179 | 78 | 105 | 53 | ＋ | 1 | 96／07 | 2 |  |
| 394 | 333 | 74 | 251 | 28 | ＋ | 1 | 96／07 | 2 |  |
| 395 | 242 | 52 | 223 | 51 | ＋ | 1 | 61／01 | 2 |  |
| 396 | 160 | 61 | 129 | 57 | － | 1 | 61／01 | 3 |  |
| 397 | 186 | 60 | 194 | 60 | － | 1 | 61／01 | 3 |  |
| 398 | 32 | 72 | 359 | 69 | ＋ | 1 | 96／07 | 4 |  |
| 399 | 359 | 69 | 79 | 24 | ＋ | 1 | 61／01 | 4 |  |
| 400 | 145 | 89 | 55 | 4 | － | 1 | 61／01 | 4 |  |
| 401 | 305 | 72 | 221 | 19 | － | 1 | 61／01 | 4 |  |
| 402 | 130 | 62 | 54 | 24 | － | 1 | 61／01 | 4 |  |
| 403 | 215 | 80 | 133 | 41 | ＋ | 1 | 61／01 | 4 |  |
| 404 | 1 | 46 | 74 | 18 | － | 1 | 96／07 | 6 |  |
| 405 | 330 | 47 | 247 | 7 | ＋ | 1 | 61／01 | 6 |  |
| 406 | 230 | 45 | 142 | 2 | ＋ | 1 | 61／01 | 6 |  |
| Data－file 6703， $\mathrm{n}=20$ |  |  |  |  |  |  |  |  |  |
| 407 | 126 | 24 | 189 | 11 | ＋ | 1 | 67／03 | 1 |  |
| 408 | 358 | 64 | 294 | 42 | ＋ | 1 | 67／03 | 1 |  |
| 409 | 192 | 68 | 280 | 6 | － | 1 | 67／03 | 1 |  |
| 410 | 190 | 56 | 116 | 22 | ＋ | 1 | 67／03 | 1 |  |
| 411 | 239 | 51 | 159 | 11 | － | 1 | 67／03 | 1 |  |
| 412 | 201 | 64 | 116 | 10 | ＋ | 1 | 67／03 | 1 |  |
| 413 | 187 | 61 | 113 | 27 | ＋ | 1 | 67／03 | 1 |  |
| 414 | 194 | 76 | 109 | 20 | ＋ | 1 | 67／03 | 1 |  |
| 415 | 192 | 74 | 113 | 32 | ＋ | 1 | 67／03 | 1 |  |
| 416 | 184 | 77 | 104 | 35 | ＋ | 1 | 67／03 | 1 |  |
| 417 | 180 | 81 | 92 | 15 | ＋ | 1 | 67／03 | 1 |  |
| 418 | 184 | 76 | 103 | 31 | ＋ | 1 | 67／03 | 1 |  |
| 419 | 182 | 74 | 105 | 40 | ＋ | 1 | 67／03 | 1 |  |
| 420 | 254 | 46 | 174 | 9 | － | 1 | 67／03 | 1 |  |
| 421 | 251 | 51 | 169 | 9 | － | 1 | 67／03 | 1 |  |
| 422 | 99 | 21 | 70 | 19 | ＋ | 1 | 67／03 | 2 |  |
| 423 | 60 | 37 | 41 | 35 | ＋ | 1 | 67／03 | 6 |  |
| 424 | 237 | 76 | 151 | 15 | － | 1 | 67／03 | 6 |  |
| 425 | 201 | 74 | 122 | 35 | － | 1 | 67／03 | 6 |  |
| 426 | 246 | 66 | 333 | 5 | ＋ | 1 | 67／03 | 6 |  |
| Data－file 7102， $\mathrm{n}=20$ |  |  |  |  |  |  |  |  |  |
| 427 | 80 | 67 | 352 | 5 | － | 1 | 71／02 | 1 |  |
| 428 | 49 | 30 | 113 | 14 | ＋ | 1 | 71／02 | 2 |  |
| 429 | 98 | 56 | 66 | 51 | ＋ | 1 | 71／02 | 2 |  |
| 430 | 90 | 70 | 90 | 70 | ＋ | 1 | 71／02 | 2 |  |
| 431 | 17 | 47 | 85 | 22 | ＋ | 1 | 71／02 | 2 |  |
| 432 | 92 | 45 | 123 | 41 | ＋ | 1 | 71／02 | 2 |  |
| 433 | 244 | 88 | 333 | 21 | ＋ | 1 | 71／02 | 2 |  |
| 434 | 85 | 70 | 133 | 61 | ＋ | 1 | 71／02 | 2 |  |
| 435 | 294 | 55 | 294 | 55 | ＋ | 1 | 71／02 | 2 |  |
| 436 | 154 | 81 | 241 | 19 | ＋ | 1 | 71／02 | 4 |  |
| 437 | 135 | 86 | 223 | 24 | ＋ | 1 | 71／02 | 4 |  |
| 438 | 87 | 71 | 6 | 25 | ＋ | 1 | 71／02 | 4 |  |
| 439 | 80 | 64 | 163 | 15 | － | 1 | 71／02 | 4 |  |
| 440 | 116 | 86 | 204 | 26 | ＋ | 1 | 71／02 | 4 |  |
| 441 | 251 | 68 | 171 | 24 | ＋ | 1 | 71／02 | 4 |  |
| 442 | 302 | 74 | 226 | 40 | － | 1 | 71／02 | 4 |  |
| 443 | 329 | 71 | 246 | 19 | － | 1 | 71／02 | 4 |  |
| 444 | 275 | 88 | 186 | 23 | ＋ | 1 | 71／02 | 4 |  |
| 445 | 128 | 86 | 41 | 35 | ＋ | 1 | 71／02 | 6 |  |
| 446 | 2 | 50 | 273 | 1 | ＋ | 1 | 71／02 | 6 |  |


|  |  |  |  |  | $\begin{aligned} & \ddot{0} \\ & \stackrel{0}{5} \\ & \ddot{n} \end{aligned}$ |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Data-file 7303, $\mathrm{n}=19$ |  |  |  |  |  |  |  |  |  |
| 447 | 112 | 52 | 170 | 35 | + | 1 | 73/03 | 1 |  |
| 448 | 91 | 55 | 147 | 38 | + | 1 | 73/03 |  |  |
| 449 | 83 | 56 | 158 | 21 | + | 1 | 73/03 | 1 |  |
| 450 | 68 | 72 | 149 | 27 | - | 1 | 73/03 | 1 |  |
| 451 | 122 | 71 | 195 | 40 | + | 1 | 73/03 |  |  |
| 452 | 63 | 71 | 147 | 18 | - | 1 | 73/03 |  |  |
| 453 | 76 | 62 | 154 | 20 | - | 1 | 73/03 | 1 |  |
| 454 | 289 | 72 | 217 | 44 | - | 1 | 73/03 | , |  |
| 455 | 69 | 78 | 154 | 23 | - | 1 | 73/03 | , |  |
| 456 | 292 | 56 | 228 | 33 | + | 1 | 73/03 | 2 |  |
| 457 | 280 | 60 | 225 | 45 | + | 1 | 73/03 | 2 |  |
| 458 | 272 | 36 | 253 | 34 | + | 1 | 73/03 | 2 |  |
| 459 | 250 | 41 | 263 | 40 | - | 1 | 73/03 | 3 |  |
| 460 | 180 | 82 | 180 | 82 | - | 1 | 73/03 | 3 |  |
| 461 | 54 | 64 | 5 | 53 | + | 1 | 73/03 | 4 |  |
| 462 | 88 | 75 | 166 | 37 | - | 1 | 73/03 | 4 |  |
| 463 | 78 | 84 | 167 | 8 | - | 1 | 73/03 | 4 |  |
| 464 | 269 | 70 | 255 | 69 | + | 1 | 73/03 | 6 |  |
| 465 | 15 | 86 | 101 | 44 | $+$ | 1 | 73/03 | 6 |  |
| Data-file 8602, $\mathrm{n}=12$ |  |  |  |  |  |  |  |  |  |
| 466 | 193 | 71 | 109 | 18 | + | 1 | 86/02 | 1 |  |
| 467 | 5 | 73 | 73 | 50 | - | 1 | 86/02 | 1 |  |
| 468 | 292 | 75 | 253 | 71 | + | 1 | 86/02 | 2 |  |
| 469 | 130 | 69 | 50 | 24 | - | 1 | 86/02 | 4 |  |
| 470 | 99 | 39 | 77 | 37 | + | 1 | 86/02 | 4 |  |
| 471 | 122 | 85 | 33 | 10 | - | 1 | 86/02 | 4 |  |
| 472 | 344 | 78 | 256 | 8 | - | 1 | 86/02 | 4 |  |
| 473 | 86 | 74 | 359 | 9 | + | 1 | 86/02 | 4 |  |
| 474 | 122 | 26 | 165 | 19 | + | 1 | 86/02 | 4 |  |
| 475 | 123 | 73 | 157 | 69 | + | 1 | 86/02 | 6 |  |
| 476 | 220 | 85 | 308 | 24 | + | 1 | 86/02 | 6 |  |
| 477 | 125 | 51 | 196 | 22 | - | 1 | 86/02 | 6 |  |
| Data-file 8605, $\mathrm{n}=14$ |  |  |  |  |  |  |  |  |  |
| 478 | 302 | 70 | 25 | 19 | + | 1 | 86/05 | 1 |  |
| 479 | 141 | 69 | 87 | 57 | + | 1 | 86/05 | 1 |  |
| 480 | 85 | 73 | 173 | 7 | + | 1 | 86/05 | 1 |  |
| 481 | 78 | 49 | 57 | 47 | + | 1 | 86/05 | 1 |  |
| 482 | 121 | 66 | 171 | 55 | + | 1 | 86/05 | 1 |  |
| 483 | 324 | 86 | 240 | 55 | - | 1 | 86/05 | 1 |  |
| 484 | 258 | 82 | 336 | 55 | + | 1 | 86/05 | 1 |  |
| 485 | 243 | 84 | 313 | 72 | + |  | 86/05 | 1 |  |
| 486 | 140 | 45 | 61 | 11 | + | 1 | 86/05 | 2 |  |
| 487 | 267 | 70 | 281 | 69 | + | 1 | 86/05 | 2 |  |
| 488 | 270 | 80 | 338 | 65 | - | 1 | 86/05 | 3 |  |
| 489 | 117 | 50 | 145 | 46 | - | 1 | 86/05 | 4 |  |
| 490 | 92 | 46 | 142 | 34 | + | 1 | 86/05 | 6 |  |
| 491 | 250 | 66 | 179 | 36 | - | 1 | 86/05 | 6 |  |
| Data-file 8607, $\mathrm{n}=16$ |  |  |  |  |  |  |  |  |  |
| 492 | 60 | 82 | 137 | 57 | - | 1 | 86/07 | 1 |  |
| 493 | 220 | 61 | 149 | 30 | + | 1 | 86/07 | 1 |  |
| 494 | 110 | 62 | 110 | 62 | + | 1 | 86/07 | 1 |  |
| 495 | 210 | 70 | 129 | 24 | + | 1 | 86/07 | 1 |  |
| 496 | 282 | 61 | 352 | 32 | + | 1 | 86/07 | 1 |  |
| 497 | 270 | 70 | 350 | 26 | + | 1 | 86/07 | 1 |  |
| 498 | 285 | 52 | 359 | 20 | + | 1 | 86/07 | 1 |  |
| 499 | 282 | 68 | 211 | 39 | - | 1 | 86/07 | 1 |  |
| 500 | 310 | 66 | 223 | 6 | - | 1 | 86/07 | 1 |  |
| 501 | 22 | 88 | 295 | 62 | - | 1 | 86/07 | 2 |  |
| 502 | 128 | 64 | 65 | 43 | + | 1 | 86/07 | 2 |  |
| 503 | 208 | 61 | 293 | 10 | + | 1 | 86/07 | 2 |  |
| 504 | 72 | 49 | 27 | 39 | + | 1 | 86/07 | 4 |  |
| 505 | 294 | 86 | 23 | 13 | - | 1 | 86/07 | 6 |  |
| 506 | 290 | 55 | 12 | 12 | - | 1 | 86/07 | 6 |  |
| 507 | 122 | 36 | 176 | 23 | - | 1 | 86/07 | 6 |  |
| Data-file 9004, $\mathrm{n}=20$ |  |  |  |  |  |  |  |  |  |
| 508 | 15 | 84 | 100 | 38 | - | 1 | 90/04 | 1 |  |
| 509 | 159 | 34 | 115 | 26 | + | 1 | 90/04 | 1 |  |
| 510 | 161 | 56 | 109 | 43 | + | 1 | 90/04 | 1 |  |
| 511 | 72 | 76 | 121 | 69 | - | 1 | 90/04 | 1 |  |
| 512 | 190 | 88 | 103 | 55 | + | 1 | 90/04 | 1 |  |
| 513 | 86 | 56 | 38 | 45 | + | 1 | 90/04 | 2 |  |
| 514 | 144 | 33 | 108 | 28 | + | 1 | 90/04 | 2 |  |
| 515 | 214 | 84 | 300 | 36 | + | 1 | 90/04 | 2 |  |
| 516 | 73 | 70 | 12 | 54 | - | 1 | 90/04 | 3 |  |
| 517 | 168 | 43 | 107 | 25 | - | 1 | 90/04 | 3 |  |
| 518 | 60 | 47 | 105 | 37 | - | 1 | 90/04 | 3 |  |
| 519 | 170 | 70 | 105 | 50 | - | 1 | 90/04 | 3 |  |
| 520 | 108 | 50 | 90 | 49 | - | 1 | 90/04 | 3 |  |
| 521 | 153 | 69 | 80 | 36 | - | 1 | 90/04 | 3 |  |


|  |  |  |  |  | $\begin{aligned} & \ddot{H} \\ & \stackrel{y}{5} \\ & \hline 0 . \end{aligned}$ |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 522 | 98 | 50 | 85 | 49 | - | 1 | 90/04 | 3 |  |
| 523 | 114 | 79 | 106 | 79 | - | 1 | 90/04 | 4 |  |
| 524 | 200 | 82 | 151 | 77 | + | 1 | 90/04 | 4 |  |
| 525 | 248 | 80 | 334 | 20 | + | 1 | 90/04 | 6 |  |
| 526 | 124 | 64 | 92 | 60 | + | 1 | 90/04 | 6 |  |
| 527 | 108 | 63 | 129 | 62 | + | 1 | 90/04 | 6 |  |
| Data-file 9201, $\mathrm{n}=20$ |  |  |  |  |  |  |  |  |  |
| 528 | 290 | 82 | 213 | 57 | - | 1 | 92/01 | 1 |  |
| 529 | 39 | 45 | 95 | 29 | - | 1 | 92/01 | 1 |  |
| 530 | 60 | 88 | 148 | 37 | + | 1 | 92/01 | 1 |  |
| 531 | 284 | 24 | 280 | 24 | + | 1 | 92/01 | 1 |  |
| 532 | 1 | 59 | 87 | 6 | - | 1 | 92/01 | 1 |  |
| 533 | 269 | 86 | 181 | 22 | - | 1 | 92/01 | 1 |  |
| 534 | 280 | 86 | 191 | 19 | - | 1 | 92/01 | 1 |  |
| 535 | 74 | 86 | 160 | 42 | + | 1 | 92/01 | 1 |  |
| 536 | 39 | 78 | 122 | 31 | - |  | 92/01 | 1 |  |
| 537 | 272 | 88 | 185 | 62 | - | 1 | 92/01 | 1 |  |
| 538 | 57 | 89 | 327 | 15 | + | 1 | 92/01 | 1 |  |
| 539 | 242 | 70 | 268 | 68 | - | 1 | 92/01 | 3 |  |
| 540 | 117 | 58 | 128 | 58 | - | 1 | 92/01 | 3 |  |
| 541 | 115 | 66 | 115 | 66 | - | 1 | 92/01 | 3 |  |
| 542 | 226 | 70 | 257 | 67 | - | 1 | 92/01 | 3 |  |
| 543 | 238 | 31 | 254 | 30 | + | 1 | 92/01 | 6 |  |
| 544 | 192 | 85 | 276 | 51 | + | 1 | 92/01 | 6 |  |
| 545 | 97 | 50 | 89 | 50 | + | 1 | 92/01 | 6 |  |
| 546 | 97 | 89 | 7 | 2 | + | 1 | 92/01 | 6 |  |
| 547 | 74 | 86 | 345 | 10 | + | 1 | 92/01 | 6 |  |
| Data-file 9202, $\mathrm{n}=10$ |  |  |  |  |  |  |  |  |  |
| 548 | 260 | 50 | 306 | 40 | - | 1 | 92/02 | 4 |  |
| 549 | 211 | 70 | 296 | 13 | - | 1 | 92/02 | 4 |  |
| 550 | 292 | 74 | 15 | 23 | + | 1 | 92/02 | 4 |  |
| 551 | 326 | 85 | 56 | 4 | + | 1 | 92/02 | 4 |  |
| 552 | 64 | 77 | 338 | 17 | + |  | 92/02 | 4 |  |
| 553 | 149 | 89 | 59 | 6 | - | 1 | 92/02 | 4 |  |
| 554 | 205 | 78 | 294 | 4 | - | 1 | 92/02 | 4 |  |
| 555 | 276 | 66 | 0 | 13 | + | 1 | 92/02 | 6 |  |
| 556 | 195 | 74 | 279 | 21 | + | 1 | 92/02 | 6 |  |
| 557 | 52 | 72 | 332 | 27 | + | 1 | 92/02 | 6 |  |
| Data-file 9204, $\mathrm{n}=23$ |  |  |  |  |  |  |  |  |  |
| 558 | 11 | 59 | 310 | 39 | - |  | 92/04 | 2 |  |
| 559 | 215 | 56 | 148 | 30 | - | 1 | 92/04 | 2 |  |
| 560 | 71 | 89 | 341 | 20 | + | 1 | 92/04 | 2 |  |
| 561 | 324 | 40 | 324 | 40 | + | 1 | 92/04 | 2 |  |
| 562 | 34 | 72 | 316 | 33 | - |  | 92/04 | 2 |  |
| 563 | 215 | 63 | 286 | 32 | + | 1 | 92/04 | 2 |  |
| 564 | 324 | 88 | 234 | 10 | + |  | 92/04 | 2 |  |
| 565 | 80 | 48 | 69 | 48 | + | 1 | 92/04 | 2 |  |
| 566 | 86 | 42 | 44 | 34 | + | 1 | 92/04 | 2 |  |
| 567 | 95 | 46 | 70 | 43 | + | 1 | 92/04 | 2 |  |
| 568 | 58 | 77 | 339 | 40 | - |  | 92/04 | 3 |  |
| 569 | 231 | 88 | 321 | 11 | - | 1 | 92/04 | 4 |  |
| 570 | 247 | 89 | 157 | 2 | + | 1 | 92/04 | 4 |  |
| 571 | 220 | 89 | 310 | 26 | + | 1 | 92/04 | 4 |  |
| 572 | 59 | 86 | 331 | 30 | + | 1 | 92/04 | 4 |  |
| 573 | 19 | 69 | 303 | 32 | + | 1 | 92/04 | 4 |  |
| 574 | 40 | 78 | 318 | 32 | + | 1 | 92/04 | 4 |  |
| 575 | 320 | 88 | 49 | 33 | + | 1 | 92/04 | 4 |  |
| 576 | 58 | 58 | 131 | 24 | - | 1 | 92/04 | 6 |  |
| 577 | 339 | 72 | 296 | 65 | - | 1 | 92/04 | 6 |  |
| 578 | 332 | 44 | 304 | 41 | - | 1 | 92/04 | 6 |  |
| 579 | 123 | 62 | 198 | 25 | - | 1 | 92/04 | 6 |  |
| 580 | 345 | 35 | 294 | 24 | - | 1 | 92/04 | 6 |  |
| Data-file 9301, $\mathrm{n}=26$ |  |  |  |  |  |  |  |  |  |
| 581 | 296 | 76 | 24 | 7 | + | 1 | 93/01 | 1 |  |
| 582 | 292 | 81 | 18 | 24 | + | 1 | 93/01 | 1 |  |
| 583 | 250 | 51 | 310 | 32 | - | 1 | 93/01 | 1 |  |
| 584 | 58 | 74 | 19 | 70 | + | 1 | 93/01 | 1 |  |
| 585 | 284 | 84 | 12 | 18 | + | 1 | 93/01 | 1 |  |
| 586 | 262 | 60 | 199 | 38 | + | 1 | 93/01 | 1 |  |
| 587 | 207 | 88 | 296 | 27 | - | 1 | 93/01 | 1 |  |
| 588 | 224 | 84 | 309 | 41 | - | 1 | 93/01 | 1 |  |
| 589 | 286 | 80 | 335 | 75 | + | 1 | 93/01 | 1 |  |
| 590 | 64 | 88 | 153 | 31 | - | 1 | 93/01 | 1 |  |
| 591 | 65 | 80 | 348 | 53 | - | 1 | 93/01 | 1 |  |
| 592 | 303 | 62 | 2 | 44 | + | 1 | 93/01 | 1 |  |
| 593 | 322 | 35 | 254 | 15 | + | 1 | 93/01 | 2 |  |
| 594 | 34 | 60 | 119 | 9 | + | 1 | 93/01 | 2 |  |
| 595 | 203 | 87 | 292 | 12 | + | 1 | 93/01 | 2 |  |
| 596 | 190 | 60 | 273 | 13 | + | 1 | 93/01 | 2 |  |
| 597 | 259 | 60 | 211 | 49 | - | 1 | 93/01 | 3 |  |


|  |  |  |  | $\mid$ | $\begin{aligned} & \ddot{0} \\ & \stackrel{0}{0} \\ & 0 \end{aligned}$ |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 598 | 63 | 81 | 357 | 68 | - | 1 | 93/01 | 3 |  |
| 599 | 248 | 64 | 278 | 61 | - | 1 | 93/01 | 3 |  |
| 600 | 175 | 57 | 248 | 25 | - | 1 | 93/01 | 3 |  |
| 601 | 194 | 68 | 268 | 35 | - | 1 | 93/01 | 3 |  |
| 602 | 34 | 60 | 70 | 54 | - | 1 | 93/01 | 3 |  |
| 603 | 206 | 85 | 294 | 27 | - | 1 | 93/01 | 3 |  |
| 604 | 92 | 86 | 8 | 56 | - | 1 | 93/01 | 3 |  |
| 605 | 340 | 89 | 340 | 89 | - | 1 | 93/01 | 6 |  |
| 606 | 303 | 75 | 221 | 29 | - | 1 | 93/01 | 6 |  |
| Data-file 9906, $\mathrm{n}=15$ |  |  |  |  |  |  |  |  |  |
| 607 | 228 | 66 | 266 | 61 | + | 1 | 99/06 | 2 |  |
| 608 | 305 | 52 | 276 | 48 | + | 1 | 99/06 | 2 |  |
| 609 | 290 | 18 | 274 | 17 | + | 1 | 99/06 | 2 |  |
| 610 | 157 | 49 | 74 | 9 | - | 1 | 99/06 | 4 |  |
| 611 | 273 | 52 | 262 | 51 | - | 1 | 99/06 | 4 |  |
| 612 | 84 | 79 | 5 | 43 | + | 1 | 99/06 | 4 |  |
| 613 | 270 | 80 | 354 | 35 | - | 1 | 99/06 | 4 |  |
| 614 | 112 | 56 | 32 | 14 | + | 1 | 99/06 | 4 |  |
| 615 | 130 | 49 | 58 | 20 | - | 1 | 99/06 | 4 |  |
| 616 | 310 | 28 | 18 | 12 | - | 1 | 99/06 | 4 |  |
| 617 | 231 | 79 | 158 | 56 | + | 1 | 99/06 | 4 |  |
| 618 | 292 | 26 | 353 | 14 | + | 1 | 99/06 | 4 |  |
| 619 | 35 | 74 | 316 | 35 | - | 1 | 99/06 | 6 |  |
| 620 | 290 | 18 | 11 | 3 | + | 1 | 99/06 | 6 |  |
| 621 | 270 | 26 | 190 | 5 | - | 1 | 99/06 | 6 |  |
| Data-file a103, $\mathrm{n}=32$ |  |  |  |  |  |  |  |  |  |
| 622 | 290 | 64 | 207 | 15 | - | 1 | A1/01 | 1 |  |
| 623 | 16 | 79 | 104 | 12 | - | 1 | A1/01 | 1 |  |
| 624 | 22 | 85 | 109 | 29 | - | 1 | A1/01 | 1 |  |
| 625 | 9 | 83 | 97 | 21 | - | 1 | A1/01 | 1 |  |
| 626 | 3 | 75 | 88 | 19 | - | 1 | A1/01 | 1 |  |
| 627 | 9 | 74 | 95 | 14 | - | 1 | A1/01 | 1 |  |
| 628 | 292 | 42 | 349 | 26 | + | 1 | A1/03 | 1 |  |
| 629 | 81 | 88 | 171 | 11 | + | 1 | A1/03 | 1 |  |
| 630 | 92 | 86 | 166 | 76 | - | 1 | A1/01 | 2 |  |
| 631 | 142 | 82 | 53 | 9 | + | 1 | A1/01 | 2 |  |
| 632 | 277 | 74 | 277 | 74 | + | 1 | A1/01 | 2 |  |
| 633 | 300 | 65 | 225 | 30 | + | 1 | A1/03 | 2 |  |
| 634 | 353 | 85 | 290 | 80 | - | 1 | A1/03 | 2 |  |
| 635 | 313 | 80 | 38 | 24 | - | 1 | A1/03 | 2 |  |
| 636 | 302 | 78 | 30 | 8 | - | 1 | A1/03 | 2 |  |
| 637 | 248 | 41 | 296 | 30 | + | 1 | A1/03 | 2 |  |
| 638 | 299 | 68 | 254 | 61 | + | 1 | A1/03 | 2 |  |
| 639 | 194 | 84 | 113 | 55 | - | 1 | A1/01 | 3 |  |
| 640 | 112 | 76 | 65 | 70 | - | 1 | A1/01 | 3 |  |
| 641 | 118 | 89 | 30 | 59 | - | 1 | A1/01 | 3 |  |
| 642 | 123 | 72 | 160 | 68 | - | 1 | A1/01 | 3 |  |
| 643 | 117 | 86 | 48 | 80 | - | 1 | A1/01 | 3 |  |
| 644 | 298 | 65 | 340 | 57 | - | 1 | A1/01 | 3 |  |
| 645 | 117 | 76 | 58 | 64 | - | 1 | A1/01 | 3 |  |
| 646 | 110 | 84 | 40 | 72 | - | 1 | A1/01 | 3 |  |
| 647 | 249 | 85 | 161 | 30 | + | 1 | A1/03 | 4 |  |
| 648 | 53 | 36 | 110 | 22 | - | 1 | A1/01 | 6 |  |
| 649 | 263 | 81 | 183 | 47 | + | 1 | A1/01 | 6 |  |
| 650 | 195 | 79 | 111 | 27 | - | 1 | A1/01 | 6 |  |
| 651 | 129 | 88 | 42 | 58 | - | 1 | A1/01 | 6 |  |
| 652 | 129 | 88 | 39 | 21 | - | 1 | A1/01 | 6 |  |
| 653 | 310 | 51 | 25 | 17 | - | 1 | A1/03 | 6 |  |
| Data-file a105, $\mathrm{n}=11$ |  |  |  |  |  |  |  |  |  |
| 654 | 108 | 65 | 20 | 4 | - | 1 | A1/05 | 1 |  |
| 655 | 108 | 73 | 24 | 19 | - | 1 | A1/05 | 1 |  |
| 656 | 113 | 34 | 113 | 34 | + | 1 | A1/05 | 1 |  |
| 657 | 185 | 46 | 139 | 35 | + | 1 | A1/05 | 1 |  |
| 658 | 72 | 89 | 342 | 16 | + | 1 | A1/05 | 1 |  |
| 659 | 55 | 82 | 131 | 60 | - | 1 | A1/05 | 3 |  |
| 660 | 114 | 75 | 84 | 73 | - | 1 | A1/05 | 3 |  |
| 661 | 95 | 66 | 95 | 66 | - | 1 | A1/05 | 3 |  |
| 662 | 199 | 75 | 281 | 27 | - | 1 | A1/05 | 3 |  |
| 663 | 116 | 85 | 41 | 70 | - | 1 | A1/05 | 3 |  |
| 664 | 165 | 84 | 254 | 10 | $+$ | 1 | A1/05 | 6 |  |
| Additional fault data (not included in calculations for chapter 6) |  |  |  |  |  |  |  |  |  |
| 1 | 200 | 65 | 113 | 6 | + | 1 | 11/01 | 1 |  |
| 2 | 58 | 67 | 347 | 36 | - | 1 | 11/01 | 6 |  |
| 3 | 248 | 37 | 329 | 7 | - | 1 | 11/02 | 6 |  |
| 4 | 208 | 62 | 127 | 17 | + | 1 | 12/01 | 1 |  |
| 5 | 242 | 80 | 298 | 72 | + | 1 | 12/01 | 2 |  |
| 6 | 203 | 88 | 284 | 77 | + | 1 | 12/01 | 4 |  |
| 7 | 208 | 59 | 289 | 13 | - | 1 | 12/01 | 6 |  |
| 8 | 232 | 85 | 321 | 11 | - | 1 | 12/11 | 1 |  |
| 9 | 245 | 65 | 166 | 23 | - | 1 | 13/01 | 1 |  |


|  |  |  |  |  | $\begin{aligned} & \ddot{0} \\ & \stackrel{0}{5} \\ & \ddot{n} \end{aligned}$ |  |  | $\begin{aligned} & \text { E. } \\ & \text { ⿹ㅡㅇ } \\ & \text { E } \\ & 0 \end{aligned}$ |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 10 | 292 | 86 | 203 | 12 | + | 1 | 13/01 | 2 |  |
| 11 | 289 | 50 | 334 | 40 | - | 1 | 13/01 | 6 |  |
| 12 | 260 | 80 | 344 | 29 | - | 4 | 13/01 | 6 |  |
| 13 | 278 | 86 | 197 | 64 | + | 1 | 13/02 | 2 |  |
| 14 | 235 | 83 | 319 | 43 | - | 2 | 13/03 | 3 |  |
| 15 | 310 | 66 | 232 | 24 | - | 2 | 13/03 | 4 |  |
| 16 | 285 | 45 | 333 | 33 | - | 1 | 14/07 | 4 |  |
| 17 | 93 | 41 | 93 | 41 | - | 1 | 14/08 | 3 |  |
| 18 | 76 | 25 | 25 | 16 | - | 2 | 14/08 | 3 |  |
| 19 | 187 | 40 | 164 | 38 | - | 1 | 14/09 | 3 |  |
| 20 | 73 | 51 | 143 | 22 | + |  | 15/07 | 2 |  |
| 21 | 330 | 47 | 53 | 8 | + | 2 | 16/04 | 4 |  |
| 22 | 100 | 69 | 29 | 39 | + | 1 | 16/07 | 2 |  |
| 23 | 100 | 69 | 71 | 67 | - |  | 16/07 | 3 |  |
| 24 | 150 | 79 | 62 | 9 | + | 1 | 17/01 | 2 |  |
| 25 | 170 | 70 | 80 | 1 | + | 1 | 17/01 | 2 |  |
| 26 | 99 | 40 | 158 | 24 | - |  | 17/01 | 4 |  |
| 27 | 288 | 78 | 201 | 16 | + | 1 | 17/02 | 4 |  |
| 28 | 299 | 86 | 299 | 86 | + | 1 | 17/03 | 6 |  |
| 29 | 65 | 73 | 106 | 67 | + | 4 | 17/06 | 6 |  |
| 30 | 259 | 45 | 195 | 24 | - | 1 | 19/02 | 1 |  |
| 31 | 270 | 36 | 270 | 36 | - | 1 | 19/02 | 3 |  |
| 32 | 22 | 89 | 110 | 64 | - | 3 | 19/02 | 4 |  |
| 33 | 298 | 59 | 19 | 14 | + | 3 | 19/04 | 1 |  |
| 34 | 50 | 89 | 321 | 18 | - | 1 | 19/05 | 2 |  |
| 35 | 18 | 42 | 58 | 34 | + | 1 | 21/01 | 2 |  |
| 36 | 156 | 76 | 236 | 36 | + | 1 | 21/01 | 4 |  |
| 37 | 156 | 86 | 69 | 41 | - | 1 | 21/01 | 6 |  |
| 38 | 177 | 85 | 90 | 26 | + | 1 | 21/03 | 1 |  |
| 39 | 24 | 73 | 104 | 30 | - |  | 21/03 | 1 |  |
| 40 | 187 | 75 | 106 | 31 | + | 1 | 21/03 | 1 |  |
| 41 | 50 | 26 | 87 | 21 | + | 1 | 21/03 | 2 |  |
| 42 | 165 | 51 | 100 | 27 | + | 1 | 21/03 | 2 |  |
| 43 | 222 | 45 | 247 | 42 | - | 1 | 21/03 | 3 |  |
| 44 | 114 | 80 | 59 | 73 | - | 1 | 21/03 | 3 |  |
| 45 | 23 | 30 | 90 | 12 | + | 1 | 21/04 | 2 |  |
| 46 | 85 | 71 | 4 | 26 | + | 1 | 21/05 | 4 |  |
| 47 | 92 | 86 | 179 | 41 | + | 1 | 21/07 | 1 |  |
| 48 | 11 | 52 | 290 | 12 | + | 2 | 25/09 | 1 |  |
| 49 | 169 | 56 | 137 | 51 | + | 3 | 25/09 | 1 |  |
| 50 | 252 | 60 | 325 | 27 | + | 1 | 25/09 | 2 |  |
| 51 | 247 | 56 | 310 | 33 | + | 1 | 25/09 | 2 |  |
| 52 | 206 | 88 | 292 | 67 | + | 2 | 25/10 | 2 |  |
| 53 | 124 | 81 | 50 | 59 | - | 3 | 25/10 | 3 |  |
| 54 | 35 | 85 | 111 | 69 | $+$ | 2 | 25/10 | 6 |  |
| 55 | 350 | 58 | 77 | 6 | - | 1 | 25/10 | 6 |  |
| 56 | 326 | 80 | 54 | 9 | - | 4 | 26/03 | 2 |  |
| 57 | 234 | 40 | 299 | 19 | - | 1 | 26/03 | 6 |  |
| 58 | 61 | 81 | 338 | 37 |  | 1 | 27/01 | 2 |  |
| 59 | 351 | 86 | 261 | 6 | + | 2 | 27/01 | 2 |  |
| 60 | 210 | 82 | 293 | 41 | + | 4 | 27/01 | 2 |  |
| 61 | 64 | 86 | 338 | 40 | + | 3 | 27/01 | 6 |  |
| 62 | 300 | 87 | 23 | 65 | - | 2 | 27/01 | 6 |  |
| 63 | 210 | 40 | 169 | 32 | + | 1 | 27/03 | 4 |  |
| 64 | 116 | 28 | 159 | 21 | + | 2 | 29/01 | 1 |  |
| 65 | 116 | 28 | 50 | 12 | - | 3 | 29/01 | 6 |  |
| 66 | 50 | 52 | 122 | 22 | - | 1 | 29/03 | 1 |  |
| 67 | 70 | 42 | 141 | 16 | - | 2 | 29/06 | 1 |  |
| 68 | 50 | 25 | 1 | 17 | + | 1 | 30/02 | 4 |  |
| 69 | 100 | 86 | 12 | 35 | + | 3 | 30/02 | 4 |  |
| 70 | 335 | 55 | 3 | 52 | + | 2 | 30/03 | 1 |  |
| 71 | 304 | 34 | 15 | 13 | - | 4 | 30/03 | 6 |  |
| 72 | 6 | 88 | 277 | 15 | - | 2 | 30/04 | 2 |  |
| 73 | 30 | 89 | 301 | 54 | + | 1 | 30/04 | 6 |  |
| 74 | 50 | 76 | 333 | 40 | - | 3 | 30/05 | 2 |  |
| 75 | 143 | 84 | 54 | 10 | + | 1 | 30/05 | 2 |  |
| 76 | 118 | 80 | 204 | 24 | - | 2 | 32/01 | 2 |  |
| 77 | 302 | 69 | 222 | 24 | - | 2 | 32/01 | 4 |  |
| 78 | 275 | 70 | 3 | 6 | - | 3 | 32/02 | 4 |  |
| 79 | 258 | 78 | 184 | 52 | - | 1 | 32/03 | 1 |  |
| 80 | 342 | 55 | 16 | 50 | - | 2 | 32/03 | 3 |  |
| 81 | 328 | 70 | 43 | 35 | + | 2 | 33/02 | 1 |  |
| 82 | 208 | 65 | 126 | 16 | + | 1 | 33/02 | 1 |  |
| 83 | 61 | 86 | 332 | 26 | - | 2 | 33/02 | 2 |  |
| 84 | 254 | 81 | 166 | 13 | + | 1 | 33/02 | 4 |  |
| 85 | 117 | 80 | 31 | 20 | + | 2 | 33/03 | 2 |  |
| 86 | 346 | 51 | 52 | 27 | + | 2 | 33/03 | 4 |  |
| 87 | 250 | 73 | 164 | 13 | + | 1 | 33/03 | 4 |  |
| 88 | 336 | 63 | 60 | 10 | + | 2 | 33/03 | 4 |  |
| 89 | 330 | 62 | 258 | 29 | + | 1 | 34/02 | 2 |  |
| 90 | 87 | 68 | 4 | 17 | $+$ | 1 | 34/02 | 4 |  |


| 号 0 0 兑 0 |  |  |  | $\mid \stackrel{E}{\square}$ | $\begin{aligned} & \ddot{0} \\ & \text { ジ } \\ & \text { שn } \end{aligned}$ |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 91 | 300 | 64 | 14 | 29 | － | 3 | 34／02 | 6 |  |
| 92 | 335 | 39 | 288 | 29 | ＋ | 3 | 34／02 | 6 |  |
| 93 | 19 | 39 | 36 | 38 | ＋ | 1 | 34／04 | 4 |  |
| 94 | 22 | 75 | 103 | 30 | － | 2 | 35／01 | 1 |  |
| 95 | 66 | 70 | 146 | 26 | ＋ | 3 | 35／01 | 2 |  |
| 96 | 60 | 48 | 72 | 47 | － | 1 | 35／01 | 3 |  |
| 97 | 320 | 39 | 6 | 29 | ＋ | 3 | 35／02 | 1 |  |
| 98 | 328 | 46 | 254 | 15 | － | 2 | 35／02 | 4 |  |
| 99 | 338 | 68 | 51 | 35 | － | 4 | 35／02 | 6 |  |
| 100 | 175 | 78 | 252 | 46 | － | 1 | 35／05 | 6 |  |
| 101 | 282 | 65 | 357 | 30 | ＋ | 2 | 39／01 | 1 |  |
| 102 | 275 | 55 | 234 | 47 | ＋ | 3 | 39／02 | 2 |  |
| 103 | 280 | 55 | 320 | 48 | － | 2 | 41／04 | 3 |  |
| 104 | 58 | 24 | 20 | 19 | ＋ | 1 | 41／04 | 4 |  |
| 105 | 64 | 70 | 10 | 58 | ＋ | 3 | 42／01 | 6 |  |
| 106 | 85 | 86 | 356 | 11 | － | 3 | 42／02 | 1 |  |
| 107 | 278 | 80 | 191 | 16 | － | 3 | 42／02 | 1 |  |
| 108 | 76 | 70 | 354 | 21 | － | 2 | 42／02 | 1 |  |
| 109 | 115 | 66 | 62 | 54 | － | 3 | 42／02 | 3 |  |
| 110 | 116 | 48 | 189 | 19 | － | 1 | 42／02 | 4 |  |
| 111 | 115 | 66 | 195 | 22 | － | 1 | 42／02 | 6 |  |
| 112 | 280 | 80 | 197 | 32 | ＋ | 3 | 42／03 | 2 |  |
| 113 | 253 | 78 | 190 | 66 | ＋ | 1 | 42／03 | 2 |  |
| 114 | 190 | 75 | 106 | 22 | － | 3 | 42／04 | 6 |  |
| 115 | 212 | 70 | 133 | 29 | － | 2 | 42／04 | 6 |  |
| 116 | 291 | 51 | 237 | 36 | ＋ | 3 | 43／03 | 2 |  |
| 117 | 291 | 51 | 217 | 18 | － | 2 | 43／03 | 4 |  |
| 118 | 60 | 60 | 134 | 25 | － | 1 | 43／04 | 1 |  |
| 119 | 328 | 66 | 41 | 33 | ＋ | 1 | 44／01 | 1 |  |
| 120 | 268 | 46 | 356 | 2 | － | 1 | 44／01 | 4 |  |
| 121 | 46 | 60 | 338 | 33 | － | 1 | 50／01 | 6 |  |
| 122 | 35 | 55 | 35 | 55 | － | 1 | 50／05 | 3 |  |
| 123 | 348 | 70 | 298 | 60 | － | 1 | 50／05 | 3 |  |
| 124 | 85 | 80 | 166 | 39 | ＋ | 1 | 51／03 | 1 |  |
| 125 | 348 | 35 | 57 | 14 | ＋ | 1 | 51／03 | 4 |  |
| 126 | 345 | 56 | 55 | 27 | ＋ | 1 | 51／03 | 4 |  |
| 127 | 328 | 35 | 30 | 19 | － | 1 | 51／03 | 6 |  |
| 128 | 334 | 62 | 37 | 41 | － | 1 | 51／03 | 6 |  |
| 129 | 60 | 80 | 343 | 52 | － | 1 | 51／04 | 2 |  |
| 130 | 105 | 76 | 193 | 9 | ＋ | 1 | 51／05 | 1 |  |
| 131 | 332 | 72 | 245 | 10 | － | 1 | 51／05 | 1 |  |


|  |  |  |  |  | $\begin{aligned} & \ddot{0} \\ & \stackrel{\rightharpoonup}{0} \\ & 0 \end{aligned}$ |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 132 | 174 | 82 | 101 | 65 | － | 1 | 95／01 | 3 |  |
| 133 | 160 | 62 | 151 | 62 | － | 1 | 95／01 | 3 |  |
| 134 | 179 | 73 | 194 | 73 | － | 1 | 95／01 | 3 |  |
| 135 | 110 | 47 | 140 | 43 | － | 1 | 95／01 | 3 |  |
| 136 | 153 | 56 | 172 | 55 | － | 1 | 95／01 | 3 |  |
| 137 | 227 | 34 | 197 | 30 | － | 1 | 95／01 | 3 |  |
| 138 | 186 | 86 | 145 | 85 | － | 1 | 95／01 | 3 |  |
| 139 | 139 | 58 | 154 | 57 | － | 1 | 95／01 | 3 |  |
| 140 | 164 | 50 | 139 | 47 | － | 1 | 95／01 | 3 |  |
| 141 | 170 | 33 | 143 | 30 | － | 1 | 95／01 | 3 |  |
| 142 | 130 | 44 | 130 | 44 | － | 1 | 95／01 | 3 |  |
| 143 | 181 | 31 | 187 | 31 | － | 1 | 95／01 | 3 |  |
| 144 | 184 | 42 | 186 | 42 | － | 1 | 95／01 | 3 |  |
| 145 | 221 | 56 | 229 | 56 | － | 1 | 95／01 | 3 |  |
| 146 | 130 | 44 | 216 | 4 | ＋ | 1 | 95／01 | 4 |  |
| 147 | 127 | 40 | 215 | 2 | ＋ | 1 | 95／01 | 4 |  |
| 148 | 130 | 33 | 214 | 4 | ＋ | 1 | 95／01 | 4 |  |
| 149 | 240 | 66 | 157 | 17 | － | 1 | 95／01 | 6 |  |
| 150 | 211 | 38 | 167 | 29 | － | 1 | 95／01 | 6 |  |
| 151 | 232 | 55 | 320 | 2 | $+$ | 1 | 95／01 | 6 |  |

Explanation：
sense of movement：

+ reverse
－normal
Population：
：SSE－NNW directed $\sigma 1+$ ENE－WSW $\sigma 3$
2： $\mathrm{E}-\mathrm{W}$ directed $\sigma 1+$ subvertical $\sigma 3$
3：WNW－ESE extension（subvertical $\sigma 1$ ）
4：SSW－NNE directed $\sigma 1+$ WNW－ESE $\sigma 3$
6：not assigned


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This research work aimed at elucidating two main questions: (1) Is there a fabric in the mid-crustal plutons that is related to the kinematics of the subduction system and (2) which deformation mechanisms were involved during the arc evolution.
The Kohistan Arc Complex (KAC), now located between the Eurasian and Indian Plates, and thus part of the NW Himalayan collisional system provides excellent outcrops to tackle the raised questions. Results of detailed mapping, structural traverses, petrological/geochemical analyses and geochronological measurements were linked with the structural analysis of shear zones to document the deformation history of the former lower part of the KAC.
The base of the arc is characterized by imbricate (meta-) gabbroic, dioritic, granitic and tonalitic rocks. Strain localisation took place continuously from magmatic emplacement to solid state deformation during cooling of the plutons and formed 3 successive sets of shear zones. The anastomosing pattern of shear zones described here represents arc-related deformation during subduction of the Tethys oceanic lithosphere below the KAC.



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[^1]:    opx calculated with "Norm" program by P. Ulmer using: OPX
    opx: CATIONS assuming stoichiometry and chargebalance

[^2]:    calculated with "Norm" program by P. Ulmer using: Sphene Norm
    CATIONS calculated on the bases of 3 cations and 10 charges - Oh-group

[^3]:    a) all zircon fractions abraded; incl = inclusions; prism = prismatic; $\mathrm{spr}=$ short-prismatic
    b) Calculated on the basis of radiogenic $\mathrm{Pb} 208 / \mathrm{Pb} 206$ ratios, assuming concordance; c) Corrected for fractionation and spike
    d) Corrected for fractionation, spike, blank and common lead (Stacey \& Kramers, 1975)

